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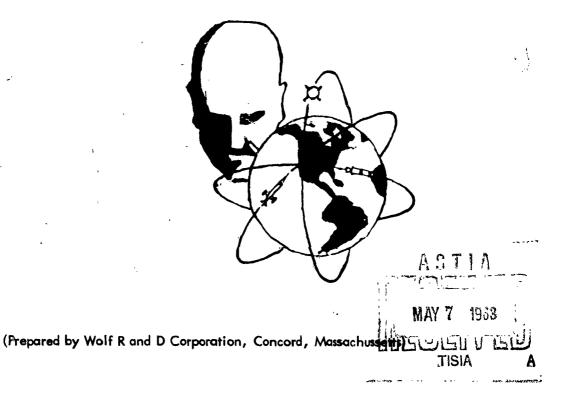
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SPADATS SYSTEM SUPPORT

TECHNICAL DOCUMENTARY REPORT ESD-TDR-63-334 JANUARY 1963

496L SYSTEM PROGRAM OFFICE ELECTRONIC SYSTEMS DIVISION AIR FORCE SYSTEMS COMMAND UNITED STATES AIR FORCE L. G. Hanscom Field, Bedford, Massachussetts



# Best Available Copy

ESD-TOR-63-334 (13) 29450 957 200 S P A D Á T S SYSTEM SUPPORT V. Incl. Illus, fables ASTIA MAY 7 TISIA The programs described in this report were developed under the following contracts: 12) Contracts. AF19(604) 5523. AF19(604) 61.1 and others AF19(604) 6344 AF19(604) 7740 AF19(628) 663 AF19(628) 2051 Wolf **Research and Development Corporation** P.O. Box 136, West Concord, Massechusetts

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PREFACE

This volume contains the complete descriptions and operating instructions for a collection of programs used at the SPADATS Center, Ent Air Force Base, Colorado Springs, Colorado, The collection is primarily useful as an aid to analysts in their task of positively identifying objects detected in space. The documentation does not purport to describe a total system, but rather those programs which were deemed necessary to complement and support an already existing one.

The major portion of these programs were originally developed under the guidance and direction of Dr. Eberhard W. Wahl. Valuable assistance in the analysis of the problems involved was also rendered by Dr. H. Beat Wackernagel, Mr. Edward F. Casey, Mr. Laurence W. Cothbert, Mr. Richard F. Jenney(WRDC), and Mr. Baruch Rosenberg (WRDC). Three of the programs (ASUM, ISUM and SSUM) were developed and written by the Air Force in Colorado Springs.

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#### 1.1 PSR, Position Situation Report (SITRPT)

#### 1.2 Function

The position and other status information on the catalogued satellites is supplied by this program. For a specified time either or both of the following types of reports can be produced:

- a) The Position Situation Report.
- b) The Satellite Situation Report.

The Position Situation Report shows the status of all satellites at a particular time each day. This report is issued in two forms:

- 1. A complete printout prepared for off-line printing through the UBC.
- 2. A truncated version which can be punched on 5-level paper tape for teletype transmission.

Because this report includes information on all known objects, it may contain classified information. The reports are appropriately marked automatically by the computer based on the classification note stored in the Information file of the SEAI tape. See block format in ISUM writeup, section 1-14.

The Satellite Situation Report contains only the information not marked as classified in the Information File.

#### 1.3 Input

In addition to the Schedule Tape control cards, this program requires at least one parameter card (P in col. 80). The first of these parameter cards specifies the following information:

- 1. The output option (col. 1):
  - a) 0 = Position Situation Report
  - b) 1 = Satellite Situation Report
  - c) 2 = Both Reports

- 2. The time at which the report is desired:
  - a) Two digits for hour (cols. 2, 3)
  - b) Two digits for minutes (cols. 4, 5)
  - c) Two digits for day of month (cols. 9, 10)
  - d) Three alphanumeric digits for the name of the month (cols.12 to 14)
  - e)' Four digits for the year (cols. 16 to 19)
  - f) Z punch to signify zebra time (col. 6)
- 3. The Satellite Situation Report output code (col. 23):
  - a) Blank or zero (0) will suppress print of debris.
     A test is made for a one punch in character two of the first word of the I-File
  - b) One (1) will supply information on all satellites.
- 4. The Position Situation Report output code (col. 24):
  - a) Blank or zero (0) perigee and apogee are printed in statute miles.
  - b) One (1) perigee and apogee in kilometers.
- 5. Parameter card indicator (P in col. 80)

Additional parameter cards can be included. These will be treated as comment cards which will be printed immediately following Part I of the Satellite Situation Report.

The Element file and Information file of the SEAI tape supply the data required to compute the position information requested.

#### 1.4 Output

The output on logical eleven (11) contains the following reports:

- 1. The Position Situation Report if the output option zero (0) or two (2) used.
- The Satellite Situation Report if the output option one (1) or two (2) used.

The Position Situation Report supplies the data listed below:

- a) Time and classification of the report.
- b) Satellite name, number and element number.
- c) Latitude and longitude west, both in degrees, at the time of the report.

- d) The orbital elements, including the inclination in degrees, the anomalistic period in minutes, and the eccentricity.
- e) The apogr and perigee in statute miles or kilometers, depending on the output code used in col. 24.
- f) The revolution number at the time of the report and the RA<sub>N</sub>,  $L_N$ ,  $T_N$  of this revolution.
- g) The classification of the satellite this field will be blank if the satellite is unclassified.

The Satellite Situation Report is comprised of the following information:

Part I - Objects in orbit inclusive or exclusive of debris depending on the code used in col. 23 of the first parameter card.

a) Time of Report

b) Satellite name, code name, source and launch date.

- c) The anomalistic period in minutes, the inclination in degrees, the apogee and perigee in statute miles.
- d) Transmitting frequencies, if any.
- e) Comments from the parameter cards which were included as input data, such as:

1961 OMICRON 3-52 FIFTY METALOBJECTS IN PLANE OF 61 OMICRON 1 AND 61 OMICRON 2 ORBITS

Part II - Object removed from orbit.

a) Satellite name, code, name and source.

b) Launch date and decay date.

#### 1.5 Processing

The first parameter card is deciphered in the CONBUF area to determine the time requested for the report. The "debris", output, and code switches are set according to the selection made on this first card.

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The elements are loaded into the EBLOC by the ELMLOD routine. The first entry of the I-File is read and matched against the first element in the EFILE. If the two files are in phase, the computations are initiated for the time requested. The appropriate report or reports are generated according to the internal switches recorded. These switch settings are a result of the paths selected by the first parameter card.

If the IFILE and EFILE do not match, a test is made on the satellite number. A satellite number in the EFILE which is smaller than that in the IFILE results in the comment "NO DATA IN IFILE." These comparisons and the comments generated by them continue until both files are again in phase or empty. This procedure is followed since both files are assumed to be arranged in ascending order by the satellite number.

#### Error Messages

- OVERFLOW AT JA = \_\_\_\_\_. Comment will be printed on off-line output and an exit made to the executive program.
- SUBROUTINE ERROR AT JA = \_\_\_\_\_. This will be printed on off-line output when an error return is made from any subroutine. An exit will be made to the executive program.

#### Formulation

Initial computations from input elements as computed in BEGIN:

$$p_{o} = h_{x_{o}}^{2} + h_{y_{o}}^{2} + h_{z_{o}}^{2}$$

$$W_{x} = h_{x_{o}} / \sqrt{p_{o}}$$

$$W_{y} = h_{y_{o}} / \sqrt{p_{o}}$$

$$W_{z} = \cos i = h_{z_{o}} / \sqrt{p_{o}}$$

$$\sin i = \sqrt{1 - \cos^{2} i}$$

$$i = \tan^{-1} \frac{\sin i}{\cos i}$$

$$\sin \Omega = W_{x} / \sin i$$

$$\cos \Omega = -W_{y} / \sin i$$

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$$\Omega_{0} = \tan^{-1} \frac{\sin \Omega}{\cos \Omega}$$

$$e_{0} = \sqrt{a_{x}^{2} + a_{y}^{2}}$$

$$a_{0} = p_{0} / (1 - e_{0}^{2})$$

$$n_{0} = k_{e} / a_{0}^{3/2}$$

$$c'' = -360 M_{0}^{2} c_{0} / \pi$$

$$q_{0} = a_{0} (1 - e_{0})$$

$$k_{e} \dot{L}_{s0} = k_{e} J^{3} \left\{ 3 - 5 e_{0}^{2} - |\cos i| (1 - 3/2 e_{0}^{2}) - sin^{2} i \cdot (4 - \frac{27}{4} e_{0}^{2}) \right\}$$

$$U_{o} = L_{o} - \Omega_{o} \text{ if } W_{z_{o}} \ge 0$$
$$U_{o} = L_{o} + \Omega_{o} \text{ if } W_{z_{o}} < 0$$

- 2. Compute  $t = (t_i t_o)$ . 1440 where  $t_i$  is time at which position report is requested. Enter XYZSB to get position at time  $t_i$
- 3. Compute  $\theta_G$  at  $t_o$ :

 $\theta_{G} = \theta_{o} + .9856472 \cdot t_{o} (Days) + 360.9856472 \cdot t_{o} (fraction)$ 

4. Enter subroutine SUBPT to get sub-latitude and sub-longitude points:

$$\lambda_{E} = \tan^{-1} (y/x) - .0043752691 \cdot t - \theta_{G}$$
  
$$\phi = \tan^{-1} \left[ \frac{U_{z}}{(1-f)^{2} \sqrt{1-U_{z}}^{2}} \right]$$

5. Compute remainder of output:

PERIOD =  $2\pi/XN$ PERIGEE = a (1 - e) - 1 APOGEE = a (1 + e) - 1

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REV = REV<sub>0</sub> + 
$$\begin{bmatrix} T/P_N \end{bmatrix}$$
 where  

$$P_N = \frac{2\pi}{n} \left\{ 1 - \frac{3}{2} J_2 \left( \frac{a_e}{p} \right)^2 \left[ 3 - \frac{e^2}{2} - \sin i^2 \left( 4 - \frac{3}{4e^2} \right) \right] \right\}$$

$$E_0 = \tan^{-1} \left[ \frac{\sqrt{1 - e^2} a_{y_n}}{e^2 + a_{x_n}} \right]$$

$$M_N = E_0 + \frac{\sqrt{1 - e^2} a_{y_n}}{1 + a_{x_n}}$$

$$\omega = \tan^{-1} \left[ \frac{a_{y_n}}{a_{x_n}} \right]$$

$$M_A = L - \omega \pm \Omega : -\text{if } W_z \ge 0, + \text{if } W_z < 0$$

$$t_N = t_i + \frac{M_N - M_A}{1440 \cdot n} \quad \text{where } \left| M_N - M_A \right| \le \pi$$

$$RA_N = \Omega + \hat{\Omega} \left[ (t_n - t_0) \cdot 1440 - t \right]$$

$$\theta_G = \theta_{G_0} + .9856472 \cdot t_n (days) + 360.9856472 \cdot t_n (fraction)$$

$$L_N = 360 - RA_N + \theta_G \text{ if } RA_N > \theta_G$$

$$L_N = \theta_G - RA_N \text{ if } RA_N < \theta_G$$

$$L_N = 0 \text{ if } RA_N = \theta_G$$

Location	Symbol	Meaning
CCLAS		1, some satellite in I-file is confidential
		0, no confidential satellites
CHAR		Number of non-blank characters in
		transmitting frequencies
CLAS		Classification of satellite, from I file
ELN	Ln	Longitude at node
EMA	M <sub>à</sub>	Mean anomaly at report time
EMN	M <sub>n</sub>	Mean anomaly at node
ENDDT		Decay date
ENDDT1		Decay date
EOESW		END OF ELEMENT Switch
		0, continue reading elements from EBLOC 1, all elements in EBLOC have been processed

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EOS		END OF SAVED ELEMENTS Switch
		0, continues reading elements from SAVE file
		i, all elements from SAVE file have been processed
ERJ		Contains octal address of location where subroutine
		or overflow error occurred
ERJH		L or R indicating which half of ERJ
FLC1		Fractional part of rotation rate of earth
		(deg/solar day) = .9856472
FLC2		Rotation rate of earth (deg/solar day) 360.9856472
FNUM		Number of lines of transmitting frequencies to
		be printed
FSQ	f <sup>2</sup>	$f^2$ where f is flattening of earth = .112381556 x 10 <sup>-4</sup>
FTNOTE		Footnote switch
		0, do not print footnote
		1, print footnote to give units of heliocentric
		satellite output
HCLEM		Array for storage of elements of heliocentric
		satellites from I file.
HCLAS		Classification of satellite
HDCLAS1		Contains first half of classification heading
HDCLAS2		Contains second half of classification heading
HELIO		0, elements available in E file
		1, position uncertain
		2, heliocentric satellite
	q	Perigee distance
HQ1	Q	Apogee distance
HQ2		Year part of object name
HSAT1		Object name
HSAT2		0, or space, always include satellite in Satellite
JUNK		Situation Report
		1, do not include in Satellite Situation Report unless
		output Code 1 is desired.
LNCH		Launch Date
LNCH1		Launch Date
KILOM		0, perigee and apogee output to be in statute miles
		1, perigee and apogee output to be in kilometers

PSR

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	LONGW	L	Longitude (west) of subsatellite point
	OBJNO		Satellite number
	OPT		Output options
			0, output is a Position Situation Report
			1, output is a Satellite Situation Report
		•	2, output is both reports
	OUTPUT		Satellite Situation Report output code
			0, or space, do not print satellites having a "1"
			punch in character 2 of the first word in I file
			1, print all satellites
	PERIOD	Pa	Anomalistic period (in minutes)
	PGCNT	_ d	Page count
	PHI	φ.	Latitude of subsatellite point
	PINCL	i	Inclination
•	RAN	RAN	Right ascension of ascending node
	REV	REV	Revolution number
•	SATI		Satellite number from I file
	SCLAS		1, at least one satellite is SECRET
	• • • •	•	0, no SECRET satellites
	SINDXI		Saves index register 1
	SSLAM	λ <sub>Έ</sub>	Longitude (east) of subsatellite point
	SSLAT	<u>с</u>	Latitude of subsatellite point
•	TCLAS	•	1, at least one satellite is TOP SECRET
	· · ·		0, no TOP SECRET satellites
	TDMON	· · · ·	Table used to convert month name to month
			number
••	TFBUT	.'	Array for output of transmission frequencies
•	TFCNT	· · · ·	Count of number of characters of transmitting
		• •	frequency when packed for output
	THGRN	θG	$\theta_{C}$ at time of report
	TI	t <sub>i</sub>	Time of report in days since 1950
	TIF	<b>1</b> 	Fractional part of TI
	T1H1		Output buffers for time of report
	T1H2		Output buffers for time of report
	T1H3		Output buffers for time of report

T1W TN t<sub>n</sub> TNI TNF TOI UCLAS UPESW

UPISW

UPSSW

Integer part of TI Time of node Time of node (integer)

Time of node (fraction)

Temporary, used in computing  $\theta_{G}$ 

0, no unclassified satellites

1, at least one unclassified satellite E file switch

0, no elements in E file for previous satellite in I file; do not pick up more elements i, read next element from E file

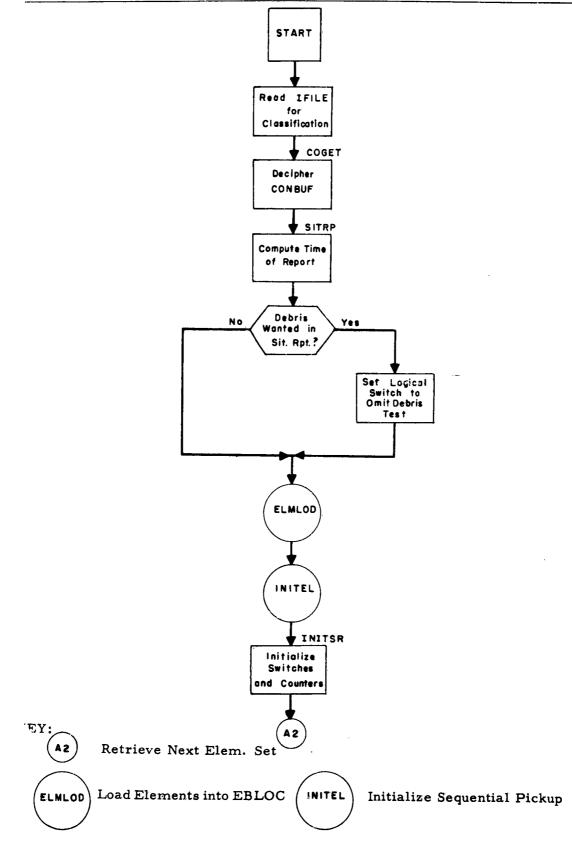
I file switch

0, no elements in I file for previous satellite from E file; do notpick up next element from I file 1, read next satellite from I file.

SAVE file switch

0, no elements in SAVE file for previous satellite; do not pick up next element from SAVE file.

1, read next element from SAVE file



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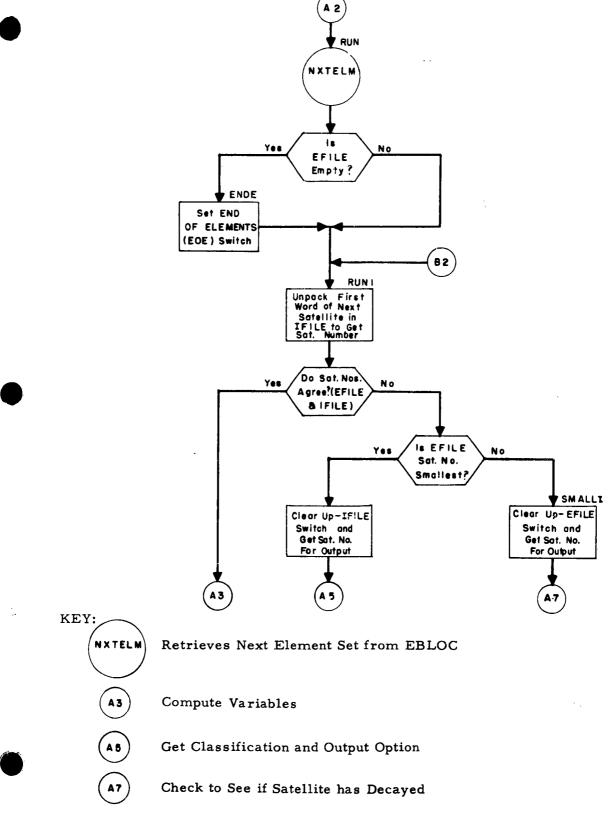
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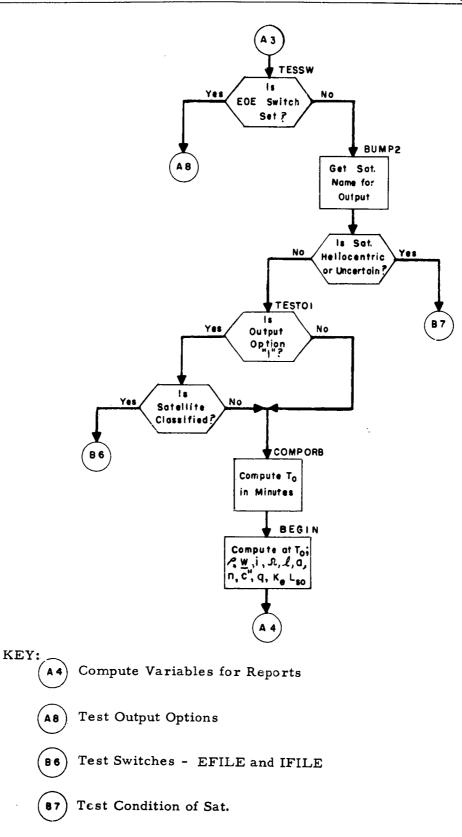




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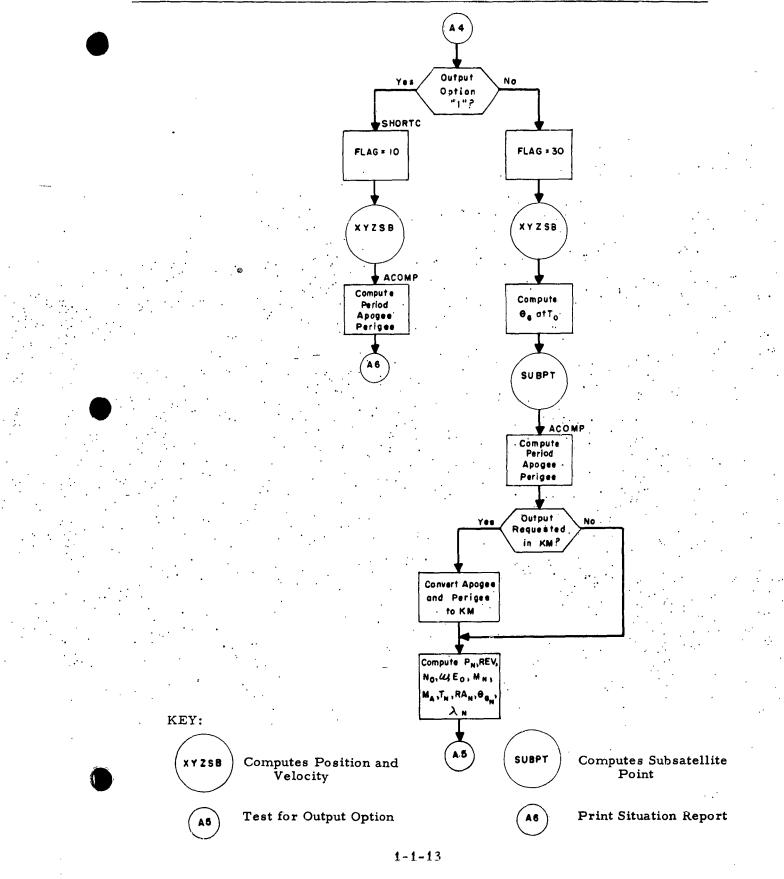
. . . . .

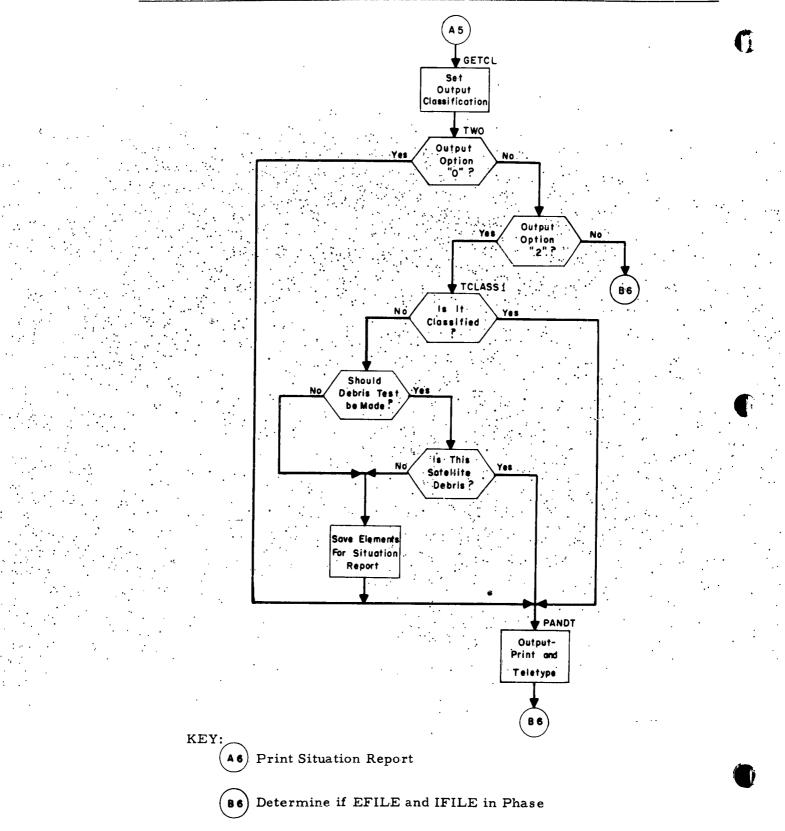
C





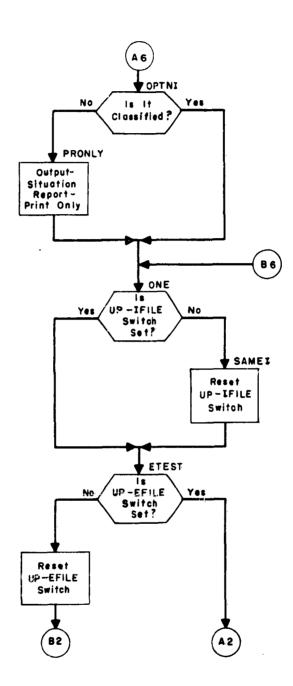








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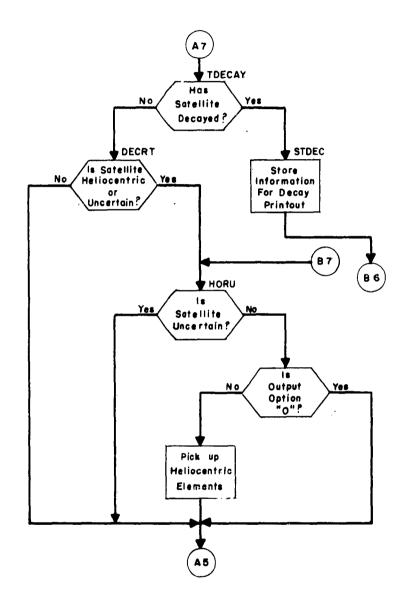


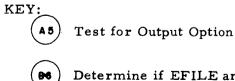
### KEY:

Retrieve Next Element Set

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(B2) Unpack First Word of IFILE to Get Sat. No.



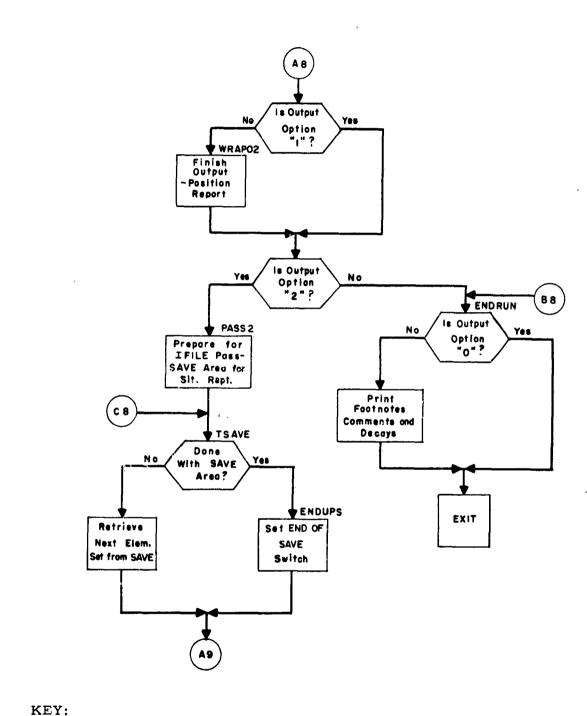


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Determine if EFILE and IFILE in Phase

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1.,





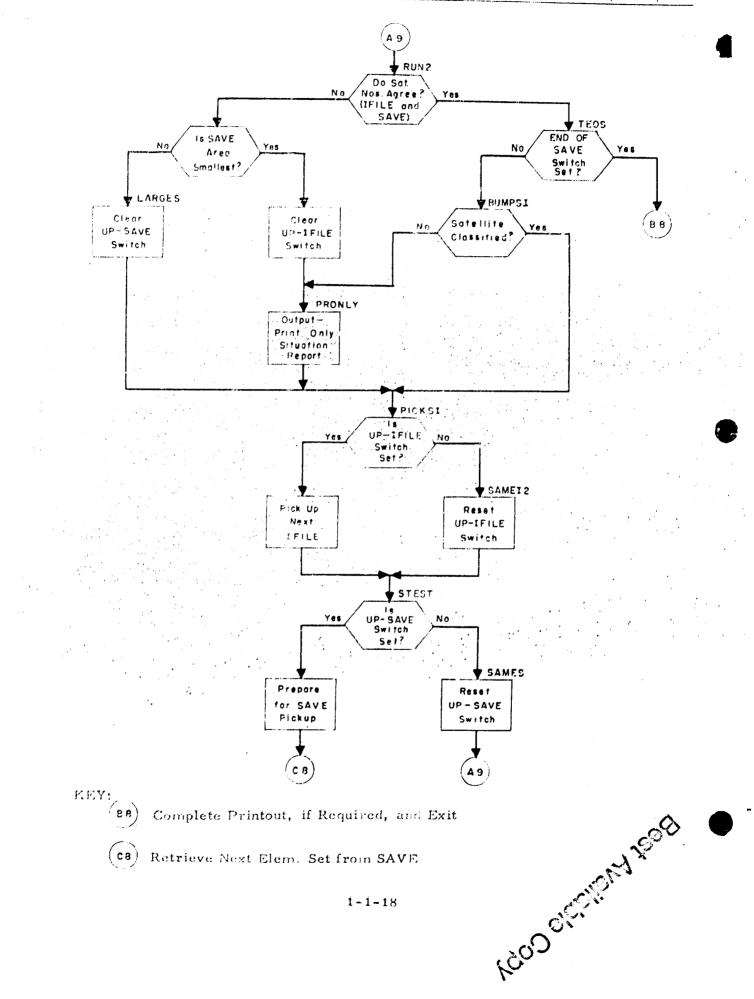
Determine if SAVE and IFILE in Phase

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PSR

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PSR



#### 2.1

REDUCT. Nodal Crossing Reduction

#### 2.2 Function

The Nodal Crossing Reduction Program, hereafter referred to as the Reduction Program, reduces observations to the last nodal crossing; computes certain differences between parameters as computed from the observation and as computed from orbital elements; may compare the differences against predetermined toler ance limits; and produces results in printed and/or teletype format.

The differences, or residuals, may be used by the analyst to determine the necessity of updating existing orbital elements, to determine new elements for satellite debris, to redefine the orbit of a "lost", satellite, or to identify the observation.

The observation may be visual, radar, Baker-Nunn, doppler, or direction finder. The first three are treated by that portion of the Reduction Program known as the General Reduction, and the last two by the Doppler Reduction and Direction Finder Reduction, respectively. The observation types may be intermixed from different sensors, or for different satellites:

Table 2.1 specifies the particular residuals computed by each of the sections of the Reduction Program.

uals	Computed by	Var	rious	Progr
	m:		Rig	htAsco

Т	ab	le	2.	1
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ension Height **Program Section** Residual Time Residual Residual Visual Х Х Radar Х х Х Baker-Nunn Х Х Doppler Х **Direction** Finder Х

Resid am Sections

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For clarification of subsequent sections the following definitions are made:

A predicted quantity (generally indicated with a prime) is the quantity calculated from orbital elements.

