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AIR FORCE SYSTEMS COMMAND UNITED STATES AIR FORCE L. G. Hanscom Field, Bedford, Massachussetts

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Iv. Incl. Illus, tables r **S** TI *•t* **r;A** *I* MAY 7 TISIA The programs described in this report were developed under Ihe following czntracts: Confracts AFI 9(604) 5523 $AF19(604) 61$, and others AF19(6G4) 6344 AF 19(604) 7740 AF 19(628) 663 AFJ9(628) 2051 \overline{A} **Wolf WOIT**
Research and Development Corporation P.O. Box 136, West Concord, Massachusetts

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PREFACE

This vulume contains the complete descriptions and operating instructions for a collection of programs used at the SPADATS Center, Ent Air Force Base, Colorado Springs, Colorado The collection is primarily useful as an aid to analysts in their task of positively identifying objects detected in space. The documentation does not purport to describe a total system, but rather those programs which were deemed necessary to complement and support an already existing one.

The major portion of these programs were originally developed under the guidance and direction of Dr. Eberhard W. Wahl. Valuable assistance in the analysis of the problems involved was also rendered by Dr. H. Beat Wackernagel, Mr. Edward F. Casey, Mr. Laurence W. C.thbert, Mr. Richard F. Jenney(WRDC), and Mr. Baruch Rosenberg (WRDC). Three **-f** the programs (ASUM, ISUM and SSUM) were developed and written by the Air Force in Colorado Springs.

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CONTENTS

Section Subsection Subsection

I PROGRAM DESCRIPTION

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ILLUSTRATIONS

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1.1 PSR., Position Situation Report (SITRPT)

1. 2 Function

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The position and other status information on the catalogued satellites is supplied by this program. For a specified time either or both of the following types of reports can be produced:

- a) The Position Situation Report.
- b) The Satellite Situation Report.

The Position Situation Report shows the status of all satellites at a particular time each day. This report is issued in two forms:

- **I.** A complete printout prepared for off-line printing through the UBC.
- 2. A truncated version which can be punched on 5-level paper tape for teletype transmission.

Because this report includes information on all known objects, it may contain classified information. The reports are appropriately marked automatically by the computer based on the classification note stored in the Information file of the SEAl tape. See block format in ISUM writeup, section 1-14.

The Satellite Situation Report contains only the information not marked as classified in the Information File.

1.3 Input

In addition to the Schedule Tape control cards, this program requires at least one parameter card (P in col. 80). The first of these parameter cards specifies the following information:

- **1.** The output option (col. **1):**
	- a) 0 = Position Situation Report
	- b) i = Satellite Situation Report
	- c) $2 = Both$ Reports
- 2. The time at which the report is desired:
	- a) Two digits for hour (cols. 2, 3)
	- b) Two digits for minutes (cols. 4, 5)
	- c) Two digits for day of month (cols. 9, 10)
	- d) Three alphanumeric digits for the name of the month (cols. 12 to 14)
	- e)' Four digits for the year (cols. 16 to 19)
	- f) Z punch to signify zebra time (col. 6)
- 3. The Satellite Situation Report output code (col. 23):
	- a) Blank or zero (0) will suppress print of debris. A test is made for a one punch in character two of the first word of the I-File
	- b) One **(1)** will supply information on all satellites.
- 4. The Position Situation Report output code (col. 24):
	- a) Blank or zero (0) perigee and apogee are printed in statute miles.
	- b) One **(1)** perigee and apogee in kilometers.
- 5. Parameter card indicator (P in col. 80)

Additional parameter cards can be included. These will be treated as comment cards which will be printed immediately following Part I of the Satellite Situation Report.

The Element file and information file of the SEAl tape supply the data required to compute the position information requested.

1. 4 Output

The output on logical eleven (11) contains the following reports:

- i. The Position Situation Report if the output option zero (0) or two (2) used.
- 2. The Satellite Situation Report if the output option one **(1)** or two (2) used.

The Position Situation Report supplies the data listed below:

- a) Time and classification of the report.
- b) Satellite name, number and element number.
- c) Latitude and longitude west, both in degrees, at the time of the report.

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- d) The orbital elements, including the inclination in degrees, the anomalistic period in minutes, and the eccentricity.
- e) The apog ϵ and perigee in statute miles or kilometers. depending on the output code used in col. 24.
- f) The revolution number at the time of the report and the RA_N, L_N , T_N of this revolution.
- g) The classification of the satellite this field will be blank if the satellite is unclassified.

The Satellite Situation Report is comprised of the following information:

Part I - Objects in orbit inclusive or exclusive of debris depending on the code used in col. **²³** of the first parameter card.

a) Time of Report

b) Satellite name, code name, source and . launch date.

- c) The anomalistic period in minutes, the inclination in degrees, the apogee and perigee in statute miles.
- d) Transmitting frequencies, if any.
- e) Comments from the parameter cards which were included as input data, such as:

1961 OMICRON 3-52 FIFTY METALOBJECTS IN PLANE OF 61 OMICRON I AND 61 OMICRON 2 ORBITS

Part II - Object removed from orbit.

a) Satellite name, code, name and source.

b) Launch date and decay date.

1.5 Processing

The first parameter card is deciphered in the CONBUF area to determine the time requested for the report. The "debris", output, and code switches are set according to the selection made on this first card.

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The elements are loaded into the EBLOC by the ELMLOD routine. The first entry of the I-File is read and matched against the first element in the EFILE. If the two files are in phase, the computations are initiated for the time requested. The appropriate report or reports are generated according to the internal switches recorded. These switch settings are a result of the paths selected by the first parameter card.

If the IFILE and EFILE do not match, a test is made on the satellite number. A satellite number in the EFILE which is smaller than that in the IFILE results in the comment "NO DATA IN IFILE." These comparisons and the comments generated by them continue until both files are again in phase or empty. This procedure is followed since both files are assumed to be arranged in ascending order by the satellite number.

Error Messages

- OVERFLOW AT JA =_______. Comment will be printed 1. on off-line output and an exit made to the executive program.
- \cdot 2. SUBROUTINE ERROR AT $JA =$ ______. This will be printed on off-line output when an error return is made from any subroutine. An exit will be made to the executive program.

Formulation

Initial computations from input elements as computed in BEGIN:

$$
P_{o} = h_{x_{o}}^{2} + h_{y_{o}}^{2} + h_{z_{o}}^{2}
$$

\n
$$
W_{x} = h_{x_{o}} / \sqrt{P_{o}}
$$

\n
$$
W_{y} = h_{y_{o}} / \sqrt{P_{o}}
$$

\n
$$
W_{z} = \cos i = h_{z_{o}} / \sqrt{P_{o}}
$$

\n
$$
\sin i = \sqrt{1 - \cos^{2} i}
$$

\n
$$
i = \tan^{-1} \frac{\sin i}{\cos i}
$$

\n
$$
\sin \Omega = W_{x} / \sin i
$$

\n
$$
\cos \Omega = -W_{y} / \sin i
$$

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$$
\Omega_o = \tan^{-1} \frac{\sin \Omega}{\cos \Omega}
$$

\n
$$
e_o = \sqrt{\frac{2}{a_x N_0} + \frac{2}{a_y N_0}}
$$

\n
$$
a_o = p_o / (1 - e_o^2)
$$

\n
$$
n_o = k_e / a_o^{3/2}
$$

\n
$$
c'' = -360 M_o^2 c_o / \pi
$$

\n
$$
q_o = a_o (1 - e_o)
$$

\n
$$
k_e \tilde{L}_{so} = k_e J \{3 - 5 e_o^2 - |\cos i| (1 - 3/2 e_o^2) \sin^2 i (4 - \frac{27}{4} - e_o^2) \}
$$

$$
U_o = L_o - \Omega_o \text{ if } W_{Z_o} \ge 0
$$

$$
U_o = L_o + \Omega_o \text{ if } W_{Z_o} < 0
$$

- 2. Compute $t = (t_i t_o)$. 1440 where t_i is time at which position report is requested. Enter XYZSB to get position at time t_i
- 3. Compute θ G at t_o:

 $\theta_{G} = \theta_{o} + .9856472 \cdot t_{o}$ (Days) + 360.9856472 $\cdot t_{o}$ (fraction)

4. Enter subroutine SUBPT to get sub-latitude and sub-longitude points:

$$
\lambda_{\mathbf{E}} = \tan^{-1} (y/x) - .0043752691 \cdot t - \theta_{\mathbf{G}}
$$

$$
\phi = \tan^{-1} \left[\frac{U_{z}}{(1-f)^{2} \sqrt{1-U_{z}}^{2}} \right]
$$

5. Compute remainder of output:

PERIOD = $2\pi/KN$ $PERIGEE = a (1-e) - 1$ APOGEE = $a(1 + e) - 1$

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$$
REV = REV_{o} + [T/P_{N}] \text{ where}
$$
\n
$$
P_{N} = \frac{2\pi}{n} \left\{ 1 - \frac{3}{2} \right\} J_{2} (\frac{a_{e}}{p})^{2} \left[3 - \frac{e^{2}}{2} - \sin i^{2} (4 - \frac{3}{4e^{2}}) \right] \left\}
$$
\n
$$
E_{o} = \tan^{-1} \left[\frac{\sqrt{1 - e^{2}} a_{y_{n}}}{e^{2} + a_{x_{n}}} \right]
$$
\n
$$
M_{N} = E_{o} + \frac{\sqrt{1 - e^{2}} a_{y_{n}}}{1 + a_{x_{n}}}
$$
\n
$$
\omega = \tan^{-1} \left[\frac{a_{y_{n}}}{a_{x_{n}}} \right]
$$
\n
$$
M_{A} = L - \omega \pm \Omega \quad \text{:-if} \quad W_{z} \geq o, \quad \text{if} \quad W_{z} < o
$$
\n
$$
t_{N} = t_{i} + \frac{M_{N} - M_{A}}{1440 \cdot n} \quad \text{where} \quad |M_{N} - M_{A}| \leq \pi
$$
\n
$$
RA_{N} = \Omega + \Omega \left[(t_{n} - t_{o}) \cdot 1440 - t \right]
$$
\n
$$
\theta_{G} = \theta_{G_{o}} + .9856472 \cdot t_{n} (days) + 360.9856472 \cdot t_{n} (fraction)
$$
\n
$$
L_{N} = 360 - RA_{N} + \theta_{G} \text{ if } RA_{N} > \theta_{G}
$$
\n
$$
L_{N} = 0 \text{ if } RA_{N} = \theta_{G}
$$
\n
$$
L_{N} = 0 \text{ if } RA_{N} = \theta_{G}
$$

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UPESW E file switch UPISW I file switch

 t_{n}

TIW Integer part of TI TN t Time of node TNI Time of node (integer) TNF TIME Of node (fraction)

TOI Temporary, used in computing θ_G

UCLAS 0, no unclassified satellites

J, at least one unclassified satellite

0, no elements in E file for previous satellite. in I file; do not pick up more elements i, read next element from E file

0, no elements in I file for previous satellite from E file; do not pick up nextelement from I file 1, read next satellite from I file.

UPSSW SAVE file switch

0. no elements in SAVE file for previous satellite; do not pick up next element from SAVE file.

f, read next element from. SAVE file

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KEY: $A2$

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Retrieve Next Element Set

Unpack First Word of IFILE to Get Sat. No. $B2$

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PSR

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المجال المجال المجا

Determine if EFILE and IFILE in Phase

بالمسابر للمسابر بد

 \sim ana araw $\prod_{i=1}^{m}$

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Determine if SAVE and IFILE in Phase

 σ and σ . Then, i.e.

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 \hat{f} is a component of the \hat{f}

 χ^2/χ^2

 $= -$

 $_{\rm PSR}$

2.1

REDUCT. Nodal Crossing Reduction

$2, 2$ Function

The Nodal Crossing Reduction Program, hereafter referred to as the Reduction Program, reduces observations to the last nodal crossing; computes certain differences between parameters as computed from the observation and as computed from orbital elements; may compare the differences against predetermined toler. ance limits; and produces results in printed and/or teletype format.

The differences, or residuals, may be used by the analyst to determine the necessity of undating existing orbital elements, to determine new elements for satellite debris, to redefine the orbit of a "lost" satellite, or to identify the observation.

The observation may be visual, radar, Baker-Nunn, doppler, or direction finder. The first three are treated by that portion of the Reduction Program known as the General Reduction, and the last two by the Doppler Reduction and Direction Finder Reduction, respectively. The observation types may be intermixed from different sensors, or for different satellites.

Table 2.1 specifies the particular residuals computed by each of the sections of the Reduction Program.

Residuals Computed by Various Program Sections

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For clarification of subsequent sections the following definitions are made:

A predicted quantity (generally indicated with a prime) is the quantity calculated from orbital elements.

A computed quantity is the quantity calculated from the observation.

A known observation is a sighting which has been identified by the sensor as a particular satellite.

An unknown observation (abbreviated UO) is a sighting which has not been identified by the sensor.

: A verified observation is a known observation which may **4..** agree with the predicted position of the specified satellite.

An unverified observation is a known observation which does not agree with the predicted position of the specified satellite.

A tagged observation is an observation, either known or unknown, which does agree with the predicted position of one or more satellites.

An untagged observation is an-observation, either known or unknown, which does not agree with the predicted position of any satellite.

The station tape is a binary magnetic tape containing station (sensor) coordinates for specified stations.

The residual of a quantity is the difference between the computed and predicted quantity.

2.3 Input

Input data may be divided into four groups, some of which are optional. The four are: 1) the switch option card, 2) the element lead card and element sets, 3) the station lead card and station cards, and 4) observation lead card(s) and observations.

2. 3. **1** Switch Option Card

The switch option card sets the program switches. Setting of a switch is indicated by any non-zero numeric punch in the appropriate card column. The switches and their functions are indicated in Table 2. 2.

2. 3. **2** Element Lead and Element Cards

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The standard element sets required for reducing known observations, as well as-the sets considered necessary in attempting to tag unknowns, must be included as a part of the input deck. The element cards must be preceded by the element lead card, identified by having only a seven (7) punch in card column **8.** At present, no more than 250 element sets are permitted. Reading of the elements is terminated by either a station or an observation lead card.

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2. 3. 3 Station Lead and Station Cards

To reduce observations, the geographic coordinates of the observing station must be available. The coordinates may be read from the station tape (if program switch 1 is not set) and/or from standard station cards. Each group of station cards must be preceded by the station lead card, identified by having only an eight (8) punch in card column 8. At present, no more than 750 stations may be used.

2. 3.4 Observation Lead and Observation Cards

The observational data, the primary input to the Reduction Program, must be in the standard observation format. Each group of observations must be preceded by an observation lead card, identified by the nine (9) punch in column 8. Column 7 of the lead card is also used to specify the tolerances to be used by the program, where applicable. Table 2. 3 speci **, !** the codes and their corresponding values. Lead cards to change the tolerances may precede observations anywhere in the observation card deck. There is no limit to the number of observations which may be processed, since they are processed individually.

A blank observation card will terminate the reading of the preceding group of observational data. A card with only a nonzero numeric punch in column 79 will terminate the program.

Tolerance Codes and Corresponding Values

2.4 Output

Reduction results may be obtained in the form of teletype paper tape and/or printed output, depending upon the setting of program switch SS9 (cf. Table 2.2). Headings are also a function of the teletype option. The output consists of the satellite inventory, output preliminary to the results, and the results themselves.

$2.4.1$ Satellite Inventory

The satellite inventory, a listing of the satellite and element numbers used, is written following the termination of the reading of element sets. It is identified by the heading SATELLITE INVENTORY FOLLOWS: $-38.38.3$

The satellite and element numbers are printed in groups seven to the line, each group separated from the next by a

Output Preliminary to Reduction Results $2.4.2...$

and the case of the state The elements and tolerances in effect for the results which will follow appear in five lines, as follows:

Line 1. The satellite number

Line 2: Epoch revolution number, epoch time, nodal period (days), its first and one half of its second deriva-

> tive with respect to time, the semi-major axis, and eccentricity

Line 3. Right ascension of the ascending node, its first and

one half of its second derivative with respect to time, nodal period (minutes), and inclination

Line 4. Blank

Line 5. The tolerances (time, right ascension of the ascending node, and height)

REDUCT

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A summary of the five lines in symbolic notation follows:

Satellite No.

 Ω

 $2.4.3$

Results for Tagged or Known Observations

The output format and quantities calculated depend upon the teletype option and upon the type of observation. If teletype output is desired, all headings are omitted and the output is placed in two lines. However, the quantities contained in the two lines are identical with those which would have been printed in one line, had the teletype option not been selected. In the following subsections the description will apply to a case in which teletype output is not requested.

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 P_{Nm}

As noted above, the headings and quantities computed are a function of the observation type. Each heading is described below. The headings will be printed each time a different group of observation type is encountered; that is, each time the observation type requires the use of a different section of the program.

$2, 4, 3, 1$ Output Common to All Observation Types

The output from each of three portions of the Reduction Program begins with two common quantities, the identification, ID, and the epoch revolution, N. The identification is composed of two eight character words which may be symbolically represented by

SATYMMDD HHMMSSs.

where

 $HH = hour of the day$ MM = minutes of the day $SS =$ seconds of the day $ss =$ hundredths of seconds

If the teletype ending sentinel, \$, appears at the end of any line of output, the observation time precedes epoch by more than four days.

3. 2 Dimensions of the Output Quantities

Unless otherwise indicated the dimensions of the output quantities are in time in days, distance in kilometers, and angular measure in degrees.

2.4.3.3 Output from the General Reduction Program

Following the epoch revolution number are the argument of latitude, U, of the satellite; computed time of nodal crossing, "T.SUB N; its residual, DELTA T; the latitude, PHI **S,** and longitude, : L **S,** of the sub-satellite point; the computed right ascension of the ascending node, RA N; its residual, DEL RA; the computed satel lite height, H(KM); its residual, DEL H; observation type, TYPE; element number, ELEM; and station number, STA. If the observation type is visual, VIS, or Baker-Nunn, B-N, the height residual will appear as 0-0. If the type is radar, RDR, the computed quantity will be printed.

2.4.3.4 Output from the Doppler Reduction Program

Following the epoch revolution number are the computed time of nodal crossing, T SUB N; its residual, DELTA T; the great circle distance, D, in nautical miles from the station to the eubsatellite point; the elevation angle, H; the slant range, **S;** the

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element number, ELEM; and the station number, SAT. The title (DOPPLER) is appended to the heading, or in case the teletype option is selected, the symbol, DOP, follows the station number.

2. 4. 3.5 Output from the Direction Finder Program

Following the epoch revolution number are the computed time of nodal crossing, T SUB N; its residual, DELTA T; the slant range, **S;** the computed satellite height, H(KM); the element number, ELEM; and the station number, STA. The comment (DIR FINDER) is appended to the heading, or in case the teletype option is selected, the symbol DF follows the station number.

2. 4. 3. 6 Observation Comments

The observation comments which may appear in the printed output and their explanations are given below. Each observation comment is preceded by the 16-digit observation identification.

Observation comments (1) -(3) may be printed for both known and unknown observations.

- (1) UNDEFINED OBSERVATION (2). Illegal equipment type specified.
- (2) NO STORED COORDINATES (2). The station coordinates of the observing station are not available.
- (3) IMPROPERLY LOGGED. An error exists in the observation, e. g. , hours greater than 24, an elevation angle greater than 90 degrees, etc.

Observation comments (4)-(13) may be printed for known observations only. Each message is followed by the station number.

(4) NO ELEMENTS IN SYSTEM (1). The elements corresponding to the satellite designated in the observation have not been read.

- (5) QPRIME LESS THAN OR EQUAL TO ONE (1). The computed perigee distance is less than or equal to one earth radius.
- (6) DID NOT MEET RESIDUAL TOLERANCES (3). The observation did not meet the specified residual tolerances.
- (7) BAKER-NUNN OBS W/+RA (5). An observation is encountered in the General Reduction Program with an equipment type that indicates Baker-Nunn, but the observation does not have the required minus overpunch in column 3i of the observation card. If the observation is unknown, it becomes an untagged UO.

If the observation is unknown, it becomes an untagged UO.

- (8) ELEVATION NEGATIVE (7). The elevation computed from the declination by the General Reduction Program is negative. If an unknown observation, it becomes an UNTAGGED UO.
- (9) PHI S GREATER THAN I (8). The computed sub-satellite latitude exceeds the inclination.
- **(10)** DF AZ NEGATIVE (6). A Direction Finder observation has a negative azimuth.
- **(11)** DF ELEVATION NEGATIVE (8). A Direction Finder observation has a negative computed elevation.
- **(12)** NON-CONVERGENT OBS (9). The computation of the satellite latitude does not converge within the specified number of iterations.
- (13) SLANT RANGE (ZERO) (0). An observation with a zero slant range has been encountered by the General Reduction Program radar portion. If the observation is unknown, it becomes an untagged UO.

(14) UNTAGGED UO. An unknown observation has been reduced against all the element sets and remains untagged within the specified tolerances. If input options specify, the observation will be written on the interim tape and rereduced with large tolerances.

2. 4. 3. 7 Miscellaneous Printed Output'

The various comments which may appear in the printed matter and their explanations follow.

(1) PN MISSING **SO'** PA USED FOR SAT XXX ELEM XXX'

The nodal period at epoch is missing from the sixth-card of a seven card set. The anomalistic period has been used in its place.

(2) SATELLITE XXXX CARD XXX OUT OF ORDER

An element card is not in correct sequence, or an extraneous card is contained in the element sets. The elements for the particular satellite will not be stored in computer memory, hence are unavailable.

• (3) . UNTAGGED UOS REDUCED W/O.T OLERANCES All observations which follow have previously appeared as untagged UO's, and have now been rereduced with large tolerances. •.tolerances. •

'(4) END OF RUN ***********

All observations have been reduced with a specified tolerance. Any results which follow have been reduced with different tolerances.'