A computed quantity is the quantity calculated from the observation.

A known observation is a sighting which has been identified by the sensor as a particular satellite.

An unknown observation (abbreviated UO) is a sighting which has not been identified by the sensor.

A verified observation is a known observation which may agree with the predicted position of the specified satellite.

An unverified observation is a known observation which does not agree with the predicted position of the specified satellite.

A tagged observation is an observation, either known or unknown, which does agree with the predicted position of one or more satellites.

An untagged observation is an observation, either known or unknown, which does not agree with the predicted position of any satellite.

The station tape is a binary magnetic tape containing station (sensor) coordinates for specified stations.

The residual of a quantity is the difference between the computed and predicted quantity.

#### 2.3 Input

Input data may be divided into four groups, some of which are optional. The four are: 1) the switch option card, 2) the element lead card and element sets, 3) the station lead card and station cards, and 4) observation lead card(s) and observations.

#### 2.3.1 Switch Option Card

The switch option card sets the program switches. Setting of a switch is indicated by any non-zero numeric punch in the appropriate card column. The switches and their functions are indicated in Table 2.2.

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<b>Program Switches and Th</b>	neir Functions
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Switch Name	Card Column	Function, if Set
SS1	1	Inhibits the use of the station tape (if set, the station data must then be supplied as card input).
SS3	3	Inhibits the requirement for an interim tape and also eliminates any check of residuals vs. tolerances.
SS8	. 8	Causes all observations to be processed as unknowns, i.e., all are reduced against all available element sets. Known satellite numbers are retained for reference. Tolerance tests will be made unless SS3 is set.
SS9	9	Causes teletype formats to be used in preparation for the TELTYP Program.
SS10	10	Eliminates the rereduction, with open tolerance gates, of untagged observations.
SS11	11	Uses the perigee distance from the element set; if not set, perigee distance is computed.

#### 2.3.2 Element Lead and Element Cards

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The standard element sets required for reducing known observations, as well as the sets considered necessary in attempting to tag unknowns, must be included as a part of the input deck. The element cards must be preceded by the element lead card, identified by having only a seven (7) punch in card column 8. At present, no more than 250 element sets are permitted. Reading of the elements is terminated by either a station or an observation lead card.

#### 2. 3. 3 Station Lead and Station Cards

To reduce observations, the geographic coordinates of the observing station must be available. The coordinates may be read from the station tape (if program switch 1 is not set) and/or from standard station cards. Each group of station cards must be preceded by the station lead card, identified by having only an eight (8) punch in card column 8. At present, no more than 750 stations may be used.

#### 2.3.4 Observation Lead and Observation Cards

The observational data, the primary input to the Reduction Program, must be in the standard observation format. Each group of observations must be preceded by an observation lead card, identified by the nine (9) punch in column 8. Column 7 of the lead card is also used to specify the tolerances to be used by the program, where applicable. Table 2.3 specify the codes and their corresponding values. Lead cards to change the tolerances may precede observations anywhere in the observation card deck. There is no limit to the number of observations which may be processed, since they are processed individually.

A blank observation card will terminate the reading of the preceding group of observational data. A card with only a nonzero numeric punch in column 79 will terminate the program.

Code	Time (days)	Right Ascension (degrees)	Height (km)				
$\begin{bmatrix} Blank \\ 0 \end{bmatrix}$	.002	20	200				
1	1	5	500				
2	. 003	360	10000				
3	. 002	360	10000				
4	. 001	360	10000				
5	. 003	5	300				
_6	. 05	2. 5	200				

Tolerance Codes and Corresponding Values

#### 2.4 Output

Reduction results may be obtained in the form of teletype paper tape and/or printed output, depending upon the setting of program switch SS9 (cf. Table 2.2). Headings are also a function of the teletype option. The output consists of the satellite inventory, output preliminary to the results, and the results themselves.

#### 2.4.1 Satellite Inventory

The satellite inventory, a listing of the satellite and element numbers used, is written following the termination of the reading of element sets. It is identified by the heading SATELLITE INVENTORY FOLLOWS.

The satellite and element numbers are printed in groups, seven to the line, each group separated from the next by a slash.

2. 4. 2 Output Preliminary to Reduction Results

The elements and tolerances in effect for the results which will follow appear in five lines, as follows:

Line 1. The satellite number

Line, 2: Epoch revolution number, epoch time, nodal period (days), its first and one half of its second deriva-

tive with respect to time; the semi-major axis, and eccentricity

Line 3. Right ascension of the ascending node, its first and

one half of its second derivative with respect to time, nodal period (minutes), and inclination

Line 4. Blank

Line 5. The tolerances (time, right ascension of the ascending node, and height)

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A summary of the five lines in symbolic notation follows:

Satellite No.

N	t	PN	с	Ь	a	Ç
Ω	Ω		ω	் ப	ü	P Nm
Δt	$\Delta RA_t$	ΔH				

2.4.3

#### Results for Tagged or Known Observations

The output format and quantities calculated depend upon the teletype option and upon the type of observation. If teletype output is desired, all headings are omitted and the output is placed in two lines. However, the quantities contained in the two lines are identical with those which would have been printed in one line, had the teletype option not been selected. In the following subsections the description will apply to a case in which teletype output is not requested.

As noted above, the headings and quantities computed are a function of the observation type. Each heading is described below. The headings will be printed each time a different group of observation type is encountered; that is, each time the observation type requires the use of a different section of the program.

#### 2.4.3.1 Output Common to All Observation Types

The output from each of three portions of the Reduction Program begins with two common quantities, the identification, ID, and the epoch revolution, N. The identification is composed of two eight character words which may be symbolically represented by

#### SATYMMDD HHMMSSs.

where

SAT	=	satellite number
Y	=	last digit of the year
MM	=	month of the year
DD	=	day of the month

2.4.3.3

HH = hour of the day
MM = minutes of the day
SS = seconds of the day
ss = hundredths of seconds

If the teletype ending sentinel, \$, appears at the end of any line of output, the observation time precedes epoch by more than four days.

#### 3.2 Dimensions of the Output Quantities

Unless otherwise indicated the dimensions of the output quantities are in time in days, distance in kilometers, and angular measure in degrees.

#### Output from the General Reduction Program

Following the epoch revolution number are the argument of latitude, U, of the satellite; computed time of nodal crossing, T SUB N; its residual, DELTA T; the latitude, PHI S, and longitude, L S, of the sub-satellite point; the computed right ascension of the ascending node, RA N; its residual, DEL RA; the computed satellite height, H(KM); its residual, DEL H; observation type, TYPE; element number, ELEM; and station number, STA. If the observation type is visual, VIS, or Baker-Nunn, B-N, the height residual will appear as 0-0. If the type is radar, RDR, the computed quantity will be printed.

#### 2.4.3.4 Output from the Doppler Reduction Program

Following the epoch revolution number are the computed time of nodal crossing, T SUB N; its residual, DELTA T; the great circle distance, D, in nautical miles from the station to the subsatellite point; the elevation angle, H; the slant range, S; the

element number, ELEM; and the station number, SAT. The title (DOPPLER) is appended to the heading, or in case the teletype option is selected, the symbol, DOP, follows the station number.

#### 2.4.3.5 Output from the Direction Finder Program

Following the epoch revolution number are the computed time of nodal crossing, T SUB N; its residual, DELTA T; the slant range, S; the computed satellite height, H(KM); the element number, ELEM; and the station number, STA. The comment (DIR FINDER) is appended to the heading, or in case the teletype option is selected, the symbol DF follows the station number.

#### 2.4.3.6 Observation Comments

The observation comments which may appear in the printed output and their explanations are given below. Each observation comment is preceded by the 16-digit observation identification.

Observation comments (1)-(3) may be printed for both known and unknown observations.

- (1) UNDEFINED OBSERVATION (2). Illegal equipment type specified.
- (2) NO STORED COORDINATES (2). The station coordinates of the observing station are not available.
- (3) IMPROPERLY LOGGED. An error exists in the observation, e.g., hours greater than 24, an elevation angle greater than 90 degrees, etc.

Observation comments (4)-(13) may be printed for known observations only. Each message is followed by the station number.

(4) NO ELEMENTS IN SYSTEM (1). The elements corresponding to the satellite designated in the observation have not been read.

- (5) QPRIME LESS THAN OR EQUAL TO ONE (1). The computed perigee distance is less than or equal to one earth radius.
- (6) DID NOT MEET RESIDUAL TOLERANCES (3). The observation did not meet the specified residual tolerances.
- (7) BAKER-NUNN OBS W/+RA (5). An observation is encountered in the General Reduction Program with an equipment type that indicates Baker-Nunn, but the observation does not have the required minus overpunch in column 31 of the observation card. If the observation is unknown, it becomes an untagged UO.

If the observation is unknown, it becomes an untagged UO.

- (8) ELEVATION NEGATIVE (7). The elevation computed from the declination by the General Reduction Program is negative. If an unknown observation, it becomes an UNTAGGED UO.
- (9) PHI S GREATER THAN I (8). The computed sub-satellite latitude exceeds the inclination.
- (10) DF AZ NEGATIVE (6). A Direction Finder observation has a negative azimuth.
- (11) DF ELEVATION NEGATIVE (8). A Direction Finder observation has a negative computed elevation.
- (12) NON-CONVERGENT OBS (9). The computation of the satellite latitude does not converge within the specified number of iterations.
- (13) SLANT RANGE (ZERO) (0). An observation with a zero slant range has been encountered by the General Reduction Program radar portion. If the observation is unknown, it becomes an untagged UO.



(14) UNTAGGED UO. An unknown observation has been reduced against all the element sets and remains untagged within the specified tolerances. If input options specify, the observation will be written on the interim tape and rereduced with large tolerances.

### 2.4.3.7 Miscellaneous Printed Output

The various comments which may appear in the printed matter and their explanations follow.

(1) PN MISSING SO PA USED FOR SAT XXX ELEM XXX

The nodal period at epoch is missing from the sixth card of a seven card set. The anomalistic period has been used in its place.

(2) SATELLITE XXXX CARD XXX OUT OF ORDER

An element card is not in correct sequence, or an extraneous card is contained in the element sets. The elements for the particular satellite will not be stored in computer memory, hence are unavailable.

(3) UNTAGGED UOS REDUCED W/O TOLERANCES All observations which follow have previously appeared as untagged UO's, and have now been rereduced with large tolerances.

(4) END OF RUN \*\*\*\*\*\*\*\*\*

All observations have been reduced with a specified tolerance. Any results which follow have been reduced with different tolerances.

(5) UO S FOLLOW

All observations which follow on the page did not correspond to any satellite, the elements of which were successfully stored in computer memory.

## (6) UNK \*\*\*\*\*\*\*\*\*

The next observation is an unknown. The message number is given. If the observation was a known observation which was treated as an unknown, the satellite number is also given.

#### 2.4.3.8 Flexowriter Output

The messages which may appear on the flexowriter, and their explanations follow.

### (1) SATELLITE XXXX CARD XXX OUT OF ORDER

An element card is not in correct sequence, or an

extraneous card is contained in the element sets. The elements for the particular satellite will not be stored in computer memory, hence are unavailable.

(2) MOUNT RIGHT STATION TAPE

The station tape is not available to the program. The program awaits the stop-go option.

(3) ILLEGAL LEAD CARD SKIPPED

An illegal lead card (cf. Sec. 2.3) is present in the input deck. The program continues according to the last valid lead card.

(4) MOUNT STATION TAPE ON 7

The program is ready to run, but a stop-go option is provided in case the station tape is required.

(5) REMOVE STATION TAPE U7

The program run is complete, but a stop-go option is provided to give time to remove the station tape.

### 2.5 Processing

The program accomplishes its function of reducing observational data to the last nodal crossing, one observation at a time. However, certain preliminary operations must be performed prior

to operating on the first observation, and as circumstances require, prior to operating on subsequent observations.

## 2.5.1 Preliminary Operations

The preliminary operations are initialization of various program switches, examination of switch options, reading of station coordinates from the station tape, reading of satellite elements from cards, reading of station data from cards, and the reading of the first observation lead card.

### 2.5.1.1 Initialization of Program Switches

Certain switches which will control the path of the program are initialized to neutral settings.

## 2.5.1.2 Reading of Switch Option

The several program switch settings are read from the switch option card (cf. Sec. 2. 3. 1). These settings determine, among other things, whether or not the station tape is to be read and the interim tape rewound.

2.5.1.3 Reading the Station Tape

If switch SS1 is set the station tape is examined; and if it is identified as the station tape, the values of station number, latitude, longitude, and height are read into the appropriate array. If the identification test fails, a STOP-GO option is provided to allow the computer operator to mount the correct tape. GO causes the program to examine the newly mounted tape for the proper identification. STOP causes the program to return control to the executive routine.

### 2.5.1.4 Reading the Element Sets

The element sets are generally read as a single group by the subroutine ELRED. The program logic also permits their

being read in separate groups, providing they are identified by an element lead card. Certain validity checks are made upon each element set. If no gross errors exist, the elements are stored in the element array for future use. If errors do exist, appropriate error messages are generated and the element set is otherwise ignored.

Reading of element sets is terminated by either a station lead card or an observation lead card. Termination of reading causes the satellite inventory, a listing of those satellites whose elements were successfully read, to be generated as the first output of the program.

### 2.5.1.5 Reading of Station Data from Cards

If a station lead card is encountered during the running of the program, the data from the station card or cards which follow the lead card are stored in the station data array. Reading of the station cards is terminated by either an element lead card or an observation lead card.

## 2.5.1.6 Reading of an Observation Lead Card

Each observation or group of observations must be preceded by an observation lead card, which also specifies the tolerances to be used, where applicable, to cause the tagging of unknown observations. Table 2.3 specifies the codes used in card column 8 of the lead card and the corresponding values of the tolerances. Following the reading of the first observation lead card the processing of observations can begin.

### 2.5.2 Preliminary Processing for all Observations

The observations are read and processed individually. Prior to the actual reduction calculations, tests are made for the validity of the observation, the station coordinates are retrieved, the appropriate elements are located and the observation type determined.

### REDUCT

### 2.5.2.1 Testing of the Observation

Certain fields from the observation card are checked for validity. These fields are month, day, hour, minute, second, elevation angle, and azimuth angle. An error in any one of the fields will cause the rejection of the observation with the appropriate comment (cf. Sec. 2. 4. 2. 6 (3)). The program proceeds to read another observation card.

### 2.5.2.2 Retrieval of Station Coordinates

If the several fields of the observation card contain no illegal values, the station number from the observation is compared to that from the previous observation. If the two numbers do not agree, the station data array is searched for the current station number. If the station number is found, the station latitude, longitude, and height are retrieved, and station dependent quantities are computed. If the current station number is not found in the station data array, the observation will be rejected with the appropriate comment (cf. Sec. 2. 4. 3. 6 (2)). The program proceeds to read another observation.

If the two station numbers do agree, the latitude and longitude are examined for zero which would indicate that the station coordinates were not previously found. The observation is then rejected with the appropriate comment.

### 2.5.2.3 Retrieval of Satellite Elements

If the observation is a known observation, the satellite number is compared to that of the previous satellite. If they agree, then the elements have already been retrieved and no search of the element array is necessary. If the satellite number is not the same as that of the previous satellite, the element array is searched via subroutine ELCAL3. If the search is successful the nodal elements

are printed (cf. Sec. 2.4.2). If the elements are not found the observation is rejected with the appropriate comment (cf. Sec. 2.4.3.6 (4)). The program proceeds to read another observation.

If the observation is an unknown observation, then the subroutine ELCAL3 will sequentially retrieve the elements from the element array until the observation has been reduced against all elements.

Subsequent to the retrieval of the elements, perigee distance for the revolution at the time of the observation is computed. If perigee distance is less than one earth radius, computations will either cease in the case of a known observation, or continue with the next element set for an unknown observation. The appropriate comment (cf. Sec. 2. 4. 3. 6 (5)) is printed for the known observation, and another observation is read.

If the perigee distance computation is satisfactory, the actual reduction computation is ready to be made on the basis of the observation type.

2.5.2.4 Determination of Observation Type

The observation type, a part of the observation itself, is used to determine which of the three subsections of the program is to be used for the reduction computations. An illegal observation type will cause the rejection of the observation with the appropriate comment (cf. Sec. 2.4.3.6 (1)).

2.5.3 Processing With the General Sighting Routine

The General Sighting program processes only observations in which azimuth and elevation angles, or right ascension and declination are given. Slant range may also be utilized. The general observation types which fulfill this requirement are visual, radar and Baker-Nunn camera observations.

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# 2.5.3.1 Special Processing of Visual Observations

Visual observations may be either azimuth and elevation angles or right ascension and declination. If the right ascension and declination are given, they are converted to azimuth and elevation angles. A negative elevation angle computed from a known observation causes an error message (cf. Sec. 2.4.3.6 (8)) to be generated and the program proceeds to process the next observation. If the observation is unknown, it becomes an untagged UO. Computation of a positive elevation angle causes processing to continue as indicated in Sec. 2.5.3.3.

# 2.5.3.12 Special Processing of Baker-Nunn Camera Observations

Baker-Nunn camera observations must have right ascension and declination given. These are converted to azimuth and elevation angles and the processing continues as indicated in Sec. 2. 5. 3. 3. If the overpunch signifying right ascension and declination is missing, an error message (cf. Sec. 2. 4. 3. 6 (7)) is generated for a known observation or the unknown observation is made an untagged UO. In either case, the program proceeds to the next observation.

3 Common Processing of Visual and Baker-Nunn Camera Observations

Neither visual nor Baker-Nunn camera observations have any range information. For this reason an iterative procedure is used to determine the sub-satellite position. The station's latitude and sidereal time are taken as first approximations to the satellite's latitude and right ascension, respectively, on the assumption that the satellite is close to the station. Failure of the calculations to converge causes an error message (cf. Sec. 2. 4. 3. 6 (12)) to be generated for known observations. The program proceeds to the next observation for the known observation or to the next element set for an unknown observation. If the iteration produces the 2.5.3.4

satellite latitude within the specified limit, the latitude is checked against the inclination angle. An unsatisfactory result causes an error message (cf. Sec. 2. 4. 3. 6 (9)) to be generated for a known observation and the program proceeds to the next observation. For an unknown observation the next element set is selected.

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If the computed latitude can be on the orbital plane, processing continues as indicated in Sec. 2.5.3.5.

## Special Processing of Radar Observations

The examination of the azimuth and elevation angle field for a radar observation is identical to that described for a visual observation (cf. Sec. 2.5.3.1). However, radar observations include a range measurement, and this is used to determine the subsatellite position. An illegal range, i. e., negative, causes the appropriate message (cf. Sec. 2.4.3.6 (13)) to be generated for a known observation, and causes an unknown observation to become an untagged UO. In either case, the program proceeds to process the next observation.

If the range is valid, the satellite's latitude and right ascension are computed. The latitude is compared with the inclination angle. If the test is satisfactory, the height residual is computed. Processing then continues as in Sec. 2. 5. 3. 5. If not an error message (cf. Sec. 2. 4. 3. 6 (9)) is generated for a known observation, and the program proceeds to process the next observation. If the observation is unknown, the program proceeds to the next element set.

### 2.5.3.5 Common Processing of all General Sighting Observations

The longitude of the satellite, right ascension of the node, the residuals of time and right ascension of the node are computed for all general sighting observations. At this point the observation

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has passed the various gross tests and the residuals, be they large or small, have been computed. The setting of switch option SS3 determines whether or not the residuals for an unknown observation should be compared with the tolerances. If the switch is not set and the observation is unknown, the comparison is made. Failing the comparison causes the program to select the next element set; passing causes the observation to be tagged. If the switch SS3 is set, the tolerances are ignored. The observation and the reduction results are then output for printing. The program proceeds to process the next observation on the same observation with a new element set, depending upon whether or not the current observation is known or unknown.

## 2.5.4 Processing With the Doppler Routine

The Doppler portion of the program processes passive track Doppler observations and Fence observations in which range rate is the only information available. Consequently, only the time residual is computed.

Preliminary calculations concerning the relative position of the ascending node and the station are made to determine certain constants for subsequent calculations. The argument of latitude, true anomaly, eccentric anomaly, and mean anomaly of the satellite are intermediate quantities used in determining the time residual, the elevation angle and the slant range. The setting of switch SS3 determines whether or not the time residual for an unknown observation should be compared with the time tolerance.

## 2.5.5 Processing With the Direction Finder Routine

AREA CREW STR. AVAILABLE

The Direction Finder portion of the program processes observations in which the azimuth angle of the closest approach is given. For this reason only the time residual is computed.

The observation will be rejected with the appropriate message (cf. Sec. 2. 4. 3. 6 (11)) if the azimuth is negative. If it is positive the time residual, the elevation angle and the slant range are computed. A negative elevation angle will cause an error message (cf. Sec. 2. 4. 3. 6 (10)) to be generated for a known observation. The program proceeds to the next observation or the next element set depending upon whether the negative elevation angle was computed from a known or an unknown observation.

If the elevation angle is positive, the setting of switch SS3 determines whether or not the time residual for an unknown observation should be compared with the time tolerance. If the comparison is made and the residual exceeds the tolerance, the program proceeds to select the next element set. Otherwise the observation is tagged and the reduction results for both the tagged unknown or known observation are output for printing. The program proceeds to process the next observation or the same observation with a new element set depending upon whether or not the current observation is known or unknown.

### 2.5.6 Optional Reprocessing of Untagged Unknowns

Unknown observations which have failed to be tagged may be reprocessed with large tolerances, depending upon the setting of switches SS3 or SS10. If neither are set, unknown observations which have not been tagged are written on the interim tape following the processing with the last available element set.

After all observations have been initially processed, the tolerances are set to large values. The observations are then read from the interim tape, one at a time, and reprocessed with all element sets.

# C

# General Sighting Reduction Program

# Input:

Orbital elements at epoch:

То	time of node	N	epoch revolution
P	nodal period in days/rev.	ω	argument of perigee
с	rate of change of period	ω	rate of change of $\omega$
d	rate of change of c	ω	rate of change of $\dot{\omega}$
Ω.	right ascension of ascending	q	perigee distance
•	node	• q	rate of change of q
Ω	rate of change of $\Omega$	ä	rate of change of q
Ω	rate of change of $\Omega$	. т <b>′</b>	epoch of decay
i	inclination of orbital plane	- 0	equation
e	eccentricity		
· .	•		

# Sighting data:

ф	Latitude of station
Ψ \	Longitude of station west
~ . TT	
H	height of station
t <sub>i</sub>	time of observation
either	a right ascension
• .	δ declination
or	Az azimuth
	h elevation
	•

# 2.6.1 Formulation for the General Sighting Reduction Program 1) $\Delta t = t_i - T_0$ 2) $RA' = RA + RA \Delta t + \frac{1}{2} RA (\Delta t)^2$ $\omega' = \omega + \dot{\omega} \Delta t + \frac{1}{2} \dot{\omega} (\Delta t)^2$ 3) $P_a = 360^\circ P/(360^\circ - \dot{\omega} P)$ 4) $a = (P_a / .058672947)^{2/3}$ 5) $P'_{a} = P_{a} + P \Delta t + 1/2 P (\Delta t)^{2}$ 6) $a' = (P'_a / .058672947)^{2/3}$ q' = a(1 - e), Go to 87) If q not given: $\Delta t' = T'_o - t_i$ If q given: $q' = q + q \Delta t' + 1/2 q' (\Delta t)$ If q' = 1: stop 8) If q' > 1 : e' = (a' - q)/aIf $e' < 0 : 0 \neq e'$ 9) 10) $e'' = \sqrt{\frac{1-e'}{1-e'}}$ 11) $\theta_z = [\theta_0 + .98565 (Day of Year of Obs)]$ + 360. 98565 (Fract. of day) + $\lambda$ ], $0 \le \theta_{y} < 360$ 12) If a given convert a from hours to degrees

REDUCT	Wolf Research and Development Corporation
2.6.2	The following Equations are for Baker-Nunn and
	Visual portions only. Radar is listed separately
13)	If this is a Baker Nunn observation: Go to 14 If not: Go to 20
14)	If $\delta < -23^{\circ}$ If $\delta > -23^{\circ}$
	$a_{obn} = .53123888^{\circ} - a$ $a_{obn} = .65918611^{\circ} - a$
	$Z_{bn} = .53138889^{\circ}$ $Z_{bn} = .65941945^{\circ}$
	$\theta_{bn} = .46215278^{\circ}$ $\theta_{bn} = .57353333^{\circ}$
	$\rho = \sin\theta_{bn} \left( \tan \delta + \tan \frac{\theta_{bn}}{2} \cos a_{obn} \right)$
16)	$a_{bn} - a_{obn} = tan^{-1} (\rho \sin a_{obn} / (1 - \rho \cos a_{obn}))$
17)	$a = -(a_{bn} + Z_{bn})$
18)	$\gamma = \cos \left( \left( a_{bn} + a_{obn} \right) / 2 \right) \cdot \tan \left( \frac{\theta_{bn}}{2} \right) \sec \left( \left( a_{bn} - a_{obn} \right) / 2 \right)$
19)	$\delta = 2 \tan^{-1}(\gamma) + \delta$
20)	If a given: Go to 21
	If not: Go to 25
21)	$H' = \theta + \alpha$ ("a" as in input)
22)	$h = \sin^{-1} (\sin \phi \sin \delta + \cos \phi \cos \delta \cos H')$
	If $h < 0$ , Return for next obs.
23)	$\sin (Az) = -\cos \delta \sin H' / \cos h$
24)	$\cos (\mathbf{A}z) = (\sin \delta - \sin \phi \sin h) / \cos \phi \cosh h$
25)	$\phi' = \tan^{-1} (.99329985 \tan \phi)$
26)	$RA_s = \theta$
	$\phi'_j = \phi'$
27)	$R_o = .9966443/[100670015 \cos^2 \phi_j]^{1/2} + H/6378.174$
28)	If $\phi'_j \ge i$ : $u = 90^\circ$ (X sign of $\phi'_j$ ), Go to 31
	If $\phi'_j < i$ : Go to 29 1-2-22

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29)	$\tilde{u'} = \sin^{-1} (\sin \phi'_j / \sin i)$
30)	If $\cos (RA_s - RA') < 0: u = 180 - u'$
	If cos $(RA_s - RA') \ge 0$ : $u = u'$
31)	$v = u - \omega'$
32)	$r = a'(1 - e'^2)/(1 + e'\cos v)$
33)	$H'' = r - R_o$
34)	a = 90 - sin <sup>-1</sup> (cos h/(1 + H"))- h
35)	$\phi'_{j+1} = \sin^{-1} (\cos a \sin \phi + \sin a \cos \phi \cos A^z)$
36)	$\Delta L = \sin^{-1} (\sin a \sin Az / \cos \phi'_{j+1})$
37)	$RA_s = (\theta + \triangle L),  0 \leq RA_s \leq 360$
38)	If $( \phi_{j+1} - \phi_j  - 10^{-4}) < 0$ : Go to 39
	If $\ge 0$ , $j + 1 \rightarrow j$ : Go to 27 unless observation does not converge within the specified number of iterations. In this case, return for next observation.
39)	If $\phi_{j+1} \ge i$ : Return for next observation If not, $\phi_{j+1} \rightarrow \phi_s$ : Go to 40
40)	$L_s = (\lambda + \Delta L), 0 \leq L_s \leq 360$
41)	If $i = 90^{\circ}$ , $\Delta \lambda' = 0$ : Go to 43
42)	$\Delta \lambda' = (\sin^{-1} (\tan \phi_s / \tan i))$
43)	If $\cos u < 0$ : $\Delta \lambda = \Delta \lambda' - \pi$
	If $\cos u \ge 0$ : $\Delta \lambda = -\Delta \lambda'$
44)	$RA_N = (RA_s + \Delta \lambda), 0 \le RA_N \le 360$

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45) 
$$E_s = 2 \tan^{-1} (e'' \tan v/2)$$
  
46)  $E_{RA} = -2 \tan^{-1} (e'' \tan (\omega'/2))$   
47)  $M_s = E_s - e' \sin E_s$   
48)  $M_{RA} = E_{RA} - e' \sin E_{RA}$   
49)  $T_n = t_i - ((M_s - M_{RA})/360) P'_a, 0 \le (M_s - M_{RA}) \le 360$   
50)  $\Delta N = (T_n - T_0)/P'_a$  (the integral portion thereof)  
51)  $T_x = T_0 + P \Delta N + c (\Delta N)^2 + d (\Delta N)^3$   
52) If  $|T_x - T_n| < P/2$ : Go to 53  
If not, then when:  $(T_x - T_n) > 0$ , set  $\Delta N = \Delta N - 1$ , Go to 53  
 $(T_x - T_n) = 0$ , Go to 53  
 $(T_x - T_n) < 0$ , set  $\Delta N = \Delta N + 1$ , Go to 53  
 $(T_x - T_n) < 0$ , set  $\Delta N = \Delta N + 1$ , Go to 53  
53)  $RA_x = RA + RA (T_x - T_0) + \frac{1}{2} RA (T_x - T_0)^2$   
54)  $\Delta T_n = T_n - T_x$   
 $\Delta RA = RA_n - R_x$ 

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Output :

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N, 
$$u^{0}$$
,  $T_{n}$ ,  $\Delta T_{n}$ ,  $\phi_{s}$ ,  $L_{s}^{0}$ ,  $RA_{n}^{0}$ ,  $\Delta RA_{n}^{0}$ , H<sup>"</sup>, Km

1-2-24

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## 2.6.3 Radar Portion of General Sighting Reduction Program

### Input:

Same as for other portions of General Sighting Reduction Program with exception that  $\rho$  (slant range) is included as part of sighting data.

Equations:

1) - 12)Same as in G. S.R.P Sames as 20 - 26 in G. S. R. P 13)-19)  $R_0 = [.996643/(1 - .00670015 \cos^2 \phi_i)]^{1/2} + H/6378.174$ 20)  $R'' = [R_0^2 + (\rho/6378.174)^2 + 2R_0(\rho/6378.174) \sin h]^{1/2}$ 21)  $a = \sin^{-1} (\rho \cosh h/6378.174 \text{ R}'')$ 22) If a < 0: Return for next observation 23) If  $a \ge 0$ : Go to 24  $\dot{\phi}_{c} = \sin^{-1} (\sin \phi \cos a + \cos \phi \sin a \cos Az)$ 24)  $\Delta L = \sin^{-1} (\sin a \sin Az / \cos \phi_s)$ 25)  $\cos \Delta L = (\cos \alpha - \sin \phi' \sin \phi'_{s} / \cos \phi' \cos \phi'_{s})$ If  $\cos \Delta L < 0$ :  $\Delta L = 180 - \Delta L$ : Go to 26 If  $\cos \Delta L \ge 0$ : Go to 26  $RA_s = (\theta + \Delta L), 0 \leq RA_s \leq 360$ 26) If  $\phi_s > i$ : Return for next observation 27) If not: Go to 28  $u' = \sin^{-1} (\sin \phi_s / \sin i)$ 28)

Output:

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29) If  $\cos (RA_s - RA') < 0$ : u = 180 - u'If  $\cos (RA_s - RA') > 0$ : u = u'30)  $v = u - \omega'$ 31)  $r = [a'(1 - e'^2)/(1 + e'\cos v)] 6378.174$ 32) - 46) same as 40 - 54 in G.S.R.P

Same as in other G.S.R. P. except that R'', observed height, is included.

Equations:

# 2.6.4 Direction Finder Reduction Program

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# Input:

Same as for General Sighting Reduction Program except that elevation - declination is not given.

Equations.	
1) - 11)	Same as in G. S. R. P
12)	$\lambda_n = \theta - R'$
13)	$\Delta \lambda = \lambda_n + \lambda$
14)	$(u + \Delta u) = \tan^{-1} (\tan \Delta \lambda / \cos i)$
15)	If $\cos \Delta \lambda \ge 0$ : Go to 17
	If $\cos \Delta \lambda < 0$ : Go to 16
16)	If $(u + \Delta u) \geq 0$ : $(u + \Delta u) = (u + \Delta u) - 180$
	If $(u + \Delta u) < 0$ : $(u + \Delta u) = (u + \Delta u) + 180$
17)	$\phi'_{\lambda} = \sin^{-1} (\sin (u + \Delta u) \sin i)$
18)	$\phi' = \tan^{-1} (.99329985 \tan \phi)$
19)	$\Delta \phi' = \phi_{\lambda}' - \phi'$
20)	If $Az < 0$ : Return for next observation
	If $Az \ge 0$ : Go to 21
21)	If $Az - 180 < 0$ : $Az_i = Az$
	If $Az - 180 \ge 0$ : $Az_i = Az - 180$
22)	If $Az_i' - 10^{-5} < 0$ : Go to 24
	If $Az'_i - 10^{-5} \ge 0$ : Go to 23

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23)	If $ \Delta \phi'  - 10^{-5} < 0$ : Go to 24 If $ \Delta \phi'  - 10^{-5} \ge 0$ : Go to 25
24)	$u = u + \Delta u$ : Go to 35
25)	$\sin \eta = \sin \Delta \lambda / \sin (u + \Delta u)$
26)	$\cos \eta = \tan \phi_{\lambda} / \tan (u + \Delta u)$
27)	$\cot \Delta u = (\cos \eta / \tan \Delta \phi') - ((\sin \eta / (\sin \Delta \phi' \tan Az'_i)))$
28)	$\Delta u = \tan^{-1} (1/\cot \Delta u)$
29)	If $\Delta u < 0$ : Go to 30
	If $\Delta u = 0$ : Go to 34
	If $\Delta u \ge 0$ : Go to 32
30)	If $\Delta u + 90 < 0$ : Go to 31
	If $\Delta u \ge 0$ : Go to 34
31)	$\Delta u = \Delta u + 180$ : Go to 34
32)	If $90 - \Delta u < 0$ : Go to 33
	If 90 - $\Delta u \ge 0$ : Go to 34
33)	$\Delta u = \Delta u - 180$
34)	$u = (u + \Delta u) - \Delta u$
35)	$v = u - \omega'$
36) - 43)	Same as 45 to 52 in G.S.R.P
44)	$\Delta T_n = T_n - T_x$
45)	$\sin \phi'_{s} = \sin u \sin i$
46)	R = $[.9966443/(100670015 \cos^2 \phi'_{\rm g}]^{1/2} + H/6378.174$

# 1-2-28

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47)	$r = a' (1 - e'^2)/(1 + e' \cos v)$
48)	H'' = (r - R) 6378.174
49)	$R_{o} = [.9966443/(100670015 \cos^2 \phi'_{s})]^{1/2}$
50) - 52)	Same as 41 - 43 in G.S.R.P
53)	$\lambda_i = \lambda_n + \Delta \lambda$
54)	$\Delta \lambda_{0} = -\lambda - \lambda_{i}$
55)	$\cos L_{o} = \sin \phi'_{s} \sin \phi' + \cos \phi'_{s} \cos \phi' \cos \Delta \lambda_{o}$
56)	$tan h = r \cos L_0 - R_0 / r \sin L_0$
57)	If $h < 0$ : Return for next observation
	If $h \ge 0$ : Go to 57
58)	$\rho = (R_0^2 + r^2 - 2 R_0 r \cos L_0)^{1/2} 6378.174$

Output:

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N,  $T_N$ ,  $\Delta t_n$ ,  $\rho$ , H, h.

# 1-2-29

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# 2.6.5 Doppler Reduction Program

# Input:

Same as General Sighting Reduction Program except that elevation - declination and azimuth-right ascension are not given.

	Equations:	·
	1) - 11)	Same as in G. S. R. P.
	12)	$\lambda_n = [RA' - \theta_G], 0 \le \lambda_n \le 360$
	13)	$\Delta \alpha = \lambda_n - \lambda \qquad \Delta \alpha = \lambda_n - \lambda$
•	14)	If $\Delta a < 0$ : Go to 15
		If $\Delta a > 0$ : Go to 17
		If $\Delta a = 0$ : Go to 19
	15)	If (∆a + 240) < 0: Go to 16
		If $(\Delta a + 240) \ge 0$ : Go to 19
	16)	$\Delta a = \Delta a + 360$ : Go to 19
	17)	If $(\Delta a - 180) > 0$ : Go to 18
	· · · ·	If (∆a - 180) ≤ 0: Go to 19
	18)	$\Delta a = \Delta a - 360$
	19)	$\Delta a_1 = \lambda_n - \lambda + 180$
	20)	If $\Delta a_1 > 0$ : Go to 21
		If $\Delta a_1 \leq 0$ : Go to 22
	21)	$\Delta a_1 = \Delta a_1,  0 \leq \Delta a_1 \leq 360$
	22)	If $(\Delta a_1 - 180) \ge 0$ : Go to 23
		If $(\Delta a_1 - 180) < 0$ : Go to 24

	23)	$\Delta a_1 = \Delta a_1 - 360$	
	24)	If $\phi \geq 0$ : Go to 25 If $\phi < 0$ : Go to 27	
	25)	If $(\Delta a + 180) \ge 0$ : Go to 30 If $(\Delta a + 180) < 0$ : Go to 26	
	26)	$\Delta a = \Delta a_1$ : Go to 31	
· · · ·	27)	If $\Delta a_1 \ge 0$ : Go to 28 If $\Delta a_1 < 0$ : Go to 29	
2	28)	If $(\Delta a_1 - 90) \ge 0$ : Go to 29 If $(\Delta a_1 - 90) < 0$ : Go to 31	• • • • • •
	29)	$\Delta a = \Delta a_1$	
	30)	$C_{13} = 1$ $C_{15} = 1$ , Go to 32	
· .	31)	$C_{13} = -1$	
	32)	$C_{15} = -1$ cos c = cos $ \phi  \cos  \Delta a $	
	33)	$\sin c = (1 - \cos^2 c)^{1/2}$	
	34)	$\beta = \sin^{-1} \frac{\sin  \phi }{\sin c}$	
		If $C_{15} \ge 0$ : Go to 36	
		If $C_{15} < 0$ : Go to 37	
	36)	If $\Delta a \ge 0$ : Go to 37 If $\Delta a < 0$ : Go to 38	
	37)	$\gamma = 180 - \beta - i$ , Go to 39	
	38)	$\gamma =  \beta - i $	

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	39)	$d = \sin^{-1} [(\sin c) (\sin \gamma)]$
	40)	$u' = \cos^{-1} \left  \frac{\cos c}{\cos d} \right $
	41)	If $\cos c < 0$ : Go to 42
		If $\cos c \ge 0$ : Go to 43
	42)	u' = 180 - u'
	43)	If $\phi \geq 0$ : Go to 44
		If $\phi < 0$ : Go to 46
	44)	If $(u'C_{13}) \ge 0$ : Go to 45
		If $(u'C_{13}) < 0$ : Go to 46
	45)	u = u', Go to 51
	46)	u = 180 -  u' : Go to 51
	47)	u'' = u' + 180
	48)	If $(u'' C_{13}) \ge 0$ : Go to 49
		If $(u'' C_{13}) < 0$ : Go to 50
,	49)	u = u'', Go to 51
	50)	u = 360 -  u'
	51)	$\mathbf{v} = \mathbf{u} - \boldsymbol{\omega}'$
	52)-59)	Same as 45 - 52 in G.S.R.P
	60)	$\Delta T_n = T_n - T_x$
	61)	$\phi' = \tan^{-1} (.99329985 \tan \phi)$
	62)	$R_{o} = .9966443/[100670015 \cos^{2} \phi']^{1/2} + H/6378.174$
	63)	$r = a'(1 - e'^2)/(1 + e' \cos v)$
	64)	$h = \tan^{-1} [ (r \cos d - R)/r \sin d]$

## 1-2-32

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65)	If $(h-90) = 0$ : Go to 66
	If $(h-90) \neq 0$ : Go to 67
66)	$\rho = r - R$ , Go to 68
67)	$\rho = 6378.174 \cdot r \cdot sin d/cos h$
68)	d = 60 d

# Output:

N, 
$$T_n$$
,  $\Delta T_n$ , d, h,  $\rho$ 

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2.7	Glossar	Y	
	Location	Symbol	Meaning
	А	a	Semi-major axis
	AAl		Column 1 overpunch from observation card
	AA2		Column 25 overpunch from observation card
	AA3		Column 31 overpunch from observation card
	ALPHA	* a	Angle used in intermediate comp. (see Fig.2.1)
	APRIME	a'	Predicted semi-major axis
	BNA	aBN	Baker-Nunn section; right ascension of sat.
	BNAO	<sup>a</sup> OBN	Baker-Nunn section; equinox constant
	BNGAM	$\gamma_{BN}$	Baker-Nunn section; angle gamma
	BNTH	θ <sub>BN</sub>	Baker-Nunn section; equinox constant
	BNZ	<sup>Z</sup> BN	Baker-Nunn section; equinox constant
	C13, C15	C <sub>13</sub> , C <sub>15</sub>	Doppler section; quadrant conversion constants
	C2PI	2π	Constant = 6.28318531
	С	с	Rate of change of period (days/rev. <sup>2</sup> )
	CLR	ρ	Slant range in earth radii
	COSET		Cosine of $\eta$ , (See Fig. 2.2)
	COSHD		Cosine of elevation
	COSI		Cosine of inclination angle
	COSP, COSPF	ર	Cosine of geocentric latitude of the station
	COSZ		Cosine of azimuth angle
	COTU		DF section; cotangent of $\Delta \mu$
	D	d	First derivative of rate of change of period
	DA		Control section; day of observation
	DA		Doppler section; predicted longitude difference between station and nodal crossing
	DA1	Δal	Doppler section; $\Delta a + \pi$
	DAY		Observation time in days of year
	DB	β	Doppler section; intermediate angle (See Fig. 2.1)
	DD	ď	Doppler section; arc distance from station to subsatellite point
	DEG		Conversion constant, radians to degrees
	DELL	Δλ	Longitude difference between station and sub- satellite point

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DELN	$\Delta \mathbf{N}$	Difference between epoch revolution and the revolution of the observation
DELRA	$\Delta RA$	Computed minus predicted satellite right ascension
DELT	$\Delta t_n$	Time difference between observation and epoch times.
DELTAT	$\Delta t$	Time residual (observed minus predicted time of nodal crossing)
DELTN	$\Delta t_n$	Time difference between observation and the preceding ascending node
DF		Direction finder
DFL	$\Delta \lambda$	DF section; right ascension difference between station and ascending node
DG	γ	Doppler section; angle gamma
DH	$\Delta H_{t}$	Height tolerance in kilometers
DI	i	Inclination angle
DLW	$\lambda'_{N}$	Doppler section; predicted longitude of ascending node
DRA	$\Delta RA_t$	Right ascension tolerance limit in degrees
DT	$\Delta t_t$	Time tolerance limit in days
DTCMP	Τ <sub>N</sub>	Predicted time of nodal passage
DU	μ	Doppler section; computed argument of latitude of satellite
E	e	Eccentricity
ELEM		Element number
ELEMR,ELE	СМW	Arrays used to store element sets
ELS	L <sub>s</sub>	Longitude of subsatellite point
EMRA	M <sub>N</sub>	Mean anomaly, node related
EMS	Ms	Mean anomaly, satellite related
EN	Nobs	Revolution number at observation time
EPRIME	e'	Predicted eccentricity at time of observation
ERA	EN	Eccentric anomaly of ascending node
ERAD		Intermediate quantity, $\frac{1-e}{1+e}$
ES	Es	Eccentric anomaly of satellite
FORDE		Contents of col. 80 of element card
FRACT		Fractional day portion of observation time

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i Ν λ <sub>s</sub> H δorh	Inclination angle Epoch revolution number Station longitude Station height
λ <sub>s</sub> H	Station longitude
H	0
H	Station height
δ or h	
	Declination or elevation angle from observation
ρ	Slant range
	Minute of observation
h	Doppler section: elevation angle
H''	Height of satellite above earth's
HA	Hour angle
	Hour of observation
	DF section: object height
	Array of satellite and element numbers
	Obs ID (satellite number and last digit of year of observation)
	. Obs ID (hour and minute of observation time)
	Switch: 1 = known observation 2 = unknown observation
	Switch: 1 = elements found 2 = elements not found
	Obs ID, (month and day of observation time)
	Obs ID, (seconds of obs $x100$ )
	Same as IIEYEl
	Same as IEYE1
	Input card number or card type
	Tag switch: 1 = UO tagged 0 = UO not tagged
	Switch set for type of General Reduction: negative for Baker-Nunn; zero for visual, positive for radar
	Upper limit of number of iterations for convergence in computing geocentric latitude of the satellite
	Tolerance switch: determined by observation lead time
	Equipment type used by observing station
	Year of observation
	Switch set to 1 for unknown obs., which indicates that residuals are to be tested against tolerance limits; otherwise set to 0
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1-2-36

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REDUCT

JCLS		Switch $1 = classified$ observation 0 = unclassified observation
JJ		Number of stations read from station tape
JTI		Input tape
J T2		Output tape
к		Switch indicating which section of program to process obs: 2 = General 3 = Doppler 4 = Direction Finder
KAT		Satellite number
KEL		Number of element sets read
KSTA		Station number
KTAP		Switch normally set to 1. Set to 2 when all UOs written on interim tape
KUFO		UO switch: -1 = known observation 0 = compute vector (General Rada 1 = vector previously computed 2 = initial setting
LEM		Element number
LEND		Card type
М		Satellite number from observation card
MLAST		Previous satellite number
МО		Month of observation
MOUT		Output indicator for type of General Reduction used
MSG		Observation message number
MT		Heading switch: 2 = no heading necessary for General 3 = No heading necessary for Doppler 6 = No heading necessary for DF
Ν	Ν	Epoch revolution number
NAT		Satellite number
NDCT		Control section: Switch set negative when necessary to obtain new obs from input tape
NOUT		Set to \$ when time of obs. precedes epoch by more than four days
NSTA		Station number from observation card
NUFO		<ul> <li>Previous obs. unknown switch. If previous observation was unknown this switch is set to non-zero</li> </ul>

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PA	Pa	Anomalistic period in days per revolution
PAPRI	ra Pa	Predicted anomalistic period
PEYEM	* a	Satellite number
PHI	$\phi_{s}$	Geodetic latitude of station
PHIPJ, PHIPJ1	. /	Geocentric latitude of satellite
PHIPL	Գյ Ժ <b>ʻ</b>	Geocentric latitude of station
PHIPR	φ's φ'	Geocentric latitude of station
PI	Ψ	Constant = 3.14159265
PI2	π/2	Constant = 1.57079633
PM	•	Nodal period at epoch in minutes per revolution
PN	P <sub>N</sub> m	Nodal period at epoch in days per revolution
Q	P <sub>N</sub>	Perigee distance in earth radii
QDOT	q ·	Rate of change of perigee distance (er/day)
	q	1/2 second derivate of perigee distance
QDDOT	q	
RA	ດ ດ້	Right ascension of ascending node at epoch
RACOMP	77	Predicted right ascension of ascending node
RAD		Conversion constant, radians to degrees
RAN	a	Right ascension of ascending node at observation time
RAS	as	Right ascension of subsatellite point
RDOT	Ω	Rate of change of RA
RDDOT	Ω	1/2 second derivative of RA
RO	Ro	Radius of earth at subsatellite point
ROP	r	Radius of earth at station
RPOLY	ά′	Predicted RA at observation time
RPPRI	$\Delta$ h	Difference between observed and predicted height
RS	R <sub>s</sub>	Distance from center of earth to given sub- satellite latitude
RSMALL	r	Radial distance of satellite from center of earth
SAT		Satellite number
SEC		Seconds of observation
SHS	h	DF section: elevation angle
SINAL		Sine of angle, d(see Fig. 2.1)

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	SINET		DF section: sine $\eta$ (see Fig. 2.2)
	SINI		Sine of inclination angle
	SINP, SINPR		Sine of geocentric latitude of station
	SINZ		Sine of azimuth angle
	SLR	ρ	Slant range
	SO	ρ	Doppler section: slant range
	SS1		Input switch to inhibit station tape
	SS3		Input switch to eliminate tolerance limit tests and to inhibit the interim tape
	SS8		Input switch to run all observations as unknowns
	SS9		Input switch if teletype output desired
	SS10		Input switch to eliminate re-reduction of untagged observations; inhibits interim tape
	SS11		Input switch if q-term supplied
	STAIN		Array of station coordinates read from station tape
	STEL		Working array of station coordinates
	TANI		Tangent of inclination angle
	THETA	θ	Sidereal Time at station
	TLHS		DF section: tangent of elevation angle
	TMO	,	Month of observation
	TNDF	Τ <sub>N</sub>	Predicted time of node
	TOD	[t <sub>0</sub> ]	Integer portion of epoch time
	TOF	t <sub>frac</sub>	Fractional part of epoch time
	TOT	t <sub>o</sub>	Epoch time
	U	μ	Argument of latitude of satellite in orbital plane
	UDELU		DF section: $\mu + \Delta \mu$
	UO		Unknown observation
	v	v	True Anomaly
	w	ω	Argument of perigee
	WDOT	ŵ	First derivative of argument of perigee
*	WDDOT	ŵ	1/2 second derivative of argument of perigee
	WPOLY	ມ	Predicted argument of perigee
	х	φs	Station latitude
-	XAT		Tens and units digits of satellite number

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XDAY	t <sub>s.</sub>	Smithsonian day at epoch time
XDEC		Five least significant columns of elevation angle or declination field
Y	λ <sub>s</sub>	Station longitude
YDAY		Observation time in Smithsonian days
YR		Year of observation
YRA		Six least significant columns of azimuth angle or right ascension field
Z	h <sub>s</sub>	Station height
ZPRI	Az	DF section: predicted azimuth angle
ZRA	Az or a	Azimuth angle or right ascension of satellite

1-2-40

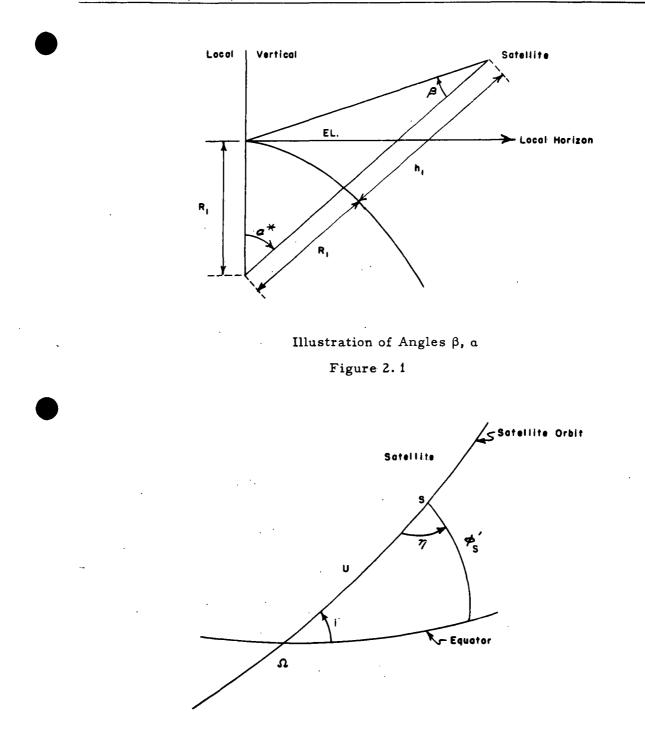
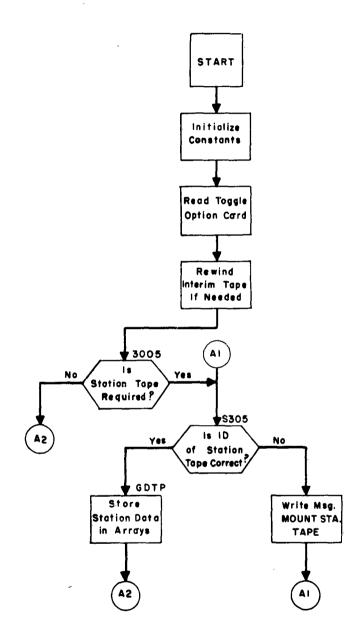


Illustration of Angle  $\eta$ Figure 2.2

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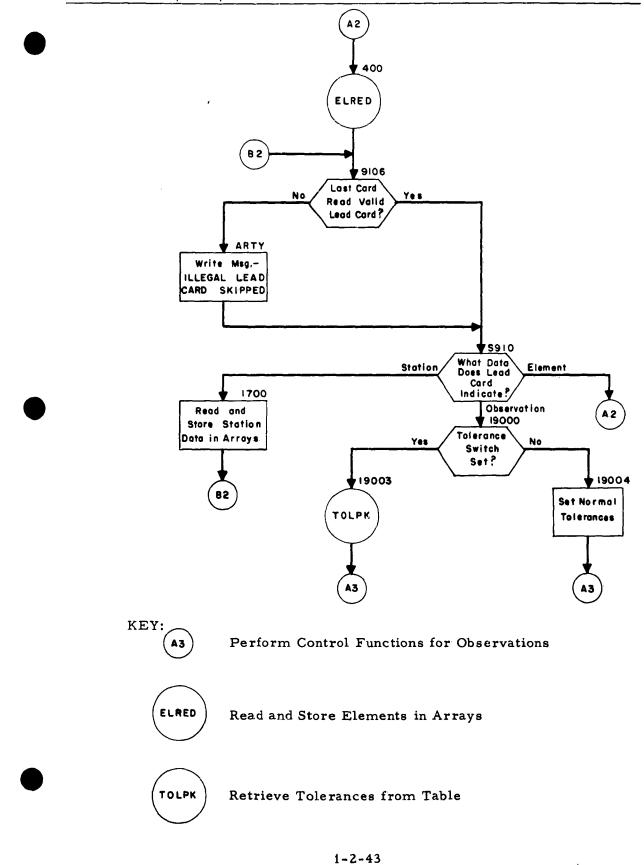


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Read and Examine Input Data

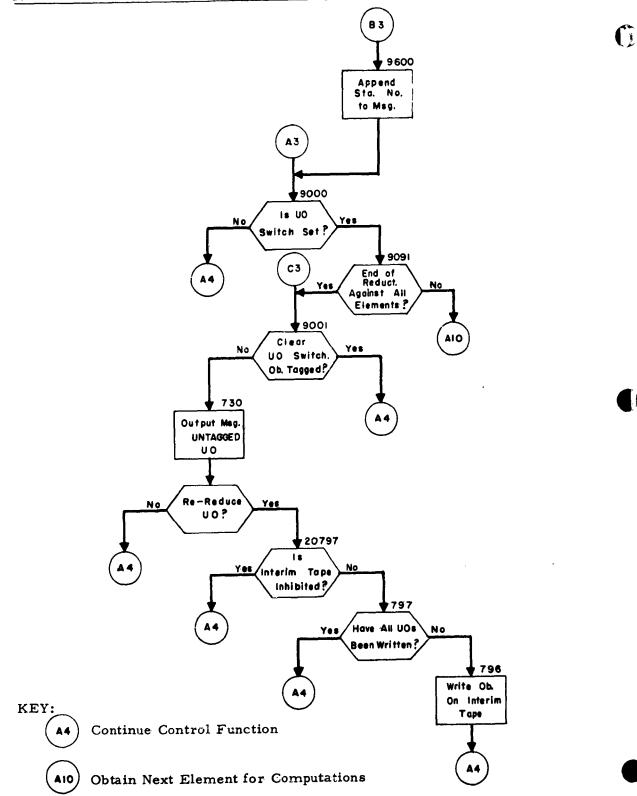
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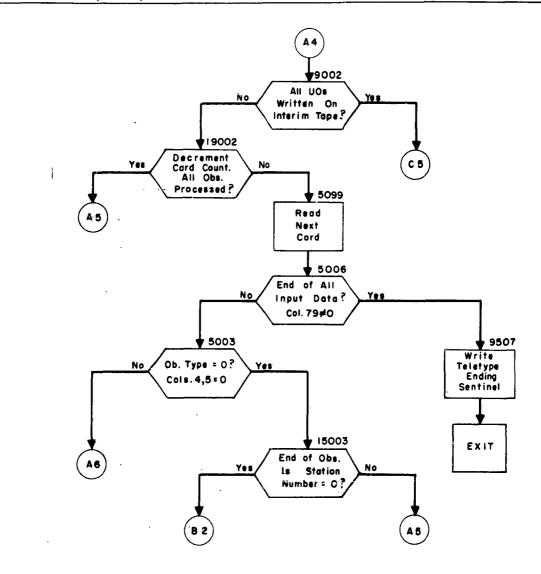
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KEY: B2 Test Validity of Lead Card

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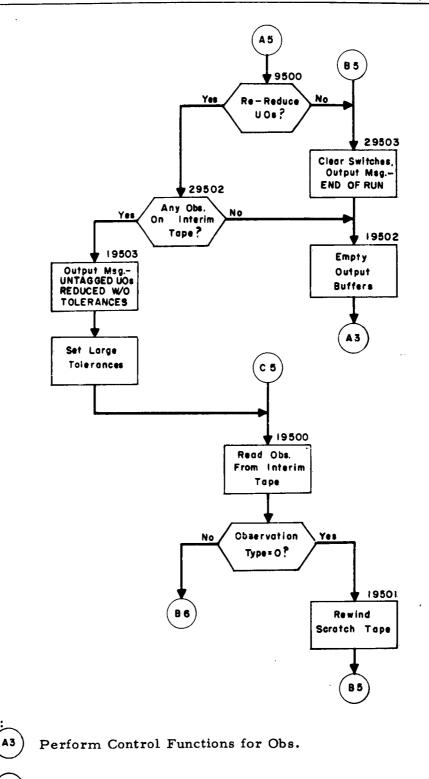
A6

- Process an Unknown Observation, If Any, From Interim Tape
- Read an Observation From Interim Tape
  - Continue Processing Observation

KEY:

B 6

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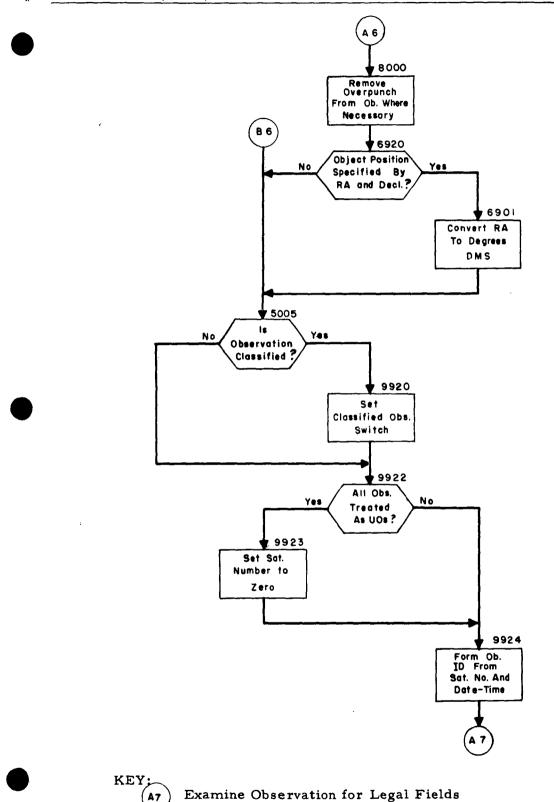


Continue Processing of Unknown Observation from Interim Tape

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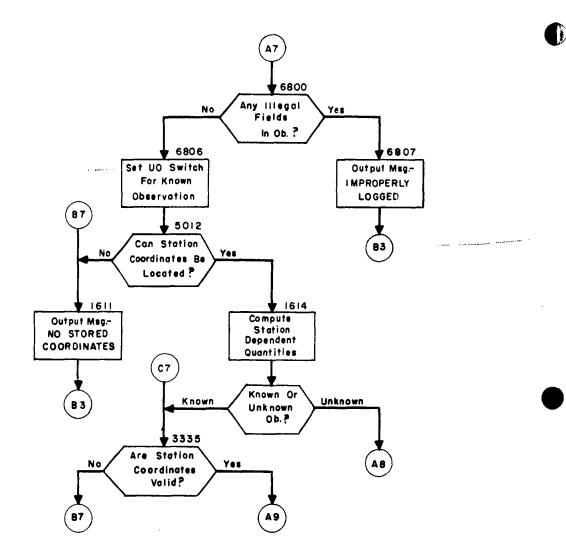


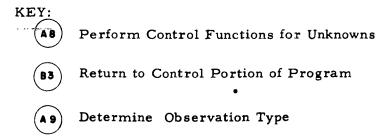
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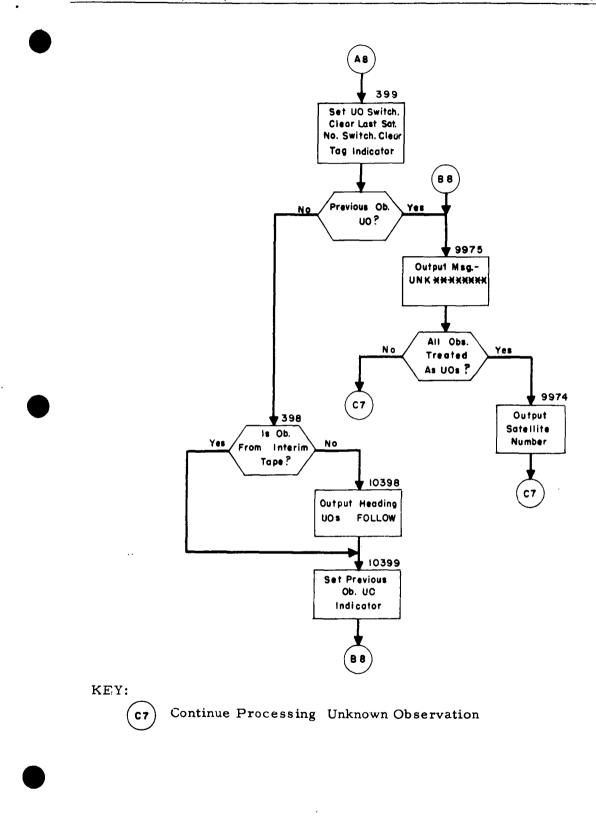