(5) UO **S** FOLLOW

All observations which follow on the page did not correspond to any satellite, the elements of which were successfully stored in computer memory.

(6) UNK***********

The next observation is an unknown. The message number is given. If the observation was a known observation which was treated as an unknown, the satellite number is also given.

2.4.3.8 Flexowriter Output

The messages which may appear on the flexowriter, and their explanations follow.

(1) SATELLITE'XXXX' CARD XXX OUT OF. ORDER

An element card is not in correct sequence, or an

extraneous card is contained in the element sets. The elements for the particular satellite will not be stored in computer memory, hence are unavailable.

(2) MOUNT RIGHT STATION TAPE

The station tape is not available to the program. The program awaits the stop-go option. .

(3) ILLEGAL LEAD CARD SKIPPED.

: An illegal lead card (cf. Sec. 2. 3) is present in the input deck.: The program continues according to the last valid lead card.

(4) MOUNT STATION TAPE ON 7

The program is ready to run, but a stop-go option is provided in case the station tape is required.

(5) REMOVE STATION TAPE.U7 **"** '

The program run is complete, but a stop-go option is. provided to give time to remove the station tape.

2. 5 Processing

The program, accomplishes its function of reducing observational data to the last nodal crossing, one observation at a time. However, certain preliminary operations must be performed prior

to operating on the first observation, and as circumstances require, prior to operating on subsequent observations.

2. 5. 1 Preliminary Operations

The preliminary operations are initialization of various program switches, examination of switch options, reading of station coordinates from the station tape, reading of satellite elements from cards, reading of station data from cards, and the reading of the first observation lead card.

"2. 5. **l1** 1 Initialization of Program Switches

Certain switches which will control the path of the program are initialized to neutral settings.

2. **5.1.** 2 Reading of Switch Option

The several program switch settings are read from the switch option card (cf. Sec. 2. 3. 1). These settings determine, among other things, whether or not the station tape is to be read and the interim tape rewound.

2 .5. 1. 3 'Reading the Station Tape

If switch SS1 is set the station tape is examined; and if it is identified as the station tape, the values of station number, latitude, longitude, and height are read into the appropriate array. If the identification test fails, a STOP-GO option is provided to allow the computer operator to mount the correct tape. GO causes the program to examine the newly mounted tape for the proper identification. STOP causes the program to return control to the executive routine.,

2. 5. 1. 4 Reading the Element Sets

The element sets are generally read as a single group by the subroutine ELRED. The program logic also permits their

being read in separate groups, providing they are identified by an element lead card. Certain validity checks are made upon each element set. If no gross errors exist, the elements are stored in the element array for future use. If errors do exist, appropriate error messages are generated and the element set is otherwise ignored.

Reading of element sets is terminated by either a station lead card or an observation lead card. Termination of reading causes the satellite inventory, a listing of those satellites whose elements were successfully read, to be generated as the first output of the program.

2. 5. **1.** 5 Reading of Station Data from Cards

If a station lead card is encountered during the running of the program, the data from the station card or cards which follow the lead card are stored in the station data array. Reading of the station cards is terminated by either an element lead card or an observation lead card.

2. 5. **1.** 6 Reading of an Observation Lead Card

Each observation or group of observations must be preceded by an observation lead card, which also specifies the tolerances to be used, where applicable, to cause the tagging of unknown observations. Table 2. 3 specifies the codes used in card column 8 of the lead card and the corresponding values of the tolerances. Following the reading of the first observation lead card the processing of observations can begin.

2. 5. 2 Preliminary Processing for all Observations

The observations are read and processed individually. Prior to the actual reduction calculations, tests are made for the validity of the observation, the station coordinates are retrieved, the appropriate elements are located and the observation type determined.

2.5.2.1 Testing of the Observation

Certain fields from the observation card are checked for validity. These fields are month, day, hour, minute, second, elevation angle, and azimuth angle. An error in any one of the fields will cause the rejection of the observation with the appropriate comment (cf. Sec. 2. 4. 2. 6 (3)). The program proceeds to read another observation card.

2. 5. **2.** 2 Retrieval of Station Coordinates

If the several fields of the observation card contain no illegal values, the station aumber from the observation is compared to that from the previous observation. If the two numbers do not agree, the station data array is searched for the current station number. If the station number is found, the station latitude, longitude, and height are retrieved, and station dependent quantities are computed. If the current station number is not found in the station data array, the observation will be rejected with the appropriate comment (cf. Sec. 2.4.3.6 (2)). The program proceeds to read another observation.

If the two station numbers do agree, the latitude and longitude are examined for zero which would indicate that the station coordinates were not previously found. The observation is then rejected with the appropriate comment.

2. 5. **2.** 3 Retrieval of Satellite Elements

If the observation is a known observation, the satellite number is compared to that of the previous satellite. If they agree, then the elements have already been retrieved and no search of the element array is necessary. If the satellite number is not the same as that of the previous satellite, the element array is searched via subroutine ELCAL3. If the search is successful the nodal elements

are printed $(cf. Sec. 2.4.2)$. If the elements are not found the observation is rejected with the appropriate comment (cf. Sec. 2. 4. 3. 6 (4)). The program proceeds to read another observation.

If the observation is an unknown observation, then the subroutine ELCAL3 will sequentially retrieve the elements from the element array until the observation has been reduced against all elements.

Subsequent to the retrieval of the elements, perigee distance for the revolution at the time of the observation is computed. If perigee distance is less than one earth radius, computations will either cease in' the case of a known observation, or continue with the next element set for an unknown observation. The appropriate comment (cf. Sec. 2.4.3.6 (5)) is printed for the known observation, and another observation is read.

-If the perigee distance computation is satisfactory, the actual reduction computation is ready to be made on the basis of the observation type..

2. 5. 2. 4 Determination of Observation Type

The observation type, a part of the observation itself, is used to determine which of the three subsections of the program is to be used for the reduction computations. An illegal observation type will cause the rejection of the observation with the appropriate comment (cf. Sec. Z. 4. 3. 6 (1)).

2. 5. 3 Processing With the General Sighting Routine

The General Sighting program processes only observations in which azimuth and elevation angles, or right ascension and declination are given. Slant range may also be utilized. The general observation types which fulfill this requirement are visual, radar and Baker-Nunn camera observations.

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2. 5.3. **1** Special Processing of Visual Observations

Visual observations may be either azimuth and elevation angles or right ascension and declination. If the right ascension and declination are given, they are converted to azimuth and elevation angles. A negative elevation angle computed from a known observation causes an error message (cf. Sec. 2. 4. 3. 6 (8)) to be generated and the program proceeds to process the next observation. If the observation is unknown, it becomes an untagged **UO.'** .Computation of a positive elevation angle causes processing to continue as indicated in Sec. 2. 5. 3. 3.

2.5. 3.12 Special Processing of Baker-Nunn Camera Observations'

Baker-Nunn camera: observations must have right ascension and declination given. These are converted to azimuth and elevation angles and the processing continues as indicated in Sec. 2. 5. 3. 3. If the overpunch signifying right ascension and declination is missing, an error message (cf. Sec. 2.4.3.6(7)) is generated for a known observation or the unknown observation is made an untagged UO. In either case, the program proceeds to the next observation.

Common Processing of Visual and Baker-Nunn Camera Observations

Neither visual nor Baker-Nunn camera observations have any range information. For this reason an iterative procedure is used to determine the sub-satellite position. The station's latitude and sidereal time are taken as first approximations to the satellite's latitude and right ascension, respectively, on the assumption that the satellite is close to the station.. Failure of the calculations to converge causes an error message (cf. Sec. 2.4.3.6 (12)) to be generated for known observations. The program proceeds to the next observation for the known observation or to the next element set for an unknown observation. If the iteration produces the

satellite latitude within the specified limit, the latitude is checked against the inclination angle. An unsatisfactory result causes an error message (cf. Sec. 2. 4. 3. 6 (9)) to be generated for a known observation and the program proceeds to the next observation. For an unknown observation the next element set is selected.

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If the computed latitude can be on the orbital plane, processing continues as indicated in Sec. 2. 5. 3. 5.

2.5. 3.4 Special Processing of Radar Observations

The examination of the azimuth and elevation angle field for a radar observation is identical to that described for a visual observation (cf. Sec. 2. 5. 3. 1). However, radar observations include a range measurement, and this is used to determine the subsatellite position. An illegal range, i. e. , negative, causes the appropriate message (cf. Sec. 2. 4. 3. 6 (13)) to be generated for a known observation, and causes an unknown observation to become : an untagged **UO.** In either case, the program proceeds to process the next observation.

If the range is valid, the satellite's latitude and right ascension are computed. The latitude is compared with the inclination angle. If the test is satisfactory, the height residual is computed. Processing then continues as in Sec. 2. 5. 3. 5. If not.an error message (cf. Sec. 2. 4. 3. 6 (9)) is generated for a known observation, and the program proceeds to process the next observation. If the observation is unknown, the program proceeds to the next element set.

2. 5. 3. 5 Common Processing of all General Sighting Observations

The longitude of the satellite, right ascension of the node, the residuals of time and right ascension of the node are computed for all general sighting observations. At this point the observation

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has passed the various gross tests and the residuals, be they large or small, have been computed. The setting of switch option **SS3** determines whether or not the residuals for an unknown observation should be compared with the tolerances. If the switch is not set and the observation is unknown, the comparison is made. Failing the comparison causes the program to select the next element set; passing causes the observation to be tagged. If the switch SS3 is set, the tolerances are ignored. The observation and the reduction results are then output for printing. The program proceeds to process the next observation on the same observation with a new element set, depending upon whether or not the current observation is known or unknown.

2. 5.4 Processing With the Doppler Routine

The Doppler portion of the program processes passive track Doppler observations and Fence observations in which range rate **e** is the only information available. Consequently, only the time residual is computed.

Preliminary calculations concerning the relative position of the ascending node and the station are made to determine certain constants for subsequent calculations. The argument of latitude, true anomaly, eccentric anomaly, and mean anomaly of the satellite are intermediate quantities used in determining the time residual, the elevation angle and the slant range. The setting of switch SS3 determines whether or not the time residual for an unknown observation should be compared with the time tolerance.

2. 5.5 Processing With the Direction Finder Routine

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The Direction Finder portion of the program processes observations in which the azimuth angle of the closest approach is given. For this reason only the time residual is computed.

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The observation will be rejected with the appropriate message (cf. Sec. **2.** 4. 3. 6 **(11))** if the azimuth is negative. If it is positive the time residual, the elevation angle and the slant range are computed. A negative elevation angle will cause an error message (cf. Sec. 2. 4. 3. 6 (10)) to be generated for a known observation. The program proceeds to the next observation or the next element set depending upon whether the negative elevation angle was computed from a known or an unknown observation.

If the elevation angle is positive, the setting of switch SS3 determines whether or not the time residual for an unknown observation should be compared with the time tolerance. If the comparison is made and the residual exceeds the tolerance, the program proceeds to select the next element set. Otherwise the observation is tagged and the reduction results for both the tagged unknown or known observation are output for printing. The program proceeds to process the next observation or the same observation with a new element set depending upon whether or not the current observation is known or unknown.

2. 5. 6 Optional Reprocessing of Untagged Unknowns

Unknown observations which have failed to be tagged may be reprocessed with large tolerances, depending upon the setting of switches SS3 oz **SSIO.** If neither are set, unknown observations which have not been tagged are written on the interim tape following the processing with the last available element set.

After all observations have been initially processed, the tolerances are set to large values. The observations are then read from the interim tape, one at a time, and reprocessed with all element sets.

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General Sighting Reduction Program

Input:

Orbital elements at epoch:

Sighting data:

2.6. **1** Formulation for the General Sighting Reduction Program 1) $\Delta t = t_i - T_0$ 2) RA = RA + RA $\Delta t + \frac{1}{2}$ RA $(\Delta t)^{2}$ $\omega' = \omega + \omega \Delta t + \frac{1}{2} \omega (\Delta t)^2$ 3) $P_a = 360^\circ P/(360^\circ - \omega P)$ 4) a = $(P_a / .058672947)^{2/3}$ 5) $P_a' = P_a + P \Delta t + 1/2 P (\Delta t)^2$ 6) $a' = (P_a' / .058672947)^{2/3}$ 7) If **q** not given: $q' = a (1 - e)$, Go to 8 If **q** given: $\Delta t' = T'_0 - t_i$ $q' = q + q \Delta t' + 1/2 \ddot{q} (\Delta t)$ 8) If **q'** = **1:** stop If q[']> 1 : e' = (a' - q)/a 9) If $e' < 0$: $0 + e'$ 10) $e'' = \sqrt{1 - e}$ \mathbf{r} . 1+ **11)** $\theta_{z} = \begin{bmatrix} \theta_{o} + 98565 \end{bmatrix}$ (Day of Year of Obs) + 360. 98565 (Fract. of day) + λ], $0 \le \theta_{\alpha} < 360$ 12) If a given convert a from hours to degrees

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E_g = 2 \tan^{-1} (e'' \tan \sqrt{2})
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\n46) $E_{RA} = -2 \tan^{-1} (e'' \tan (\omega/2))$
\n47) $M_g = E_g - e' \sin E_g$
\n48) $M_{RA} = E_{RA} - e' \sin E_{RA}$
\n49) $T_n = t_i - ((M_g - M_{RA})/360) P_a', 0 \le (M_g - M_{RA}) \le 360$
\n50) $\Delta N = (T_n - T_0)/P_a'$ (the integral portion thereof)
\n51) $T_x = T_0 + P \Delta N + c (\Delta N)^2 + d (\Delta N)^3$
\n52) If $|T_x - T_n| \le P/2$: Go to 53
\nIf not, then when: $(T_x - T_n) > 0$, set $\Delta N = \Delta N - 1$, Go to 53
\n $(T_x - T_n) = 0$, Go to 53
\n $(T_x - T_n) < 0$, set $\Delta N = \Delta N + 1$, Go to 53
\n53) $RA_x = RA + RA (T_x - T_o) + \frac{1}{2} RA (T_x - T_o)^2$
\n54) $\Delta T_n = T_n - T_x$
\n $\Delta RA = RA_n - R_x$

Output:

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02. 6. 3 Radar Portion of General Sighting Reduction Program

Input:

Same as for other portions of General Sighting Reduction Program with exception that ρ (slant range) is included as part of sighting data.

Equations:

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1) - 12) Same as in G. S.R. P 13)-19) Sames as 20 **-** 26 in G. S. R. P 20) **R**_O = $[.996643/(1-.00670015 \cos^2 \phi_i)]^{1/2}$ + H/6378.174 **Z1**) $R'' = [R_0^2 + (\rho / 6378.174)^2 + 2 R_0 (\rho / 6378.174) \sin h]^{1/2}$ 22) $a = \sin^{-1} (\rho \cos h/6378.174 R'')$ **23)** If a < 0: Return for next observation If $a \ge 0$: Goto 24 24) $\phi_e = \sin^{-1} (\sin \phi \cos \alpha + \cos \phi \sin \alpha \cos \alpha)$ 25) $\Delta L = \sin^{-1} (\sin \alpha \sin Az / \cos \phi_c)$ cos $\Delta L = (\cos \alpha - \sin \phi' \sin \phi'_{s}/\cos \phi' \cos \phi'_{s})$ If $\cos \Delta L < 0$: $\Delta L = 180 - \Delta L$: Go to 26 If $cos \Delta L > 0$: Go to 26 26) **RA_S** = (θ + Δ L), $0 \leq$ **RA**_S \leq 360 **27)** If $\phi_{\rm s}$ > i: Return for next observation If not: Go to 28 28) $u' = sin^{-1}(sin \phi_s/sin i)$

Output:

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29) If $\cos (RA_s - RA) < 0$: u = 180 **-** u' If cos $(RA_s - RA') > 0$: u = u' $\overline{30}$ v = u - w['] 31) $r = \left[a' (1 - e^{t^2}) / (1 + e' \cos v) \right] 6378.174$ 32) -46) same as 40 - 54 in G.S.R.P

> Same as in other G. S. R. P. except that R["], observed height, is included.

2. 6.4 Direction Finder Reduction Program

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Input:

Same as for General Sighting Reduction Program except that elevation - declination is not given.

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Output:

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$2.6.5$ Doppler Reduction Program

Input:

Same as General Sighting Reduction Program except that elevation - declination and azimuth-right ascension are not given.

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Output:

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Illustration of Angle η Figure 2. 2

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Continue Processing of Unknown Observation from Interim Tape

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3. 1 ROC, Radar Orbit Computation

3. 2 Function

The orbital elements of a satellite are computed from either the radar observations of the elevation angle, azimuth angle and slant range, or directly from the geocentric rectangular coordinates and velocities. The programis useful in obtaining orbital elements for newly launched satellites, in generating a new set of elements for satellites that have not been recently sighted, or in correcting the elements of a satellite.

A predicted orbit can be based on as few as three consecutive observations recorded along a single tracking run **of** a big radar, such as Millstone. An odd number of observations must be supplied. The move observations given the better the results. The midpoint of the radar run is first determined. Then the other points are picked up in pairs, on either side of the midpoint, and effectively averaged. The time span between individual pairs should be greater than forty (40) seconds. The geocentric rectangular coordinates of the midpoint are then determined and the elements computed.

If the geocentric rectangular coordinates and velocities of a point are used as input to the program, the averaging procedure is skipped and the orbital elements are computed directly.

3. 3 Input

A request card precedes either input option **(1)** or (2). The format of this card is as follows:

Cols. 1-3 Option **(1)** - Number of observation cards to be read. Must be an odd integer between 3 and 999.

Option (2) - field ignored

- Cols. 6-17 Alphanumeric satellite name or identification used in output.
- Cols. 18-35 Alphanumeric date used as part of identification for output.

Col. 41 **1** - element cards requested as output

Col. 42 **1** - Option **(1)** - x, y, z, *k,* y, z

printed for the midpoint of the observations.

- Option (2) - field ignored

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Representation of a Radar Run

Fig. 3.1

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Cols. 72-80 Must be blank.

Input option (1) should be used if orbital elements are to be computed from radar observations. Option (2) should be selected if the orbital elements are to be computed from the geocentric rectangular coordinates and velocities of a point in space. The input decks should be as follows:

- 1. Radar observations supplied.
	- 1. Request card
	- **2.** Fiandard station card
	- 3. An odd number of observation cards (3 to999) in order of increasing time.
	- 4. A card with a nine (9) punch in column 79,used to signal the end of the data deck.
- 2. $x, y, z, \dot{x}, \dot{y}$, and \dot{z} supplied.
	- 1. Request card
	- 2. Rectangular coordinates in the following format:

Cols. 1-14 x in kms. Cols. 15-28 y in kms. Cols. $29-42$ z in kms. Cols, 43-46 Year of epoch

3. Velocity components in the following format:

Cols. $1-14$ \dot{x} in kms. /sec. Cols. $15-28$ y in kms./sec. Cols. 29-42 \dot{z} in kms./sec.

- Cols. 43-56 Day of year of observation of coordinates.
- 4. Card with a nine (9) punch in column 79, used to terminate the data deck.

The quantities $x, y, z, \dot{x}, \dot{y}, \dot{z}$, and day of year must have a decimal point punched somewhere in the field.

3. 4 Output

There are two output options (A) and (B) in addition to the normal output describedunder (C).

> **A.** This option may be requested only under input option (1). If selected, the geocentric rectangular

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coordinates and velocity components, at the mean time of the input observations, will be printed. The x, y, and z coordinates in kilometers will be preceded by the labels $XX=$, $YY=$, $ZZ=$, and the velocity components in kilometers per second will be preceded by the notations XDER=, YDER=, and ZDER. The time of the observation in the day of year notation will be labled TIME=.

- B. This output option if selected will cause a punch out of the standard six card element set. The following quantities will be punched: satellite number (as read in from the observation card), element card number, satellite name, eccentricity, inclination, year of epoch, time of epoch in day of year notation, nodal period at epoch, right ascension of ascending node, argument of perigee, and perigee distance. All other fields of the six card element set are left blank since the quantities can not be computed.
- C. The normal output will be printed in the format described below.

The first line will contain the heading RADAR ORBIT COMPUTATION followed by the station name as read from the station card. The second line has the label satellite followed by the alphanumeric name as supplied on the request card. This is followed by the satellite number which was read from the observation cards. The date supplied on the request card appears on the third line of output. The elements which have been computed are then listed under appropriate headings as follows: the same-major axis in km. and earth radii, the right ascension of the ascending node in degrees, the radius vector in km. , eccentricity and its squared term, the nodal period in days and minutes, the argument of perigee in . degrees, the velocity in km. per sec., the perigee

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distance in km. and in earth radii, the time of the last perigee pass in days, the inclination in degrees, the apogee distance in km. and in earth radii, and the time of the last nodal crossing in days.

3.5 Processing

If input option (B) is used, the elements are computed directly from the x, y, z, \dot{x} , \dot{y} , and \dot{z} supplied as input. Otherwise, these coordinates are computed from the station coordinates and the radar observations. An odd number of consecutive observations, at least three, must be supplied. The midpoint, or middle observation, of a radar run is chosen. The other points are selected in pairs, one from each side of the midpoint. The time span between the individual observations in a pair should be greater than 40 seconds.

For each observation, the three topocentric coordinates are converted into the geocentric rectangular coordinates x, y , and z . Next, the velocity components \dot{x}, \dot{y} , and \dot{z} are found for each pair of observations and are averaged together for the final velocity components at the midpoint.

The orbital elements are then computed from the postion and velocity components. The magnitude of the radius vector and the velocity are computed directly from the rectangular coordinates and velocity components. The semi-major axis, nodal period, eccentricity, perigee and apogee distance are computed next, as well as, the inclination and the right ascension of the ascending node. The true, eccentric, and mean anomalies are found and the time of perigee and nodal crossing are computed. Output is then written as specified in the output section.

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3.6 Formulation

If input option 2 used go to D, otherwise compute A, B, C.

A.
$$
\phi' = \tan^{-1}(0.99329985 \tan \phi)
$$

$$
R = \frac{.9966443}{\sqrt{1 - .00670015 \cos^2{\phi'}}} + \frac{H}{6378.174}
$$

B. Determine rectangular coordinates for each observation (i).

$$
r_{i} = \sqrt{R^{2} + \rho_{i} (\rho_{i} + 2 \cdot R \sin h)}
$$
\n
$$
\beta_{i} = [\sin^{-1}(R \cos h_{i}/r_{i})] \ (0^{0} \leq \beta \leq 90^{0})
$$
\n
$$
\alpha_{i}^{*} = \left| \frac{\Pi}{2} - \beta_{i} - h_{i} \right|
$$
\n
$$
\phi_{s}^{\prime} = \sin^{-1}(\cos \alpha_{i}^{*} \sin \phi + \sin \alpha_{i} \cos \phi \cos Az_{i})
$$
\n
$$
\Delta \lambda_{s} = \sin^{-1}(\sin \alpha_{i}^{*} \sin Az_{i}/\cos \phi_{s}^{\prime})
$$
\n
$$
\lambda_{s} = \lambda - \Delta \lambda_{s}
$$
\n
$$
\theta_{G} = [\theta_{0} + .9856472 \text{ day}_{i} + 360.9856472 \text{ fract. day}_{i} - \lambda_{s}], (0^{0} \leq \theta_{G} < 360^{0})
$$
\n
$$
x_{i} = r_{i} \cos \phi_{s}^{\prime} \cos \theta_{G}
$$
\n
$$
y_{i} = r_{i} \cos \phi_{s}^{\prime} \sin \theta_{G}
$$
\n
$$
z_{i} = r_{i} \sin \phi_{s}^{\prime}
$$
\n
$$
\underline{r}_{i} = (x_{i}^{2} + y_{i}^{2} + z_{i}^{2})^{3/2}
$$

C. Compute velocity components by effectively averaging individual \dot{y}_i , \dot{z}_j , components of each pair for final x, \dot{y} , \dot{z} components. Then compute velocity and \dot{r} . Initially indices are as follows: I = 1 signifying first observation, n is total number of observations,
m is midpoint observation.

$$
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$$

$$
\tau_{12} = (t_m - t_{\ell}) (0.07436574/60)
$$
\n
$$
\tau_{23} = (t_n - t_m) (0.07436574/60)
$$
\n
$$
\tau_{13} = (t_n - t_{\ell}) (0.07436574/60)
$$
\n
$$
G_1 = \tau_{23}^2 / \tau_{12} \tau_{23} \tau_{13}
$$
\n
$$
G_2 = (\tau_{23}^2 - \tau_{12}^2) / \tau_{12} \tau_{28} \tau_{13}
$$
\n
$$
G_3 = \tau_{12}^2 / \tau_{12} \tau_{23} \tau_{13}
$$
\n
$$
H_1 = \tau_{23} / 12
$$
\n
$$
H_2 = (\tau_{23} - \tau_{12}) / 12
$$
\n
$$
H_3 = \tau_{12} / 12
$$
\n
$$
d_1 = G_1 + H_1 / \underline{r}_{\ell}
$$
\n
$$
d_2 = G_2 + H_2 / \underline{r}_m
$$
\n
$$
d_3 = G_3 + H_3 / \underline{r}_n
$$
\n
$$
\dot{x} = -d_1 x_{\ell} + d_2 x_m + d_3 x_n
$$
\n
$$
\dot{y} = -d_1 y_{\ell} + d_2 y_m + d_3 y_n
$$
\n
$$
\dot{z} = -d_1 z_{\ell} + d_2 z_m + d_3 z_n
$$
\nUpdate ℓ and n. $\ell = \ell + 1$, n = n

 $n - 1$. If $l = m$, go on. Otherwise return to C and compute \dot{x} , \dot{y} , \dot{z} for another pair of observations. m \therefore $=\sum_{i=1}^{\infty} \frac{i}{x_i}$ *(l, for l pairs*)

$$
x = \frac{1}{j} \times \frac{1}{j} / \ell, \text{ for } \ell \text{ pair}
$$
\n
$$
\dot{y} = \sum \dot{y} \frac{1}{j} / \ell
$$
\n
$$
\dot{z} = \sum \dot{z} \frac{1}{j} / \ell
$$

D. Compute elements

$$
\underline{r} = \sqrt{x^2 + y^2 + z^2}
$$

$$
V = \sqrt{\dot{x}^2 + \dot{y}^2 + \dot{z}^2}
$$

 \sim

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$$
\begin{aligned}\n\dot{\underline{r}} &= (x \dot{x} + y \dot{y} + z \dot{z}) / \underline{r} \\
C_5 &= \left[2(631.353746)^2 / \underline{r} \right] \cdot V^2 \\
a &= ((631.353746)^2 / C_5)/6378.174 \\
P_n &= .058672947 (a) \frac{3}{2} \\
e &= \sqrt{1 - (V^2 - \underline{r}^2) \underline{r}^2 / 631.353746^2} \\
p &= a (1 - e)\n\end{aligned}
$$
\n
$$
Q = a (1 + e)
$$
\n
$$
Q = a (1 + e)
$$
\ni = [cos⁻¹ ((xy - yx)/631.353746 \sqrt{p})], (-90^{\circ} \le i \le + 90^{\circ}) \\
RA = tan⁻¹ (\frac{y \underline{z} - \underline{z} \underline{y}}{x \underline{z} - z \underline{x}}) \\
sin \mu = \frac{z}{\underline{r} \sin i}\n
$$
cos \mu = (x + (z \sin RA \cot i)/\underline{r} \cos RA \\
\mu = f_{\text{qual}}(cos \mu, \sin \mu)
$$
\n
$$
cos v = (p/\underline{r} - 1)/e \\
v = f_{\text{qual}}(cos v, \underline{r}) \\
\omega = [\mu - v], (0^{\circ} \le \omega < 360^{\circ}) \\
\tan \frac{E_1}{2} = \sqrt{\frac{1 - e}{1 + e}} \tan \frac{\dot{v}}{2} \\
M_1 = E_1 - e \sin E_1 \\
\Delta t_1 = M_1 \cdot P_n / 2\Pi \\
T_p = t_{\text{obs}} - \Delta t_1
$$

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$$
\tan \frac{E_2}{2} = \sqrt{\frac{1-e}{1+e}} \tan \frac{\omega}{2}
$$

$$
M_2 = E_2 - e \sin E_2
$$

$$
\Delta t_2 = M_2 \cdot P_n / 2\Pi
$$

$$
T_n = T_p - \Delta t_2
$$

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$$
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$$

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3. 7 Glossary

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Illustration of Relations between Quantities

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Fig. 3.2

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- QUAC(X, Y) places an angle in the proper quadrant $(0 \text{ to } 2\pi)$ after taking the absolute value of the arc sine of the angle, $(|\sin^{-1}Y|)$. The following arguments must be given: X **=** cosine of the angle $Y = sine of the angle$
- QUAL(X, Y) places an angle in the proper quadrant (0 to 2π) after taking the arc cosine of the angle, $(cos⁻¹X)$. The following arguments must be given: $X = \text{cosine of the angle}$ $Y = sine of the angle$
- $QUAT(Y, Z)$ places an angle (the eccentric anomaly) in the proper quadrant (0 to 2π) after taking the arc tangent of the angle times two, $(2 \tan^{-1} Z)$. The following arguments must be given: $Y = sine of the true anomaly (v)$ $Z = \text{tangent of the eccentric anomaly over two, } (\tan \frac{E}{Z})$
- SMITH(X) retrieves θ_0 for the year of computation, given the argument X which is the last digit of the year.

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4. 1 LOCVEC, Vector Coordinates for Lockheed

4. 2 Function

The program computes the predicted position and velocity for Lockheed from nodal element sets. The output is written on magnetic tape for printing and for punching five-channel paper tape for teletype transmission. The teletype data is received at Sunnyvale Tracking Center where it becomes binary input for a **C D** C 1604 computer.

4. 3 Input

The input data is ordered as follows:

- (1) Up to five hundred request cards
- (2) A card with a numeric punch (1-9) in col. 8. This indicates the end of the request deck.
- (3) Standard six or seven card element sets of the satellites in the request deck or all satellites (up to 500 sets).
- (4) A blank card which signals the termination of the element deck.

The elements sets do not have to be in the same order as the satellites appearing in the request deck.

The following information should be included on the request card:

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4.4 Outp

4.4.1 Printed **C**

There is one output line for each satellite. The order is the same as that of the request deck. Each line contains: (a) the satellite number; (b) the Sunnyvale Tracking Center number or, if preceded by a nine (9), the SPADATS number; (c) the element number; (d) the month, day, hour of epoch; (e) seconds since the start of the epoch month; (f) epoch revolution number; (g) the x, y, z coordinates in feet, and the velocity components in feet per second.

4.4.2 Teletype

The basic message consists of a header and thirty-one words of five-level teletype punch. A visual header and a listable teletype header are included at the start of each message. The visual portion identifies the tape as a SPADATS vector with the characters "SP" and the four digit vehicle number, The listable header contains the word "SPADATS" followed by the vehicle number, the month, day and the nearest GMT hour of the last update of vectors. Tape contents and format are presented on the following pages.

4.5 Processing

The request cards are read into an array. This array is checked as each element set is read into core, and the element sets, for which vectors are required, are stored.

The nodal elements to be used are converted to **N,** M, sets as required by the Analytical Integration Routine (XYZSB). The required position and velocity components are obtained from this routine.

LOCVEC converts the position coordinates to feet and the velocity components to feet per second. These quantities are then written on the output tape for off-line printing through the UBC. The conversion to the required teletype format, described in the output section, is then achieved and written on the tape for off-line punching on five-channel paper tape. This teletype data becomes input for a CDC 1604 computer at the Sunnyvale Tracking Center.

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TAPE CONTENTS AND FORMAT

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- include the parity bits. The check sum is formed by tabulating all 24 bit words on the tape and adding all bits. The overflow beyond the 24th bit is folded under and added to the previous sum to get the final check sum.
- Note 2. Words 2 and 3 give the identification number. Any 8 legal BCD characters are permissible. It is suggested that these words be used to indicate the Spacetrack object designation (i.e. 1961 **A1** for 1961 Alpha 1,or 1960 B 2 for 1960 Beta 2).
- Note 3. The need for more than one vector at any time for one vehicle is not apparent in this case. Therefore, these 3 bits should always indicate 1. **1()**

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- Note 4. Machine time is the number of seconds elapsed since 00 Hrs., GMT, on the first of the month.
- Note 5. In Figure **1** the D's indicate data bits and the P's indicate parity bits. Note that there is both horizontal and vertical odd parity. Note the repeating pattern of data word start bits. There must be a tape start header and a tape stop pattern. If there are 10 different vehicles on one physical length of tape there must be 10 tape start headers and 10 tape stop patterns, one for each vehicle. The tape must be transmitted three times in succession.

LOCVEC

Wolf Research and Development Corporation

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Fig. 4.2

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Satellite No., Sunnyvale sat. no., element no., Output: $t_{\text{day of month'}}$, t_{hour} , t_{secs} , N_o , $x, y, z, \dot{x}, \dot{y}, \dot{z}$

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4.7 Glossary

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Other terms may be found in description of the subroutine XYZSB.

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5.1 CCOE, Cartesian. Coordinates From Orbital Elements

55,2 Function

From a given element set, this program calculates the Cartesian coordinates x, y, z and the components of velocity $\frac{dx}{dt}$, $\frac{dy}{dt}$, $\frac{dz}{dt}$ of a satellite for a specified length of time a given intervals. The output can be expressed either in kilometers and kilometers per second, or in earth radii and earth radii per minute.

5..3 Input

The data deck is comprised of: 1) a standard 6 or 7 card element set, 2) as many request cards as needed and, 3) a blank card. This sequence is repeated for all satellites requiring this computation. The final sequence should be followed by a second blank card to terminate the program.

The following information should be punched on the request card:

¹**-** output in earth radii

-Decimal points may be punched anywhere in each of the first three fields.

5,4 Output

The output consists of Cartesian coordinates and their related components of velocity for the time increments requested. The dimensions given are in kilometers and kilometers per second, or earth radii and earth radii per minute, depending on the output option chosen. The sentinels required by the TELTYP program are included so that conversion to teletype tape is possible.

5.5 Processing

After reading a standard six or seven card element set and a request card, the variables x, y, z, x, y, and z are calculated and printed according to the specified output option. The time is incremented by the time step specified on the request card. Unless the stop time has been exceeded, the program computes new variables for the incremented time.

When the stop time specified on a request card has been exceeded, the program assumes that the next card is a request card for the same satellite for another time interval. If the card is blank, the program returns to the element read section for another standard element set. A return to the executive routine is made, if a second blank is encountered. Otherwise the requests for the new satellite are processed as above.

5.5.1 Error Messages

- 1. NEGATIVE NODAL PERIOD, SATNO ______. Reading continues until next case is found
- 2. ELEMENT CARDS OUT OF ORDER. Same procedure as under (1)
- 3. MISSING NODAL PERIOD. Continues processing that case. P_a is used for P_n .
- 4. SUBROUTINE ERROR EXIT FROM OCTAL___ Subroutine or irrecoverable input output error. Exits to system if GO option taken. The message will be repeated if the STOP option is used. A dump should be given if the computer is in a non-interruptable mode of operation.

5.6 **Formulation**
\n1.
$$
P_a = \frac{360 P_o}{360 - P_o d}
$$

\n2. $a_o = \begin{bmatrix} P_a \\ .058672947 \end{bmatrix}$
\n3. $t = T_I - t_o$
\n4. $xx = \frac{t}{P_o}$

5. $N =$ Integer part of xx

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where $\hat{r} = \hat{r}^{\dagger}_{\text{max}}$

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26. $b_y = a(N) \sqrt{1 - e^2(N)} (-\sin \omega \cdot \sin \Omega + \cos \omega \cdot \cos \Omega \cdot \cos i)$

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6. 1 RESPLT, Residual Plot

6. **²**Function

RESPLT reduces observations against the N, M element sets and produces punched card output which can be used on the EAI Data Plotter. The residuals which can be plotted are: **(1)** the difference in time between the predicted and observed positions versus the revolution and, (2) the vector magnitude difference versus the revolution.

The plots are especially useful as an aid in the analysis of element sets. Large residuals may indicate that the current element sets need updating or that a piece has broken off the satellite. Maneuvers,not otherwise easily detected, may be revealed when these residuals are presented graphic form (cf. Fig. 6. 2).

6. 3 Input

The program requires observation data to be entered as card input. Element data and sensor data may be entered as card input or read from the SEAl tape.

The observations must be obtained from standard observation cards which have previously been sorted by satellite number. These

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cards must have an association status punched in column 80 (cf. Fig. 6.6). If untagged observations (satellite number = zero) are to be reduced against a particular element set, a standard observation card containing the number of the desired element set and a blank sensor number should be placed in the observation deck preceding the untagged observations. The number of observation cards allowed is limited only by the space allocated to OBLOC.

There are six input options which may be specified on the SPSJOB card. Each option specifies the source of the sensor and element data, i. e., input cards or SEA1 tape.

Fig. 6. **3**

Element data may be obtained from standard 7 card element sets with an E in column 80. Since the executive routine reads the element sets into EBLOC, the number of element sets is limited by the size of that block. Element data may be obtained from the SEAl tape if element sets are not included in the input deck.

Sensor data may be obtained from standard sensor cards with an S in column 80. The number is restricted only by the size of SBLOC. Sensor data may be obtained from the SEAl tape if no sensor cards are included in the input deck.

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6.4 Output

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Printed output consists of the reduced observations sorted by revolution number within a satellite number in groups of 100. A heading line, which describes the output quantities in the columns below, is printed for each satellite. The satellite number, observation number, revolution number, time in minutes since epoch, vector magnitude in kilometers, number of revolutions since epoch, element number, association indicator, and station number are printed for every revolution number.

The punched card output can be used to plot residuals on the EAI Data Plotter using $K + E$ 10 x 10 cm. graph paper. The pen command for the Data Plotter is contained in column 55. The data deck is divided up as follows:

- **0** 1) A card with a command to stop the plotter (7 in col. 55). The operator can set the origin at this time.
- 2) A set of cards which will enable the plotter to draw the x and y axes (6 cards).
- 3) A card which will stop the plotter so that the above can be tried again if necessary.
- 4) A set of cards which will plot the visual characters S, **E,** R, T and the information required to identify the plots. S will be followed by the satellite number, E will be followed by the element number, R will be followed by the epoch revolution number, and T will be followed by the the time of epoch in days.
- 5) A card which will stop the plotter (7 in col. 55).
- 6) The data cards containing the residual information to be plotted. There is one card for each line of printed output.

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IBM 523 Summary Punch

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6. 4. 1 EAI Data Plotter Operating Instructions **4** ¹

An IBM 523 Summary Punch is used to read the punched cards for input to the plotter. Board wiring diagrams are included for each of the two possible graphs (cf. Fig. 6. 4 and Fig. 6. 5).

The residuals which may be plotted are data time versus data revolution and vector magnitude versus data revolution. Coordinates and pen commands for both graphs appear on the same data cards.

Note that it is possible to plot both graphs on the same piece of graph paper by reading the data cards through the **523** Summary Punch twice. This is accomplished by reading all of the punched cards for the first graph. The axes and visible information as well as the points on the graph will be plotted. If part six (6) of the data deck is reread, the second set of the residuals may be plotted on the same graph paper.

The plotter pen color and **523** board must be changed before rereading the data cards. Two separate graphs may be produced by reading all punched cards twice and changing the graph paper and **523** board.

For the proper plotting of all data, the origin on the graph paper should be five (5) centimeters right and nine (9) centimeters up from the bottom left corner of the graph paper.

Graph scaling is as follows. The revolution number from epoch is the x-axis for both graphs. The epoch revolution number lies at $x = 0$. Each centimeter represents 20 revolutions, so that the possible values for plotting lie in the range of -100 and + 400 revolutions. Vector magnitude in kilometers is the y-axis for one graph. Each centimeter represents 500 kilometers and the possible range of values is 0 to 4500 kilometers. Delta time in

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minutes is the y-axis for the other graph. Each centimeter represents five (5) minutes and the possible range of values is -45 to +45 minutes.

6. 5 Processing

The program reduces observations against element sets to obtain time and vector magnitude residuals which may be plotted against the revolution number. Up to **100** observations may be processed, one at a time. The association status is checked to determine whether the observation is an angles observation only or a radar observation.

Association Status	Observation Kind	Association Category	Will RESPLT Accept?
	Radar	Associated	Yes
2	Radar	Doubtful	Yes
3	Radar	Unassociated	Yes
4	Angles Only	Associated	Yes
5	Angles Only	Doubtful	Yes
6	Angles Only	Unassociated	Yes
7	Range Rate	Associated	No.
8	Range Rate	Unassociated	No.
9	Radar	Special Unassoc.	Yes
0 or Δ			No.

Fig. 6.6

The proper year constants are obtained and time of epoch is converted to days since 1950. The observation counter is increased by one and the association indicator, observation number, and sensor number from the observation are stored for output. The sidereal time at the time of epoch is computed. The semi-latus rectum and the components of the unit vector perpendicular to the orbit plane are found to compute the inclination, the right ascension

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of the ascending node, the eccentricity, and the semi-major axis at epoch.

The mean argument of latitude, the mean angular motion, a drag coefficient, and the perifocal distance are computed. The time since epoch, the sidereal time at the station, and the x, y components of the station vector are found. The subroutine XYZSB is used to compute the predicted position and velocity plus intermediate quantities.

Observation type (OTYPE) will be referred to in subsequent paragraphs as illustrated in Fig. 6. 7

Fig. 6.7

If the observation is an angles only observation, the range is computed. If $\text{OTYPE} = 1$, the predicted azimuth and elevation are found. If OTYPE = 5, the predicted right ascension and declination are found.

If the observation is a radar observation, the predicted range is computed. The predicted azimuth and elevation are computed. If OTYPE = 3 or 4, the predicted range rate is computed.

For all observations, the residuals are computed to find the vector magnitude. Unless 100 observations have been processed, the program returns for the next observation.

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When all of the observations have been processed for a single satellite the program generates data for punched card output for the plotter. The punched cards contain all necessary plotter control functions. Punched cards are generated to plot the axes, visual satellite, element, epoch revolution, and epoch day numbers. These cards are followed by the data cards. The program then returns to process the next group of observations.