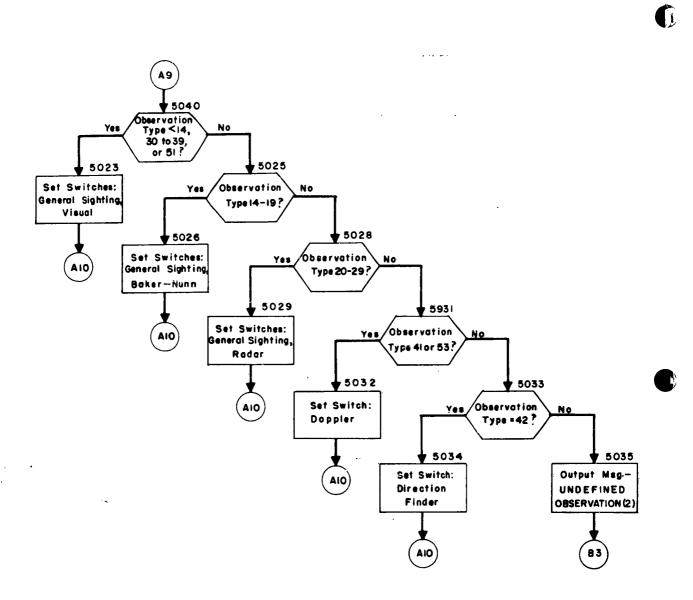
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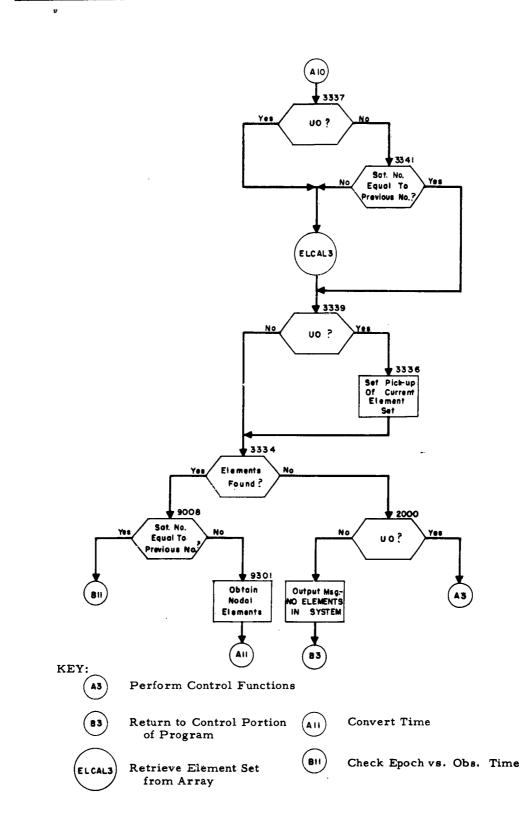
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(AIO) Obtain Element Set for Computations

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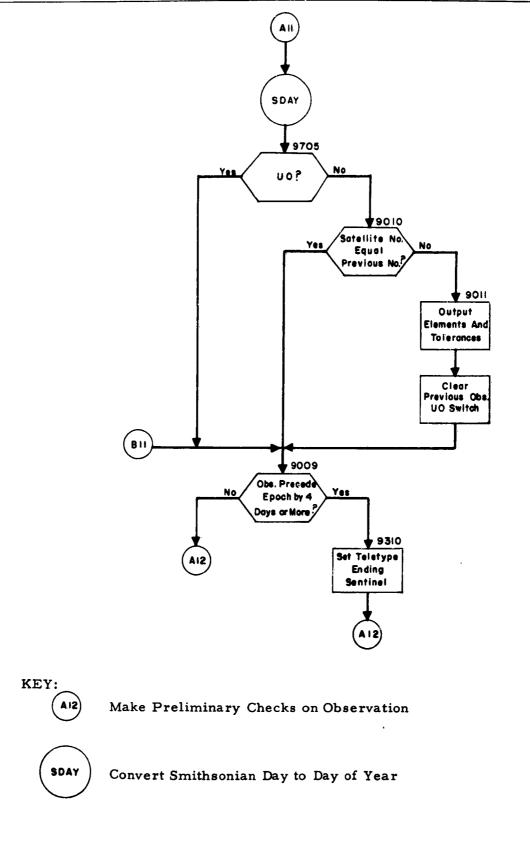


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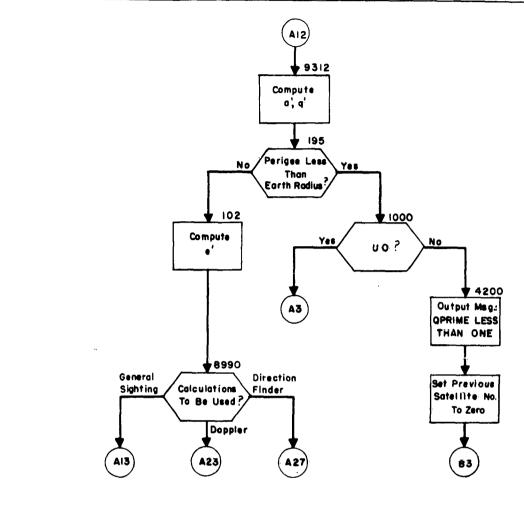
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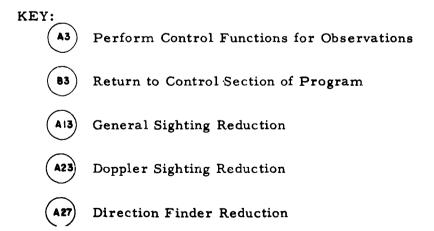
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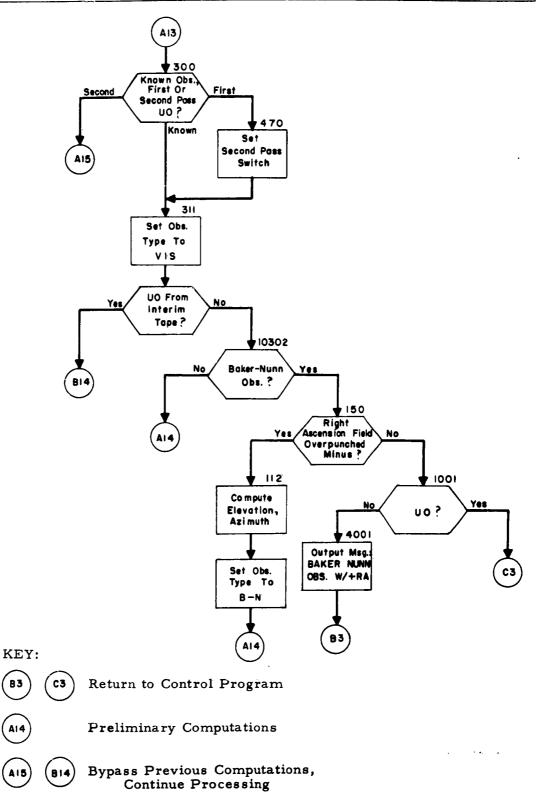
the time details and the little of the in shells with the "





1-2-53

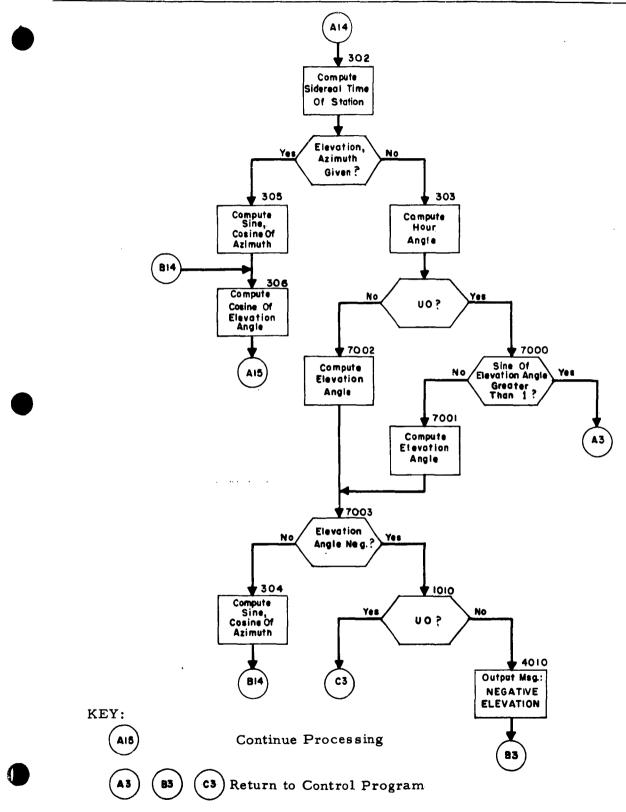
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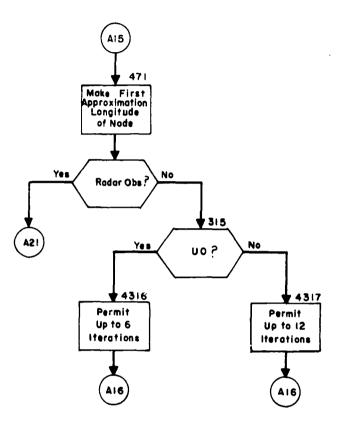


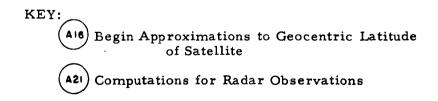
## 1-2-55

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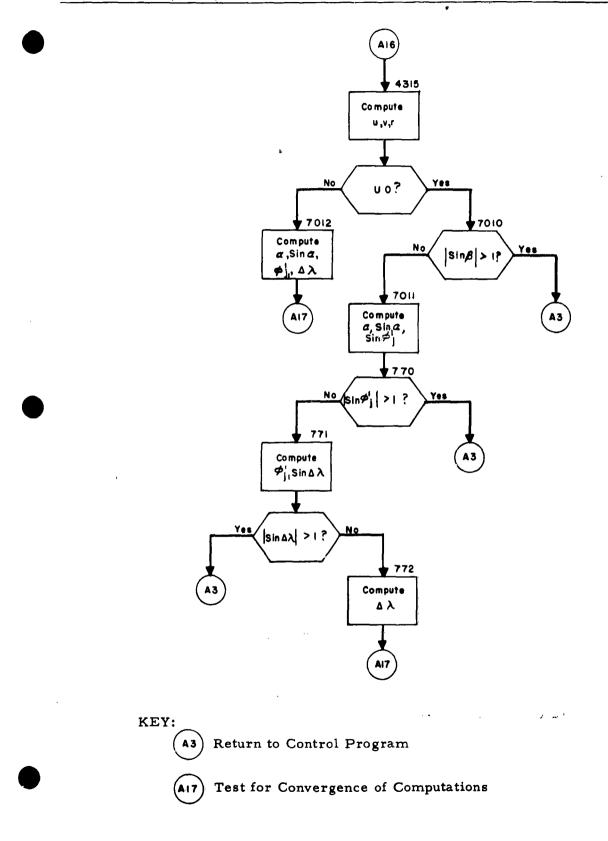
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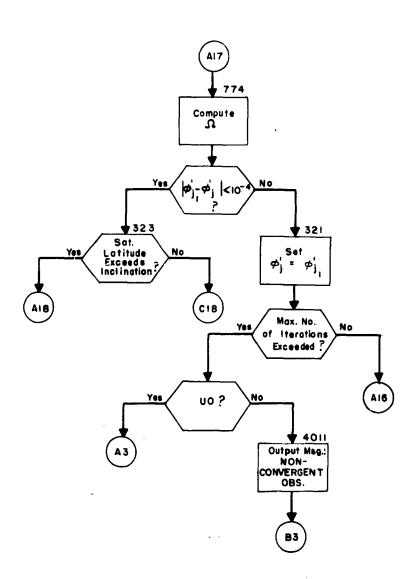


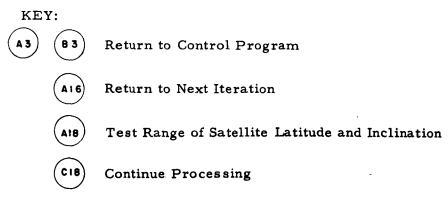
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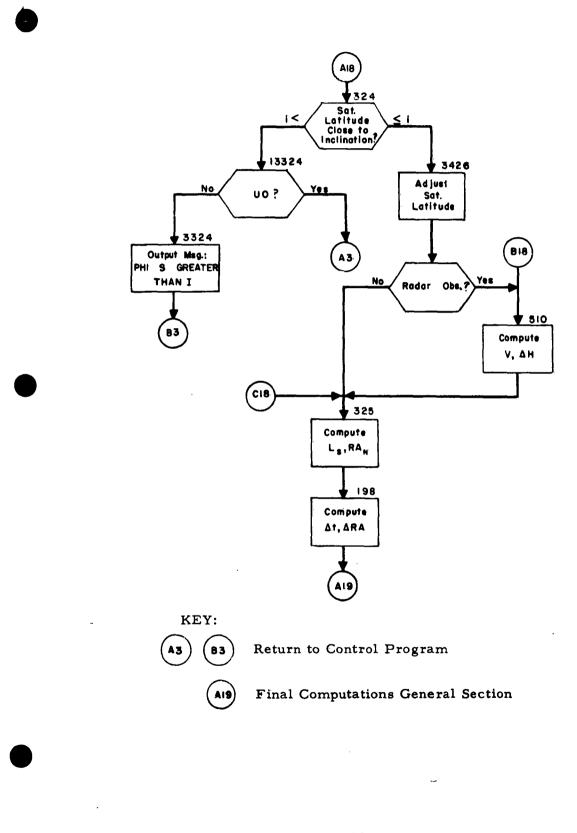






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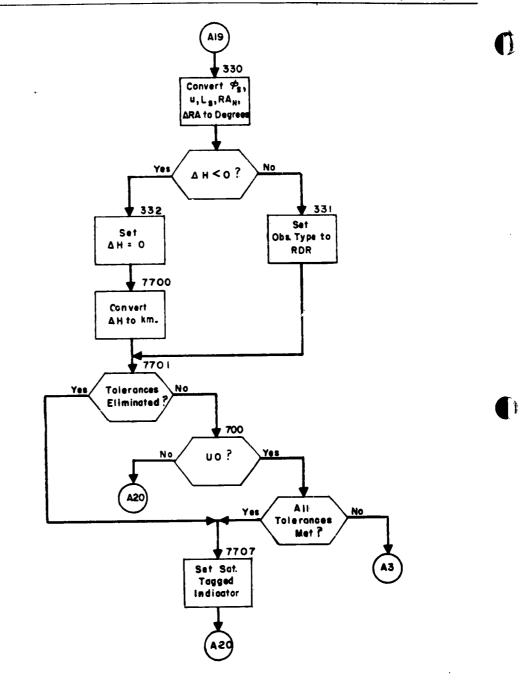


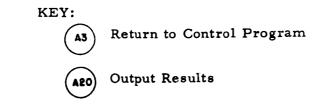
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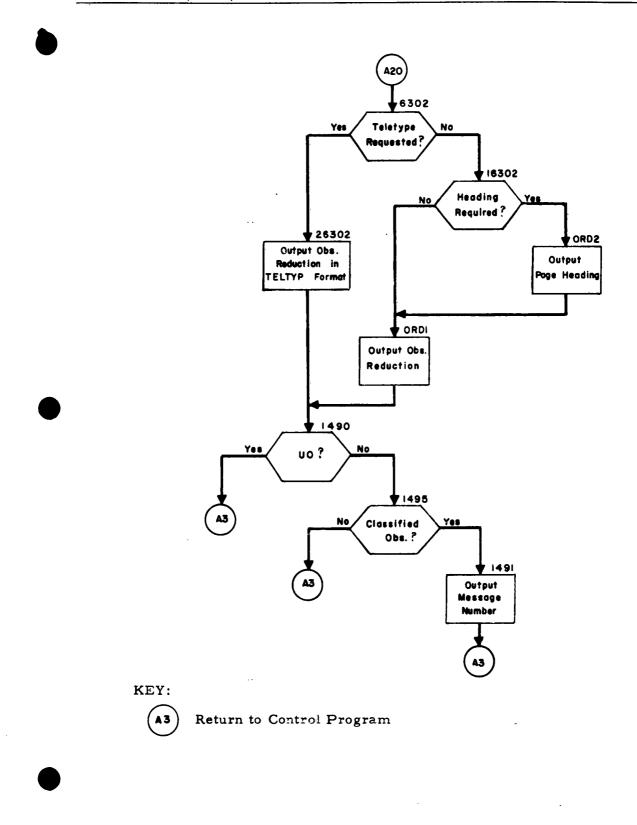
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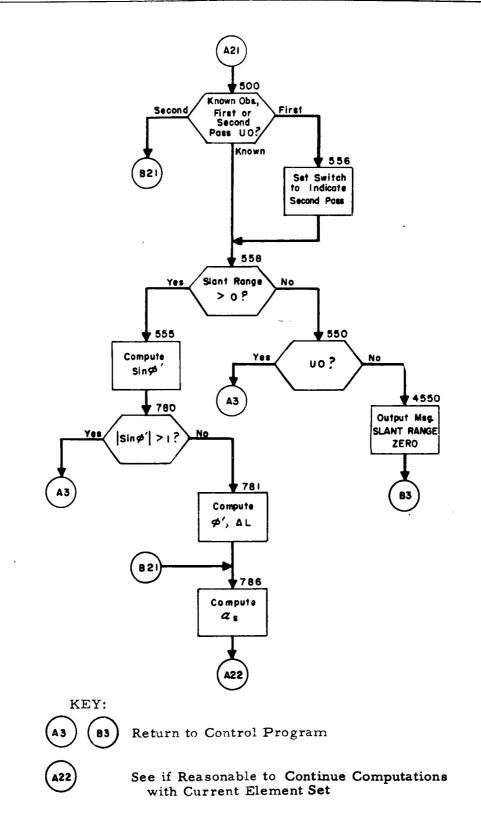
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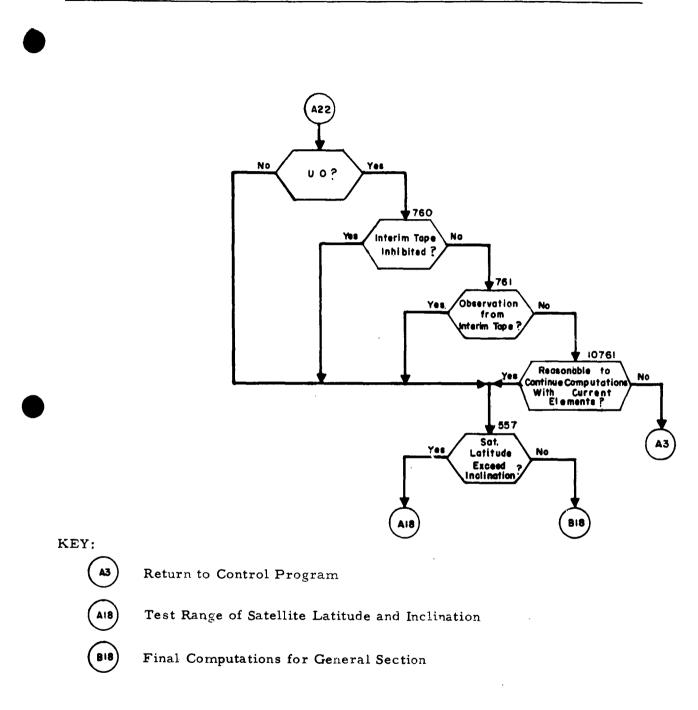
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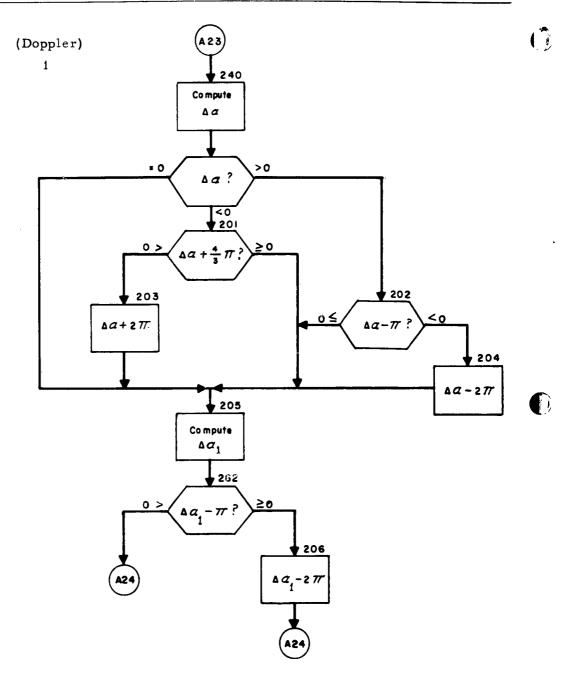
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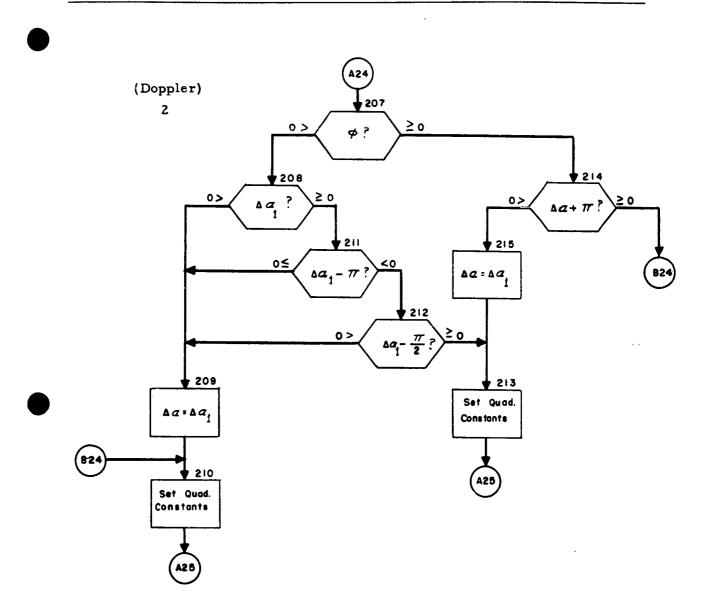


Set Final Quadrant Constants



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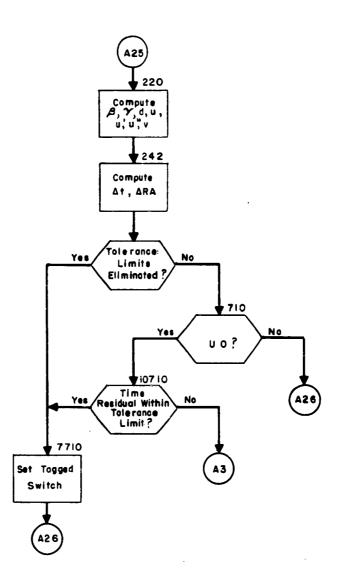


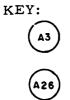


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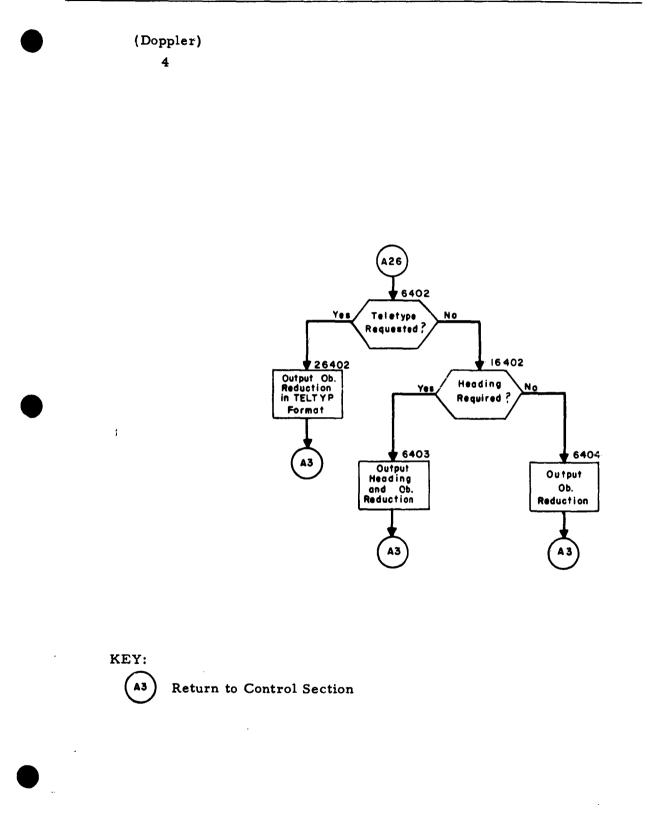


Return to Control Section

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Output for Doppler

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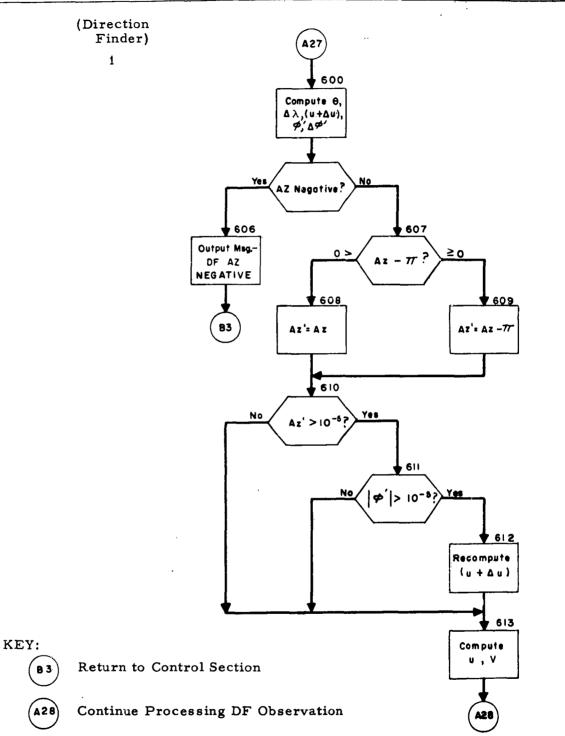




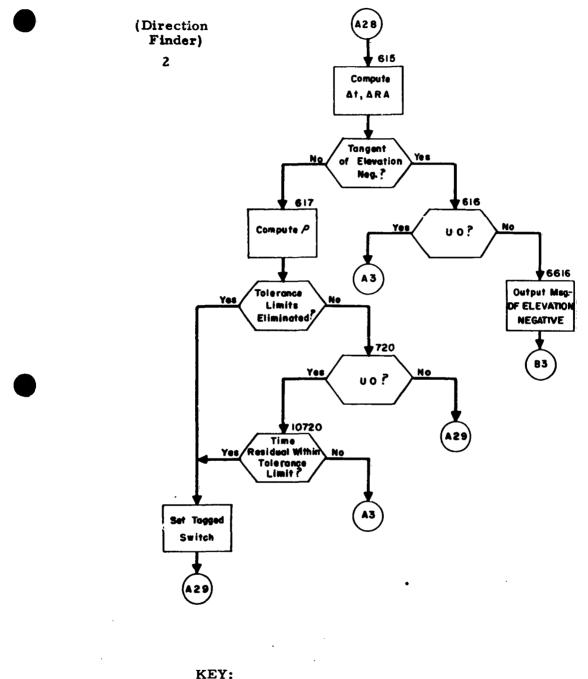
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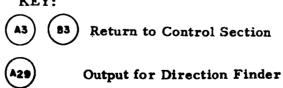
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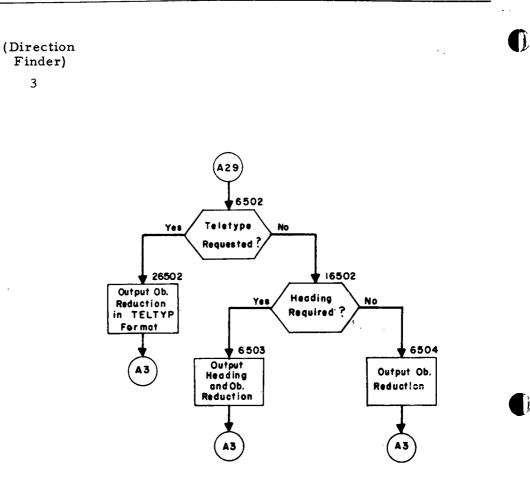
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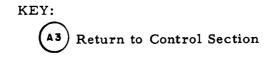




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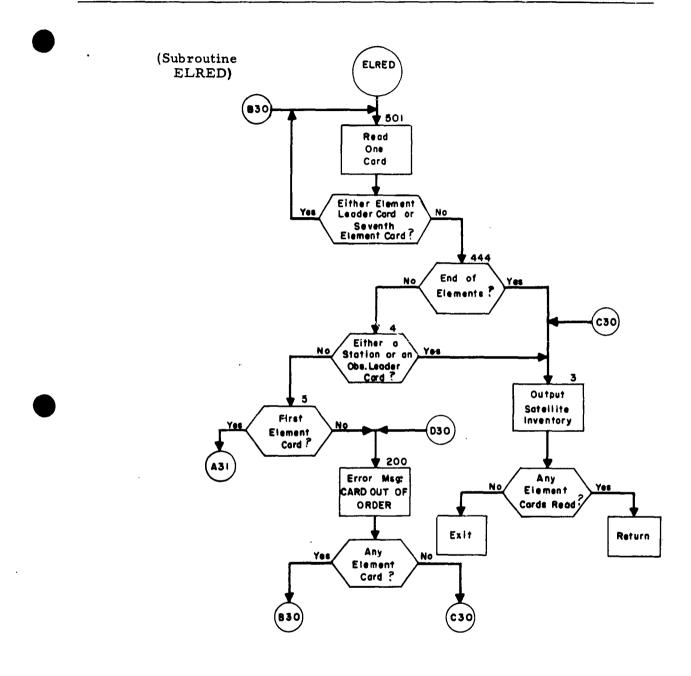
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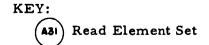


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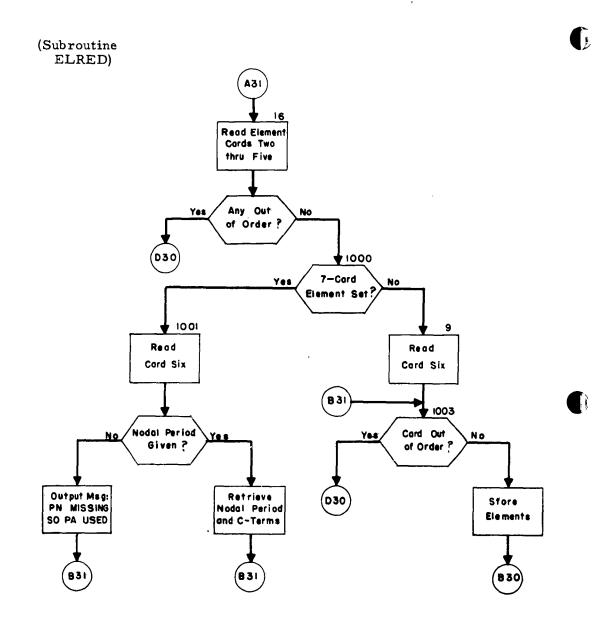


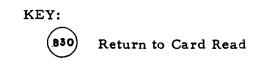
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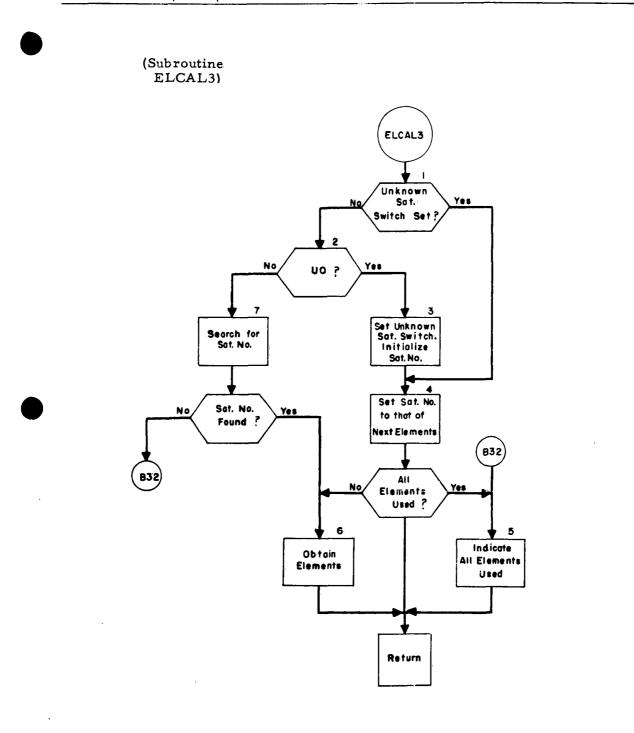
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#### 3.1 ROC, Radar Orbit Computation

#### 3.2 Function

The orbital elements of a satellite are computed from either the radar observations of the elevation angle, azimuth angle and slant range, or directly from the geocentric rectangular coordinates and velocities. The program is useful in obtaining orbital elements for newly launched satellites, in generating a new set of elements for satellites that have not been recently sighted, or in correcting the elements of a satellite.

A predicted orbit can be based on as few as three consecutive observations recorded along a single tracking run of a big radar, such as Millstone. An odd number of observations must be supplied. The move observations given the better the results. The midpoint of the radar run is first determined. Then the other points are picked up in pairs, on either side of the midpoint, and effectively averaged. The time span between individual pairs should be greater than forty (40) seconds. The geocentric rectangular coordinates of the midpoint are then determined and the elements computed.

If the geocentric rectangular coordinates and velocities of a point are used as input to the program, the averaging procedure is skipped and the orbital elements are computed directly.

#### 3.3 Input

A request card precedes either input option (1) or (2). The format of this card is as follows:

Cols. 1-3 Option (1) - Number of observation cards to be read. Must be an odd integer between 3 and 999.

Option (2) - field ignored

- Cols. 6-17 Alphanumeric satellite name or identification used in output.
- Cols. 18-35 Alphanumeric date used as part of identification for output.

Col. 41 1 - element cards requested as output

Col. 42 1 - Option (1) -  $x, y, z, \dot{x}, \dot{y}, \dot{z}$ 

printed for the midpoint of the observations.

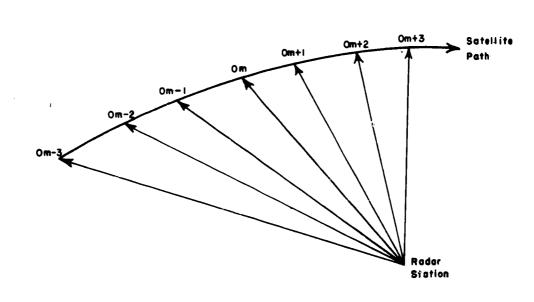
- Option (2) - field ignored

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Representation of a Radar Run

Fig. 3.1

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### Cols. 72-80 Must be blank.

Input option (1) should be used if orbital elements are to be computed from radar observations. Option (2) should be selected if the orbital elements are to be computed from the geocentric rectangular coordinates and velocities of a point in space. The input decks should be as follows:

- 1. Radar observations supplied.
  - 1. Request card
  - 2. Slandard station card
  - 3. An odd number of observation cards (3 to999) in order of increasing time.
  - 4. A card with a nine (9) punch in column 79, used to signal the end of the data deck.
- 2.  $x, y, z, \dot{x}, \dot{y}$ , and  $\dot{z}$  supplied.
  - 1. Request card
  - 2. Rectangular coordinates in the following format:

 Cols.
 1-14
 x in kms.

 Cols.
 15-28
 y in kms.

 Cols.
 29-42
 z in kms.

 Cols.
 43-46
 Year of epoch

3. Velocity components in the following format:

Cols.1-14x in kms./sec.Cols.15-28y in kms./sec.Cols.29-42z in kms./sec.

- Cols. 43-56 Day of year of observation of coordinates.
- 4. Card with a nine (9) punch in column 79, used to terminate the data deck.

The quantities  $x, y, z, \dot{x}, \dot{y}, \dot{z}$ , and day of year must have a decimal point punched somewhere in the field.

3.4 Output

There are two output options (A) and (B) in addition to the normal output describedunder (C).

A. This option may be requested only under input option (1). If selected, the geocentric rectangular

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coordinates and velocity components, at the mean time of the input observations, will be printed. The x, y, and z coordinates in kilometers will be preceded by the labels XX=, YY=, ZZ=, and the velocity components in kilometers per second will be preceded by the notations XDER=, YDER=, and ZDER. The time of the observation in the day of year notation will be labled TIME=.

- B. This output option if selected will cause a punch out of the standard six card element set. The following quantities will be punched: satellite number ( as read in from the observation card ), element card number, satellite name, eccentricity, inclination, year of epoch, time of epoch in day of year notation, nodal period at epoch, right ascension of ascending node, argument of perigee, and perigee distance. All other fields of the six card element set are left blank since the quantities can not be computed.
- C. The normal output will be printed in the format described below.

The first line will contain the heading RADAR ORBIT COMPUTATION followed by the station name as read from the station card. The second line has the label satellite followed by the alphanumeric name as supplied on the request card. This is followed by the satellite number which was read from the observation cards. The date supplied on the request card appears on the third line of output. The elements which have been computed are then listed under appropriate headings as follows: the same-major axis in km. and earth radii, the right ascension of the ascending node in degrees, the radius vector in km., eccentricity and its squared term, the nodal period in days and minutes, the argument of perigee in degrees, the velocity in km. per sec., the perigee

1-3-4

distance in km. and in earth radii, the time of the last perigee pass in days, the inclination in degrees, the apogee distance in km. and in earth radii, and the time of the last nodal crossing in days.

## 3.5 Processing

If input option (B) is used, the elements are computed directly from the x, y, z,  $\dot{x}$ ,  $\dot{y}$ , and  $\dot{z}$  supplied as input. Otherwise, these coordinates are computed from the station coordinates and the radar observations. An odd number of consecutive observations, at least three, must be supplied. The midpoint, or middle observation, of a radar run is chosen. The other points are selected in pairs, one from each side of the midpoint. The time span between the individual observations in a pair should be greater than 40 seconds.

For each observation, the three topocentric coordinates are converted into the geocentric rectangular coordinates x, y, and z. Next, the velocity components  $\dot{x}$ ,  $\dot{y}$ , and  $\dot{z}$  are found for each pair of observations and are averaged together for the final velocity components at the midpoint.

The orbital elements are then computed from the postion and velocity components. The magnitude of the radius vector and the velocity are computed directly from the rectangular coordinates and velocity components. The semi-major axis, nodal period, eccentricity, perigee and apogee distance are computed next, as well as, the inclination and the right ascension of the ascending node. The true, eccentric, and mean anomalies are found and the time of perigee and nodal crossing are computed. Output is then written as specified in the output section.



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## 3.6 Formulation

If input option 2 used go to D, otherwise compute A, B, C.

A. 
$$\phi' = \tan^{-1} (.99329985 \tan \phi)$$

$$R = \frac{.9966443}{\sqrt{1 - .00670015 \cos^2 \phi'}} + \frac{H}{6378.174}$$

B. Determine rectangular coordinates for each observation (i).

$$r_{i} = \sqrt{R^{2} + \rho_{i} (\rho_{i} + 2 \cdot R \sin h)}$$

$$\beta_{i} = \left[ |\sin^{-1} (R \cos h_{i}/r_{i})| \right] \quad (0^{\circ} \le \beta \le 90^{\circ})$$

$$a_{i}^{*} = \left| \frac{\Pi}{2} - \beta_{i} - h_{i} \right|$$

$$\phi_{s}^{\prime} = \sin^{-1} (\cos a_{i}^{*} \sin \phi' + \sin a_{i} \cos \phi' \cos Az_{i})$$

$$\Delta\lambda_{s} = \sin^{-1} (\sin a_{i}^{*} \sin Az_{i}/\cos \phi_{s}')$$

$$\lambda_{s} = \lambda - \Delta\lambda_{s}$$

$$\theta_{G} = \left[ \theta_{o} + .9856472 \text{ day}_{i} + 360.9856472 \text{ fract. day}_{i} - \lambda_{s} \right], (0^{\circ} \le \theta_{G} < 360^{\circ})$$

$$x_{i} = r_{i} \cos \phi_{s}' \cos \theta_{G}$$

$$y_{i} = r_{i} \cos \phi_{s}' \sin \theta_{G}$$

$$z_{i} = r_{i} \sin \phi_{s}'$$

$$\underline{r}_{i} = (x_{i}^{2} + y_{i}^{2} + z_{i}^{2})^{3/2}$$

c.

Compute velocity components by effectively averaging individual  $\dot{x}_j$ ,  $\dot{y}_j$ ,  $\dot{z}_j$ , components of each pair for final  $\dot{x}$ ,  $\dot{y}$ ,  $\dot{z}$  components. Then compute velocity and  $\dot{r}$ . Initially indices are as follows: l = 1 signifying first observation, n is total number of observations, m is midpoint observation.

ROC



$$\tau_{12} = (t_{m} - t_{\ell}) (.07436574/60)$$
  

$$\tau_{23} = (t_{n} - t_{m}) (.07436574/60)$$
  

$$\tau_{13} = (t_{n} - t_{\ell}) (.07436574/60)$$
  

$$G_{1} = \tau_{23}^{2}/\tau_{12}\tau_{23}\tau_{13}$$
  

$$G_{2} = (\tau_{23}^{2} - \tau_{12}^{2})/\tau_{12}\tau_{28}\tau_{13}$$
  

$$G_{3} = \tau_{12}^{2}/\tau_{12}\tau_{23}\tau_{13}$$
  

$$H_{1} = \tau_{23}/12$$
  

$$H_{2} = (\tau_{23} - \tau_{12})/12$$
  

$$H_{3} = \tau_{12}/12$$
  

$$d_{1} = G_{1} + H_{1}/\underline{r}_{\ell}$$
  

$$d_{2} = G_{2} + H_{2}/\underline{r}_{m}$$
  

$$d_{3} = G_{3} + H_{3}/\underline{r}_{n}$$
  

$$\dot{x} = -d_{1}x_{\ell} + d_{2}x_{m} + d_{3}x_{n}$$
  

$$\dot{y} = -d_{1}y_{\ell} + d_{2}y_{m} + d_{3}y_{n}$$
  

$$\dot{z} = -d_{1}z_{\ell} + d_{2}z_{m} + d_{3}z_{n}$$
  
Update  $\ell$  and n.  $\ell = \ell + 1$ ,  $n = n$ 

Update l and n. l = l + 1, n = n - 1. If l = m, go on. Otherwise return to C and compute  $\dot{x}$ ,  $\dot{y}$ ,  $\dot{z}$  for another pair of observations.  $\dot{x} = \frac{\Sigma}{2} \dot{x} / l$  for l pairs

$$\dot{\mathbf{y}} = \sum \dot{\mathbf{y}}_{j} / \ell$$
$$\dot{\mathbf{z}} = \sum \dot{\mathbf{z}}_{j} / \ell$$

D. Compute elements

$$\underline{\mathbf{r}} = \sqrt{\mathbf{x}^2 + \mathbf{y}^2 + \mathbf{z}^2}$$
$$\mathbf{V} = \sqrt{\mathbf{\dot{x}}^2 + \mathbf{\dot{y}}^2 + \mathbf{\dot{z}}^2}$$

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$$\begin{aligned} \dot{\underline{r}} &= (x\dot{x} + y\dot{y} + z\dot{z})/\underline{r} \\ C_5 &= \left[ 2(631, 353746)^2/\underline{r} \right] \cdot V^2 \\ a &= ((631, 353746)^2/C_5)/6378.174 \\ P_n &= .058672947 (a)^{3/2} \\ e &= \sqrt{1 (V^2 - \dot{\underline{r}}^2) \underline{r}^2/631.353746^2 a} \\ p &= a (1 - e^2) \\ q &= a (1 - e^2) \\ q &= a (1 - e) \\ O &= a (1 + e) \\ i &= \left[ \cos^{-1} ((x\dot{y} - y\dot{x})/631.353746 \sqrt{p}) \right], (-90^{\circ} \leq i \leq + 90^{\circ}) \\ RA &= \tan^{-1} (\frac{y\dot{z} - z\dot{y}}{x\dot{z} - z\dot{x}}) \\ sin \mu &= \frac{z}{\underline{r} \sin i} \\ cos \mu &= (x + (z \sin RA \cot i)/\underline{r} \cos RA \\ \mu &= f_{qual} (\cos \mu, \sin \mu) \\ cos v &= (p/\underline{r} - 1)/e \\ v &= f_{qual} (\cos v, \underline{\dot{r}}) \\ \omega &= \left[ \mu - v \right], (0^{\circ} \leq \omega < 360^{\circ}) \\ tan \quad \frac{E_1}{2} &= \sqrt{\frac{1 - e}{1 + e}} tan \quad \frac{\dot{v}}{2} \\ M_1 &= E_1 - e \sin E_1 \\ \Delta t_1 &= M_1. P_n/2fi \\ T_p &= t_{obs} - \Delta t_1 \end{aligned}$$

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$$\tan \frac{E_2}{2} = \sqrt{\frac{1-e}{1+e}} \tan \frac{\omega}{2}$$
$$M_2 = E_2 - e \sin E_2$$
$$\Delta t_2 = M_2 \cdot P_n / 2\Pi$$
$$T_n = T_p - \Delta t_2$$

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## 3.7 Glossary

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	Location	Symbol	Meaning
	AMEAN	M1	Mean anomaly
	AMEANN	M <sub>2</sub>	Mean anomaly
	ANGLE	i	Inclination in degrees
	APOGEE	Qo	Apogee in km above the earth
	APOGER	Q	Apogee distance in earth radii
	ARGPER	ω	Argument of perigee in degrees
	AXIS	a	Semi-major axis in km.
	AZIM	Az	Azimuth from observation card
	BETA	β	See Figure 3.2 angle $\beta$
	BUT	То	Time of epoch (day of year)
	COP	0	Constant 631.353746 squared
	COSRA		Cosine of right ascension of <b>as</b> cending node at epoch
	COSU		Cosine of argument of latitude
	COSV		Cosine of true anomaly
	CUT		Time of epoch
	DAT	d <sub>1</sub>	Coefficient used to compute $\dot{x}, \dot{y}, \dot{z}$
	DAY	-	Day of observation
	DECLA	φ <sup>′</sup> s	Geocentric latitude of satellite
	DELONG	Δλs	Difference in longitude between station and subsatellite point
	DELTAN		Difference in time from perigee to node
	DELTAT		Time of obs minus time of perigee
	DET	d <sub>2</sub>	Coefficient used to compute $\dot{x}, \dot{y}, \dot{z}$
	DOT	d_3	Coefficient used to compute x, y, z
	E	e	Eccentricity
	ECCA	E <sub>1</sub>	Eccentric anomaly
	ECCAN	E <sub>2</sub>	Eccentric anomaly .
	ELAPSE		Day of year of observation
	ELEV	h	Elevation from observation
	ESOUA	e <sup>2</sup>	Eccentricity squared
	FRACT		Fractional day of observation
	GAT	Gl	Coefficient used to compute $\dot{x}, \dot{y}, \dot{z}$
	GEOCL	φ'	Geocentric latitude of station

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GEOGL	ф	Station latitude
GET	G <sub>2</sub>	Coefficient used to compute $\dot{x}, \dot{y}, \dot{z}$
GOT	G	Coefficient used to compute x, y, z
GRI	5	Time of observation minus time at node
HALB	n	Number of pairs of observations
HAT	н <sub>1</sub>	Coefficient used to compute x, y, z
HEIGHT	-	Station height in meters.
HET	H <sub>2</sub>	Coefficient used to compute x, y, z
HOT	H <sub>3</sub>	Coefficient used to compute x, y, z
HOUR	-	Hour of observation
HW		Hour after hour of first observation
IDCI		Input tape
IDCO		Output tape
IFLAG		End of input indicator
ISSW1		Output option= 1 for punched element cards
ISSW2		Output option= 1 for Cartesian coordinate output
ITYPE		Input card type, non-zero for end input
IYR	·	Year of observation
KNAVE		Time of epoch
$\mathbf{L}$		Observation card number being processed
М		Midpoint observation (number)
МО		Month of observation
N		Obs paired with L (number)
		Number obs cards to be read
NATNO		Satellite number
NK		Overpunched column l from obs. card
NL		Last two digits of satellite number from obs. card
PAK	P	Semi-latus rectum
PERIG	q.	Altitude at perigee in km.
PERIOD	P	Nodal period at obs. time in days
PERIOM	Pl	Period in minutes
PERIR	q	Perigee distance in earth radii

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RADIUS	R	Earth radius at station
RASC	¢	Angle a <sup>*</sup> , see Figure 3.2
RAT	<u>r</u> <sup>3</sup>	Array of $\underline{r}^3$
RAXIS	ar	Semi-major axis in earth radii
RITASC	RA	Right ascension of ascending node in degrees
RK		Constant 6378.174 km./e.r.
ROÕT	r	Distance from center of earth to satellite
RT		Constant.07436574/60= Ke /60
RX		Constant .017453295 deg/rad.
SATNO		Satellite number
SCAN		Time of ob. in seconds
SDAT		Smithsonian date of ob.
SECOND		Seconds of observation
SIDE	θg	Sidereal time
SINRA	8	Sine of right ascension of ascending node
SINU		Sine of argument of latitude
SLANT	ρ	Slant range from observation
SLONG	λ	Longitude of subsatellite point
SMIN		Minute of observation
SMONTH		Month of observation
SNAME1		Alphanumeric satellite name
STAME )		
STEME		
STOME		Alphanumeric station name
STUME		
TAME	<sup>+</sup> 12	Elapsed time between obs. and middle obs.
TAZ		
TEZ }		Alphanumeric data of chaquistion
TIZ		Alphanumeric date of observation
TEME	<b>*</b> 23	Elapsed time between middle and last obs.
THE		Tangent of $E_1/2$

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THEN		Tangent of $E_2/2$
TIME		Array of obs. times
TINOD		Time of node in days
TINODS		Time of node in Smithsonian days
TIPER		Time of perigee in days
TIPERS		Smithsonian time of perigee in days
TOBS		Day of year of middle observation
TOBSS		Day of Smithsonian year of middle obs.
TOME	<b>τ</b> 13	Elapsed time between obs. and last obs.
TRUE	v	True anomaly
U	u	Arg. of lat. of sat. in orbit plane
VECDOT	• r	First derivative of $r$ in km/sec.
VECTOR	r	Vector magnitude in km.
VELOX	v	Velocity in km/sec.
VK		Computed 7.90529382/no. of pairs of obs.
W		Day after day of first observation
WLONG	λο	Station longitude
Х		Array of computed x-coordinates
XDER	$\Sigma \dot{x}_i$ vk	Yields final x velocity component
XDOT	ż	x velocity component
XDUMP	$\Sigma \dot{x}$	Summation of individual $\dot{\mathbf{x}}$ components
XX	x	$\mathbf{x}$ coordinate of $\mathbf{r}$ on km.
Y		Array of computed y-coordinates
YCONST	θο	Sidereal time at start of year of obs.
YDER	Σyj·vk	Yields final y velocity component
Y DO T	ý	y velocity component
YDUMP	Σÿ	Summation of individual y components
YEAR		Last digit of year of observation
YR		Year of observation
YY	У	y coordinate of $\underline{r}$ in km.
Z		Array of computed z-coordinates
ZDER	$\Sigma \dot{z}_i \cdot vk$	Yields final z velocity component
ZDOT	· z	z velocity component
ZDUMP	$\Sigma \dot{z}$	Summation of individual $\dot{z}$ -components
ZZ	Z	z coordinate of <u>r</u> in km.

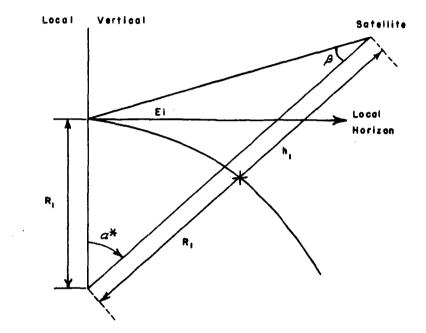
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## Illustration of Relations between Quantities

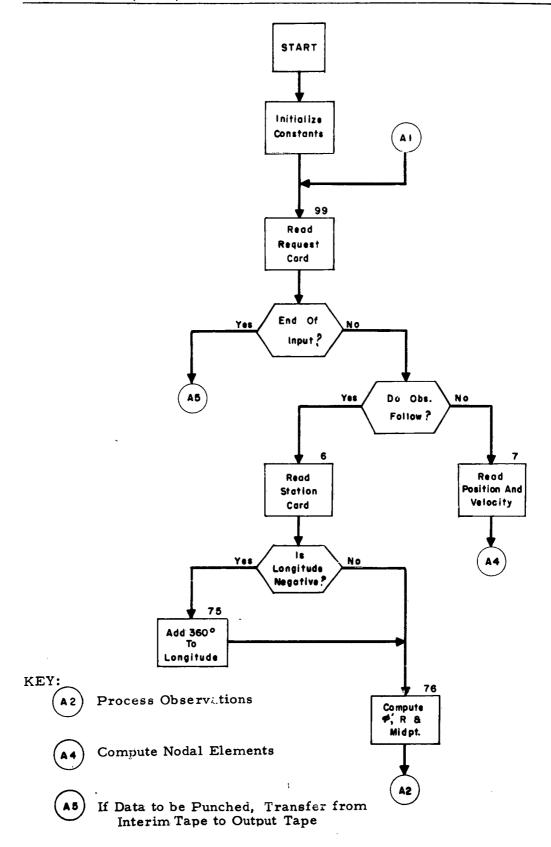
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Fig. 3.2

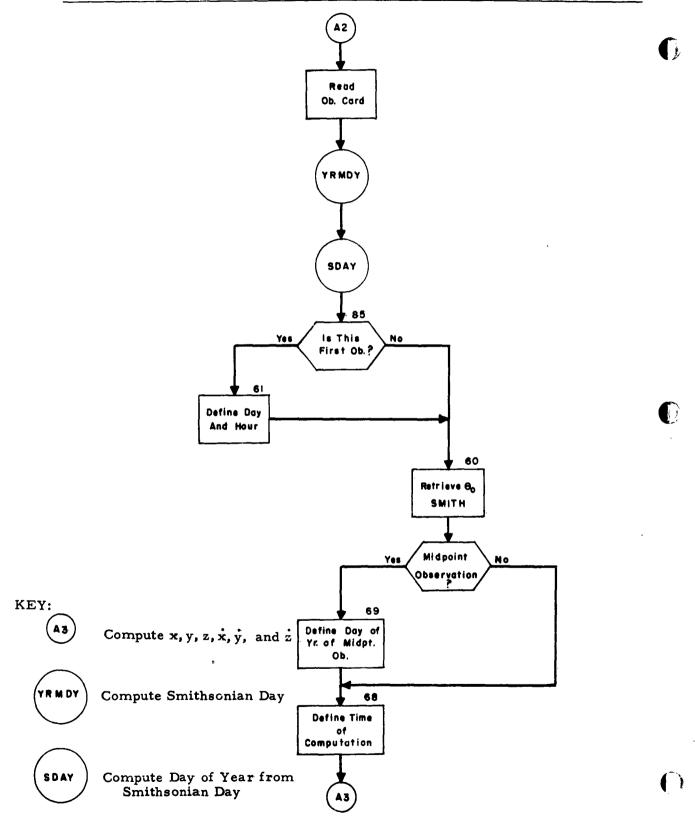
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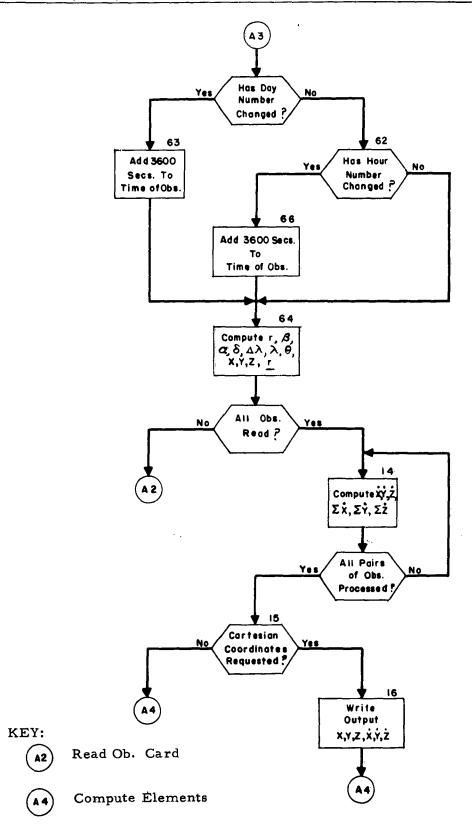
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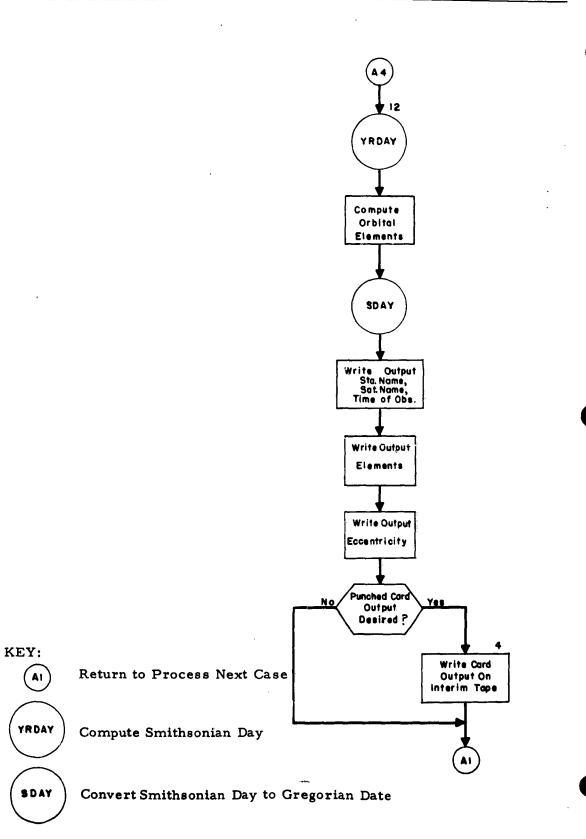


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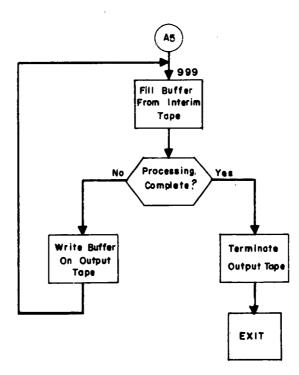
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- QUAC(X, Y) places an angle in the proper quadrant (0 to 2π) after taking the absolute value of the arc sine of the angle, (| sin<sup>-1</sup>Y|). The following arguments must be given: X = cosine of the angle Y = sine of the angle
- QUAL(X, Y) places an angle in the proper quadrant (0 to 2π) after taking the arc cosine of the angle, (cos<sup>-1</sup>X). The following arguments must be given: X = cosine of the angle Y = sine of the angle
- QUAT(Y, Z) places an angle (the eccentric anomaly) in the proper quadrant (0 to  $2\pi$ ) after taking the arc tangent of the angle times two, (2 tan<sup>-1</sup>Z). The following arguments must be given: Y = sine of the true anomaly (v) Z = tangent of the eccentric anomaly over two, (tan  $\frac{E}{Z}$ ) SMITH(X) - retrieves  $\theta_0$  for the year of computation, given the
- argument X which is the last digit of the year.

1-3-20

## 4.1 LOCVEC, Vector Coordinates for Lockheed

## 4.2 Function

The program computes the predicted position and velocity for Lockheed from nodal element sets. The output is written on magnetic tape for printing and for punching five-channel paper tape for teletype transmission. The teletype data is received at Sunnyvale Tracking Center where it becomes binary input for a CDC 1604 computer.

### 4.3 Input

The input data is ordered as follows:

- (1) Up to five hundred request cards
- (2) A card with a numeric punch (1-9) in col. 8. This indicates the end of the request deck.
- (3) Standard six or seven card element sets of the satellites in the request deck or all satellites (up to 500 sets).
- (4) A blank card which signals the termination of the element deck.

The elements sets do not have to be in the same order as the satellites appearing in the request deck.

The following information should be included on the request card:

Cols.	1-3	The SPADATS satellite number		
Col.	8	Must be blank		
Cols.	9-12	The STC (Sunnyvale Tracking Center)		
		satellite number, or if the SPADATS		
		satellite number is used, it should be		
		preceded by a nine (9) punch in col. 9.		
Cols. 13-20		The alphanumeric name of the satellite.		
Cols. 21-24		The alphanumeric abbreviation of the		
		country of orgin.		

#### 1-4-1

### 4.4 Output

#### 4.4.1 Printed

There is one output line for each satellite. The order is the same as that of the request deck. Each line contains: (a) the satellite number; (b) the Sunnyvale Tracking Center number or, if preceded by a nine (9), the SPADATS number; (c) the element number; (d) the month, day, hour of epoch; (e) seconds since the start of the epoch month; (f) epoch revolution number; (g) the x, y, z coordinates in feet, and the velocity components in feet per second.