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O 6. 6 Formulation

1. $\theta_G = \theta_o + t_{days}$ (rotation of earth) deg/solar day

2. Compute semi-latus rectum.
\n
$$
P_0 = h_{xo}^2 + h_{yo}^2 + h_{zo}^2
$$

3. Compute x, y, z components of W (unit vector \perp orbit plane).

$$
W_{xo} = h_{xo} / \sqrt{p_0}
$$

$$
W_{yo} = h_{yo} / \sqrt{p_0}
$$

$$
W_{zo} = h_{zo} / \sqrt{p_0}
$$

4. Compute inclination.

$$
i = \pi/2 \text{ if } W_{ZO} = 0, \text{ otherwise;}
$$

$$
i = \tan^{-1} \left(\sqrt{1 - W_{ZO}}^2 / W_{ZO} \right)
$$

5. Compute right ascension of ascending node at epoch.

$$
\Omega_o = \tan^{-1} \left[\left(W_{xo} / \sin i \right) / \left(-W_{yo} / \sin i \right) \right]
$$

6. Compute eccentricity and semi-major axis at epoch.

$$
e_o = \sqrt{a_{xNo}^2 + a_{yNo}^2}
$$

 $a_o = p_o/1 - e_o^2$

7. Compute mean argument of latitude and mean angular motion at epoch.

$$
U_o = L_o - \Omega_o
$$
 if W_{zo} positive, otherwise $u_o = L_o + \Omega_o$
 $n_o = k_e/a_o^{3/2}$

8. Compute a drag coefficient, perifocal distance, and K_L $c'' = -c(n^2)360\pi^2$ $q_o = a_o (l - e_o)$

$$
K_{e}L_{so} = \frac{K_{e}I_{ae}^{2}}{P_{o}^{7/2}} \left[3 - 5e_{o}^{2} - (4 - \frac{27}{4}e_{o}^{2})\sin^{2} i - (1 - \frac{3}{2}e_{o}^{2})/\cos i/\right]
$$

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9. Compute time since epoch and sidereal time at station.

$$
T = (t_{obs} - t_0) 1440 + time of day in minutes since midnight.\n
$$
\theta_g = \theta_G + \lambda + .0043752691 T
$$
$$

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10. Compute observed position from the observation. \bar{x} = cos θ $\bar{x}/\cos \theta$

$$
\frac{\overline{Y}}{\underline{Y}} = \sin \theta_{\text{s}} \frac{\overline{X}}{\cos \theta_{\text{s}}}
$$

Use subroutine XYZSB to compute predicted position, r, and velocity, \mathbf{r} , plus intermediate quantities.

If angles only observations, go to 17.

Equations 11-16 for radar observations.

11. Compute x, y, z components obs. unit vector from observer to obj. with respect to horizon.

 L_{xh} = - cos Az cos h $L_{\rm yh}$ = sin Az cosh

 L_{ab} = \sinh

12. Compute x, y, z components S (unit vector from observer to south).

$$
S_{\mathbf{x}} = \sin \phi \cos \theta_{\mathbf{s}}
$$

$$
S_y = \sin \phi \sin \theta_g
$$

- S_{7} = $-cos \phi$
- 13. Compute x, y, z components Z (unit vector from observer to zenith).

 $Z_{\mathbf{x}}$ = cos ϕ cos $\theta_{\mathbf{x}}$ $Z_V = \cos \phi \sin \theta_g$

 Z_{7} = sin ϕ

14. Compute x, y, z components E (unit vector from obs. to east). $E_x = -\sin \theta_g$

 $E_y = \cos \theta_g$

 $E_z =$

15. Compute L_0 (obs. unit vector from obs. to object).

$$
\underline{\mathbf{L}} = \mathbf{L}_{\mathbf{x}\mathbf{h}} \underline{\mathbf{S}} + \mathbf{L}_{\mathbf{v}\mathbf{h}} \underline{\mathbf{E}} + \mathbf{L}_{\mathbf{z}\mathbf{h}} \underline{\mathbf{Z}}
$$

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16. Compute ob. r, vector directed to object, and **P.**

$$
\underline{\mathbf{r}} = \rho \underline{\mathbf{L}} - \underline{\mathbf{R}}
$$

$$
\mathbf{r} = \sqrt{\mathbf{x}^2 + \mathbf{y}^2 + \mathbf{z}^2}
$$

$$
\beta = |\sin^{-1}(\frac{1}{\mathbf{r}} (\underline{\mathbf{R}} \cdot \underline{\mathbf{W}}))|
$$

Go to 30.

Equations 17-29 for angles only observations.

17. Compute x, y, z components of range vector p.

$$
\rho = \underline{r} + \underline{R}
$$

\n
$$
\rho_c = \sqrt{\rho_x^2 + \rho_y^2 + \rho_z^2}
$$

\nIf Az-E1 type observation, go to 18.

If RA-Dec type observation, go to 20.

18. Compute predicted elevation.

$$
L_{xh} = \frac{\rho_x \cos \theta_s \sin \phi + \rho_z \sin \theta_s \sin \phi - \rho_z \cos \phi}{\rho}
$$

$$
L_{yh} = \frac{-\rho_x \sin \theta_s + \rho_y \cos \theta_s}{\rho}
$$

$$
L_{zh} = \frac{\rho_x \cos \theta_s \cos \phi + \rho_z \sin \theta_s \cos \phi + \rho_z \sin \phi}{\rho}
$$

$$
h_c = \tan^{-1} \left(\frac{L_{zh}}{\sqrt{1 - L_{zh}^2}} \right)
$$

19. Compute predicted azimuth.

$$
A z_c = \tan^{-1} \left(\frac{L_{\rm yh}}{-L_{\rm xh}} \right)
$$

Go to 22.

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20. Compute predicted right ascension.

$$
\underline{L} = \rho / \rho
$$

$$
a_{c} = \tan^{-1} \left(\frac{L}{L_{\chi}} \right)
$$

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21. Compute predicted declination.

$$
\delta_c = \tan^{-1}\left(\frac{L_z}{1 - L_z^2}\right)
$$

Go to 27.

22. Compute x, y, z components obs. unit vector from observer to object with respect to horizon.

 L_{xh} = cos δ cos α $L_{vh} = \cos \delta \sin \alpha$ $L_{\rm zh}$ = $\sin \delta$

- 23-26. Same as equations 12-15. Go to **28.**
	- 27. Compute x, y, z components \overline{L}_0 .

 L_x = cos δ cos α $L_y = \cos \delta \sin \delta$ L **=** sin6

z

28. Compute observation range.

$$
\rho = (X \cdot W_{xo} + Y \cdot W_{yo} + Z \cdot W_{zo})/(L_x \cdot W_{xo} + L_y \cdot W_{yo} + L_z \cdot W_{zo})
$$

29. Compute predicted range.

$$
\rho_c = \sqrt{\rho_x^2 + \rho_y^2 + \rho_z^2}
$$

30. Compute Δu , change in mean argument of latitude.

$$
\Delta u = \tan^{-1}\left(\frac{U}{\frac{U}{x} \cdot W}\right)
$$

31. Compute Δt . $\Delta t = (\Delta u \underline{R}^2)/(K_e \sqrt{p})$

If angles only observation, go to 36.

- 32. Same as equation 17.
- 33-34. Same as equations 18-19.

35. If Az-El, Range Rate Obs. (OTYPE = 4), compute predicted range rate.

 $\frac{1}{X} = -\frac{1}{Y} \dot{\theta}$ $\frac{1}{\underline{Y}} = \frac{1}{X} \dot{\theta}$

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 $\overline{z} = 0$ $\dot{\rho}_c = \underline{L} \cdot (\dot{\rho} + \dot{R})$ 36. Compute range residual. Δ **r** = $\rho - \rho_c$ 37. If Az-El type observation, (OTYPE \neq 5). $\Delta h = h - h_c$, $\Delta a = |Az - Az_c|$ If RA-Dec type observation, (OTYPE = 5).
 $\Delta h = \delta - \delta_c$, $\Delta a = |a - a_c|$ 38. If $\Delta h > \pi$, $\Delta h = 2\pi - \Delta h$ If $\Delta a > \pi$, $\Delta a = 2 \pi - \Delta a$ If angles only observation, go to 40. 39. Compute vector magnitude. $V_{mag} = \sqrt{\Delta r^2 + (\rho \Delta h)^2 + (\rho \Delta a \cos h)^2}$ Go to 41. 40. $V_{\text{mag}} = \sqrt{(\rho_c \Delta h)^2 + (\rho_c \Delta a \cos h)^2}$ 41. Compute revolution and delta revolution number. $N = N_0 + \frac{t}{2\pi}$ t **AN=n** 21r Process next observation.

6. 7 File Definitions

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 $\label{eq:2.1} \frac{1}{\sqrt{2}}\int_{0}^{\infty}\frac{1}{\sqrt{2\pi}}\left(\frac{1}{\sqrt{2\pi}}\right)^{2\alpha} \frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac{1}{\sqrt{2\pi}}\frac$

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Empty Output Buffers and Exit to Exec.

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7. 1 PREPINT, Satellite Situation Report From Nodal Elements

7. **2** Function

The purpose of PREPINT is to supply the sub-satellite point and related data of all satellites at specified times. The longitude and time at the last ascending node are also computed for each satellite.

A maximum of 500 satellites may be included in each report. However, positions will be computed only for those satellites included in the request deck (See 7. 3). If any comments (such as "IN HELIOCENTRIC ORBIT") appear on the request card of a satellite, these comments are printed out and the program does not attempt to determine the sub-satellite point.

During one run of PREPINT, as many as **10** reports may be produced, one report for each time card (See 7. 3). If a report time is 1200 hours, the output heading will be WEEKLY SATELLITE SITUATION REPORT. A daily report is issued at all other times but both types of report are identical other than heading.

7.3 Input

The input to PREPINT is arranged as follows: **1)** request cards $(500); 2) request deck termination; 3) time cards $(10);$$ 4) time deck terminator; 5) standard 6 or 7 card element sets; and 6) element deck terminator.

(1) The request cards contain the following information:

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Cols. 9-20 Satellite name

Cols. **21-62** Comments to be printed on report if Col.7 contains a "1" punch. (A maximum of 50 satellites may have comments.)

All other columns on this card are not used.

(2) The request deck is terminated by a card with a " 2" punch in Col. 7

(3) The time cards contain the times at which reports are desired. They are punched as follows:

> Cols. 1-4 Year Cols. 6-7 Month number Col. 8 Blank Cols. 9-10 Day of month Cols. $11-17$ Hours and minutes (HHMM \cdot MM) Col. 20 Output unit indicator = 0, output in statute mil **⁼**1, output in kilometers

All other columns are not used.

(4) The time deck is terminated by a card with a " 7" punch in Col. **8.**

(5) Both standard 6 card element sets and standard 7 card element sets may be included in the element deck.

(6) A blank card is used as an end of input indicator.

The ordering of the satellites on the output is determined by the order of the request deck, not by the element set order. More than 500 sets of elements may be read in since the program stores only the elements for satellites in the request deck.

7.4 Output

The heading of each satellite situation report contains the date and time for the information following it. One line of data is

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printed for each satellite. If comments were read from the request card, these comments are printed out beside the satellite number. Otherwise, the following information is printed out for each satellite:

1) Identifying information: satellite name and number, and element set number.

2) Sub-satellite point at report time: latitude and longitude west, in degrees.

3) Inclination, i, in degrees and nodal period, $P_{N'}$, in minutes.

4) Distances to apogee and perigee in statute miles or kilometers, depending on Col. 20 of the time card.

5) Revolution number at report time and $T_{\mathbf{N}}, \Omega_{\mathbf{N}},$ and $\lambda_{\mathbf{N}}$ for this revolution.

6) Eccentricity, e, satellite height in statute miles or kilometers.

7) Satellite latitude and longitude west in tenths of a degree.

The sentinels required by the TELTYP Program are supplied to enable transmission of the above output including all information from **1)** to 4).

7. 5 Processing

The entire input deck is read in and the data stored in arrays before any computations are made. First, a maximum of 500 request cards are read in. The satellite identification and comments, if any, from the request deck are saved. Next, the report times and output options are read from the time cards and stored. Finally, the element deck is read in and the elements for all satellites in the request deck are stored. For all element sets which are out of order, the satellite and card numbers are printed out via the flexowriter with an appropriate message. After the element deck terminator is read in, a check is made to see if there were any

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element sets out of order. If there were cards out of order, the $\| \cdot \|$ program returns control to the executive program. Otherwise, computations for the first report are started.

After retrieving the first report time from the time array, the teletype sentinels and report headings are written on the output tape. Next, the first satellite number is retrieved from the request deck array. If comments appeared on the request card for this satellite, they are written on the output tape and the next satellite number is picked up from the request deck array. For satellites having no comments, the elements are retrieved from the element array for computation of the sub-satellite point at report time. First, the revolution number at report time is computed and all elements are updated to the time of the ascending node for this revolution. If the satellite has decayed prior to report time, an appropriate message is written on the output tape and the next satellite number is retrieved from the request deck array. However, if the satellite is still in orbit, the sub-satellite point at report time and other data for output are computed. The output is converted to the proper units and written on the output tape. The next satellite number is retrieved from the request deck array and the computations continue until all requested satellites have been processed.

The next report time is retrieved from the time array and the same procedure followed until reports have been completed for all requested times. The program then returns control to the executive program.

7. **5.** 1 Error Messages

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1. SATELLITE CARD OUT OF ORDER

Program continues reading entire element deck and (writes this message for every card out of order. After all elements have been read in, program exits to executive routine.

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2. SUBROUTINE ERROR EXIT FROM OCTAL

Subroutine or irrecoverable input-output error. Program exits to executive routine if the GO option is taken. If the STOP option is specified, the message is retyped. A dump should be taken if possible.

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$$
E_N = -2 \tan^{-1} \left(\sqrt{\frac{1 - \theta N}{1 + \theta N}} \tan \frac{\omega N}{2} \right)
$$

\n18. $M_N = E_N - \theta_N \sin E_N$
\n19. $M(t) = M_N + \frac{2 \Pi \cdot \Delta t}{P_N}$
\n20. Solve for $E(t)$: $M(t) = E(t) - \theta_N \sin E(t)$
\n21. $v = 2 \tan^{-1} \left(\sqrt{\frac{1 + \theta N}{1 - \theta N}} \tan \frac{E(t)}{2} \right)$
\n22. $\mu = v + \omega_N + \omega_N \Delta t + \frac{1}{2} \omega_N \Delta t^2$
\n23. $\phi' = \sin^{-1} (\sin i \cdot \sin \mu)$
\n24. $\phi = \tan^{-1} \left(\frac{\tan \phi'}{\sqrt{9329985}} \right)$
\n25. $R_0 = \frac{.9966443}{\sqrt{1 - .00670015 \cdot \cos^2 \phi}}$
\n26. $r = \frac{a_N (1 - \theta_N^2)}{(1 + \theta_N)^{\cos V}}$
\n27. $H = r - R_0$
\n28. $\beta = \sin^{-1} \left(\frac{\cos i}{\cos \phi} \right)$
\n29. $\Delta \lambda' = \sin^{-1} \left(\frac{\tan \phi'}{\tan i} \right)$
\n30. $\Delta \lambda = \Delta \lambda' + \Delta t (\Omega_N - 6.3003883) + \frac{1}{2} \Omega_N \Delta t^2$
\n31. $\lambda_N = \theta_G - \Omega_N$, $0 \le \lambda_N < 360$
\n32. $\lambda_S = \lambda_N - \Delta \lambda$, $0 \le \lambda_S < 360$
\n33. $\text{Apose} = (2a_N - q_N - 1) \cdot K \text{ where K is km/ e.r. or sm/ e.r.}$ conversion factor
\nPerigee = $(q_N - 1) \cdot K$
\n $\frac{[107.08829)(1 + \theta_N \cos v)}{\sin \beta} \left(\frac{6.3003883 \cos \omega \cos^2 \omega}{1 + \theta_N \cos^2 \omega} \$

Course = tan⁻¹ $\frac{(107.08829)(1+e_N \cos v) \sin \beta - (6.3003883 \cos \phi' \sin(1-e_N^2))}{\sqrt{\frac{a_N(1-e_N^2)}{a_N(1-e_N^2)}}}$ 34.

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7.7 Glossary

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PERIOD Nodal period in minutes PNOD Array for storage of nodal periods (P_{N_0}) PNODL $P_{N_{\alpha}}$ Nodal period in days/rev. PRDDO **Array for storage of 1/2** ω PRDDOT $1/2 \ddot{\omega}$ One half 2nd derivative of argument of perigee Q q Perigee distance in earth radii QA Array for storage of q values QDDO Δ Array for storage of $1/2$ q' values $QDDOT$ 1/2 q Second derivative of argument of perigee QDO Array for storage of q values QDOT q First derivative of argument of perigee R $\triangle N$ Revolutions since epoch RADDO $\begin{array}{c} \text{Array for storage of } 1/2 \Omega \text{ values} \end{array}$ RADDOT $1/2 \Omega$ One half 2nd derivative of right ascension of the ascending node RADOT **Array for storage of** Ω RASDOT Ω First derivative of right ascension of the ascending node RAX 57.2957195 (deg/rad) RIGHT Right ascension of ascending node in degrees RITAS Array for storage of **0** values RITASC Ω Right ascension of the ascending node RJ r Distance from center of earth to satellite RO R_0 Radius of earth at subsatellite point **SNAME 1** Satellite name **SNAME2** 1 SNAME A^l ^{Arrays} for storage of satellite names **SNAME** Bj \texttt{SOLMN} M_N Mean anomaly at the node SOLMS M(t) Mean anomaly at time of report SUBLN λ_N Longitude of the node in degrees TIM Time of report in hours and minutes TIME Array of times at which reports are to be is sued TME Time of report in hours and minutes TIZ Year of report TN t_N Time of nodal crossing for revolution number of report

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8. **1** MAKETAPE, Make Input Tape for TELTYP

8. **2** Function

The program MAKETAPE produces a magnetic output tape incorporating the message sentinels required by the TELTYP program. Only one message is produced by the program. An optional control feature will produce the message broken into 90-line segments.

8. 3 Input

Input data originates from the Schedule Tape. This input is moved by the system from the Schedule Tape to the System Data Tape (logical 0).

Two control cards are used in addition to the data cards containing the message to be converted. The following should be in cols. 17-24:

1. TELEFORM

2. FINDATA

8. 3. 1 Description of Control Cards

1. TELEFORM

This card, if present, will precede the data deck. When encountered, it will signal the program to break the message into 90-line segments.

2. FINDATA

This must be the last card of the data deck. The program exits to the executive program when this card is intercepted.

8.4 Output

The output tape (logical **1i)** created by this program is in a form acceptable for subsequent conversion by the TELTYP program. A hard copy of the message with its sentinels may be obtained by printing through the UBC using data select one. For a discussion of the sentinels used, the writeup of the TELTYP program should be consulted.

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8.5 Processing

MAKETAPE reads data into core from logical tape zero. Each card image is scanned for the control information described under the input section. If TELEFORM is found, an internal switch is set to cause line counting. If the program is in the line counting mode, each group of 90 lines will be preceded by the begin sentinel and followed by the end sentinel required by the TELTYP program.

All data read into core will be written onto the output tape, with the exception of right-adjusted blank fields. The program will exit to the executive routine when the FINDATA control card is intercepted.

It is generally expected that the TELTYP program will be used immediately following this program or, at least, before logical tape eleven is wrapped up by the system.

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Reads Next Input Card

9.1 XYZIAR, Look Angle Report From x, y, z, Coordinates

9.2 Function

The program is used for predicting the position of a satellite in terms of the tracking coordinates of a particular station. The program is especially useful for deep space probes.

The position of the satellite is defined in terms of the right ascension, declination, azimuth, elevation, and slant range at the prediction time, for a particular station. The elevation and illumination angles of the sun are also computed to determine if the satellite is visible.

Input options exist to consider the restrictions imposed by the type of observing equipment used. No data will appear for a requested prediction time unless these limitations have been met.

The ephemeris data, used in the calculations of the look angles, may be read from an ephemeris tape previously written by either the Unified Encke Differential Correction Program (accuracy, approximately ten lunar distances), or by the Interplanetary Program. If the ephemeris tapes are not available, the data may be read in from punched cards. Predictions for more than one station may be based on the same ephemeris data.

9. 3 Input

An input set consists of a standard station card for the observing station, a request card, and the ephemeris information from an ephemeris tape or punched cards containing the ephemeris data followed by a blank card. Each input set results in a look angle schedule for the station specified. If look angle schedules are required for the same satellite but different observing stations, additional pairs of cards, composed of a station card and a request card, may be added to the input deck. If new ephemeris data is to be entered from punched cards, the station card of the next input set must contain a negative station number. Any number of input sets may be entered. The last set must be followed by a blank card which is in addition to the blank card following the ephemeris data cards.

The format of the request card follows: columns 1 to 4 contain the year of the desired predictions; columns 5 to 7 contain the day of year; columns 8 and 9 contain the hour; columns 10 and 11 contain the minutes; columns 12 to 17 contain the base, or starting time, for the desired predictions.

The time increment used is determined by the ephemeris data. Columns 25 to 34 contain the maximum time increment or the time range for predictions. This time is expressed in minutes from the base time of the request. Any ephemeris data beyond this time range will be ignored. If columns 25 to 34 are blank, the program will calculate predictions for all ephemeris data supplied. Program option switches are specified in columns 44 to 47. A 1 punch in the respective column will set an internal switch. A 1 punch in column 44 indicates that the ephemeris data will come from the ephemeris tape for this set of input. If column 44 is blank the program expects ephemeris data from ephemeris cards. A 1 punch in column 45 indicates that only visible passes are desired. If a 1 is punched in column 46, negative elevations will be acceptable. A 1 is punched in column 47 if punched ephemeris cards are desired as output.

9.3.1 Ephemeris Cards

The ephemeris cards contain the time increment from the base time specified on the request card and the inertial geocentric coordinates of the satellite. All data is in floating point. Columns 1 to 14 contain the time increment, columns 15 to 28 contain the x-coordinate, columns 29 to 42 contain the y- coordinate, and columns 43 to 56 contain the z-coordinate in earth radii. Ephemeris cards are orderedby increasing time increment. A satellite identification card with the alphanumeric satellite name punched in columns I to 16 must precede the ephemeris cards.

9.3.2 Ephemeris Tape

The ephemeris tape can be obtained from two sources: 1) the Unified Encke Differential Correction Program and; 2) the Interplanetary Program. Both programs were written by Aeronutronic a division of the Ford Motor Company. In the former case, logical tape 10 is the desired tape and in the latter, the required tape is logical 9. The format for both tapes follows:

- Block 1. Contains the alphanumeric satellite name in the first two words.
- Block 2 thru N. Contains a time increment from the base time and x, y, z, \dot{x} , \dot{y} , \dot{z} in the inertial geocentric coordinate system at that particular time. Each value is contained in one full computer word and is in the floating point format. Eighteen groups of these seven values plus two zero words make up each one hundred and twenty-eight word block. The tape is considered to be terminated when the first word of a seven word set is filled with Z's.