### 4.4.2 Teletype

The basic message consists of a header and thirty-one words of five-level teletype punch. A visual header and a listable teletype header are included at the start of each message. The visual portion identifies the tape as a SPADATS vector with the characters "SP" and the four digit vehicle number; The listable header contains the word "SPADATS" followed by the vehicle number, the month, day and the nearest GMT hour of the last update of vectors. Tape contents and format are presented on the following pages.

### 4.5

# Processing

The request cards are read into an array. This array is checked as each element set is read into core, and the element sets, for which vectors are required, are stored.

The nodal elements to be used are converted to  $\underline{\mathbb{N}}, \underline{\mathbb{M}}$ , sets as required by the Analytical Integration Routine (XYZSB). The required position and velocity components are obtained from this routine.

LOCVEC converts the position coordinates to feet and the velocity components to feet per second. These quantities are then written on the output tape for off-line printing through the UBC. The conversion to the required teletype format, described in the output section, is then achieved and written on the tape for off-line punching on five-channel paper tape. This teletype data becomes input for a CDC 1604 computer at the Sunnyvale Tracking Center.

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Word No.	Content & Format
1	Check sum 2 <sup>24</sup> (see note 1).
2&3	Identification number, 8 BCD characters (see note 2).
4	Revolution number for vector, 19 bits, $D_{23} - D_5$ fixed integer, scaled $2^5$ .
	Bits D <sub>4</sub> -D <sub>0</sub> are all zeros.
5	Bits D <sub>23</sub> –D17 are all zeros.
	Bits D16 must be set equal to "1".
	Bits D15-D14 must be zeros.
	Bits D <sub>13</sub> —D <sub>0</sub> are vehicle numbers (0001 to 9999) fixed integer.
	This number will be supplied by the STA for each vehicle. (i.e. 1204 or 2203, etc.)
6	D <sub>23</sub> -D <sub>16</sub> , 8 bits all zeros.
	D <sub>15</sub> -D <sub>13</sub> , 3 bits fixed integer, number of vectors being transmitted for this particular vehicle.(This will probably always be 1. See further note 3)
	D <sub>12</sub> -D <sub>0</sub> are all zeros.
7	D <sub>23</sub> - D <sub>0,</sub> 24 bits all zeros.
8	Source, 4 BCD characters.
9	D <sub>23</sub> —D <sub>0</sub> 24 bits all zeros.
10	NF 11 11 2 11 11 11
H	И И И И И И
12	11 11 11 11 11
13	
14	11 11 11 11 11 11 11

# TAPE CONTENTS AND FORMAT

1-4-3

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15	D <sub>23</sub> -D <sub>20</sub> , 4 bits all zeros .
	D <sub>19</sub> -D <sub>0</sub> , 20 bits fixed integer, number of data points (or number of observations) used in fit.
16 & 17	t (machine time) see note 4.
18 & 19	x ]
20 & 21	Y position vectors; 48 bits, floating point.
22 & 23	z
24 & 25	× ]
26 & 27	$\mathbf{\dot{Y}}$ velocity vectors; 48 bits, floating point.
28 & 29	ż
30 & 31	σ (standard deviation of data)
Note I.	Word 1 is a check sum of all bits on the tape. It does not
	include the parity bits. The check sum is formed by tabu-

The overflow beyond the 24th bit is folded under and added to the previous sum to get the final check sum.

lating all 24 bit words on the tape and adding all bits.

- Note 2. Words 2 and 3 give the identification number. Any 8 legal BCD characters are permissible. It is suggested that these words be used to indicate the Spacetrack object designation (i.e. 1961 A1 for 1961 Alpha 1, or 1960 B 2 for 1960 Beta 2).
- Note 3. The need for more than one vector at any time for one vehicle is not apparent in this case. Therefore, these 3 bits should always indicate 1.

### 1-4-4

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- Note 4. Machine time is the number of seconds elapsed since 00 Hrs., GMT, on the first of the month.
- Note 5. In Figure 1 the D's indicate data bits and the P's indicate parity bits. Note that there is both horizontal and vertical odd parity. Note the repeating pattern of data word start bits. There must be a tape start header and a tape stop pattern. If there are 10 different vehicles on one physical length of tape there must be 10 tape start headers and 10 tape stop patterns, one for each vehicle. The tape must be transmitted three times in succession.



				FURMAL OF DALA LAR	S S					
HEADER OR TAPE START	DATA WORD START (DWS)	DATA WORD 1	 N ≦ D	DATA WORD 2	o≱v	DATA WORD 3	ω≷ν	DATA WORD 4	σ≥ν	etc.
		D <sub>23</sub> D D D D <sub>3</sub> P		ΟΟΟΟΡ		ΟΟΟΟΟ		σσσσσδ		
•	•	۵.	•		•	<b>ں</b> ۔	•		•	•
						2 4 2 4 2 4 2 4 2 4 2 4 2 4 2 4 2 4 2 4		о с о с о с о с о с		
		LAST DATA DATA WORD 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	TAPE STOP							
	1			Data words are 24 b		Data words are 24 bits with horizontal and vertical parity.	ertica	ıl parity.		
				F 1g. 4.1	<b>4.</b> L					

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LOCVEC

Wolf Research and Development Corporation

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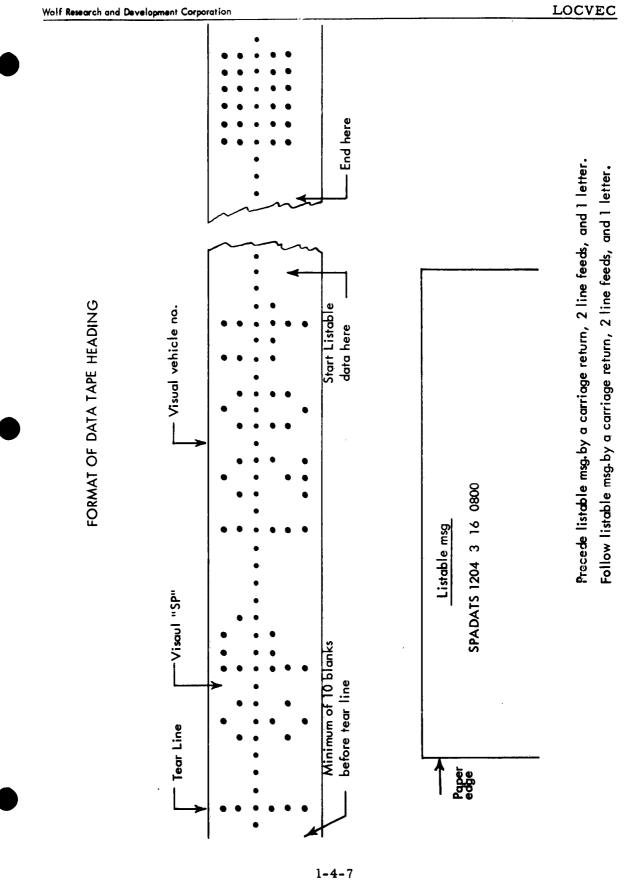


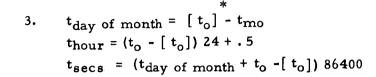
Fig. 4.2

LOCVEC

. 6	For	rmulation
	А.	Convert nodal elements to $\overline{N}$ , $\overline{M}$ elements
	1.	a = q/(1-e)
	2.	$\mathbf{p} = \mathbf{q} \left( 1 + \mathbf{e} \right)$
	3.	$n_0 = .07436574/a^{3/2}$
	4.	$a_{xn} = e \cos \omega$
	5.	$a_{yn} = e \sin \omega$
	6.	$E_{\rm N} = -2 \tan^{-1} (\sqrt{(1-e)/(1+e)} \cdot \tan \omega/2)$
	7.	$M = E_N - e \sin E_N$
	8.	$L = \omega + M + \Omega$
	9.	$C'' = (-360 n_0^2 c \pi^2)$ where $C'' = drag.$ coeff; c is rate of change of period.
	10.	$k_e L_{so} = \frac{.000120717}{p^{7/2}} \left[ 3 - 5e^2 - \left  \cos i \right  \left( 1 - \frac{3}{2} e^2 \right) - \sin^2 i \left( 4 - \frac{27}{4} e^2 \right) \right]$
	в.	Use Analytical Integration Routine (XYZSB) to compute predicted position, <u>r</u> , and velocity, <u>r</u> , and intermediate quantities. Equations for this routine may be found in Aeronutronic Publication U-1691 page 4-78.
	с.	Upon exit
	1.	<ul> <li>x = 20925725.863 x Convert earth radii to feet</li> <li>y = 20925725.863 y</li> <li>z = 20925725.863 z</li> </ul>
	2.	$\dot{x} = 25935.85859142 \dot{x}$ Convert earth radii/kemin to feet/second $\dot{y} = 25935.85859142 \dot{y}$

1-4-8

 $\dot{z} = 25935.85859142 \dot{z}$ 



Output: Satellite No., Sunnyvale sat. no., element no., <sup>t</sup>day of month, <sup>t</sup>hour, <sup>t</sup>secs, <sup>N</sup>o, <sup>x</sup>, <sup>y</sup>, <sup>z</sup>, <sup>x</sup>, <sup>y</sup>, <sup>z</sup>

\* [to] = Integer value of t<sub>o</sub>

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# 4.7 Glossary

Location	Symbol	Meaning
ADDCON		Parity indicator
AO	a	Semi-major axis at t <sub>o</sub>
AXNO	a xn	<sup>a</sup> xn components in orbit plane
AYNO	ayn	$a_{yn}$ at to of a
BODY	y II	Intermediate buffer used when output quantities converted to Baudot
С	С''	Drag coefficient
CIR		Constant = $360 (2\pi \text{ in degrees})$
CON		Constant = .05862947
COSI	cos i	Cosine of inclination
CTERM		Array of rate of change of period
DATA		31 word output buffer for data in Baudot code
DTR		Constant = .017453292519 (degrees to radians)
E	е	Eccentricity
EE		Array of eccentricities
ELAPS		Seconds in month of epoch revolution number
EO	e 2	Eccentricity at t
EOSQ, ESQ	e <sup>2</sup>	Eccentricity squared
EXPON		Exponent of floating point number
FLNU		Floating point number
FORDE		Alphanumeric E, column 80 in 7 card element sets
FPS		Constant = 25935.85859142 (earth radii/ kemin to feet/sec)
FT		Time of epoch in fractional days
HEADER		Table containing heading for paper tape output
ID		Element card number
IEL		Element number
IHOUR		Hour of epoch
ILE		Array of element numbers from request cards
IR		Epoch revolution number
ISAT		Satellite number
ISTC		Sunnyvale satellite number
IT		Time of epoch in integer days

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	J		Index equals number of element sets matched with request card satellite number
	JTI		Input tape
	JTO		Output tape
	LAPD		Day of month of epoch
	MANTIS		Mantissa of floating point number on conversion to Baudot
	МО		Array of total numbers of days since first of year by month
	MON, MONT	гн	Month of epoch
	N		Index equals number of request cards with elements available
	NAT		Satellite number
	NATNO		Array of satellite numbers from request cards
	N3PIO2		Constant - $3\pi/2$
	NSTC		Year of launch from request cards
	NSTCNO		Array of years of launch, from request cards
	NUMBODO		Table of Baudot code for digits
	P	P	Nodal period at epoch - $F_n$
	P72	$(P_n)^{7/2}$	$P_{n}^{7/2}$
	PA		Array of nodal periods at epoch in days
	PDT	с	C-rate of change of period in days/rev
	PDOT		Binary Baudot rate of change of period
	PI	π	Constant = $\pi$
	PI2	2π	Constant = $2\pi$
	PIOV2	π/2	$Constant = \pi/2$
	PTSIX	q <sub>o</sub>	Perigee distance
	QQ		Array of perigee distances
	RAN		Array of right ascensions
	REV	Ν	Epoch revolution number
	REVE		Array of epoch revolution numbers
	REVNO		Binary Baudot epoch revolution number
•	RIN		Constant = 20925725.863 (earth radii to feet)
	SAVĖ		Temporary to save contents of XRI and XR2
	SAVI		Temporary to save contents of XR3 and XR4

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SECPOC		Time of epoch in binary Baudot code seconds in month
SET		Switch set to indicate which half paper tape output buffer used
SINI	sin i	Sine of inclination
SIND	sin µ	Sine of argument of latitude
SOURCE		Country of origin of satellite from request card
SOURCEA		Array of countries of orgin
SMAN		Sign of floating point number
SNAME		Name of satellite from request card
SNAMEA		Array of names of satellite
SYMBNUM		Table of digits which yield visual symbols when punched on paper tape
Т		Time interval in minutes since time of epoch
TA	i	Inclination
TAE		Array of inclinations
TAPEOUT		128 word holding buffer for paper tape output
TEMP		Table of TTY functions
TENM6		$Constant = 10^{-6}$
TIM		Array of epoch times
 TIME	to	Time of epoch in day of year
 TWOPI	U	Constant = $2\pi$
w	ω	Argument of perigee in radians
WDOT	ώ	First derivative of argument of perigee
WN	+ · <del>•</del>	Array of arguments of perigee
WZ	cos i	Cosine of inclination
x	x	x-coordinate of <u>r</u>
<b>X</b> 1		Binary Baudot x-coordinate of r
X2		Binary Baudot x-velocity component
X23SAV		Temporary to save contents of XR2 and XR3
X4SAV		Temporary to save contents of XR4
XDOT	x	x-velocity component of r
XJGRCF		Constant = $1.20717 \times 10^{-2}$
XKE		$Constant - k_e = .0743674$
XLO	λ <sub>m</sub>	Mean longitude at t <sub>o</sub> in radians
XM302		Constant = 3/2

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XM4CO3		Constant = 4C/3
XMZ	sin i	Sine of inclination
XNO		Mean angular motion at t
XNODEO		Right ascension in radians
Y	у	y-coordinate of <u>r</u>
Y 1		Binary Baudot y-coordinate of <u>r</u>
¥2		Binary Baudot y-velocity component
YDOT	· y	y-velocity component of $\dot{\mathbf{r}}$
Z	z	z-coordinate of <u>r</u>
ZI		Binary Baudot z-coordinate of <u>r</u>
Z2		Binary Baudot z-velocity component of r
ZDOT	z	z-velocity component of $\dot{\mathbf{r}}$

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Other terms may be found in description of the subroutine XYZSB.

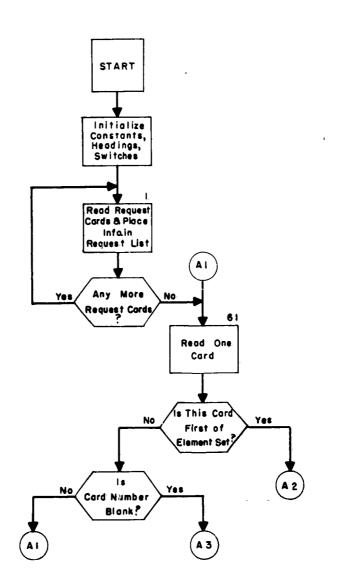
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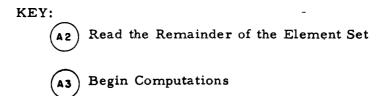
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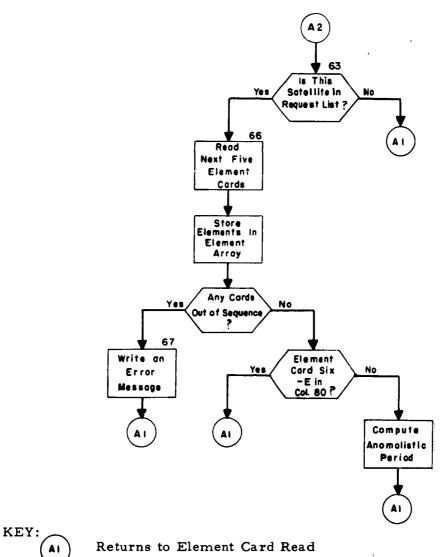
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Returns to Element Card Read

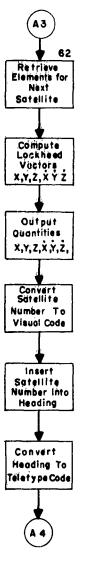
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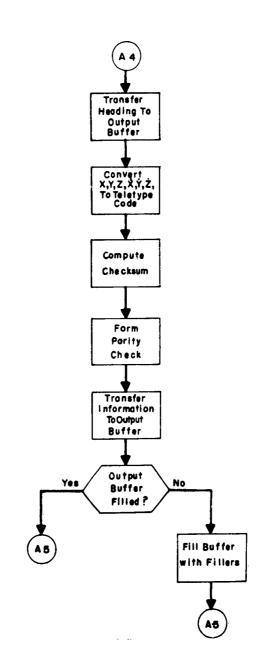




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Continues with Teletype Conversions







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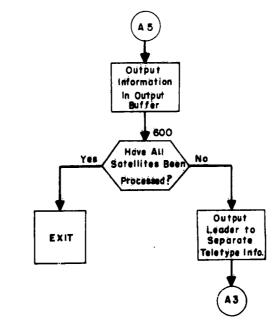
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#### CCOE, Cartesian Coordinates From Orbital Elements

### Function

From a given element set, this program calculates the Cartesian coordinates x, y, z and the components of velocity  $\frac{dx}{dt}$ ,  $\frac{dy}{dt}$ ,  $\frac{dz}{dt}$  of a satellite for a specified length of time at given intervals. The output can be expressed either in kilometers and kilometers per second, or in earth radii and earth radii per minute.

### 5.3 Input

The data deck is comprised of: 1) a standard 6 or 7 card element set, 2) as many request cards as needed and, 3) a blank card. This sequence is repeated for all satellites requiring this computation. The final sequence should be followed by a second blank card to terminate the program.

The following information should be punched on the request card:

Time start (in days)
Time step (in days)
Time stop (in days)
Blank - output in kms.

1 - output in earth radii

Decimal points may be punched anywhere in each of the first three fields.

#### 5.4 Output

The output consists of Cartesian coordinates and their related components of velocity for the time increments requested. The dimensions given are in kilometers and kilometers per second, or earth radii and earth radii per minute, depending on the output option chosen. The sentinels required by the TELTYP program are included so that conversion to teletype tape is possible.

### 5.5 Processing

After reading a standard six or seven card element set and a request card, the variables x, y, z, x, y, and z are calculated and printed according to the specified output option. The time is incremented by the time step specified on the request card. Unless the stop time has been exceeded, the program computes new variables for the incremented time.

When the stop time specified on a request card has been exceeded, the program assumes that the next card is a request card for the same satellite for another time interval. If the card is blank, the program returns to the element read section for another standard element set. A return to the executive routine is made, if a second blank is encountered. Otherwise the requests for the new satellite are processed as above.

## 5.5.1 Error Messages

- NEGATIVE NODAL PERIOD, SATNO \_\_\_\_\_. Reading continues until next case is found
- ELEMENT CARDS OUT OF ORDER. Same procedure as under (1)
- MISSING NODAL PERIOD. Continues processing that case.
   P<sub>a</sub> is used for P<sub>n</sub>.
- 4. SUBROUTINE ERROR EXIT FROM OCTAL\_\_\_\_\_. Subroutine or irrecoverable input output error. Exits to system if GO option taken. The message will be repeated if the STOP option is used. A dump should be given if the computer is in a non-interruptable mode of operation.

$$\begin{array}{rcl} 6 & \underline{Formulation} \\ 1. & P_a = & \frac{360 \ P_o}{360 - P_o} \underbrace{a} \\ 2. & a_o = \left[ \frac{P_a}{.058672947} \\ 3. & t = T_I - t_o \\ 4. & \mathbf{xx} = \frac{t}{P_o} \end{array} \right]$$

5.

N = Integer part of xx

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searc	h and Develop	ment Corporation CCOE
	6.	$\mathbf{t}_1 = \mathbf{t} - \mathbf{t}_0$
	7.	$\Omega(t) = \Omega_0 + \dot{\Omega}_0 t_1 + \frac{1}{2} \dot{\Omega}_0 t_1^2$
	8.	$\omega(t) = \omega_0 + \dot{\omega}_0  t_1 + \frac{1}{2} \dot{\omega}_0  t_1^2$
	9.	$\mathbf{t}_{n}(\mathbf{N}) = \mathbf{T}_{o} + \mathbf{P}_{o}\mathbf{N} + \mathbf{c}\mathbf{N}^{2} + \mathbf{d}\mathbf{N}^{3}$
	10.	$P_n(N) = P_0 + 2 cN + 3 dN^2$
	11.	$P_{a}(N) = \frac{360 P_{n}(N)}{360 - P_{n}(N)\omega}$
	12.	$a(N) = \left[\frac{P_a(N)}{.058672947}\right]^{2/3}$
	13.	$e(N) = 1 - \frac{a_0}{a(N)} (1 - e_0)$
	14.	$E_n = 2 \tan^{-1} \left[ \sqrt{\frac{1 - e(N)}{1 + e(N)}} \tan \left( \frac{\omega(t)}{2} \right) \right]  0 \le E_n \le 2\pi$
	15.	Is $(t - t_{\omega}(N)) < 0$ ?
		Yes: $N = N - 1$ No: Go to 16 Go to 9
	16.	If $(P_a(N) - (t - t_{\omega}(N))) \leq 0$ If $(P_a(N) - (t - t_{\omega}(N))) > 0$ N = N + 1 Go to 17 Go to 9
	17.	$M(t) = \frac{2\pi}{P_a(N)} \cdot (t - t_{\omega}(N))$
	18.	$\mathbf{E}'(\mathbf{t}) = \mathbf{M}(\mathbf{t})$
	19.	$E(t) = M(t) + e(N) \cdot sin(E'(t))$
	20.	Is $ E(t) - E'(t)  \ge 10^{-7}$
	21.	Calculates sines and cosines of $\Omega$ , $\omega$ , i
	22.	$a_x = a(N) (\cos \omega \cdot \cos \Omega - \sin \omega \cdot \sin \Omega \cdot \cos i)$
	23.	$a_y = a(N) (\cos \omega \cdot \sin \Omega + \sin \omega \cdot \cos \Omega \cdot \cos i)$
	24.	$a_z = a(N) \cdot \sin \omega \cdot \sin i$
	25.	$b_x = -a(N) \sqrt{1 - e^2(N)} (\sin \omega \cdot \cos \Omega + \cos \omega \cdot \sin \Omega \cdot \cos i)$
	26.	$b_y = a(N) \sqrt{1 - e^2(N)} (-\sin \omega \cdot \sin \Omega + \cos \omega \cdot \cos \Omega \cdot \cos i)$

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27. b <sub>z</sub> =	$a(N) \sqrt{1 - e^2(N)}$	$\cos \omega$ . sin i
28. c <sub>l</sub> =	cos (E(t)) - e(N	()
29. c <sub>2</sub> =	sin(E(t)) - e(N)	
30. c <sub>3</sub> =	$\frac{-2\pi}{Pa(N)}$ · sin (I	E(t))
	1 - e (N) · cos	(E(t))
	2π	
31. c <sub>4</sub> =	$\frac{2\pi}{Pa(N)} \cdot \cos(E)$	(t))
	$1 - e(N) \cdot cos (H)$	C(t))
	6378.174 (C <u>i</u> a <sub>x</sub> +	
33. y =	$6378.174 (C_1 a_y)$	+ C <sub>2</sub> b <sub>y</sub> )
34. z =	6378.174 ( $C_1 a_z$	+ C <sub>2</sub> b <sub>z</sub> )
35. x =	.07382146 (C <sub>3</sub> a <sub>x</sub>	+ C <sub>4</sub> b <sub>x</sub> )
36. ý =	.07382146 (C3ay	+ C <sub>4</sub> b <sub>y</sub> )
37. ż =	.07382146 (C <sub>3</sub> a <sub>z</sub>	+ C <sub>4</sub> b <sub>z</sub> )
38. $t = t +$	-	-
39. t > T		
	r · Goto6	
	Return for next	case
5.7	Glossary	
Location	Symbol	Meaning
А		Semi-major axis (on 7 card element sets only)
AMT	M(t)	Mean anomaly at time t
AN	a(N)	Semi-major axis at revolution N
AO	ao	Semi-major axis at epoch
C2 PI		2·PI
CGT	cosω	Cosine of argument of perigee
CHT	$\cos (\Omega(t))$	Cosine of right ascension at time t
CI	cos i	Cosine of inclination
со	с	Rate of change of period
CSPTRK		Nodal c term
DELT	t <sub>l</sub>	Time from epoch
DN	-	Number of complete revolutions from epoch

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DO	C.	First derivative of rate of change of period
EAA	E(t)	Approximation to eccentric anomaly
EAG	En	Eccentric anomaly of ascending node
EAT	E(t)	Eccentric anomaly at time t
EN	e(N)	Eccentricity at revolution N
EO	eo	Eccentricity
ΕT	to	Epoch time
GDOT	ш	First derivative of argument of perigee
GO	ω <sub>o</sub>	Argument of perigee
GT	ω(t)	Argument of perigee at time t
HDOT	ά	First derivative of right ascension
HGDDOT	•• ພ	Second derivative of Argumentof perigee
HHDDO T	 Ω	Second derivative of right ascension
НО	Ω <sub>o</sub>	Right ascension of ascending node
HT	Ω(t)	Right ascension at time t
IDN	N	Number of complete revolutions from epoch
IELNO		Element number
IFORD		Column 80 on fifth element card
ISATNO		Satellite number
		1, Output in earth radii and earth radii/min
KM		0, Output in kilometers and kilometers/sec
MAGIN		Input tape number, 0
MAGOUT		Output tape number, 11
NAME 1		First six characters of satellite name
NAME 2		Second six characters of satellite name
PAE	Pa	Anomalistic period of epoch
PAN	Pa(N)	Anomalistic period for revolution N
PI		π, (3. 1415926536)

HHDDOT	Ω	Second derivative of right ascension
но	Ω <sub>o</sub>	Right ascension of ascending node
HT	Ω(t)	Right ascension at time t
IDN	Ν	Number of complete revolutions from epoch
IELNO		Element number
IFORD		Column 80 on fifth element card
ISATNO		Satellite number
		1, Output in earth radii and earth radii/min.
KM		0, Output in kilometers and kilometers/sec.
MAGIN		Input tape number, 0
MAGOUT		Output tape number, 11
NAME 1		First six characters of satellite name
NAME 2		Second six characters of satellite name
PAE	Pa	Anomalistic period of epoch
PAN	Pa(N)	Anomalistic period for revolution N
PI		π, (3. 1415926536)
PNN	P <sub>n</sub> (N)	Nodal period for revolution N
PNODL	Pn	Nodal period of epoch (if 7 card element set)

PO

RHO

SGT

SHT

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ROOT

Pn

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 $\cos(\Omega(t))$ 

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Nodal period of epoch

Sine of argument of perigee

Sine of right ascension at time t

Temporary storage

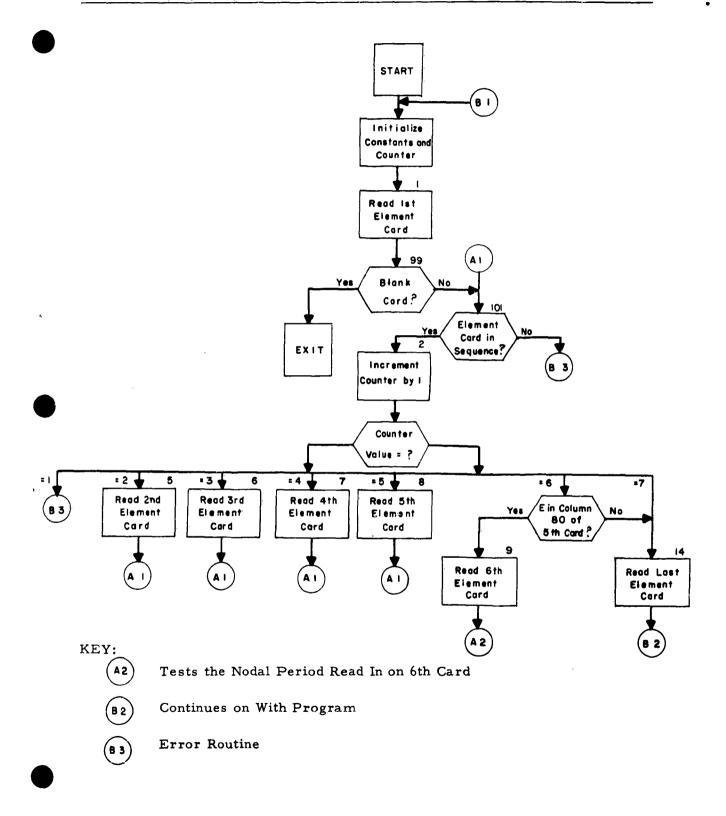
Wolf Research and Development Corporation

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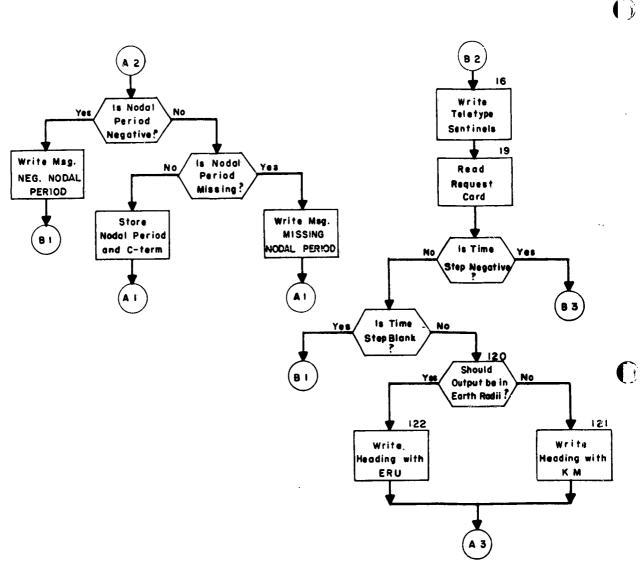
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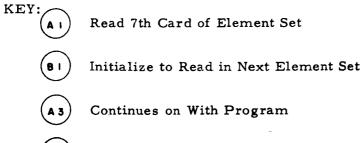
SI	Sin i	Sine of inclination
Т	τ <sub>I</sub>	Initial time in days
ТА	i	Inclination
TANN	$t_{\omega}(N)$	Time of perigee passage of revolution N
TNN	t <sub>n</sub> (N)	Time of nodal crossing for revolution N
TSTOP	т <sub>F</sub>	Final time in days
x	x	x component of position vector of satellite
xx	_	Revolutions from epoch
XDOT	$\frac{dx}{dt}$	x component of velocity
Ŷ	ÿ	y component of position vector of satellite
YDOT	dy dt	y component of velocity
Z	Z	z component of position vector of satellite
ZDOT	$\frac{dz}{dt}$	z component of velocity

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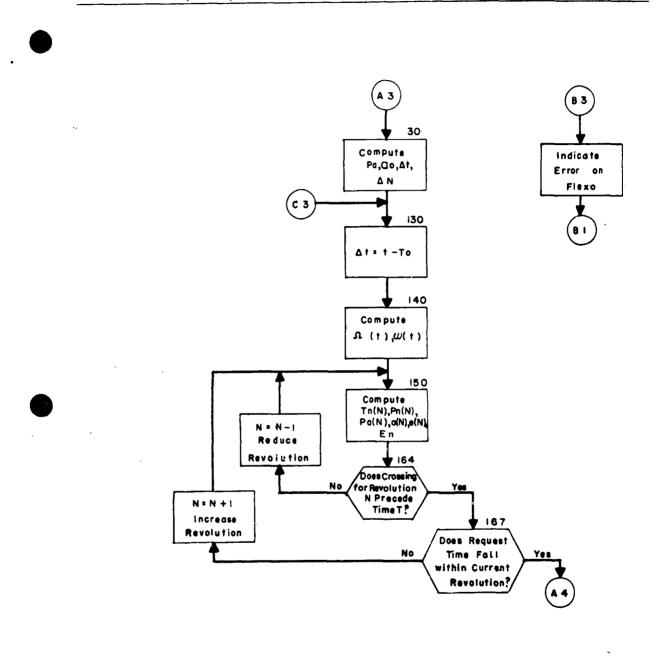
Error Routine

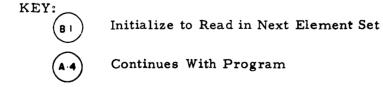
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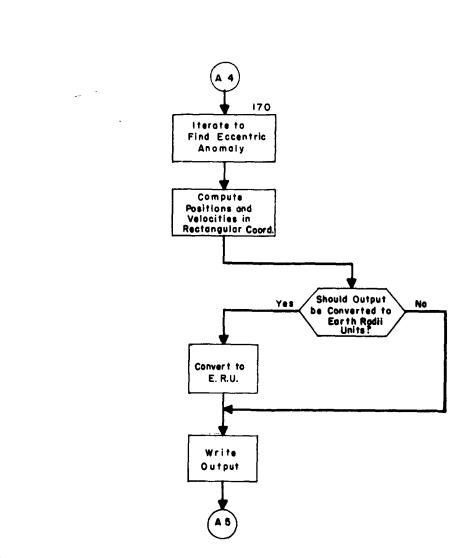




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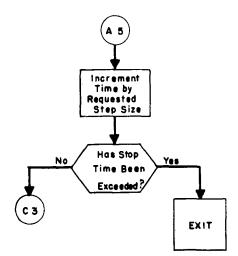
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Continues With Program

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### 6.1 RESPLT, Residual Plot

### 6.2 Function

RESPLT reduces observations against the N, <u>M</u> element sets and produces punched card output which can be used on the EAI Data Plotter. The residuals which can be plotted are: (1) the difference in time between the predicted and observed positions versus the revolution and, (2) the vector magnitude difference versus the revolution.