9.4 Output

If a 1 is punched in column 47 of the XYZLAR request card, the ephemeris data will be punched on cards in the same format as the input ephemeris data when the binary ephemeris tape is used for input. The satellite name appears on the first card.

Printed output contains all TELTYP control functions, therefore, transmission of the output is possible. The first line is the alphanumeric satellite name. The second line is the comment LOOK ANGLES FOR followed by the alphanumeric station name. The third and fourth lines are heading lines describing the output data for each prediction time. Each data line contains the day of year, hour, minute, and fraction of minute of the search point, the predicted right ascension, declination, azimuth, and elevation in degrees and slant range in kilometers. These quantities define the position of the satellite. The elevation and illumination angles of the sun complete the data line and determine if the satellite will be visible at the time of the prediction.

9.5 Processing

XYZLAR predicts the right ascension, declination, azimuth, and elevation angles for a given satellite as well as the elevation and illumination angles of the sun at the requested prediction time. Restrictions on the observing capability of the station, specified on the request card, are considered. Any restrictions that are not met will cause computations to cease for that particular prediction time,

and no printed output will appear. The program will begin processing the next ephemeris position. In addition, when and if the **47** maximum time increment is exceeded, the program will ignore the remaining ephemeris data of that input set. The next request is then processed.

If an input set utilizes ephemeris cards, all of the ephemeris data is written in binary on logical tape seven before processing begins. However, an input set may originally include an ephemeris tape mounted on logical seven and in that case, processing will begin immediately.

The program adds the specified base time from the Request card, and the time increment from the ephemeris data to obtain the time of the search point. The sidereal time at Greenwich at prediction time and the coordinates of the station in a fixed system are found in order to compute the slant range. The right ascension and declination of the satellite, the sidereal time at the station, the hour angle, zenith distance, and elevation angle are computed for the prediction time.

If negative elevation angles are acceptable, the program continues on to compute the azimuth. If, however, negative elevation angles are not acceptable, a test is made and if the angle is found to be negative the program returns to process the next group of data.

Visibility at prediction time is determined by computing the elevation and illumination angles of the sun. If the elevation angle is less than -4 degrees, and the illumination angle is greater than -4 degrees, then the satellite is visible. If only visible passes are required, these conditions must be met.

Output quantities at each prediction, or search point, time include the time, right ascension, declination, azimuth, elevation, and slant range of the satellite, and the elevation and illumination angles of the sun.

The program then processes the.next group of data. If the station number is positive, the program computes another search ephemeris for the same satellite in the manner stated above, but for this new station. If the station number is found to be negative new ephemeris information is assumed. If the station card is blank, control is returned to the executive program.

Punched card output, if specified, is under control of a PATH switch which eliminates multiple punching of the same ephemeris data. This switch also controls the production of an ephemeris tape from a set of ephemeris cards. This switch equals one (1) for the first time through the program for a particular set of ephemeris data, and two (2) for each subsequent pass through the program for this same data. It can be reset to one by reading a negative station number. Thus, ephemeris cards are punched or an ephemeris tape made only on the first pass through the program for that particular set of ephemeris data.

9.5.1 Error Message

If the program is unable to read the ephemeris tape, the comment EPHEMERIS TAPE TROUBLE will be written on the output tape. In addition, the comment PROGRAM TERMINATED DUE TO EPHEMERIS TAPE TROUBLE will appear on the Flexo, and control will return to the executive routine.

 \mathbf{C} 9. 6 Formulation **1.** $\phi' = \tan^{-1}(0.99329985 \tan \phi)$ 2. R = $[.9966443/(1-.00670015 cos² ϕ)^{1/2} + H/6378174] 6378.174$ 3. Compute station coordinates in rotating geocentric system. $x_{\text{stat}} = R \cos \lambda \cos \phi'$ y_{stat} = R sin λ cos ϕ' $z_{\text{stat}} = R \sin \phi'$ 4. Convert starting or base time to days and fractional days. t_a = days + hours /24 + minutes/1440 + secs/86400 5. Find time of computed point. $t_c = t_a + t_i$ If $t_i > t_{max}$ return to process next input set. 6. **r** = $x^2 + y^2 + z^2$ 7. $\theta_C = \theta_0 + .98564735 t_d + 360.985647 t_f$ $0 < \theta_C < 360^\circ$ 8. Compute station coordinates in inertial geocentric system \bar{x} = x_{stat} cos θ_G - y_{stat} sin θ_G \overline{y} = x_{stat} sin θ _G + y_{stat} cos θ _G \overline{z} = z_{stat} 9. $\rho = \left[(x-\overline{x})^2 + (y-\overline{y})^2 + (z-\overline{z})^2 \right]^{1/2}$ where x, y, and z are geocentric coordinates from the ephemeris data. 10. $\cos \alpha = (x - \overline{x})/[(x - \overline{x})^2 + (y - \overline{y})^2]^{1/2}$ $a = \tan^{-1} \left[(y - \overline{y})/(x - \overline{x}) \right]$ If cos a negative, $a = a + \pi$. Otherwise, $0 \le a \le 360^{\circ}$ 11. $\sin \delta = (z - \overline{z})/\rho$ $\delta = \sin^{-1} (\sin \delta)$

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- 29. Pick up next group of ephemeris data, Go to 5.

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9.7 Glossary

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i0. 1 POSE, Point Search Ephemeris

10.2 Function

When a radar picks up an unknown object, one may have for some reason, the suspicion that it could have been an unidentified satellite. If this is the case, one likes to search for it. All orbital elements, however, are completely unknown. Disregarding the recession of the line of the nodes and the precession of the line of apsides, which is permissible over one or two revolutions, once can expect that the object (if it should be a satellite) will come close to the same point in inertial space again, one revolution later.

The program computes a search ephemeris for a given station for this point in space.

10. 3 Input

The input consists of: (1) a request card, (2) a standard observation card, (3) one or more standard station cards and (4) an option card.

The information contained on the request card is as follows:

- Cols. 9-13 T-stop in minutes (DDDD.) the largest orbital period expected.
- Col. 14 Iout- the output option desired
	- $1 = \text{long form}$
	- $0 = short form$

(explained in output section)

A decimal point should be punched within fields one, two and three.

1-10-1

The option card is used to initiate the read-in of a new request deck or to terminate the program. If a **-1** is punched in cols. 1-4, the program will read in a request card, observation card, station card(s) and an option card. If a zero or blank field is read from the option card the program will be terminated.

10.4 Output

The output consists of a look-angle schedule for each station requested. The long form of output includes the following data:

- a) Station information: latitude, longitude and height
- b) Observed time in days, hours, minutes and seconds (GMT)
- c) Observed elevation, azimuth and slant range
- d) Value of θ_{C} sidereal time at Greenwich
- e) Value of $\theta_{\rm s}$ sidereal time at the station
- f) Distance of station from center of the earth (CAPR)
- g) Geocentric station latitude
- h) Station 's Cartesian coordinates where x and y are in the equatorial plane and x points to the Greenwich meridian.
- i) Right ascension and declination of station
- j) Cartesian coordinates of station 's position where x and **^y** are in the ecuatorial plane and x points to the vernal equinox.
- k) Cartesian coordinates of the point or satellite where x and y are in the equatorial plane and x points to the vernal equinox.

In addition to the above the short and long forms of output list the look-angle information:

- a) Time of crossing day, hour and minutes of Zebra time.
- b) Elevation and azimuth angles
- c) Slant range in kilometers

The sentinels required by the TELTYP program will be supplied to enable transmission of the short form, or look-angle section, of the output.

1-10-2

10.5 Processing

The first request card, observation card and one station card are read at the start of the program. The x, y , and z coordinates of the station and the object are then calculated. These values are written on the output tape if a long form of output is requested.

The assumption made is that the object will appear in the same place one revolution later. Therefore, the look-angle, or search ephemeris quantities are calculated for the time t, which is equal to the observation time plus the smallest orbital period expected. If the point is visible at the station the quantities are written out on tape.

The orbital time is then updated by the time increment and compared to the largest orbital time expected. If this period is within the limits, the search ephemeris is computed for this new time. The process continues until the time limits have been ex ceeded.

After the search ephemeris has been computed for the largest orbital period expected the next station card is read into core. If the station number is greater than zero, the look-angles for this same object are computed for this new station. A negative station number indicates that a new request deck is to be read into core and the processing initiated for another point in space. A blank or zero $\mathscr P$ station number terminates all processing and the program returns to the executive program.

10. 5. **1** Error Messages

1. THE SLANT RANGE IS MISSING.

The program exits to the executive routine.

2. SUBROUTINE ERROR EXIT FROM OCTAL Subroutine or irrecoverable input-output error. Exits to executive program if the GO option is taken. The message is retyped if the STOP option is specified. 9 A dump should be taken if possible.

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11. 1 ORPS, Orbital Plane Search

11. 2 Function

ORPS is used for predicting search times for observing satellites at a given observing station. Limitations of the observing station, imposed by the type of equipment used, are considered. For example, the results must meet the criteria for maximum slant range, minimum elevation angle, and visibility.

The program computes the XYZ search point coordinates of a satellite from which its position may be defined in terms of right ascension, declination, azimuth, elevation, and slant range. The elevation and illumination angles of the sun are also computed at the search point or prediction time.

11.3 Input

A set of input consists of a standard six or seven card element set, a standard station card for the observing station, and a request card. The set may be repeated as often as desired, but the last set must be followed by a blank card to terminate the program.

The format of the request card is as follows:

1) columns **1-10** contain the requested start time, in day of year, for predictional data for a given observing station and satellite; 2) columns 11-20 contain the time increment in minutes between search points; 3) columns 21-30 contain the stop time in day of year; 4) the search azimuth angle in degrees of the station is contained in columns 31-40; 5) columns 41-47 contain the maximum slant range in kilometers acceptable to the observing station; 6) columns 48-57 contain the minimum elevation angle in degrees acceptable to the observing station; 7) if only visible passes are desired, a one **(1)** is punched in column 58.

Quantities specified in the above paragraph depend upon the limitations of 'the equipment type used by the observing station.

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11.4 Output

ORPS may be used with the TELTYP program if teletype transmission of the output is desired. Therefore, all TELTYP control functions are included with the output.

The first line is the heading "ORBITAL PLANE SEARCH AND XYZ LOOK ANGLE PROGRAM". The next line is the satellite and element numbers of the satellite for which the predictions were requested. The next line is the comment "ORBITAL ELEMENTS". This indicates that the nodal elements used in the computations appear in the next three rows, five columns to a row. The elements that are printed out by row are the day of year of epoch, fractional day of epoch, nodal period at epoch, rate of change of period in days/ rev^2 at epoch, rate of change of c in days/ rev^3 , first and second derivatives of the nodal period, the right ascension of the ascending node at epoch, first and half second derivatives of the right ascension, the argument of perigee at epoch, first and half second derivatives of the **^C** argument of perigee, eccentricity, and the inclination.

The next three lines consist of the parameters specified on the parameter input card. The start time in days of year, the time increment in minutes, and the stop time in days of year are printed. The search azimuth of the station in degrees and the criteria for maximum slant range in kilometers and minimum elevation angle indegrees are printed.

The year constants are printed including the sideral time at the start of the year, the longitude of the sun,and the difference between the longitude of the sun, and the argument of perigee of the sun.

The next line is a heading line, and the station name, latitude, longitude, and height appear directly below. The satellite and element numbers are printed next.

The main output of the program consisting of the time of search point, predicted position, and visibility quantities is preceded by two heading lines. The output appears in order of increasing time of search point in increments specified bv the parameter input card.

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Each data line includes the year, month,day, hour, and minute of the search point, right ascension, declination, azimuth, elevation, and slant range of the satellite, and the elevation and illumination angles of the sun. Unless the specified criteria for maximum slant range, minimum elevation angle, and visibility are met, no output for the time of search point will apppear.

I1i. 4. **1** Comments with Output

1. PROGRAM ASSUMES PN = PA FOR SATELLITE NO. XXX. This comment indicates that the nodal period at epoch in days is missing from the sixth card of a seven card element set. The anomalistic period is used in its place.

2. ELEMENT CARDS ARE OUT OF ORDER FOR SAT NO XXXX ELEM NO XXXX. This comment indicates that a card was out of order or missing in an element set. The program returns to read another element set. SATELLITE XXX CARD XXX OUT OF ORDER appears simultaneously on the Flexowriter.

11.5 Processing

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ORPS predicts the right ascension, declination, azimuth and elevation angles for a given satellite and the elevation and illumination angles of the sun, in a specified time range, for a given station. The limitations of the observing station are considered. The program consists of two sections, the Orbital Plane Search section and the XYZ Look Angle section.

The Orbital Plane Search section computes the xyz point coordinates from the satellite elements and-station coordinates. The geocentric latitude of the station, difference in longitude from station to search point, longitude at the search point, radius of earth at station, and the x, y, z coordinates of the station in a rotating system are computed. The time of search point is converted from day of year to month, day, hour, and minute. The sidereal time at Greenwich at the search point time, siderealtime at the search point, right ascension of ascending node at search point, argument of latitude of satellite in the orbit plane, the argument of perigee at the prediction time, the true anaomaly, semi-major axis at epoch, anomalistic period at the prediction time, semi-major axis at the predicted time, eccentricity at the prediction time, and the distance of the satellite from the center of the earth are computed. Then the x, y, z search point coordinates are found. 1-11-3

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The XYZ Look Angle section defines the position of the satellite in terms of right ascension, declination, azimuth, elevation angle, and slant range from the x, y, z search point coordinates. The **x,** y, z coordinates of the station in a fixed system are found to compute the slant range. If the predicted slant range is greater than the maximum specified slant range, the program increments the time of search point and returns to the Orbital Plane Search section to compute the x, y, z search point coordinates for the incremented time. Otherwise, the right ascension of the satellite, declination of the satellite, sidereal time at the station, hour angle of the satellite, zenith angle, and elevation angle are computed for the search point or prediction time. If the specified minimum elevation angle is greater than the predicted elevation, the program increments the time of search point and returns to the Orbital Plane Search section. Otherwise, the azimuth is computed.

The program determines if the satellite will be visible at search point time by computing the longitude of the sun at the prediction time, right ascension, declination, and elevation of the sun, and the angle η (see illustration). If only visible passes are required, the program tests the elevation and illumination angles of the sun. If the former is less than -4 degrees and the latter is greater than -4 degrees than the pass will be visible. If not, the program increments the time and returns to the Orbital Plane Search section.

Output quantities at each search point time include the time, right ascension, declination, azimuth, elevation, and slant range of the satellite, and the elevation and illumination angle of the sun. The program then increments the time of search point, and returns to the Orbital Plane Search section unless the stop time has been exceeded.

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\n $\cos \mu = \sin \mu / \tan \mu$ Go to 17.
\n16. $\alpha_{st} = \theta_G + \lambda$
\n $\cos \mu = \cos (\alpha_{st} - \Omega_t) \cos \phi$
\n $\mu = \tan^{-1} (\sin \mu / \cos \mu)$
\n17. If $\cos \mu$ is negative $\mu_0 = \mu + \Pi$, otherwise $\mu_0 = [\mu]_{(0, 360)}$
\n18. $\omega_t = \omega + \omega (t_c - T_0) + \frac{1}{2} \omega (t_c - T_0)^2 \quad 0 \le \omega_t < 360^\circ$
\n19. $V = \mu_0 - \omega_t$ $0 \le V < 360^\circ$
\n20. $a_0 = (P_a / .058672947)^{2/3}$
\n21. $P_{at} = P_a + \dot{p} (t_c - T_0) + \frac{1}{2} \dot{p} (t_c - T_0)^2$
\n22. $a_t = (P_{at} / .058672947)^{2/3}$
\n23. $e_t = 1 - (a_0 / a_t) (1 - e)$
\n24. $r = a_t (1) - e_t^2 / (1 + e_t \omega sV)$
\n25. $x = 6378.145 \text{ r} (\cos \mu_0 \cos \Omega_t - \sin \mu_0 \sin \Omega_t \cos i)$
\n $y = 6378.145 \text{ r} (\cos \mu_0 \sin \Omega_t + \sin \mu_0 \cos \Omega_t \cos i)$
\n $z = 6378.145 \text{ r} (\sin \mu_0 \sin i)$
\n11. 6.2 Formulation - XYZ Look Angle Section
\n26. $\overline{x} = x_{stat} \cos \theta_G - y_{stat} \sin \theta_G$
\n $\overline{y} = x_{stat} \sin \theta_G + y_{stat} \cos \theta_G$
\n $\overline{z} = z_{stat}$
\n27. $\rho = \left\{ (x - \overline{x})^2 + y$

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\cos a = (x - \overline{x})/[(x - \overline{x})^2 + (y - \overline{y})^2]^{1/2}
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\n $a = \tan^{-1}[(y - \overline{y})/(x - \overline{x})]$
\n29. If $\cos a$ negative $a = a + \pi$, otherwise $a = [a]_{(0, 360)}$
\n30. $\sin \delta = (z - \overline{z})/[(x - \overline{x})^2 + (y - \overline{y})^2 + (z - \overline{z})^2]^{1/2}$
\n31. $\theta_{st} = \theta_{st} - a$ $0 \le HA_s < 360^\circ$
\n33. $\overline{z}_\beta = \cos^{-1}(\sin \phi \sin \delta + \cos \phi \cos \delta \cos HA_s)$
\n34. $h = \pi/2 - z_\beta$
\nIf $h_{\min} > h$, go to 47.
\n35. $\sin Az = (\cos \delta \sin HA_s)/\sin \beta$
\n $\cos Az = (-\sin \delta \cos \phi + \cos \delta \sin \cos HA_s)/\sin z_d$
\n $Az = \tan^{-1}(\sin Az/\cos Az)$
\nIf $\cos Az$ negative, $Az = Ax + \pi$ $0 \le Az < 360^\circ$
\n36. t_0 (t) = $L_0 + .98564735 t_a + 1.91665 \sin(.98564735 t_a - C_{14})$
\n37. $a_0 = t_0$ (t) -2. 46682 sin (2 t_0 (t))
\n38. $\delta_0 = \tan^{-1}(.4336608 \sin a_0)$
\n39. $\theta(t) = \theta_G + \lambda$
\n40. $h_0 = \sin^{-1}[\sin \phi \sin \delta_0 + \cos \phi \cos \phi \cos (\theta(t) - a_0)]$
\n41. $a_{gc} = \tan^{-1} (y/x)$ $0 \le a_{gc} < 360^\circ$
\n42. $SD_{gc} = \sin^{-1} (z/6378.145 \text{ r})$
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- 44. $I = \xi \eta$ Test illumination angle of sun (I). If I< -4 degrees, pass not visible, go to 47. If $I \geq 4$ degrees go to 46.
- 45. $I = \xi \eta$
- 46. Output time, α, δ, Az, h, ρ, h_Ω, I.
- 47. Increment time:

 $t_c = t_c + \Delta t$ If $t_c \leq t_e$ return to 8, Orbital Plane Search Section. Otherwise return to start of program to process next set.

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 $\label{eq:2.1} \frac{1}{\sqrt{2}}\int_{0}^{\infty}\frac{dx}{\sqrt{2\pi}}\,dx$

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12.1 ASUM, Observation Acquistion File Summary

12. 2 Function

This program will provide a complete listing of the current Observation Acquistion File.

12.3 Input

The input is comprised entirely of the information written on the A-File of the SEAI tape.

12.4 Output

The output consists of the sensor information required to obtain look angles for a given satellite. The output is sequential according to satellite number and includes the following:

Satellite number Sensor number Sensor name Pass code (all or visual) Format request (short or complete) Type (all, 3-point, Baker-Nunn.) Minimum azimuth Maximum azimuth Minimum elevation Maximum elevation Maximum range Grid, or step, size in minutes

12.5 Processing

The program reads the Sensor File to obtain the required sensor information. The Acquisition File is then located and decoded. Sensor information is retrieved from the sensor storage area when the need arises. If the information is not available the following message is printed:

NOT IN FILE

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The data is written onto the output tape under control of counters which allow fifty lines of output per page. Output is single spaced. It is in satellite number order and in the sensor number order called for by the Acquisition File. A. double space separates the satellites.

12. 5. 1 Error Messages

When tape reading difficulties occur the following message will appear on the flexo:

TAPE TROUBLE

TYPE GO TO RETRY JOB. TYPE STOP TO TERMINATE.

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Fig. iZ. i **1-12-3**

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 $\label{eq:1} \mathcal{L}_{\text{max}} = \frac{1}{\sqrt{2}} \sum_{i=1}^{n} \$

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Returns to Read More from A-File

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- 13.1 SSUM, Sensor File Summary
- 13. 2 Function

This program prepares a complete listing of the Sensor File.

13. 3 Input

The input is comprised entirely of the Sensor File w . it.en on the SEAI tape.

13.4 Output

A listing of the sensor data contained in the S-File is provided. The following information is included as part of the output:

Sensor number Sensor name Latitude in degrees Longitude in degrees Height in earth radii $X/cos \theta$ or $- (C+H) cos \theta$ in earth radii Z or $-(S + H) \sin \phi$ in earth radii Accuracy digit for azimuth, elevation and range Classification code Sensor type

13.5 Processing

The SEAl tape is read to locate the Sensor File. After locating this file the program reads 1386 words into core storage. If the ending sentinel is found before the buffer area is filled, the program will process as much information as it has been able to retrieve. If more information is still to be read from the Sensor File the program returns for a second pass after processing'the first buffer full of data.