The plots are especially useful as an aid in the analysis of element sets. Large residuals may indicate that the current element sets need updating or that a piece has broken off the satellite. Maneuvers, not otherwise easily detected, may be revealed when these residuals are presented graphic form (cf. Fig. 6.2).

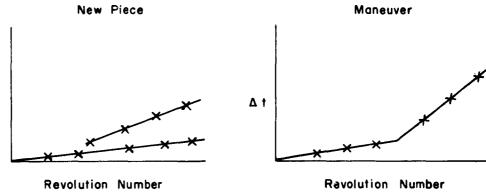
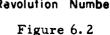


Figure 6.1



### 6.3 Input

Δt

The program requires observation data to be entered as card input. Element data and sensor data may be entered as card input or read from the SEAI tape.

The observations must be obtained from standard observation cards which have previously been sorted by satellite number. These

cards must have an association status punched in column 80 (cf. Fig. 6.6). If untagged observations (satellite number = zero) are to be reduced against a particular element set, a standard observation card containing the number of the desired element set and a blank sensor number should be placed in the observation deck preceding the untagged observations. The number of observation cards allowed is limited only by the space allocated to OBLOC.

There are six input options which may be specified on the SPSJOB card. Each option specifies the source of the sensor and element data, i. e., input cards or SEAI tape.

Option	Observation	Element	Sensor
0	Ob. Cards	SEAI Tape	SEAI Tape
1	Ob. Cards	Element Cards	SEAI Tape
2	Ob. Cards	Element Number Cards	SEAI Tape
3	Ob. Cards	SEAI Tape	Sensor Cards
4	Ob. Cards	Element Cards	Sensor Cards
5	Ob. Cards	Element Number Cards	Sensor Cards

Fig. 6.3

Element data may be obtained from standard 7 card element sets with an E in column 80. Since the executive routine reads the element sets into EBLOC, the number of element sets is limited by the size of that block. Element data may be obtained from the SEAI tape if element sets are not included in the input deck.

Sensor data may be obtained from standard sensor cards with an S in column 80. The number is restricted only by the size of SBLOC. Sensor data may be obtained from the SEAI tape if no sensor cards are included in the input deck.

### 6.4 Output

Printed output consists of the reduced observations sorted by revolution number within a satellite number in groups of 100. A heading line, which describes the output quantities in the columns below, is printed for each satellite. The satellite number, observation number, revolution number, time in minutes since epoch, vector magnitude in kilometers, number of revolutions since epoch, element number, association indicator, and station number are printed for every revolution number.

The punched card output can be used to plot residuals on the EAI Data Plotter using  $K + E 10 \times 10$  cm. graph paper. The pen command for the Data Plotter is contained in column 55. The data deck is divided up as follows:

- A card with a command to stop the plotter (7 in col. 55).
   The operator can set the origin at this time.
- A set of cards which will enable the plotter to draw the x and y axes (6 cards).
- A card which will stop the plotter so that the above can be tried again if necessary.
- 4) A set of cards which will plot the visual characters S, E, R, T and the information required to identify the plots. S will be followed by the satellite number, E will be followed by the element number, R will be followed by the epoch revolution number, and T will be followed by the the time of epoch in days.
- 5) A card which will stop the plotter (7 in col. 55).
- The data cards containing the residual information to be plotted. There is one card for each line of printed output.

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# IBM 523 Summary Punch

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# VMag(y) vs $\triangle \text{Rev}(x)$

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1-6-5

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# 6. 4.1 EAI Data Plotter Operating Instructions

An IBM 523 Summary Punch is used to read the punched cards for input to the plotter. Board wiring diagrams are included for each of the two possible graphs (cf. Fig. 6.4 and Fig. 6.5).

The residuals which may be plotted are data time versus data revolution and vector magnitude versus data revolution. Coordinates and pen commands for both graphs appear on the same data cards.

Note that it is possible to plot both graphs on the same piece of graph paper by reading the data cards through the 523 Summary Punch twice. This is accomplished by reading all of the punched cards for the first graph. The axes and visible information as well as the points on the graph will be plotted. If part six (6) of the data deck is reread, the second set of the residuals may be plotted on the same graph paper.

The plotter pen color and 523 board must be changed before rereading the data cards. Two separate graphs may be produced by reading all punched cards twice and changing the graph paper and 523 board.

For the proper plotting of all data, the origin on the graph paper should be five (5) centimeters right and nine (9) centimeters up from the bottom left corner of the graph paper.

Graph scaling is as follows. The revolution number from epoch is the x-axis for both graphs. The epoch revolution number lies at x = 0. Each centimeter represents 20 revolutions, so that the possible values for plotting lie in the range of -100 and + 400 revolutions. Vector magnitude in kilometers is the y-axis for one graph. Each centimeter represents 500 kilometers and the possible range of values is 0 to 4500 kilometers. Delta time in

minutes is the y-axis for the other graph. Each centimeter represents five (5) minutes and the possible range of values is -45 to +45 minutes.

## 6.5 Processing

The program reduces observations against element sets to obtain time and vector magnitude residuals which may be plotted against the revolution number. Up to 100 observations may be processed one at a time. The association status is checked to determine whether the observation is an angles observation only or a radar observation.

Association Status	Observation Kind	Association Category	Will RESPLT Accept?
1	Radar	Associated	Yes
2	Radar	Doubtful	Yes
3	Radar	Unassociated	Yes
4	Angles Only	Associated	Yes
5	Angles Only	Doubtful	Yes
6	Angles Only	Unassociated	Yes
7	Range Rate	Associated	No
8	Range Rate	Unassociated	No
9	Radar	Special Unassoc.	Yes
O or <b>D</b>			No

Fig. 6.6

The proper year constants are obtained and time of epoch is converted to days since 1950. The observation counter 's increased by one and the association indicator, observation number, and sensor number from the observation are stored for output. The sidereal time at the time of epoch is computed. The semi-latus rectum and the components of the unit vector perpendicular to the orbit plane are found to compute the inclination, the right ascension

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of the ascending node, the eccentricity, and the semi-major axis at epoch.

The mean argument of latitude, the mean angular motion, a drag coefficient, and the perifocal distance are computed. The time since epoch, the sidereal time at the station, and the x, y components of the station vector are found. The subroutine XYZSB is used to compute the predicted position and velocity plus intermediate quantities.

Observation type (OTYPE) will be referred to in subsequent paragraphs as illustrated in Fig. 6.7

OTYPE	Observation Kind	Quantities From Obs. Card
1	Angles Only	Azimuth-Elevation
2	Radar	Azimuth-Elevation, Range
3	Radar	Azimuth-Elevation, Range, Range Rate
4	Radar	Azimuth-Elevation, Range Rate
5	Angles Only	Right Ascension-Declination

Fig. 6.7

If the observation is an angles only observation, the range is computed. If OTYPE = 1, the predicted azimuth and elevation are found. If OTYPE = 5, the predicted right ascension and declination are found.

If the observation is a radar observation, the predicted range is computed. The predicted azimuth and elevation are computed. If OTYPE = 3 or 4, the predicted range rate is computed.

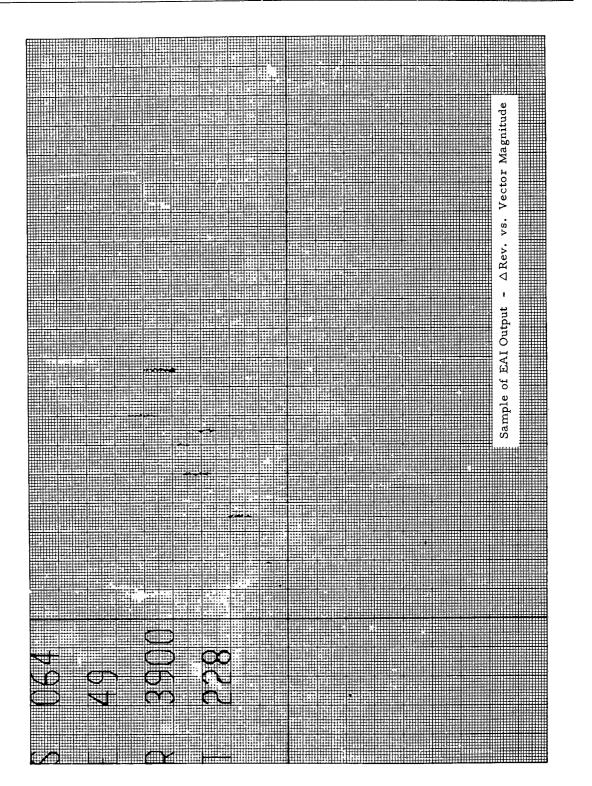
For all observations, the residuals are computed to find the vector magnitude. Unless 100 observations have been processed, the program returns for the next observation.

When all of the observations have been processed for a single satellite the program generates data for punched card output for the plotter. The punched cards contain all necessary plotter control functions. Punched cards are generated to plot the axes, visual satellite, element, epoch revolution, and epoch day numbers. These cards are followed by the data cards. The program then returns to process the next group of observations.

### 1-6-9

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1-6-10

6.6

### Formulation

- 1.  $\theta_G = \theta_0 + t_{days}$  (rotation of earth) deg/solar day
- 2. Compute semi-latus rectum.  $p_o = h_{xo}^2 + h_{yo}^2 + h_{zo}^2$
- 3. Compute x, y, z components of <u>W</u> (unit vector  $\perp$  orbit plane).

$$W_{xo} = h_{xo} / \sqrt{p_o}$$
$$W_{yo} = h_{yo} / \sqrt{p_o}$$
$$W_{zo} = h_{zo} / \sqrt{p_o}$$

4. Compute inclination.

i = 
$$\pi/2$$
 if W = 0, otherwise;  
i =  $\tan^{-1}\left(\sqrt{1 - W_{zo}^2/W_{zo}}\right)$ 

5. Compute right ascension of ascending node at epoch.

$$\Omega_{o} = \tan^{-1} \left[ (W_{xo}/\sin i) / (-W_{yo}/\sin i) \right]$$

6. Compute eccentricity and semi-major axis at epoch.

$$e_{o} = \sqrt{a_{xNo}^{2} + a_{yNo}^{2}}$$
$$a_{o} = p_{o}/1 - e_{o}^{2}$$

7. Compute mean argument of latitude and mean angular motion at epoch.

$$U_o = L_o - \Omega_o \text{ if } W_{zo} \text{ positive, otherwise } u_o = L_o + \Omega_o$$
  
 $n_o = k_e / a_o^{3/2}$ 

8. Compute a drag coefficient, perifocal distance, and  $K_e L_{so}$ .  $c'' = -c_o (n_o^2) 360 / \pi^2$  $q_o = a_o (1 - e_o)$ 

$$K_{e}L_{so} = \frac{K_{e}Ja_{e}^{2}}{\frac{p_{o}^{7/2}}{p_{o}}} \left[ 3 - 5e_{o}^{2} - (4 - \frac{27}{4}e_{o}^{2})\sin^{2}i - (1 - \frac{3}{2}e_{o}^{2})/\cos i \right]$$

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9. Compute time since epoch and sidereal time at station.

T = 
$$(t_{obs} - t_{o})$$
 1440 + time of day in minutes since midnight.  
 $\theta_{s} = \theta_{c} + \lambda + .0043752691 T$ 

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10. Compute observed position from the observation.  $\overline{X} = \cos \theta \quad \overline{X} / \cos \theta$ 

$$\frac{\underline{X}}{\underline{Y}} = \cos \theta_{s} \frac{\underline{X}}{\underline{C}} \cos \theta_{s}$$
$$\frac{\underline{Y}}{\underline{Y}} = \sin \theta_{s} \frac{\underline{X}}{\underline{C}} \cos \theta_{s}$$

Use subroutine XYZSB to compute predicted position,  $\underline{r}$ , and velocity,  $\underline{r}$ , plus intermediate quantities.

If angles only observations, go to 17.

Equations 11-16 for radar observations.

11. Compute x, y, z components obs. unit vector from observer to obj. with respect to horizon.

 $L_{xh} = -\cos Az \cosh h$ 

 $L_{vh} = sin Az cos h$ 

 $L_{zh} = \sinh h$ 

Compute x, y, z components S (unit vector from observer to south).

$$S_x = \sin \phi \cos \theta_s$$

$$S_{y} = \sin \phi \sin \theta_{s}$$

$$S_z = -\cos \phi$$

Compute x, y, z components Z (unit vector from observer to zenith).

 $Z_{x} = \cos \phi \cos \theta_{s}$ 

 $Z_v = \cos \phi \sin \theta_s$ 

 $Z_{z} = \sin \phi$ 

14. Compute x, y, z components  $\underline{E}$  (unit vector from obs. to east).  $\underline{E}_{v} = -\sin \theta_{v}$ 

$$E_y = \cos \theta_s$$

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 $E_z = 0$ 

15. Compute  $\underline{L}_{0}$  (obs. unit vector from obs. to object).

$$\underline{\mathbf{L}} = \mathbf{L}_{\mathbf{xh}} \underline{\mathbf{S}} + \mathbf{L}_{\mathbf{yh}} \underline{\mathbf{E}} + \mathbf{L}_{\mathbf{zh}} \underline{\mathbf{Z}}$$



16. Compute ob.  $\underline{r}$ , vector directed to object, and  $\beta$ .

$$\frac{\mathbf{r}}{\mathbf{r}} = \rho \underline{\mathbf{L}} - \underline{\mathbf{R}}$$
$$\mathbf{r} = \sqrt{\mathbf{x}^{2} + \mathbf{y}^{2} + \mathbf{z}^{2}}$$
$$\beta = \left| \sin^{-1} \left( \frac{1}{\mathbf{r}} (\underline{\mathbf{R}} \cdot \underline{\mathbf{W}}) \right) \right|$$

Go to 30.

Equations 17-29 for angles only observations.

17. Compute x, y, z components of range vector  $\boldsymbol{\rho}.$ 

$$\frac{\rho = \mathbf{r} + \mathbf{R}}{\rho_{c} = \sqrt{\rho_{x}^{2} + \rho_{y}^{2} + \rho_{z}^{2}}}$$
If Az-El type observation, go to 18.

If RA-Dec type observation, go to 20.

18. Compute predicted elevation.

$$L_{xh} = \frac{\rho_x \cos \theta_s \sin \phi + \rho_z \sin \theta_s \sin \phi - \rho_z \cos \phi}{\rho}$$

$$L_{yh} = \frac{-\rho_x \sin \theta_s + \rho_y \cos \theta_s}{\rho}$$

$$L_{zh} = \frac{\rho_x \cos \theta_s \cos \phi + \rho_z \sin \theta_s \cos \phi + \rho_z \sin \phi}{\rho}$$

$$h_c = \tan^{-1} \left(\frac{L_{zh}}{\sqrt{1 - L_{zh}^2}}\right)$$
Compute predicted azimuth.

19. Compute predicted azimuth.

$$Az_{c} = \tan^{-1} \left( \frac{L_{yh}}{-L_{xh}} \right)$$

Go to 22.

20. Compute predicted right ascension.

$$\frac{\mathbf{L}}{\mathbf{a}_{c}} = \frac{\rho}{\rho}$$
$$\mathbf{a}_{c} = \tan^{-1} \left( \frac{\mathbf{L}}{\mathbf{L}_{x}} \right)$$

21. Compute predicted declination.



$$\delta_{c} = \tan^{-1} \left( \frac{L_{z}}{1 - L_{z}^{2}} \right)$$

Go to 27.

22. Compute x, y, z components obs. unit vector from observer to object with respect to horizon.

 $L_{xh} = \cos \delta \cos a$  $L_{yh} = \cos \delta \sin a$  $L_{zh} = \sin \delta$ 

- 23-26. Same as equations 12-15. Go to 28.
  - 27. Compute x, y, z components  $\overline{L}_{2}$ .

 $L_{x} = \cos \delta \cos \alpha$  $L_{y} = \cos \delta \sin \alpha$ 

 $L_{z} = sin\delta$ 

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28. Compute observation range.

$$= (\mathbf{X} \cdot \mathbf{W}_{\mathbf{x}\mathbf{o}} + \mathbf{Y} \cdot \mathbf{W}_{\mathbf{y}\mathbf{o}} + \mathbf{Z} \cdot \mathbf{W}_{\mathbf{z}\mathbf{o}}) / (\mathbf{L}_{\mathbf{x}} \cdot \mathbf{W}_{\mathbf{x}\mathbf{o}} + \mathbf{L}_{\mathbf{y}} \cdot \mathbf{W}_{\mathbf{y}\mathbf{o}} + \mathbf{L}_{\mathbf{z}} \cdot \mathbf{W}_{\mathbf{z}\mathbf{o}})$$

29. Compute predicted range.

$$\rho_{c} = \sqrt{\rho_{x}^{2} + \rho_{y}^{2} + \rho_{z}^{2}}$$

30. Compute  $\Delta u$ , change in mean argument of latitude.

$$\Delta u = \tan^{-1} \left( \frac{\underline{U}}{\underline{U}_{x} \cdot \underline{W}} \right)$$

31. Compute  $\Delta t$ .  $\Delta t = (\Delta u \underline{R}^2) / (K_e \sqrt{p})$ 

If angles only observation, go to 36.

- 32. Same as equation 17.
- 33-34. Same as equations 18-19.

35. If Az-El, Range Rate Obs. (OTYPE = 4), compute predicted range rate.

 $\dot{\overline{\mathbf{X}}} = -\overline{\mathbf{Y}}\dot{\boldsymbol{\theta}}$  $\dot{\overline{\mathbf{Y}}} = \overline{\mathbf{X}}\dot{\boldsymbol{\theta}}$ 

1-6-14

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 $\overline{Z} = 0$  $\dot{\rho}_{c} = \mathbf{L} \cdot (\dot{\rho} + \dot{\mathbf{R}})$ 36. Compute range residual.  $\Delta \mathbf{r} = \rho - \rho_c$ 37. If Az-El type observation, (OTYPE  $\neq$  5).  $\Delta h = h - h_c$ ,  $\Delta a = |Az - Az_c|$ If RA-Dec type observation, (OTYPE = 5).  $\Delta h = \delta - \delta_c, \quad \Delta a = \left| a - a_c \right|$ 38. If  $\Delta h > \pi$ ,  $\Delta h = 2\pi - \Delta h$ If  $\Delta a > \pi$ ,  $\Delta a = 2 \pi - \Delta a$ If angles only observation, go to 40. 39. Compute vector magnitude.  $V_{mag} = \sqrt{\Delta r^2 + (\rho \Delta h)^2 + (\rho \Delta a \cosh)^2}$ Go to 41. 40.  $V_{mag} = \sqrt{(\rho_c \Delta h)^2 + (\rho_c \Delta a \cosh)^2}$ 41. Compute revolution and delta revolution number.  $N = N_0 + \frac{t}{2\pi}$  $\Delta N = \frac{t}{2\pi}$ 6.7 File Definitions Symbol Meaning ALPHA Azimuth in radians or right ascension ASTAT R = association indicator = 1 - 9 (BCD)

File Core Image 0 48 bit flt. pt. number 0 0 0 0 0 0 0 0 R 48 bit floating point  $\mathbf{E}$ E 48 bit floating point Expiration of bulletin, in days since  $\mathbf{E}$ 48 bit floating point S Floating point If other than  $\Delta$  or 0, indicates that

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AXNO

AYNO

CAPZ

CFI

BLEXP

<sup>a</sup>xN

<sup>a</sup>vN

1950.0

cyc/sec.<sup>2</sup>

Z, in earth radii

1-6-15

Process next observation.

RODOT contains max. freq. shift in

0 0 0 0 0 0 0 CF

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со	c, in days/rev. <sup>2</sup>	E	48 bit floating point
CTYPE	Coordinate type for LA	S	0000000@
DAYNO	Day number (relative to 1950.0)	0	48 bit flt. pt. number
DELTA	Elevation in radians, or declination	0	48 bit flt. pt. number
ELNO	Element set no. (binary)	E	Binary integer
EFOCH	T days since 1950	E	48 bit floating point
EQTYP	CL = class. T = type of equipment	0	$CL \Delta \Delta \Delta \Delta \Delta T T$
нхо	h xo	E	48 bit floating point
нүо	h yo	E	48 bit floating point
HZO	h zo	E	48 bit floating point
ISTOP	Inclination correction indicator	E	0 0 0 0 0 0 0 I
MSGNO	5 character message number	Ö	ΔΔΔΜΜΜΜ
OALT	H, in earth radii	S	Floating point
OBRITE	Brightness in apparent magnitudes	0	Same as tape record
OBSNO	Observation number	0	ΔΔΔΧΧΧΧΧ 👝
OTYPE	Observation type	0	0 0 0 0 0 0 0 OT
	<pre>0 = range rate only 1 = azimuth and elevation 2 = azimuth, elevation, range 3 = azimuth, elevation, range,     range rate 5 = right ascension and declination</pre>		
PHANG	$C_{p}$ phase angle coeff. in degrees	E	Floating point
PHIRD	$\phi$ in radians	S	Floating point
RANGE	Slant range in earth radii	0	48 bit flt. pt. number
REVF	Final revolution number	E	Floating point
REVI	Epoch rev. no. or initial rev. no.	E	Floating point
RODOT	Range rate in earth radii/kemin	ο ΄	48 bit flt. pt. number
SATN	Satellite number (BCD)	E	ΔΔΔΔΒΒΒ
SATNM	Satellite name	E	ΝΝΝΝΝΝΝ
SATNM2	Satellite name	E	ΝΝΔΔΔΔΔΔ
SATNO	Satellite number	0	ΔΔΔΔΔ S S S
SATOB	Number of observations	E	48 bit floating point

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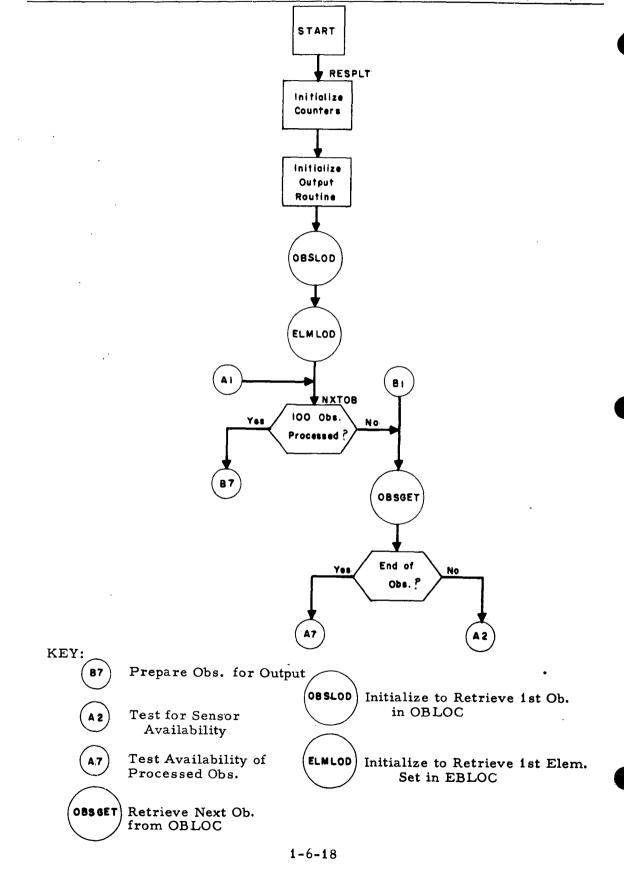
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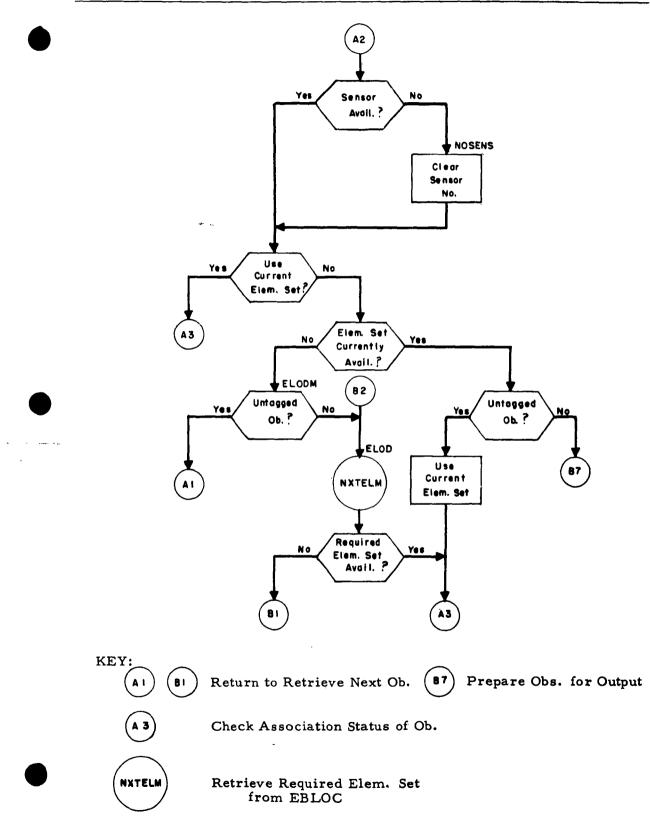
SATRM	Rms in kilometers	E	48 bit floating point
STAID	Station number	0	ΔΔΔΔΝΝΝ
	Station number	S	ΔΔΔΔΒΒΒ
STBRT	Standard brightness	E	Floating point
STGAR	F-sign bit on-classified	S	F @ A H p
STNM	Sensor code	S	ΝΝΝΝΝΝΝ
STNM2	Sensor name (BCD)	S	ΝΝΝΝΝΝΝ
STY <b>P</b> E	Sensor type	S	00000AAA
TOY	Year of $T_0 - 1$ BCD character	E	0 0 0 0 0 0 0 0 Toy
UTIME	Time of day (min. since midnight)	0	48 bit flt. pt. number
XLAMBA	$\lambda$ radians east	S	Floating point
XLO	L <sub>o</sub> radians	E	48 bit floating point
XOVCT	$X/\cos\theta$ earth radii	S	Floating point

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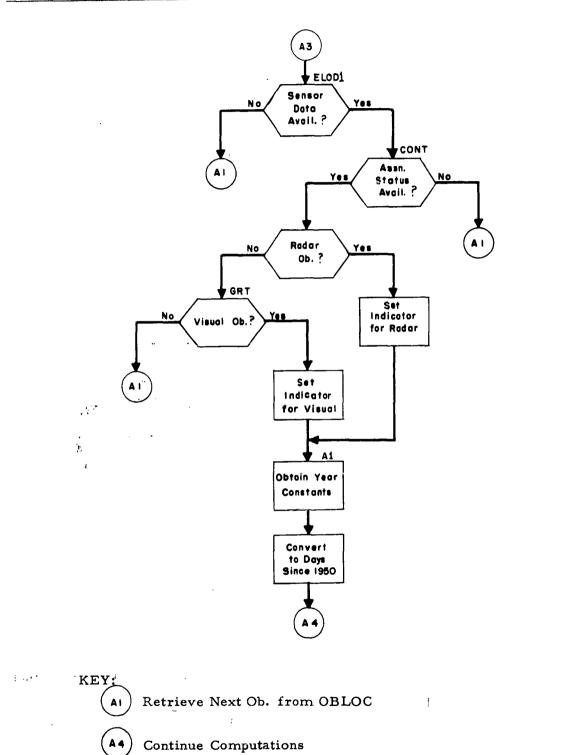


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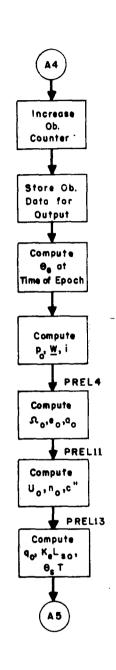
### 1-6-20

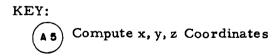
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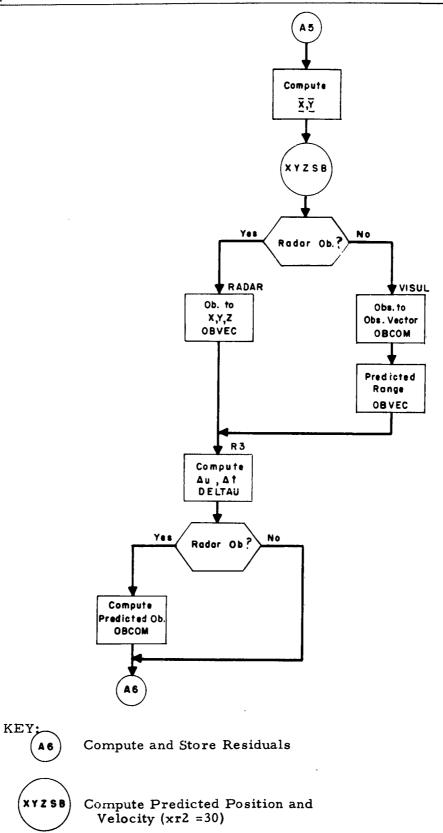
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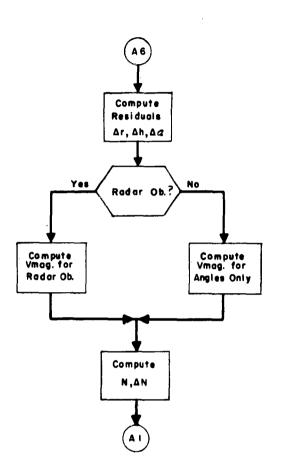


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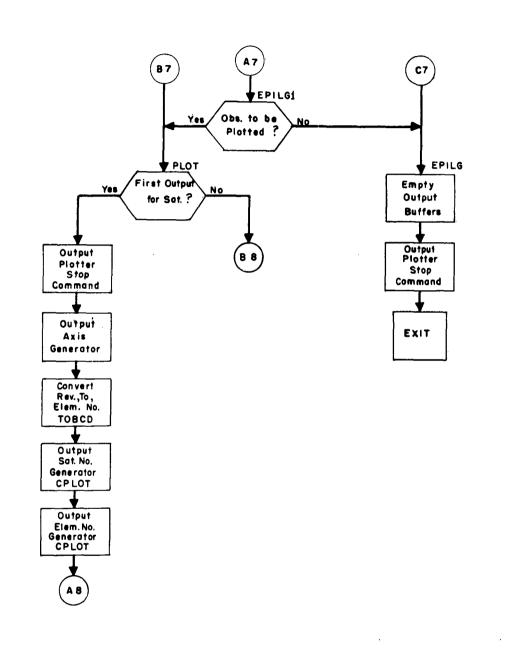
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) Continue Plot Header Generation

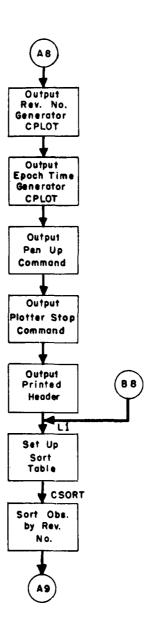
Begin Ob. Sort by Rev. No.