Counters are used to provide for nine groups of five sensors for each page of printed output. Headings and a page count are written on the top of each page. The order of the S-File determines the order of the printout.

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- 14. **1** ISUM, Information File Summary
- 14. 2 Function

The program generates a listing of all or specified sections of the Information File of the SEAl tape. A "box score" or tally of the number of satellites still orbiting can be obtained.

14. 3 Input

The input consists primarily of the I-File of the SEAl tape. Control cards included in the Job deck contain the information required to set up the options available in the program. The following data is punched in cols. $1 - 3$:

- a) ALL used to obtain a complete listing of the I-File.
- b) BOX used to obtain a tally of the number of satellites orbiting.
- c) NNN a three digit satellite number for each satellite, required if the ALL option is not used.
- d) END this signals the end of the control data and is always required.

14.4 Output

The I-File of the SEAl tape is written for off-line printing. Information printed out includes:

Satellite number Satellite name Launch date Launch site Booster country Payload country

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14.5 Processing the contract of the contract o

The control information described in the input section is read from the system data tape (logical zero). If all satellites are requested, the entire I-File is read and prepared for output purposes. If individual satellites are requested, then the I-File is scanned and only the information desired is written onto the output tape. If a box-score of the orbiting satellites is called for, then this is written out following the printout of the I-File.

14.5. **1** Error Messages

If irrecoverable tape errors are encountered while reading the SEAI tape the following message is written on the Flexo:

PROGRAM TERMINATED DUE TO POOR BLOCK MARKS, SPROCKET ERROR, **SI** OR SZ ERROR, OR DISABLED UNIT ON LOGICAL 4.

A parity error will cause a single reread and a second failure will result in the program issuing a minus one read order with the following comment: "

MINUS ONE READ ORDER ISSUED TO OVERRIDE PARITY ERROR

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15. 1 TELTYP , Magnetic Output Tape to Teletype Tape Conversaon

15.2 Function

The TELTYP Program is used to convert an output tape, written by other programs, to Baudot code. The magnetic output tape used as input to this program must contain the sentinels described in the input section. These sentinels are used to identify the messages to be converted by the TELTYP program. The output tape is searched for the beginning sentinels. Conversion then proceeds from BCD to Baudot code until an end of message sentinel is located. With the completion of one message, succeeding units are sought, identified and converted.

After completely processing the input tape the TELTYP program writes the teletype code onto the input tape. This new information follows the last output previously written onto that tape. A special data select character is used to mark this output as information to be punched from magnetic tape to 5-level paper tape via the Universal Buffer Controller (UBC).

15. 3 Input

Input to the TELTYP program is comprised entirely of the output tape (logical 11) previously written by any one program, or several programs, processed by the computer.

The TELTYP program must be called upon to perform its function immediately following the programs desiring this optional output. The output tape must not be wrapped up or rewound by the operator or system under which the programs are functioning. If the tape has been rewound,the teletype conversion will not be performed and a message will be written on the Flexo indicating the status of the input tape.

The following sentinels will be searched for on the input tape which must be addressable as logical tape 11 (eleven):

1. Data select i

2. 14 B' s

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- 3. $*$ \$
- 4. A block of U's.

15. 3. 1 Description of the functions of the sentinels.

1. Data select one

This sentinel is searched for at the beginning of each line on the input tape. If a data select one character is found, printed output is assumed and the processing continues. If any other data select character is present, other types of output are indicated and the line is ignored. The next line is then brought in for processing.

2. BBBBBBBBBBBBBB (14 B's)

This sentinel marks the beginning of a message to be processed into Baudot code. There will be as many sentinels of this type on the input tape as there are messages to be converted.

3. *\$ (an asterisk followed by a dollar sign)

This signals the end of the message being converted. The message is wrapped up when this sentinel is found and the sentinel described under 3. 1. 1 is then searched for by the program.

4. **128** words (i block) of U's

This sentinel is written on the input tape by the TELTYP program itself before the input tape is rewound and processing is started. It signals the end of the input tape and all processing ceases once it has been found.

When this sentinel has been located the program will rewind the intermediate tape used to store teletype output and copy its contents onto the input tape. The program then exits to the executive program.

15.4 Output

The output consists of the required conversion to Baudot code and is written on logical tape 11 (eleven) following the information previously written on that tape. Processing of other programs requiring tape **I i** (eleven) for output, but not requiring teletype conversion, can follow the termination of this program.

The 5-level paper tape output can be obtained through use of of the **UBC** if data select 4 is used.

15.5 Processing

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The TELTYP program first tests the input tape to determine if data is available for conversion to Baudot code. If the tape is in a rewound status a comment is made on the Flexo and control is returned to the executive program.

After it has been determined that data is available, a block of U's are written on the tape following the last piece of information recorded. The tape is then rewound and processing starts.

A search is made for an output line of information written under data select one. Once this requirement has been met, an attempt is made to locate the second sentinel (14 B's). If the B's are not located the program returns to the search for data select one. If the B's are located before the end sentinel (one block of U's) is reached, conversion of the message which follows is initiated by the setting of a logical switch.

In converting to Baudot code each character is examined. Numeric and alphanumeric characters are converted directly upon entry to a dispatch table. The appropriate shifts are determined and placed in the output buffer as required. Illegal characters are treated as blanks and a space code is supplied to the output buffer. Line feed can be called for by the insertion of an 8-5 punch (octal 15). Two carriage returns and a line feed are supplied by the program at the end of each line of information. An octal 32 (asterisk) is treated the same as any end-of-line indicator (octal 77). Each new line is examined for the data select one character.

The character by character evaluation and conversion continues until an end of message indicator is met. Two indicators are used: (a) an asterisk followed by a dollar sign, or (b) an absolute stop code which normally signifies the end of an output

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tape. The message will be terminated by either signal. Letter **4** shifts will be used to fill out the output area and an end of message indication will be given to the operator via the Flexo. The program then returns to search for the next message on the input tape.

This process continues until the block of U's is found. At that time the intermediate tape, which is used to store the converted message, is rewound and copied onto the input tape. The block of U's is written over by the information being added to the tape. An exit is then made to the executive program.

15.5. **1** Error Indications

- i. The following messages will be typed on the Flexoanda return will be made to the Exec.:
	- (a) TAPE FOR STANDARD TTY CONVERSION ALREADY REWOUND. JOB TERMINATED.
	- (b) NON-RECOVERABLE ERROR IN READING INPUT TAPE. UNABLE TO RESTORE TO ORIGINAL POSITION. REMOVE WITHOUT WRAPUP. JOB TERMINATED.

The first message was discussed under the input and processing sections. The second message is self-explanatory.

2. Other error messages are as follows:

(a) SUBROUTINE ERROR EXIT FROM OCTAL LOAD COMPACT INTO UPPER CORE AND TAKE DUMP.

This comment will be made only if the computer jumps to location zero or three. A stop-go option is given. Go will send control to the executive program. Stop will print out the message again. The operator is requested to give a dump only if the computer is in a non-interruptable mode.

(b) ERROR IN READING INPUT TAPE. TYPE GO TO ACCEPT OR STOP TO TERMINATE.

This is typed out whenever parity or sprocket errors are encountered.

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(c) BAD INTERMEDIATE TAPE. CHANGE TAPE AND TYPE GO.

Converted messages can be lost when the tape change is made.

(d) SCRATCH OR INPUT TAPE DID NOT REWIND, TYPE GO TO TRY AGAIN OR STOP TO TERMINATE JOB.

This message is self-explanatory.

(e) EM

This indicates that the end of message indicator has been encountered.

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Test Characters in Input Buffer

Fill Input Buffer from Input Tape

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16. **1** BMEWSPT, BMEWS Paper Tape Conversion

16.2 Function

BMEWSPT scans output messages from the DIP computer in search of possible SPADATS messages. Other messages of variable length are ignored. Any possible SPADATS message is examined for consistency. If the message appears to be valid, certain control bits are stripped, and the message is converted to standard observation format and put out as a binary tape. This binary output tape is in the format of a system $TTY\triangle IN$ tape and is used as direct input to ORCON.

16.2.1 Definitions

The following terms will be used throughout the BMEWSPT description:

- BMEWS Character The format of this character changes during preparation of magnetic input tape. See BMEWS word. It originally consists of 6 information bits preceded by a parity (P) bit during transmission to SPADATS.
- BMEWS Group Consists of three 24 bit BMEWS words (input tape format). There are three BMEWS groups in a SPADATS message.
- BMEWS Message Any message output by the DIP which may or may not be a SPADATS message.
- BMEWS Word Consists of three BMEWS characters. The format of a BMEWS word changes during preparation of the magnetic input tape. Details appear in that section. The format of a BMEWS word is shown in Figure 16. 1.

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Fig. 16. 1

Chad-less Paper Tape - Original paper tape received by SPADATS from the DIP. BMEWS words appear as three 7 bit characters. Control Bit - An **S,** P, or **E** bit.

- DIP Computer used to process all BMEWS data. Q-point data (observations) are outputted on chad-less paper tape and received by SPADATS.
- E-bit Error bit. The second bit of the 20 bit BMEWS word appearing on five channel paper tape. Set to one if transmission errors are detected between the DIP and SPADATS.

Filler Words - Philco 8 character words of all filler characters $(32)_{8}$.

- Five Channel Paper Tape Intermediary between chad-less paper tape and magnetic input tape. BMEWS words appear as four 5 level characters.
- I-bit-Information data bit. Any bit contained in a SPADATS message. except for control bits.
- P-bit Parity bit. The first bit of the 20 bit BMEWS word appearing on five channel paper tape.
- P -bit Parity bit. Original parity bit associated with a BMEWS character.
- S-bit Sixth level bit. The sixth level (most significant) bit added to each of the four 5 level characters on the five channel paper tape when preparing the magnetic input tape.
- SPADATS Message Three groups of three BMEWS words each of which contain Q-point data.
- TTYAIN Tape System input tape for teletype data which is processed by ORCON Tape ID is $70TTY\Delta IN$

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TYPE **25** Standard Observation - Observation in standard format from a station using a moving beam antenna with T tracking capability. Data readout is automatic. Zebra - Time - Greenwich Mean Time. $\mathcal{L}_{\rm eff}$

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16. 3 Preparing Magnetic Input Tape

Fig. 16. Z

SPADATS receives chad-less paper tape which is direct output from the DIP computer. The chad-less paper tape is converted by equipment to five channel paper tape by decoding BMEWS words.

BMEWS words are decoded into four 5-level characters. The P* bit of each of the three BMEWS characters is checked and deleted. The remaining 18 bits become the least significant of the **20** on the five channel paper tape. If the check of the P^* bits detected transmission errors, the E-bit is set to one. A new parity bit is generated and becomes the P-bit. An odd parity for the 20-bit group results.

In preparing the input tape from the paper tape, the reader must be in 5-level binary mode. The equipment adds an S-bit to each of the four 5-level characters. Therefore, the required six bit character will appear in computer storage. At least one block of filler words must be added via the simulator.

16. 3. **1** Zebra Time from Flexowriter

Since SPADATS messages do not contain the zebra year, month, and day of the Q-point data, these times must be obtained before conversion to standard observation format. Normally, there is a restriction that input data cannot be more than one day old. Then the year, month, and day may be obtained from the accounting clock. However, if the accounting clock is not working properly, or is set $\left(\int \right)$ to local time, or if the input data is more than one day old, the correct zebra time must be obtained from the flexowriter.

1-16-4

The program first interrogates the accounting clock and converts the time to BCD. The time is typed out in the format of YR-MM-DD-HHMM ZEBRA. If the time is correct, the operator types a carriage return and the program starts processing the messages.

If the time is incorrect the operator types the requested YR-MM-DD-HHMM followed by a carriage return. The corrected time is then typed out for approval. See 16. 4. 1-2 for further details.

16. 3. **²**Input Tape

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A BMEWS word will appear on the prepared input tape and in computer storage as 24 bits. Since there are 9 BMEWS words in a SPADATS message, a total of at least $4\frac{1}{2}$ Philco words are required.

The input tape may or may not contain messages of varying lengths other than SPADATS messages. The program will ignore extraneous data. A SPADATS message consists of three BMEWS groups. Each BMEWS group contains 72 bits. The format of each group including control bits is illustrated in tables following this section. Note, however, that a BMEWS message in core does not necessarily start at the beginning of a Philco half word. The beginning character of a BMEWS message may be any one of the 8 characters in a Philco word, therefore, the program examines each character separately.

If the control bits are ignored a BMEWS message will appear on the input tape and in core storage as shown in Figures 16.3 and 16.4.

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Fig. 16.4

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16.4 Output

BMEWSPT output is a magnetic tape containing Q-point data converted to standard observation format. The output tape is in the system $TTY\Delta IN$ tape format described below and is suitable for input to ORCON. Unused portions of all blocks are filler words.

'70TTYIN] The first word of the first block is the tape ID, $70TTY\Delta IN.$ This indicates to the executive routine that the tape is an ORCON input tape.

 $|\texttt{BEGINOBS}|$ The first word of the second block is the ORCON control word BEGINOBS. This indicates to ORCON that the following input will be in standard observation format.

Q-point The Q-point data converted to type 25 standard

Observations observations are written 12 observations per bloc Observations observations are written 12 observations per block. There are no restrictions on the number of blocks used for the observations which are all classified unknowns.

ENDOFOBS $\begin{bmatrix} 1 & 1 \ 1 & 1 \end{bmatrix}$ The first word of the second to last block is the ORCON control word ENDOFOBS. This is a signal for ORCON indicating the end of the standard observation input.

> The first word of the next to last block is the ORCON control word ENDA DATA. This indicates to ORCON that there is no more input from the $TTY\Delta IN$ tape.

The last block contains 120 words of Z 's

Each generated observation will occupy the equivalence of ten Philco computer words on the output tape. Each observation will be in the standard observation format shown in figure 16. 6.

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BMEWSPT

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16.5 Processing

BMEWSPT obtains the zebra year, month, and day of observation from the accounting clock and/or from corrections entered via the Flexowriter. The two header blocks containing the tape identification 70TTYAIN and the ORCON control word BEGINOBS as their first words respectively are written on the output tape. The input data is then transfered to the data storage area. If the data storage area does not contain any zero characters, the message, "MAY BE CODE MODE" is typed on the Flexo. This indicates that the input tape may have been prepared from the five channel paper tape under code mode, which ignores blanks, rather than binary made which converts blank paper tape to zero characters. This is only a possible error message.

The program then searches for a SPADATS message. Each character in core is examined until one is found which has 100 as the last three bits. This is tentatively assumed to be the SPADATS **(** identifier. The next three information bits are checked and should indicate the first group count, 000. If they are, the bits corresponding to the SPADATS identifiers and group counts for the second and third groups are examined. If they agree with the expected values, the site and sequence numbers from the first group are compared with the corresponding values in the second and third groups. If there is agreement and if the site number is **I** or 2, it is assumed that a SPADATS message has been located. If any test fails, the search for the SPADATS identifier continues with the character following the original character with 100 as its last three bits.

After a SPADATS message is found, the Q-point data is converted to a type **25** standard observation. The sequence number is converted to BCD. The site number is converted to station number by a table look up. The zebra time is converted to BCD. Year, month, day are obtained from the accounting clock or Flexowriter while hours, minutes, and seconds are obtained from the Q-point data. The credence is ignored. **4)**

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The number of consolidated reports is used in the determination of the accuracy. The first three of the six 'bits form a binary number, 0-7. This number is subtracted from 7 to give the accuracy. A resultant accuracy of 0 corresponds to a large number of reports. An accuracy of **8** or 9 may result from the detection of one or more parity errors, respectively. An error count is made up of the sum of the 9 E-bits and the number of bad parity checks. A non-zero error count results in an 8 or 9 accuracy.

If the azimuth is positve, there is a direct conversion from radians to circles. If it is negative, the magnitude is substracted $from 2\pi$ and converted to circles.

Elevation is converted to circles. If the result is greater than 0. 25, one-half circle is added to the azimuth and the elevation is subtracted from 0. 5. The resultant elevation and azimuth values are converted to degrees.

Range and range rate are converted to kilometers and kilometers/second respectively. Because the output range rate has nine characters, the integer and fractional parts are processed separately.

The resultant observations are written out on the output tape, 12 observations per block. After all data has been processed, and all observations written on the output tape, two blocks containing the ORCON control words ENDOFOBS and ENDA DATA as their first words respectively are written on the output tape. The last block written on the output tape contains 120 words of z' s. The output tape is now ready to be processed by ORCON.

1-16-11

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16. 4. **1** Flexowriter Messages

1. MOUNT SCRATCH ON UNIT 7 AND INPUT ON UNIT 0.

This message is typed at the start of the program. After typing this message, the program goes to the STOP-GO routine. The program will begin when the operator types GO.

2. YR-MM-DD-HHMM ZEBRA

The zebra time at the start of the program is typed for the operator' s approval. One or more of the zebra groups may be altered by the operator. The groups are separated by hyphens. See 16.3. **t.**

a. TRY AGAIN

This message is typed if the operator makes an error in typing in the requested time. Program waits for another type in of the time and retests.

b. ILLEGAL DATE-TIME. RETRY

Typed if an illegal group in the zebra time is found such as(T) a month> 12. The operator will then type in a new zebra time which will be tested.

- 3. WRITE ERRORS
	- READ ERRORS
	- PROC ERRORS

Appropriate comment is typed if tape errors occurred and could not be corrected. Control returns to the executive routine after typeout.

4. INPUT MAY BE CODE MODE

This message indicates that the paper tape may have been converted to the magnetic input tape under code mode instead of binary mode. The program tests one block for zero characters. If no zero characters are found, the above message is typed on the Flexo. The program then continues normally.

1-16-12

CARLIN COMMUNICATION

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Punched card output, if specified, is under control of a PATH switch which eliminates multiple punching of the same ephemeris data. This switch also controls the production of an ephemeris tape from a set of ephemeris cards. This switch equals one (1) for the first time through the program, for a particular set of ephemeris data, and two (Z) for each subsequent pass through the program for this same data. It can be reset to one by reading a negative station number. Thus ephemeris cards are punched, or an ephemeris tape made, only on the first pass through the program for that particular set of ephemeris data.

9.5.1 Error Message

If the program is unable to read the ephemeris tape, the comment EPHEMERIS TAPE TROUBLE will be written on the output tape. In addition, the comment PROGRAM TERMINATED DUE TO EPHEMERIS TAPE TROUBLE will appear on the Flexo, and control will return to the executive routine.

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Detailed Examination of the Data Word

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Examination of Possible Message

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Detailed Examination of the Data Word

Transfer of New Data to Data Storage Area

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Output of Date-Time to Flexowriter

Acceptance of New Date-Time from Flexowriter

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SECTION 2 SUBROUTINES AND FUNCTIONS

This section contains descriptions of subroutines and functions which are used by many of the programs described in Section 1. All, with the exception of XYZSB, have been written with the linkage required by FORTRAN or ALTAC. The description of XYZSB, which was written by Aeronutronic, a division of the Ford Motor Company, has been included to serve as a convenient reference within this volume.

The functions described are used to perform intermediate calculations and produce a single valued result. The subroutines perform a series of calculations and may produce more than one result.

1. 1 CONYRB, Year Constants (Subroutine)

1. 2 Purpose

CONYRB is used to obtain Theta Greenwich, the celestial longitude of the sun, and the difference between the longitude of the sun and the argument of perigee of the sun at the start of a year.

1. 3 Use

Call CONYRB (IX, XI, X2, X3, X4, X5, X6, X7, X8)

1.4 Input

IX must be a fixed point integer equal to the last digit of the year. No further arguments are necessary for entry to the subroutine.

1.5 Output

Upon exit from the subroutine, the arguments represent the following constants:

- IX $(Fixed point) = last digit of the year.$
- $X1$ (Floating point) = year.
- X2 (Floating point) **=** Theta Greenwich at start of year.
- X3 (Floating point) **=** celestial longitude of the sun.
- X4 (Floating point) = . 017202789
- X5 (Floating point) **=** difference between the longitude of the sun and argument of perigee of the sun.
- X6 (Floating point) **=.** 0334502.
- X7 (Floating point) **=.** 043053055.
- X8 (Floating point) = . 43365539.

2.1 MHOLY, Remove Hollerith Overpunch (Function)

2.2 Purpose

NHOLY is used to remove an overpunched sign and to convert the argument to the absolute value of the fixed point integer.

2.3 Use

NHOLY (IX)

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2.4 Input

The argument (IX) contains any left justified BCD character which may include an overpunched sign.

2.5 Output

NHOLY yields the absolute value of the fixed point integer contained in the left most BCD character of the original argument.

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3. **1** PROPR, Properize Argument (Function)

3. 2 Purpose

PROPR is used to properize an angle, i. e. , find the value of the angle between 0 and 2π . If x is any angle in radians, PROPR = $[X]_{(o, 2\pi)}$

3. 3 Use

PROPR (X)

3.4 Input

The angle (X) must be in floating point radians. There is no restriction on its value.

3. 5 Output

PROPR yields the value of the angle (X) in the range $0 \leq x < 2 \pi$.

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\begin{cases}\n= 1 \text{ if } B - 8 \ge 0 \text{ (leap year add 1 day)} \\
= 0 \text{ if } B - .8 < 0\n\end{cases}
$$

\nJ = year day - 59 - C
\nIf $J \le 0$, $h = 30$. 1; otherwise $h = 32.3 - C$
\nmonth = $\frac{\text{year day}}{30.6} + 2$
\nE = [30.6 month - h]
\nIf year day $\le E$ subtract 1 from the month and recompute E.

Otherwise day number in month = year day - E.

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5.1 **SHOLY, Restore Hollerith Sign (Function)**

5. 2 Purpose

SHOLY is used to interpret the original overpunched BCD argument in terms of an algebraic **+** or -

5. 3 Use

SHOLY (X, Y)

5.4 Input

SHOLY requires two arguments. The first argument (X) is the overpunched BCD character. The second argument (Y) is the unsigned value of the first argument (X) in floating point.

5.5 Output

SHOLY yields the signed floating point value of the second argument.

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6. 1 SMITH, Obtain Theta Greenwich (Function)

6. 2 Purpose

SMITH is used to obtain Theta Greenwich in degrees for the start of the year.

6. 3 Use

SMITH (X)

6.4 Input

The argument (X) is the last digit of the year expressed in floating point.

6. 5 Output

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SMITH yields the proper Theta Greenwich (X) in degrees.

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7. 1 UNHOLY, Remove Hollerith Overpunch (Function)

7.2 Purpose

UNHOLY is used to remove an overpunched sign and to convert the argument to an unsigned floating point number.

7.3 Use

UNHOLY (X)

7.4 Input

The argument (X) contains any left justified BCD character which may include an overpunched sign.

7.5 Output

UNHOLY yields the absolute value of the floating point number contained in the left-most BCD character of the original argument. UNHOLY uses the function NHOLY.

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10. 1 XYZSB, Analytical Integration Routine

This routine was written by Aeronutronic. A complete description of the subroutine, including equations, appears in the Aeronutronic Publication U-1691, pages 4-78.

10. 2 : Perpose

XYZSB computes the predicted position, r, and velocity, \dot{r} , at some given time, t. The N, M orbital elements at some epoch time, t_{0} , must be supplied upon entry to the routine (c. f. Fig. 10.1).

10. 3 Calling Sequence

10.4 Input

Index register two (XRZ) must be set equal to 10, 20 or 30 upon entry. The significance of this setting will be explained under the output section (10.5) and will be referred to as the exit flag.

The following quantities must be available in the indicated locations upon entry into this routine: (See Assign Table).

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Index register two (2) must be set equal to **10,** 20 or 30 upon entry. The significance of this setting will be explained below under Output and will be referred to as the exit flag.

10. 5 Output

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1. (XR2) = **10,** 20 or 30

If the exit flag is equal to 10, 20 or 30 the following quantities are available upon exit from this routine in the indicated locations:

2. $(XR2) = 30$

If the exit flag is equal to 20 or 30 the following is also output:

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 $(XR2) = 30$ $3.$

If the exit flag is equal to 30 the following is also output:

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XYZSB

Wolf Research and Development Corporation

PROJECTION OF ORBIT ON CELESTIAL SPHERE, WITH ORIENTATION UNIT VECTORS AND ANGLES DISPLAYED

Fig. 10.1

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Element Card Formats:

All programs accept a standard 7-card element set as input. The formats for these cards appear on the following pages. Those columns which are shaded are not used. In all floating point numbers, the decimal point may be placed anywhere in the field. Numbers requiring an exponent indicator such as c_a , d_a , and c_n , must be in one of the following formats:

+ . DDDDDDDDE + XX

- **+ .DDDDDDDDA +** X
- **+** . DDDDDDDD + XXX

All programs except **ASUM, SSUM,** ISUM, PSR, and RESPLT will also accept a 6-card element set. Fields which differ from those of the 7-card element set are noted in the rightmost column of the card format description.

Element Card (1 of 7)

Element Card (1 of 7)

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Elements

Element Card (2 of 7)

Elements

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Element Card (3 of 7)

Element Card (3 of 7)

Element Card (4 of 7)

Element Card (4 of 7)

Element Card (5 of 7)

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Element Card (5 of 7)

Element Card (6 of 7)

Element Card **(6** of **7)**

Element Card **(7** of **7)**

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Contract Contract

Sensor

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Station or Sensor Card

Station or Sensor Card

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Observation

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COLUMN **10** (ACCURACY)

Either accuracy or signal strength may be indicated in column i0, coded according to the following:

If Type, in columns 4 and 5, is 31 or greater, column 10 contains signal strength. If Type is 30 orless, column 10 contains accuracy.

COLUMNS 55 - 63 (CROSS SECTION-FREQUENCY/ MAGNITUDE)

The block containing columns 55 through 63 is a dual purpose block where cross section and frequency, or magnitude and time interval are indicated. In order to specify cross section and frequency, a minus is used in column 58. No sign is used in column 58 when this block contains magnitude and time interval.

- A. Cross section, given in square meters, is listed in columns 55 through 57. To indicate less than one square meter cross section, use appropriate numbers and a minus in column 55 thus in effect putting a decimal point before column 55. For larger valueswhere three digits would not be sufficient, use a plus in column 55 to represent ten times the indicated value (adding a zero to the value listed).
- B. Frequency in megacycles, is listed in columns 58 through 63 with the decimal point understood to be located between columns 60 and **61.** In rare cases it might be desirable to increase the range of frequency given either side of the decimal point. To do this, use a minus in column 63 to move the point one place to the left, or a plus in column 63 to move the point one place to the right.

Observation Card (Part 2 of 3)

$3 - 3 - 2$

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COLUMN 70 (EQUINOX)

Column 70 contains year of Equinox as specified by the following:

 $0 = year of date$ $1 = 1900$ $2 = 1920$ $3 = 1950$ $4 = 1975.$ $5 = 2000$ $6 = 1850$ $7 = 1855$ $8 = 1875$ $9 = 1960$

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First Card of Schedule Tape Input Deck

Schedule Tape ID Card

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First Card of Each Job Deck

JOB Card

 $3 - 4 - 2$

Optional Comment Card

REM Card

Used with Schedule Tape Mode Programs See Operating Summary Table for List of SPSJOB Cards

SPSJOB Card

Field **Cols.** 1 **1-8 ENDA CASE 2 9** 11,8,2 Punched **3 10 - 80** Blank

Used with Schedule Tape Mode Programs

* **END CASE** Card

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Used to Run Manual Schedule Tape Mode Programs

See Operating Summary Table for List of RUN Cards

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Required if RUN Card Used

END DATA Card

Required to Terminate Each Job Deck

END OF JOB Card

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Last Card of Schedule Tape Input Deck

Schedule Tape **END** Card

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Assign Deck

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Assign Deck

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COGEF USED IN COMP C SUB DELTA A 59 YANGE RATE DUS IN UC

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CONSTANT 3 P1 OVER 2
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X COMP OF R BAR DUT AT EPOCH TAPE TABLE
T/O ROUTINE
TAPE TABLE
TAPE TABLE
TIME SINCE EPOCH
TAPE ID CHLCK ROUTINE
TAPE TABLE SET-UP SWITCH
TEMP THETA DOT
GREENWICH SIDEREAL TIME AT
THETA GREENWICH SUB ZERO
THETA = SIDEREAL TIME
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TTYTAP,
TWOPL,
TYPL, XKE,
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XKMPER, X1HF50,
X15T5G,
X2OV3,
X3PI02, XINCL,
XJGRLF,
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ERROR PROCEDURES AND COMMENTS

Irrecoverable Errors:

- 1) IRRECOV ERROR
- 2) PROGRAM ERR PROC (See PROC ERRORS Table)
- 3) PROGRAM ERR lOPS (See lOPS ERRORS Table)

Irrecoverable errors comments are accompanied by a typeout of the contents of the A register and Q register:

 $A REG =$

 Q REG $=$

The program will then jump to location **3.***

Recoverable Errors:

Recoverable errors are accompanied by the STOP-GO option, wherein the operator may retry the job by typing GO or may type STOP to effect a jump to location **3.***

 \bullet Location 3 will generally cause the typeout of the Flexo comment:

SUBROUTINE ERROR EXIT FROM OCTAL XXXXX, LOAD

COMPACT INTO UPPER CORE AND TAKE DUMP.

Following the message, a stop or go loop controls the console such that if STOP is typed the message will be repeated. Typing GO returns control to the executive program terminating the job. If possible, a dump should always be taken.

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PROC ERRORS

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SPSJOB OPTION TABLE

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S) ġ, I = Input, O = Output, S = Special, X = Scratch, UV= Write King, <u>ZIV</u>I =
* = TELTYP Program May Follow, [= Repeat if Necessary, ()= Optional

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SPS Operating Summary

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SYMBOL KEY: I = Input, O = Output, S = Soecial, X = Scratch, \bigotimes = Write Ring, $\overline{\bigotimes}$ = No Write Ring, \circledcirc = 11,8,2 Punch
* = TELTYP Program May Follow, [= Repeat if Necessary, () = Optional

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SPS Operating Summary

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SYSTEM OPERATING MODES

The definition of terms as used in the descriptions included in the Input-Output section under Input Deck are as follows:

A. Schedule Tape Mode

This term implies that the program is completely **SPS** system oriented. The assign deck listed in the Assign Table has been used to bind the program to the system. An SPSJOB card must be used to call the program from the system library tape. All elements, observations, sensors, as well as the parameter cards are brought in by the executive routine after consulting the input-output options contained on the SPSJOB card and an internal Job Table. Only system supplied routines are used in the execution of input-output requirements. Upon completion the program will exit to the system location called EXEND. Toggle 24 must be on and 1-23 off to run under this mode.

Manual Schedule Tape Mode

The programs of this type require a RUN card to call the program from the system librarytape. If data follows, it will be moved to logical tape zero (0) by the executive program. The program will read data from this tape.

PROC and lOPS, which are supplied at compilation time by the ALTAC compiler, are, with few exceptions, the input-output routines used. These input-output routines have been modified and are interruptable by the real-time devices available. Variable formatting of input-output requests, as used in FORTRAN and ALTAC, as well as Macro capabilities, have been maintained. as a result of the modifications made.

The system assign deck is not used and the program contains the minimum amount of information required to bind it to the system. Upon completion of the program a return is made to the location called MANEXIT.

Toggle 24 must be on and 1-23 off to run under this mode.

SCHEDULE TAPE FORMAT

Each new job deck starts at the beginning Fillers and the state of a block, to simplify interrupt.

END CASE Card END CASE Card Blocks 2-6 show the layout of an SPSJOB run in Schedule Tape Mode.

RUN Card Block 7 shows a job run under the Manual ENDOFJOB Card Schedule Tape Mode, without data.

Data $\begin{array}{|c|c|c|c|}\n\hline\n\end{array}$ Blocks 8-9 show a job run under the Manual

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Request Card

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POSITION SITUATION REPORT

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REDUCT (Nodal Crossing Reduction)

INPUT DECK: Manual Schedule Tape Mode

INPUT DATA:

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The first card of the input data is an option switch card controlling the processing and the output. This card is followed by the standard element sets, station cards (if any), and observation cards, with each group preceded by a lead card.

The deck should be in the following order:

- i. Option Switch card:
	- Col. i: a numeric punch (1-9) will cause the Station Tape (logical 7) to be inhibited.
	- Col. 3: a numeric punch, **(f)** inhibits the interim tape (logical 8), (2) eliminates tolerance limit tests.

Col. 8: a numeric punch causes all observations

- **-**to be treated as unknowns, reducing all against the available element sets.
- Col. 9: a numeric punch modifies headings, spacings, and format length to allow subsequent processing by TELTYP to obtain output for teletype transmission.
- Col. **10:** a numeric punch eliminates a re-reduction of unknown obs. which have remained untagged after a reduction against the selected tolerance limits. If the rereduction is not eliminated, the untagged obs. , which have been written out on the interim tape, are tested against open tolerance limits ($\Delta t = 99.0$ days, $\Delta RA =$ 15. 0 degrees, $\Delta h = 99999$. 0 kilometers).

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- Col. **If:** a numeric punch directs the program to use the perigee distance (q term) of the element cards. If this term is not used, perigee distance will be computed by the formula: $q = a (1-e)$.
- 2. Element lead card: a 7 punch in col. 8.
- 3. Standard 6 or 7 card element sets (up to 250).
- 4. Station lead card: a 9 punch in col. 8.
- 5. Standard station cards (up to 750). *
- 6. Observation lead card: and 8 punch in col. 8 with the tolerance limit key punched in col. 7:**

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- 7. Observation cards (any number).
- 8. A blank card to terminate the preceding group. group of observation cards.
- 9. End of input card: a 9 punch in col. 79.
- *The Station Tape (logical 7) may be used in place of, or in addition to, the card groups 4 and 5.
- ** Observation lead cards (any number) may be inserted to vary the tolerance limits applied to the reduction of the following observation cards.

S OUTPUT:

Output 'begins with the satellite inventory. Seven groups, composed of satellite number and element number, are printed per line.

The next output is the satellite number (first line), selected elements (two lines), and the tolerance limits used in the observation output following (one line).

Observation output will occur in one of three formats, depending on which part of the program has done the reduction, as in the following table:

General Reduction Output:

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Doppler Reduction:

ID = Observation identification

 $N =$ Epoch revolution number

T SUB N **=** Computed time of nodal crossing

DELTA T = Time residual in days (difference between predicted and observed)

- $D = Arc distance in nautical miles from station to subsatellite point.$
- $H =$ Elevation, in degrees
- $S =$ Slant range, in kilometers
- ELEM = Element number
	- STA = Station number

Direction Finder Reduction:

ID = Observation identification $N =$ Epoch revolution number T SUB N = Computed time of nodal crossing DELTA $T =$ Time residual in days (difference between predicted and observed) H **=** Elevation, in degrees $S =$ Slant range, in kilometers H(KM) **=** Object height, in kilometers ELEM = Element number $STA = Station number$

If the output is to be processed by TELTYP for teletype transmission, a modified version of the above output will be printed. In general, each line, as described above, will be broken into two lines of output. Each line will be followed by an asterisk (*), and the abreviation, DOP, for doppler observations, or DF, for direct finder observations will follow the station number.

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SATELLITE INVENTORY FOLLOWS.*
51 19/ 60 19/ 73 6/ 77 $\overline{41}$ 05810221 01582700 NO ELEMENTS IN SYSTEM(1). $300*$

UO S FULLOW* UNK*********** 6878 $0*$ 05110208 00221500 $2189 - 39.00631465 - 0.00538*$ $1256.21.81$ $3099.$ $19.$ 25 DOP* 1267 39,00684022 0.02643* 06010208 00221500 25 DOP* 824, 27,84 $2022. 19.$ 63 38.99523414 $0.00552*$ 07310208 00221500 $7425. 6. 2500p*$ $4179. -33.18$ -65 38.99127295 0.02301* 07710208 00221500 5885. $4.$ 3216 , -24.20 $2S$ DOp* UNK*********** Mik 05110221 01592900 2348 225.08 52.03144556 -0.01601* 286.79 .94.19 19322.3 17807.4 RDR -31.301 320.43 $19.$ $300*$ $1404222.7252.03242214 - 0.00864*$ 06010221 01592900 290.35 -77.61 19322.3 17408.9 RDR $-31,301$ 320.43 $19.$ $300 272$ 324.99 52.02669415 -0.02319* 07310221 01592900 -31.301 320.43 157.62 30.31 19322.3 19057.3 RDR $6\,$ $300 +$ 07710221 01592900 145 325,03 52.02714149 0.00002+ $\ddot{\mathbf{A}}$ $-31,301$ $320,43$ 157.52 4.05 19322.3 19117.7 RDR $300 *$ $1877 - M*$ UNK*********** 2348 180.54 52.04140303 -0.00005* 05110221 01592900 -0.394 303.74 304.01 -76.97 1116.7 -421.4 RDR $19.$ $300 *$ 06010221 01592900 1404 180.51 52.04354025 0.00248* $304.04 : -63.92$ 1116.7 -1093.3 RDR $*0.394$ 303.74 19. $300 *$ 272 359,56 52,02072186 -0,02916* 07310221 01592900 $124,56$ -2.75 1116.7 871.8 RDR -0.394 303.74 $300 +$ 6. 07710221 01592900 145 359.56 52.02126079 -0.00586+ -0.394 303.74 124.56 -28.92 1116.7 923.7 RDR $4.$ $300 *$ 00410223 15084950 NO ELEMENTS IN SYSTEM(1). 39 k

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REDUCT

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11.636 -3.354 -0.6000 355.095 2.785 0.0000 112.62 49.98* \bullet 0.00200 20.00000 200.00000+ 06010220 19424540 1401 81.38 51.80628582 -0.00016*
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 $\begin{array}{cccc} 756 & 47.55804 & 0.062440 & -0.746 - 0.06 & 0.000 + 0.00 & 1.04231 & 0.00639* \\ 143.762 & -3.657 & -0.0016 & 60.997 & -0.428 & -0.0002 & 89.91 & 64.89* \end{array}$ $73+$ 0.00200 20.00000 200.00000* 07310223 15093600 313 157.92 54.60451877 0.00024* 39_o 19.898 267.15 117.58 -0.33 305.6 38.2 RDR $6.$ 07310223 16384100 314 155.21 54.66685240 0.00030*
22.315 243.44 117.53 =0.16 270.1 4.2 RDI 4.2 RDR 6. $39*$

 $77+$ $\begin{array}{ccccccccc} \sqrt{5} & 48.61977 & 0.062087 & -0.212-005 & -0.618-008 & 1.03837 & 0.00817* \\ 166.095 & -3.689 & -0.0047 & 20.377 & -0.470 & -0.0006 & 89.41 & 65.01* \end{array}$ 0.00200 20.00000 200.00000+ 07710223 16551900 188 $155,88$ 54.67849040 0.00037* 21.739 265.66 143.54 -0.04 208.0 -7.6 RDR 4. 39 + END OF RUN **********

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(Radar Orbit Computation)

INPUT DECK: Manual Schedule Tape Mode

INPUT DATA: Two options for input are accepted by the program: i) radar observations, 2) geocentric rectangular coordinates and velocities.

Option **1:**

1. Request Card:

Col. 1-3 Number (odd) of observation cards to be read, from 3 to 999. 6-17 Alphanumeric satellite

name or identification, for output.

18-35 Date of computation

41 1 punch request element cards as output. 42 **1** punch prints x, y, z, **x, yj** • for the mean of the observations.

72-80 Are blank

2. Standard Station Card

- 3. Standard Observation Cards, in order of increasing time. (any odd number from 3 to 999)
- 4. 9 punch in col. 79: card used to end data deck.

Option 2:

1. Request card:

Col. 6-17 Alphanumeric satellite name or identification, for output.

- 18-35 Date of computation
- 41 1 punch requests element cards as output
- 72-80 Must be blank

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2. Rectangular Coordinate Card:

'I! The field must have a decimal punch.

OUTPUT: The output consists of the normal Radar Orbit Computation plus two options as selectedby the input request card.

> The Radar Orbit Computation is comprised of headings (program title, station name, alphanumeric satellite identification, and the date) and output. Each datum of output follows its label: semi-major axis in km. , earth radii, right ascension of ascending node in degrees, radius vector in km. , eccentricity squared, nodal period in days and minutes, argument of perigee in degrees, velocity in km./sec. , perigee distance in km. and in earth radii, time of last perigee pass in days, inclination in degrees, apogee in km. and earth radii, and the time of the last nodal crossing in days..

Output Option A, which is possible only with input option (1), produces a printing of the geocentric rectangular coordinates x, y, z (headed $XX =$, $YY =$, $ZZ =$) and the velocity components (headed $XDER =$, $YDER =$, $ZDER =$). The time of the observation is given under the heading TIME=.

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Output Option B produces punched cards in the standard 6 card element set format. The data punched are the satellite number, element card number, satellite name, eccentricity, inclination, year of epoch, time of epoch, nodal period at epoch, right ascension, argument of perigee, and perigee distance. All other fields of the standard set are left blank.

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If teletype transmission is desired, the program TELTYP may be used.

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MILLSTONE MASS RADAR ORBIT COMPUTATION

SATELLITE NO. ROC TEST SATE₁₁ JTE

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MARCH 17,1961 THE BELOW INFORMATION COMPUTER FOR

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LOCVEC

(Vector Coordinates for Lockheed)

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CCOE

(Cartesian Coordinates from. Orbital Elements)

INPUT DECK: Manual Schedule Tape Mode.

- INPUT DATA: (1) Set of Standard Elements (6 or 7 cards).
	- (2) Request cards as many as required:

Cols. i - 10: Start time (days of year).

- Ii ZO: Time increment (days of year).
- 21 30: End time (days of year).
	- 31: Blank, output in kilometers per sec.
		- i, output in earth radii per min.
- (3) One blank card
- (4) The above may be repeated any number of times. To terminate the program a second blank card must be used.

OUTPUT: Logical tape **11** contains the position and velocity components with the output in kilometers and kilometers per second or in earth radii and earth radii per minute. The information is in order of increasing time (in days) for each time step during the range requested.

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Columns:

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Note: Decimal points may be punched anywhere in each of the first three fields.

> Request Card **CCOE**

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JOB CCOEr2 RUN . **CCOEBZ, DATA** 11620. 0.09786993 4172 1 58 ALPHA 33.20999976 46.54090166 1961 4174 21961 48.54090166 4172 3 0.07414181 209.17569126 194.0101.947 1.05549147 $7.226332757...$ 4172 $4-1.2992010$ -67 -4.99362775 $4172 - 5 - 0$. $-L = 4287259E - 04 - 066286715E - 04$ 4172 6 1.00000000 36..000009059 -1. $-C₀$ $-$ 0 \leq 60.000000.00100000 60.01000 $\mathcal{L}^{\mathcal{L}}$

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Wolf Research and Development Corporation RESPLT

RESPLT

(Residual Plot)

INPUT DECK: Schedule Tape Mode

An SPSJOB card is required. The following options of data input and output are specified in columns 17 and 18:

a) Input option, column 17:

b) Output option, column 18:

 $0 = print (DS 1)$ and punch cards $(DS. 2)$

INPUT DATA: Element data and sensor or station data may come either from standard cards or from the SEAI Tape depending on the input options chosen. The number of input cards is limited by the size of the **0,** E and S blocks of the system.

> Observational data always originates from standard observation cards, which have been ordered by satellite number and contain an association indicator in col. 80.

If untagged observations (a zero sat. no.) are to be reduced, the first of the untagged observations must contain the satellite number of the element set against which the observations are to be reduced. If this procedure is inconvenient, a card may be inserted preceding

the untagged observations. This card must contain the satellite number to be used and an association indicator in col. 80. In this case, the sensor number must be blank.

The association indicator is necessary on all observation cards to indicate whether the observation is an angles only or a radar observation.

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OUTPUT: Two types of output are provided: 1) the printed output of the reduced observations, and 2) the punched card output of the reductions, with pen commands, required as input to the EAI Data Plotter.

- **1)** The printed output consists of the reduced observations, sorted by revolution number, for a given satellite. Headings are printed for the data columns which contain: the satellite number, the observation number, the revolution number, the time (minutes/100) since epoch, the vector magnitude in kilometers, the revolution number **(N/1000)** since epoch, the element number, the association indicator, and the station number.
- 2) The punched card output of the reduction is used as input on an EAI Data Plotter which is assumed to have line plotting capabilities. The required

commands are punched in column 55. The first card of output will stop the plotter (7 in col. 55). This is done to enable the operator to change graph paper and position the origin. The cards which draw the axes follow the stop command (six cards). The graph labeling cards are next and cause the plotter to draw the characters **S,** E, R and T. These characters are followed by the satellite number, element number, epoch revolution number and the epoch time in days. A stop command follows the above data (7 in col. 55). (The deck may be broken at this point if the axes and labeling is not desired). The data cards containing the information to be plotted follow the stop command and continue until the next stop command is reached.

Two types of plots can be obtained from the data cards: **1)** the time difference versus the revolution, and 2) the vector magnitude versus the revolution. Two IBM **523** board wiring diagrams have been supplied in Section 1 under the program description. The scaling is such that both plots can be obtained on the same piece of graph paper. To achieve this result, the IBM 523 board must be changed when the final stop command is executed. The deck should be broken as described above and the data cards reprocessed. It is recommended that a symbol pen or color change be used in order to distinguish the two plots.

The graph scaling is as follows:

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- 1) **x** represents the epoch revolution number
- 2) each centimeter represents 20 revolutions

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- 3) y_0 is the origin of either a) the vector magnitude, or b) Δt .
	- a) each centimeter represents 500 kilometers and under b) each centimeter represents 5 minutes.
- 4) x_0 should be 5 centimeters from the left edge of the lined graph paper and y_0 should be 9 centimeters from the bottom.
- 5) full scale is then:
	- a) $x_R = -100$ to $+400$ revolutions from epoch

 $e^{-i\mathbf{a}}$

- b) $y_{VM} = 0$ to 4500 kilometers
- c) $y_t = -45$ to $+45$ minutes.

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PREPINT الرائيس (Satellite Situation Report frorn Nodal Elements) INPUT DECK: Manual Schedule Tape Mode INPUT DATA: The input deck should be in the following order: **f.** Request Cards (up to 500) in the following format: Cols. $1 - 3$: Satellite number 7 Comment indicator defined as follows 0, no comments on this card i, read comments contained in cols. $21 - 62$ 9 - 20: Satellite name 21 - 62: Comments to be printed if col. 7 contains a i punch. (A max. of 50 sats. may have comments) 2. Request deck terminator: a 2 punch in col. 7. 3. Time cards (times for which reports should be issued) in the following format: Cols. I - 4 Year 6 - 7 Month number 9 - 10 Day of month i **I** - 17 Hours and minutes (HHMM. MM) 20 Output unit indicator, defined as follows: 0, output in statute miles 1, output in kilometers 4. Time Deck terminator: a 7 punch in col. 8. 5. Standard 6 or 7 card element sets. Only the sets required are stored, therefore more than 500 sets may be included in the input deck. 6. Input deck terminator: a blank card. OUTPUT: Logical tape Ii contains the satellite name and number, element number, latitude, longitude, inclination, period,

apogee distance, perigee distance, revolution number, time of node, right ascension of node, longitude of node,

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eccentricity, height of satellite, and course. The distances may be in kilometers or statute miles. The dimensions are determined by the option selected on the Time card (col. 20).

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Request Card PREPINT

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22 59 IOTA 22 59 IOTA **1** 59 IOTA
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43 60 ZETA 1 43 60 ZETA 1 45 60 ETA **1** 46 60 ETA 2
47 60 ETA 3 47 60 ETA 3
49 60 IOTA 1 49 60 IOTA **1** 61 LAMBDA 2 **1u7** 61 NU **116** 61 OMICRON1 167 61 RHO 4 16:3 61 SIGMA **?** 170 61 UPSILON
182 61 OMEGA 1 182 61 OMEGA **1** 61A BETA 18b 61 SIGMA 3 2 1961 **10 11** 1200 ~ 10 1961 **10 11** 1200 1 1961 **10 11** 1430 1961 **10 11** 1430 1 7 4214 1 58ALPHA 14740. 0.09500069 33.18999958
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4214 3 201.61903020 107.42337540 1.0547645 4214 3 201.61903020 107.42337540 1.05476451
 4214 4 -4.96605333 7.42265104 0. 4214 4 -4.96605333 7.42265104 0.
4214 5 0. -3.780903E-004 5.651232E-004 0. 4214 5 0. -3.780903E-004 5.651232E-004 **0.** 421 6 0.07371200 -1.968034E-007 4214 7 5 78 158 BETA 2 13950. 0.18984776 34.24499,89 ⁵**78** 21961 278.25430390 1961 278.25430390 ⁵**7b** 3 0.09295403 193.98439300 94.50760521 1.10183595 5 78 4-1.484110E-007 -3.02018139 4.41471416 **0.** 5 78 5 **O -** -9.897075E-005 1.446693E-004 0. 5 78 5 0. ⁻ -9.897075E-005 1.446693E-004 0.
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108936014 083 $i1149$ 3 0.08698856 32.53621422 190.73210390 11149 4 4.736775E-008 -3.51236494 5.28651193 **0.** 11149 5 **U.** 4.308964E-005-6.485486E-O05 0. 12110 1 59 ALPHA 2 **101UU.** 0.18352360 32.66509133 11119 6 1.00000000 30.00000 0.01525878 10.000000000 12110 21961 279.61322930 1961 279.61322930 12110 3 0.08993200 255.21911570 214.33844310 **-** 1.08631547 1211U 4 1.999180E-008 -3.30525151 4.99311847 0. 12110 5 0. 1.568518E-005-2.3695O1E-005 0.

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MAKETAPE

(Make Input Tape for TELTYP)

INPUT DECK: Manual Schedule Tape Mode.

- INPUT DATA: (I) TELEFORM punched in cols. 17-24. Used if the message is to broken into 90-line segments.
	- (2) Data cards to be converted to teletype.
	- (3) FINDATA punched in cols. 17-24. Used to terminate the data deck.

OUTPUT: Logical tape **iI** is input to TELTYP. See writeup of program. The tape can be printed using data select one. The output will consist of a listing of the input data cards with the sentinels required by the TELTYP program.

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س $\ddot{ }$ σ_i o $\tilde{3}$ \sim $\ddot{ }$ **TIMES** \circ ∞ \circ 4 \mathbf{r} \circ o ∞ \mathbf{v} ω \sim $\ddot{ }$ ∞ \overline{r} o, n, $\ddot{ }$ $\tilde{\mathcal{E}}$ \overline{a} ICK BROWN FOX JUMPED OVER THE LAZY DUG S BACK 0 I \circ E QUICK SROWN FOX JUMPED OVER THE LAZY DOG S BACK UICK BROWN FOX JUMPED OVER THE LAZY DOG S BACK O ∞ $\ddot{ }$ \sim QUICK BROWN FOX JUMPED OVER THE LAZY DOG S BACK o, L٦ ∞ QUICK BROWN FOX JUMPED OVER THE LAZY DOG S BACK \overline{r} o $\overline{0}$ \overline{a} $\ddot{ }$ CK BROWN FOX JUMPED OVER THE LAZY DOG S BACK 0 σ \overline{r} ن
م $\overline{4}$ ω m BROWN FOX JUMPED OVER THE LAZY DOG S BACK 0 1 $\overline{\mathsf{c}}$ ∞ \overline{r} \circ $\ddot{ }$ \overline{N} L ω BROWN FOX JUMPED OVER THE LAZY DOG S BACK 0 1 ∞ \circ σ 4 \overline{N} \circ $\ddot{ }$ \overline{r} \mathfrak{c} ω $\frac{1}{\circ}$ K BROWN FOX JUMPED OVER THE LAZY DOG S BACK \bullet \circ $\ddot{ }$ \mathbf{r} \tilde{c} \sim WN FOX JUMPED OVER THE LAZY DOG S BACK 0 1 r ω o $\ddot{ }$ \sim \overline{C} o \mathfrak{g} $\ddot{}$ ω \sim FOX JUMPED OVER THE LAZY DOG S BACK 0 1 ROWN FOX JUMPED OVER THE LAZY DOG S BACK OWN FOX JUMPED OVER THE LAZY DOG S BACK \tilde{c} 4 \sim u٦ FOX JUMPED OVER THE LAZY DOG S BACK 0 1 o $\overline{4}$ \sim ω n. $\tilde{\mathcal{E}}$ \sim X JUMPED OVER THE LAZY DOG S BACK 0 1 S BACK \downarrow Ω OX JUMPED OVER THE LAZY DOG S BACK O \overline{a} 4 \sim $3 +$ APED OVER THE LAZY DOG S BACK 0 1 2 $\tilde{\mathcal{E}}$ \overline{C} UMPED OVER THE LAZY DOG S BACK 0 1 ω \sim JUMPED OVER THE LAZY DOG S BACK JUMPED OVER THE LAZY DOG S BACK THE LAZY DOG S BACK 0 1 2 PED OVER THE LAZY DOG S BACK O S BACK 0 1 N FOX JUMPED OVER THE LAZY DOG \circ ED OVER THE LAZY DOG S BACK OVER THE LAZY DOG OVER

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XYZLAR

(Look Angle Report from x, y, z, Coordinates)

INPUT DECK: Manual Schedule Tape Mode

INPUT DATA: The input data is composed of sets, or units, each of which consists of a standard station card, a request card, and optional ephemeris data. Ephemeris data is required only in those data units which initiate a new vehicle or satellite output. The source of ephemeris data is either punched cards or an ephemeris tape (unit 7).

> The order of cards for an input data unit using ephemeris cards and initiating a new satellite is as follows:

- **(1)** Standard Station Card
- (2) Request card specifying time parameters for the look-angle computations and input/output options.

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- (3) Vehicle identification card containing the alphanumeric vehicle name in columns 1-16. (4) Ephemeris cards, in order of increasing time increment over base time. Cols. 1-14 time increment, in minutes, from base time Cols. 15- 28 x coordinate of vehicle position Cols. 29-42 y coordinate of vehicle position Cols. 43-56 z coordinate of vehicle position
- (5) Blank card to terminate the ephemeris cards.
- (6) Blank card to terminate program.

This set or unit of input data **(1** -5) will generate one schedule of look-angles for the vehicle data contained on card types 3 and 4. Additional schedules for different stations are obtained by adding station and request cards, in pairs, after the blank card (5).

To obtain schedules for vehicles or satellites in addition to that of the first input unit, another ephemeris card set $(1-5)$ is required. A minus sign $(-)$ must be punched preceding the station number (cols. 1-4) -of the station card. Schedules for different stations result from using station and request cards, in pairs, as above.

The order of cards for an input data unit using the ephemeris tape is as follows:

- **(1)** Standard Station Card.
- (2) Request card, as above, with two exceptions:
	- Col. 44 **1 =** ephemeris tape to be used Col. $47 \qquad 1 =$ punch ephemeris data $0 = no$ ephemeris card output
- (3) Blank card to terminate program.

Prediction angles for additional stations for a given vehicle are obtained **by** adding additional station card-request card pairs before the blank card **(3).**

OUTPUT: The basic output is a schedule of predicted lookangles for a particular station at specific times. The report headings are the satellite name; the label LOOK-**ANGLES** FOR, followed **by** the station name; and headings for the sighting coordinates and time data.

> Each data line consists of a day of year, hour, minute, and fraction of minute of the search point; the predicted right ascension, declination, azimuth, and elevation in degrees; the predicted slant range in kilometers; and the elevation and illumination angles of the sun. The sun data indicates if the satellite will be visible at the time and location of prediction.

In addition to the look-angle data output, the pro gram will also produce punched cards containing the satellite' s ephemeris data in the format of the ephemeris card input. The punched cards may be obtained only when the ephemeris information originates from ephemeris tape input. **A 1** punch in column 47 of the first request card for a given satellite will produce one set of ephemeris card output.

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Request Card

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(Point Search Ephemeris)

INPUT DECK: Manual Schedule Tape Mode

INPUT DATA: (1) Request card:

Cols $1 - 5$: T-start in minutes (DDDD.) (minimum period expected)

Cols $6 - 8$: $\triangle t$ in minutes (DD.) (time increment used to increase period)

Cols 9-13: T-stop in minutes (DDDD.)

(maximum period expected)

Col 14 output options:

 $1 =$ long form requested

²= short form requested

- (2) Standard Observation Card
- (3) Standard Station Card(s)

(4) Option Card:

Cols $1-4$: -1 = read another unit of input data

 $(cards 1-4)$

0 or blank = terminate program

OUTPUT: Logical tape ii contains a table of look-angles, for each station requested, in either long or short format. The short form of output consists of lookangle information, time of crossing in days, hours, and minutes of Zebra time; elevation and azimuth angles; and slant range in kilometers. The long form includes, in addition, the Cartesian coordinates of the station and object.

> The TELTYP program may be used if teletype transmission is required.

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Decimal points may be punched anywhere Note: in each of the first three fields.

> Request Card ${\tt POSE}$

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POINT SEAPCH PROGRAM

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STATION LATITUDE = 76.5438 LONGITUDE = 291.6892 EAST HEIGHT = 349. METERS
 ORSERVED TIME = 58. DAYS 10. HOURS 17. WINUTES 47.40 SECONDS GMT
ELEVATION = 64.0000 AyIMUTH = 2.9000 RANGE = 303.0 KM/11
 THETAG = 311.22012 DEGREES
THETAS = 242.90932 DEGREES
 CAPR = 6358.216 KM
GFOCENTPIC STATTIN LATITUDE = 1.33477 PADIANS
 XT = 549.491 TTA = -1381.549 ZFTA = 6181.929 KM
DECL = 1 3509 RA = 0 9951 RANIANS<br>CAPX = -677.104 CAPY = 1323.708 CAP2 = 5181.929 KM<br>THE COOPDINATES OF THE POINT TV THE INEPIIAL REFERENCE=SYSTEM APE
x = -641.181 y = -1258.236 \overline{2} = -6477.534 \overline{1} =BRP66888RBR88888
                         -616\DeltaFHILF GREENLANDSHEWS FIE<br>
FERA TTHE FLEY AZTH R<br>
DAY HR MIN ANG, ANG,<br>
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                                    511 +538*\frac{66}{58} 11 52.79 21.4 293.6
                                    655 +593<sub>4</sub>56.12 2, 79 19, 7 293.6
                                    721 +750+5\bar{6} 12 7.79 18.1 293.7
                         293.9
  56 12 12.79 17.4
                                    779+294.1
  56 12 17.79 16.3
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ORPS

0 (Orbital Plane Search)

INPUT DECK: Manual Schedule Tape Mode

INPUT DATA: The input data is in units consisting of an, element set, a station card, and a request card. Any number of units may occur. The last unit must be followed by a blank card to terminate the program. Each unit will consist of:

- **(1)** Standard element set (6 or 7 cards)
- (2) Standard station card
- (3) Request card:
	- Cols. **1-10:** Start time of search (days of year)
		- **11-** 20: Time increment between search points (minutes)

21-30: End time of search (days of year)

- 31-40: Observing station's search azimuth angle (degrees)
- 41-47: Station' s maximum slant range (kilometers)
- 48-57: Station' s minimum elevation angle (deg:rees)

58: 1 punch-visible passes only.

(4) Blank card, will terminate the program. The card should follow the last repetition of **(1)** to (3).

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OUTPUT: The search point data for a satellite is preceded by the satellite elements, request information, year constants used in the calculations, and the station data. Each line of the search point data consists of the year, month, day, hour ana minute; the right ascension, declination, azimuth and elevation in degrees; the slant range in

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kilometers; and the elevation and illumination angles of the sun in degrees. The information is in order of increasing time during the range requested. If only visible passes were requested, a check is made and only visible points will appear in the output.

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Request Card ORPS

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ORPS52 JOB. **RUN** ORPSB2, DATA 33 13 160 GAMMA 3
33 13 21960 1100..024445879 51.287999 176.37720871 1960 176.37720871 33 13 3 .06497800 291.27649 166.96225 1.04471998 33 13 4-5.3074610E-09- 4.9070522 3.7491586 $-4.1834998E - 05$ 33 13 5-0. $E - -0$. $E -0.$ $E -0.$ $E 33 13 60.$ $0 \mathbb{R}^2$. O \bullet $0 \bullet$ 0300+107330+0616000 140 5 030001TRINDAD 159 $10.191.33333$ $120.$ 190.916670 4000. $0 \bullet$

ENDDATA

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ASUM

(Acquisition File Summary)

INPUT DECK: Manual Schedule Tape Mode INPUT DATA: SEA1 Tape (A-File) OUTPUT: The output consists of the information contained in the A-File of the SEAI tape. The output is sequential according to satellite number and includes satellite number, sensor number, sensor name, pass code (all or visual), format request (short or complete), type (all, 3-point, Baker-Nunn), minimum azimuth, maximum azimuth, minimum elevation, maximum elevation, maximum range, and step size in minutes.

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SSUM

(Sensor File Summary)

INPUT DECK: Manual Schedule Tape Mode

INPUT DATA: SEAl Tape (S-File)

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OUTPUT: A listing of the sensor data contained in the S-File is provided. The output includes sensor number, sensor name, latitude in degrees, longitude in degrees, height in earth radii, $x / \cos \theta$ or $-(C + H)$ cos θ in earth radii, z or -(S + H) sin ϕ in earth radii, accuracy digit for azimuth, elevation and range, classification, teletype request code, and sensor type.

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ISUM

(Information File Summary)

INPUT DECK: Manual Schedule Tape Mode

INPUT DATA: (I) SEAI Tape (I-File)

(2) Request cards with any of the following entries punched in columns 1-3:

- (a) Three decimal digit satellite number. Any number of cards may be used to request specific satellites.
- (b) ALL. The ALL punch requests all satellite information, replacing request card type (a).
- (c) BOX. The BOX punch will provide a box score of still orbiting satellites as additional output.
- (d) END. The END punch signals the end of request deck, and must be used as part of the input deck.

OUTPUT: The I-File of the SEAl Tape is written for off-line printing. Information printed out includes satellite number, satellite name, launch date, launch site, booster country and payload country.

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INFORMATION SUMMARY 09-10-62 09-10-62

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SAT. NUMBER 1984 CLASSIF ICATION COPED SAT. NAME 61.06

THTERNATIL CODE 61.7ETA INTERNAT'L CODE LAJ•NCH DATE 18FES **61** LAUNCH SITE LAUNCH AGENCY BOOSTER COUNTRY PAYLOAD **OON7RY** MISSION WFIGWT **(WG) SHAPE** LENGTH (MTR) HEIGHT (MTR) WLDTH (MTR) DIAMETER (MTR) MEAN DRAG VARTANCE RAD X-SEC (6 MT) **VARIANCE** MEAN REFtECT. VARIANCE **At** TUMB DATE & RATE TUMBL'TNG MODE STABTL IZATION WANEUV. CHAR. TRANS. FREQS

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DECAY DATE **DETERMINED** LIFFTIME (YRS) HELIO. ELMS. E A Q₁ 02 $\overline{\mathbf{I}}$

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INFORMATION SUMMARY

SAT, NUMBER CLASSIFICATION CODED SAT, NAME INTERNAT'L CODE COMMON NAME LAUNCH DATE L'AUNCH SITE LAUNCH AGENCY BOOSTER COUNTRY PAYLOAD COUNTRY MISSION WEIGHT (KG) **SHAPE** LENGTH (MTR) HEIGHT (MTR) WEDTH CHTRY DIAMETER (MTR) **MEAN DRAG** VARIANCE RAD X-SEC (6 MT) VARTANCE MEAN REFLECT. VARIANCE TUMB DATE & RATE TUMBL'TNG MODE STABTL IZATION NANEUV. CHAR. TRANS. FREQS

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TELTYP

(Teletype Output Conversion)

INPUT DECK: Manual Schedule Tape Mode.

INPUT DATA: Tape 11 is scanned for sentinels marking the beginning and ending of messages to be converted to teletype code. All appropriately marked messages will be converted.

Sentinels used:

- **I.** BBBBBBBBBBBBBB (14 B's): this sentinel marks the beginning of a message to converted.
- 2. $*$ $\frac{1}{2}$: the asterisk followed by a dollar sign marks the end of a message to be converted.

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OUTPUT: Tape 11 contains the data converted to Baudot code following output previously written on the tape.

BMEWSPT

(BMEWS Paper Tape)

INPUT DECK: Manual Schedule Tape Mode.

INPUT DATA: Tape 0 should contain: the **0** point data from the DIP as described in the writeup.

OUTPUT: Tape 7 will contain observation information in the standard format required for input to ORCON.

BMEWSPT

(BMEWS Paper Tape)

- INPUT DECK: Manual Schedule Tape Mode.
- INPUT DATA: Tape 0 should contain the **0** point data from the DIP as described in the writeup.

OUTPUT: Tape 7 will contain observation information in the standard format required for input to ORCON.