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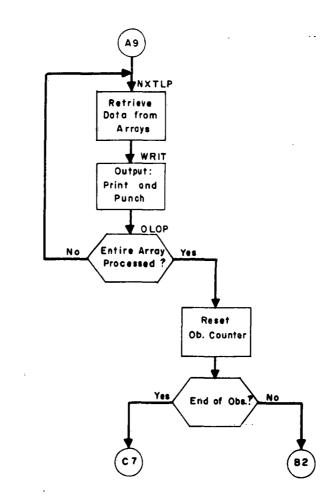
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Retrieve Ob. Data from Array Storage

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Empty Output Buffers and Exit to Exec.

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### 7.1 PREPINT, Satellite Situation Report From Nodal Elements

### 7.2 Function

The purpose of PREPINT is to supply the sub-satellite point and related data of all satellites at specified times. The longitude and time at the last ascending node are also computed for each satellite.

A maximum of 500 satellites may be included in each report. However, positions will be computed only for those satellites included in the request deck (See 7.3). If any comments (such as "IN HELIOCENTRIC ORBIT") appear on the request card of a satellite, these comments are printed out and the program does not attempt to determine the sub-satellite point.

During one run of PREPINT, as many as 10 reports may be produced, one report for each time card (See 7.3). If a report time is 1200 hours, the output heading will be WEEKLY SATELLITE SITUATION REPORT. A daily report is issued at all other times but both types of report are identical other than heading.

# 7.3 Input

The input to PREPINT is arranged as follows: 1) request cards ( $\leq 500$ ); 2) request deck terminator; 3) time cards ( $\leq 10$ ); 4) time deck terminator; 5) standard 6 or 7 card element sets; and 6) element deck terminator.

(1) The request cards contain the following information:

	= 1, read comments from cols. 21 - 62.
Col. 7	Comment indicator = 0 or $\Delta$ , no comments;
Cols. 1-3	Satellite number

### 1-7-1

### Cols. 9-20 Satellite name

Cols. 21-62 Comments to be printed on report if Col.7 contains a "1" punch. (A maximum of 50 satellites may have comments.)

All other columns on this card are not used.

(2) The request deck is terminated by a card with a "2" punch in Col. 7

(3) The time cards contain the times at which reports are desired. They are punched as follows:

Cols. 1-4 Year
Cols. 6-7 Month number
Col. 8 Blank
Cols. 9-10 Day of month
Cols. 11-17 Hours and minutes (HHMM·MM)
Col. 20 Output unit indicator = 0, output in statute mile
= 1, output in kilometers

All other columns are not used.

(4) The time deck is terminated by a card with a "7" punch in Col. 8.

(5) Both standard 6 card element sets and standard 7 card element sets may be included in the element deck.

(6) A blank card is used as an end of input indicator.

The ordering of the satellites on the output is determined by the order of the request deck, not by the element set order. More than 500 sets of elements may be read in since the program stores only the elements for satellites in the request deck.

### 7.4 Output

The heading of each satellite situation report contains the date and time for the information following it. One line of data is

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printed for each satellite. If comments were read from the request card, these comments are printed out beside the satellite number. Otherwise, the following information is printed out for each satellite:

1) Identifying information: satellite name and number, and element set number.

2) Sub-satellite point at report time: latitude and longitude west, in degrees.

3) Inclination, i, in degrees and nodal period,  $P_N$ , in minutes.

4) Distances to apogee and perigee in statute miles or kilometers, depending on Col. 20 of the time card.

5) Revolution number at report time and  $T_{\rm N}^{}, \Omega_{\rm N}^{},$  and  $\lambda_{\rm N}^{}$  for this revolution.

6) Eccentricity, e, satellite height in statute miles or kilometers.

7) Satellite latitude and longitude west in tenths of a degree.

The sentinels required by the TELTYP Program are supplied to enable transmission of the above output including all information from 1) to 4).

### 7.5 Processing

The entire input deck is read in and the data stored in arrays before any computations are made. First, a maximum of 500 request cards are read in. The satellite identification and comments, if any, from the request deck are saved. Next, the report times and output options are read from the time cards and stored. Finally, the element deck is read in and the elements for all satellites in the request deck are stored. For all element sets which are out of order, the satellite and card numbers are printed out via the flexowriter with an appropriate message. After the element deck terminator is read in, a check is made to see if there were any

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element sets out of order. If there were cards out of order, the program returns control to the executive program. Otherwise, computations for the first report are started.

After retrieving the first report time from the time array, the teletype sentinels and report headings are written on the output tape. Next, the first satellite number is retrieved from the request deck array. If comments appeared on the request card for this satellite, they are written on the output tape and the next satellite number is picked up from the request deck array. For satellites having no comments, the elements are retrieved from the element array for computation of the sub-satellite point at report time. First, the revolution number at report time is computed and all elements are updated to the time of the ascending node for this revolution. If the satellite has decayed prior to report time, an appropriate message is written on the output tape and the next satellite number is retrieved from the request deck array. However, if the satellite is still in orbit, the sub-satellite point at report time and other data for output are computed. The output is converted to the proper units and written on the output tape. The next satellite number is retrieved from the request deck array and the computations continue until all requested satellites have been processed.

The next report time is retrieved from the time array and the same procedure followed until reports have been completed for all requested times. The program then returns control to the executive program.

### 7.5.1 Error Messages

1. SATELLITE \_\_\_\_ CARD \_ \_\_ OUT OF ORDER

Program continues reading entire element deck and writes this message for every card out of order. After all elements have been read in, program exits to executive routine.

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# 2. SUBROUTINE ERROR EXIT FROM OCTAL

Subroutine or irrecoverable input-output error. Program exits to executive routine if the GO option is taken. If the STOP option is specified, the message is retyped. A dump should be taken if possible.

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7.6	Forn	nulation		
	1.	$P_{a_o} = \frac{P_{N_o}}{1 - \frac{\omega P_{N_o}}{2\pi}}$		
,	2.	$a_o = \left(\frac{P_{a_o}}{.058672947}\right)^{2/3}$		
	3.	$q_o = a_o (1 - e_o)$		
	4.	$\Delta t_{N} = P_{N_{0}} \Delta N + C_{0} \Delta N^{2} + d_{0} \Delta N^{3}$		
	5.	$P_N = P_{N_0} + 2 C_0 \Delta N + 3 d_0 \Delta N^2$		
	6.	$C_N = C_0 + 3 d \Delta N$		
	7.	$\dot{\mathbf{q}}_{\mathbf{N}} = \dot{\mathbf{q}}_{0} + \dot{\mathbf{q}} \Delta \mathbf{t}_{\mathbf{N}}$		
	8.	$q_N = q_0 + \dot{q}_N \bigtriangleup t_N + \frac{1}{2} \ddot{q}_0 \bigtriangleup t_N^2$		
	9.	$\omega_{\rm N} = \omega_{\rm o} + \dot{\omega}_{\rm o} \triangle t_{\rm N} + \frac{1}{2} \ddot{\omega}_{\rm o} \triangle t_{\rm N}^2 ,$	$0 \leq \omega_{\rm N} < 2\Pi$	Ċ
	10.	$\Omega_{\rm N} = \Omega_{\rm o} + \dot{\Omega}_{\rm o} \bigtriangleup t_{\rm N} + \frac{1}{2} \ddot{\Omega}_{\rm o} \bigtriangleup t_{\rm N}^2$ ,	$0 \leq \Omega_{N} < 2\Pi$	
	11.	$P_{a_{N}} = \frac{P_{N}}{1 - \frac{\omega_{N}P_{N}}{2\Pi}}$		
	12.	$a_{N} = \left(\frac{P_{a_{N}}}{.058672947}\right)^{2/3}$		
	13.	$\mathbf{e}_{\mathbf{N}} = 1 - \frac{\mathbf{q}_{\mathbf{N}}}{\mathbf{a}_{\mathbf{N}}}$		-
	14.	$\dot{\omega}_{\rm N} = \frac{(.086917)(a_{\rm N})^{-7/2}(5\cos^2 i - 1)}{(1 - e_{\rm N}^2)^2}$		
	15.	$\hat{\Omega}_{N} = \frac{(17383)(a_{N})^{-7/2} (\cos i)}{(1 - e_{N}^{2})^{2}}$		
	16.	$q_N = a_N (1 - e_N)$		

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17. 
$$E_{N} = -2 \tan^{-1} \left( \sqrt{\frac{1 + e_{N}}{1 + e_{N}}} \tan \frac{\omega_{N}}{2} \right)$$
18. 
$$M_{N} = E_{N} - e_{N} \sin E_{N}$$
19. 
$$M(t) = M_{N} + \frac{2\Pi + \Delta t}{P_{N}}$$
20. Solve for  $E(t) : M(t) = E(t) - e_{N} \sin E(t)$ 
21. 
$$v = 2 \tan^{-1} \left( \sqrt{\frac{1 + e_{N}}{1 - e_{N}}} \tan \frac{E(t)}{2} \right)$$
22. 
$$\mu = v + \omega_{N} + \dot{\omega}_{N} \Delta t + \frac{1}{2} \dot{\omega}_{N} \Delta t^{2}$$
23. 
$$\phi' = \sin^{-1} (\sin i \cdot \sin \mu)$$
24. 
$$\phi = \tan^{-1} \left( \frac{\tan \phi'}{.99329985} \right)$$
25. 
$$R_{o} = \frac{.9966443}{\sqrt{1 - .00670015 \cdot \cos^{2} \phi}}$$
26. 
$$r = \frac{a_{N} (1 - e_{N}^{2})}{(\mathbf{i} + e_{N}) \cos v}$$
27. 
$$H = r - R_{o}$$
28. 
$$\beta = \sin^{-1} \left( \frac{\cos i}{\cos \phi'} \right)$$
29. 
$$\Delta \lambda' = \sin^{-1} \left( \frac{\tan \phi'}{\tan i} \right)$$
30. 
$$\Delta \lambda = \Delta \lambda' + \Delta t \left( \dot{\Omega}_{N} - 6.3003883 \right) + \frac{1}{2} \ddot{\Omega}_{N} \Delta t^{2}$$
31. 
$$\lambda_{N} = \theta_{G} - \Omega_{N}, \qquad 0 \le \lambda_{N} < 360$$
32. 
$$\lambda_{S} = \lambda_{N} - \Delta \lambda, \qquad 0 \le \lambda_{S} < 360$$
33. 
$$Apogee = (2a_{N} - q_{N} - 1) \cdot K \text{ Where K is km/ e. r. or sm/ e. r. conversion factor Perigee = (q_{N} - 1) \cdot K$$
34. 
$$Course = \tan^{-1} \left[ \frac{\left( \frac{107.08829}{1 + e_{N} \cos v} \right) \sin\beta - \left( \frac{6.3003883 \cos\phi' n (1 - e_{N}^{2} - \frac{1}{\sqrt{\frac{\alpha_{N}(1 - \frac{1}{\sqrt{\frac{\alpha_{N}(1 - e_{N}^{2} - \frac{1}{\sqrt{\frac{\alpha_{N}(1 - e_{N}^{2} - \frac{1}{\sqrt{\frac{\alpha_{N}(1 - \frac$$

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# 7.7 Glossary

Location	Symbol	Meaning
AA		Comment from request card
AAl		Array for storage of comments AA
ANGL		Array for storage of inclination angles (i)
ANGLE	i	Inclination angle
ANOMA	Pa	Anomalistic period in days
APOGEE	-	Apogee distance
AXIS	а	Semi-major axis
AZIB	β	Angle β, see diagram
BB		Comment from request card
BBl		Array for storage of comments BB
С	с	Rate of change of period in days/rev. $^2$
CA		Array for storage of c terms
СС		Comment from request card
CC1		Array for storage of comments CC
CHECK	$\sin \phi'$	Sine of geocentric latitude of subsatellite point
COURSE		Course in radians
D	d	First derivative of c in days/rev. <sup>3</sup>
DA		Array for storage of d terms
DAE		Array for storage of epoch times (t <sub>o</sub> )
DAP		Array for storage of decay times $(t'_0)$
DAYE	to	Time of epoch in days of year
DAY	-	Day of month of report
DAYN		Number of whole days in $t_N$ .
DAYP	t <sub>ó</sub> '	Epoch time of decay equation
DD		Comment from request card
DDl		Array for storage of comments DD
DELLOT	$\Delta \lambda'$	Delta longitude of node excluding precession of node for this revolution
DELONG	$\Delta \lambda$	Difference in longitude from node to subsatellite point
DELTAT	$\Delta t$	Time from ascending node of updated revolution to report time $(t_{REPORT} - t_N)$
DELTTN	$\Delta t_N$	Time elapsed from epoch to last ascending node prior to report time $(t_N - t_0)$

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DEPOCH		Time of epoch in Smithsonian days
DTIM		Array for storage of report times in Smithsonian days
DTIME		Report time in Smithsonian days
E	е	Eccentricity
El		Array for storage of eccentricity values (e)
EE		Comment from request card
EE1		Array for storage of comments EE
EN	$\mathbf{E}_{\mathbf{N}}$	Eccentric anomaly at the node
EPOK		Array for storage of epoch revolution numbers $(N_0)$
EPOKR	No	Epoch revolution number
ERAD	-	$\frac{1}{1+e}$
ES	E(t)	Eccentric anomaly at time of report
FEPOCH		Time of epoch in fractional part of day
FF		Comment from request card
FF1		Array for storage of comments FF
FRACT		In fractional part of t <sub>N</sub>
FTIM		Array of fractional days of report times
FTIME		Fractional part of day of report
FORDC	с	Nodal c term from 7 card element set
FORDE		Column 80 of element cards (E in col. 80 indicates 7 card element set)
FORDP	PNo	Nodal period from 7 card element set
GEGL		Latitude of the subsatellite point in degrees
GEOCL	φ <b>΄</b>	Geocentric latitude of subsatellite point
GEOGL	φ	Geodetic latitude of subsatellite point
GG		Comment from request card
GGl		Array for storage of comments GG
н	λ	Longitude of the node
HLS	λ <sub>S</sub>	Longitude of the subsatellite point
HT	н	Height of satellite above earth

1-7-9

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ID		Element card number
IELNO		Element number
ILN		Array for storage of element numbers
IT		Number of times at which reports are to be issued
KAZ		Day for which report is issued
KIL		Array for storage of output options
KILO		Output option (km or sm) for report
KILOM		Output option $\begin{cases} = 1 \text{ output in statute miles} \\ = 0 \text{ output in kilometers} \end{cases}$
MAGIN		Input tape number
MAGOUT		Output tape number
MESS		"Out of Order" message for flexowriter
МО		Month number of report
Ν		Number of satellites requested for report
Nl		Array of comment indicators = 0 no comments = 1 print comments for this satellite
NAT		Satellite number
NATNO		Array for storage of satellite numbers
NEOGL		Latitude of subsatellite point in degrees
NLS		Longitude of subsatellite point in degrees
NN		Number of satellites having comments on request card
NO		Request card Request card switch = 0 no comments on card = 1 store comments from card = 2 end of request deck
NREV	N	Revolution number at time of report
NRSC		Number of element sets which are out of order
NYEAR		Year of report
NZIB		Course in degrees
PEDOT		Array for storage of $\omega$ values
PERDOT	ம்	First derivative of argument of perigee
PERI		Array for storage of $\omega$ values
PERIG	ω	Argument of perigee
PERGEE	•	Perigee distance

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Nodal period in minutes PERIOD Array for storage of nodal periods (PNo) PNOD  $P_{N_o}$ Nodal period in days/rev. PNODL PRDDO Array for storage of  $1/2 \ddot{\omega}$ PRDDOT 1/2 ພິ One half 2nd derivative of argument of perigee Q Perigee distance in earth radii q QA Array for storage of q values Array for storage of 1/2  $\dot{q}$  values ODDO  $1/2 \ddot{q}$ Second derivative of argument of perigee ODDOT QDO Array for storage of q values q First derivative of argument of perigee QDOT R  $\Delta N$ Revolutions since epoch RADDO Array for storage of  $1/2 \Omega$  values  $1/2 \Omega$ One half 2nd derivative of right ascension of RADDOT the ascending node RADOT Array for storage of  $\hat{\Omega}$ Ω RASDOT First derivative of right ascension of the ascending node 57.2957195 (deg/rad) RAX RIGHT Right ascension of ascending node in degrees RITAS Array for storage of  $\Omega$  values RITASC Ω Right ascension of the ascending node RJ Distance from center of earth to satellite r Ro RO Radius of earth at subsatellite point SNAME1 Satellite name SNAME2 SNAMEA Arrays for storage of satellite names SNAME B  $M_N$ SOLMN Mean anomaly at the node SOLMS M(t) Mean anomaly at time of report SUBLN Longitude of the node in degrees λ<sub>N</sub> TIM Time of report in hours and minutes TIME Array of times at which reports are to be issued TME Time of report in hours and minutes TIZ Year of report ΤN Time of nodal crossing for revolution tN number of report

#### 1-7-11

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U	μ	Argument of latitude of satellite at time of report
VEK	<b>v</b>	True anomaly at time of report
YCONS		Array for storage of $\theta_0$ values
YCONST	θο	Sidereal time at Greenwich at beginning of year of report
YEAE		Array for storage of epoch years
YEARE		Year of epoch
YEARP		Year of epoch of decay equation
YEAP		Array for storage of epoch years of decay
YEAR		Year of report
YR		Last digit of year of report
ZATE ZETE		Satellite name

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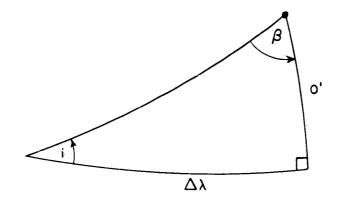
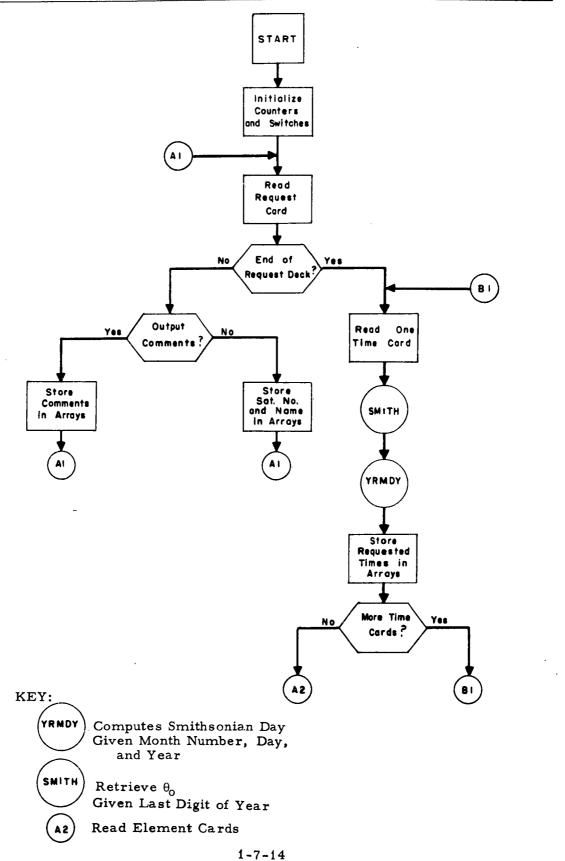


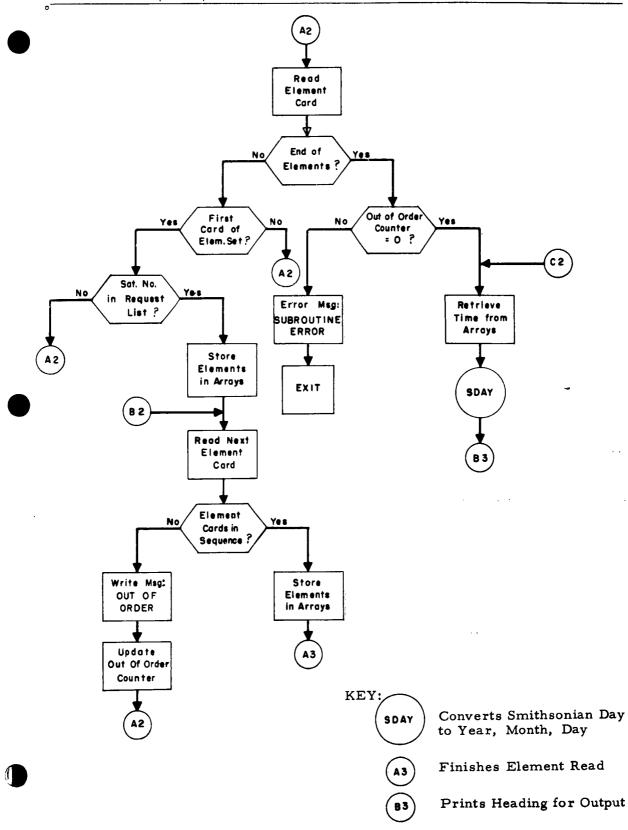
Illustration of Angle  $\beta$ 

Fig. 7.1



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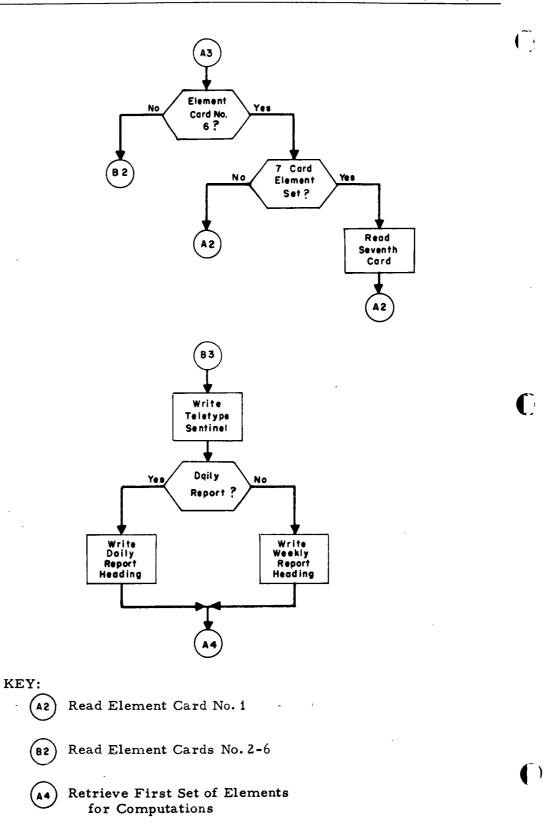


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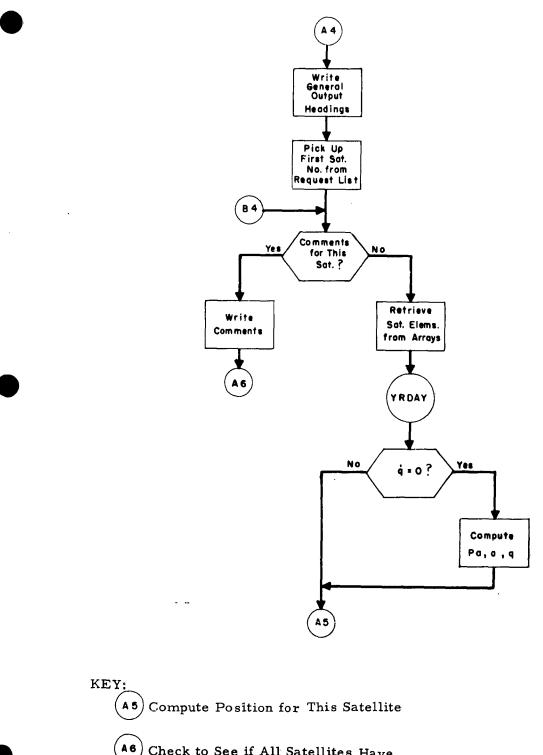


# 1-7-16

PREPINT

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6) Check to See if All Satellites Have Been Processed

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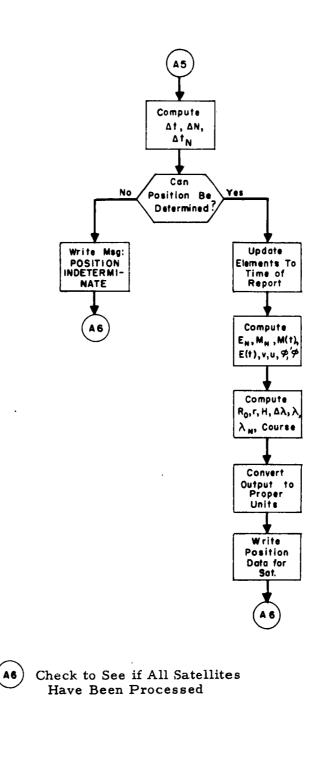
KEY:

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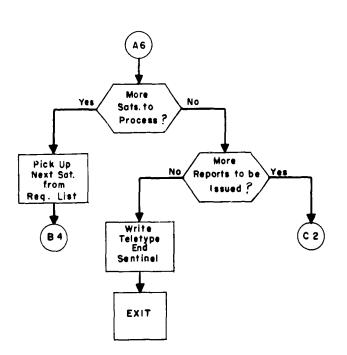
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B4)

Start Processing Next Satellite

c2) Retrieve Time for Next Report

#### 8.1 MAKETAPE, Make Input Tape for TELTYP

8.2 Function

The program MAKETAPE produces a magnetic output tape incorporating the message sentinels required by the TELTYP program. Only one message is produced by the program. An optional control feature will produce the message broken into 90-line segments.

# 8.3 Input

Input data originates from the Schedule Tape. This input is moved by the system from the Schedule Tape to the System Data Tape (logical 0).

Two control cards are used in addition to the data cards containing the message to be converted. The following should be in cols. 17-24:

1. TELEFORM

2. FINDATA

### 8.3.1 Description of Control Cards

#### 1. TELEFORM

This card, if present, will precede the data deck. When encountered, it will signal the program to break the message into 90-line segments.

2. FINDATA

This must be the last card of the data deck. The program exits to the executive program when this card is intercepted.

#### 8.4 Output

The output tape (logical 11) created by this program is in a form acceptable for subsequent conversion by the TELTYP program. A hard copy of the message with its sentinels may be obtained by printing through the UBC using data select one. For a discussion of the sentinels used, the writeup of the TELTYP program should be consulted.

## 8.5 Processing

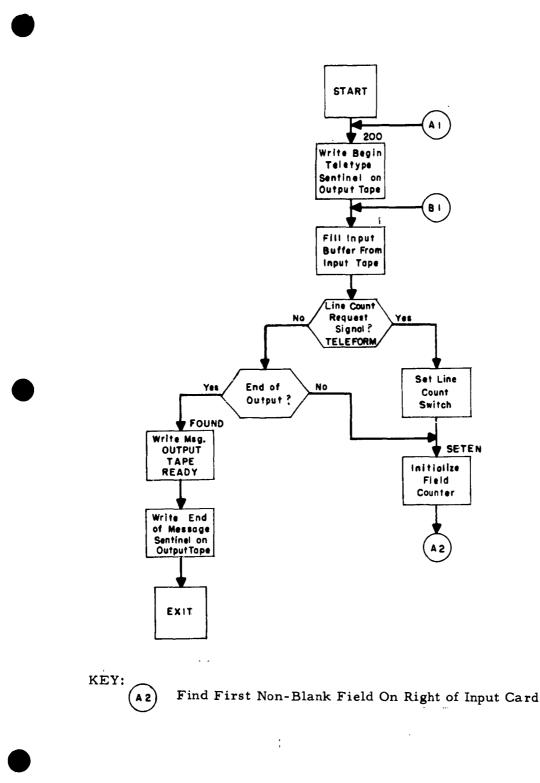
MAKETAPE reads data into core from logical tape zero. Each card image is scanned for the control information described under the input section. If TELEFORM is found, an internal switch is set to cause line counting. If the program is in the line counting mode, each group of 90 lines will be preceded by the begin sentinel and followed by the end sentinel required by the TELTYP program.

All data read into core will be written onto the output tape, with the exception of right-adjusted blank fields. The program will exit to the executive routine when the FINDATA control card is intercepted.

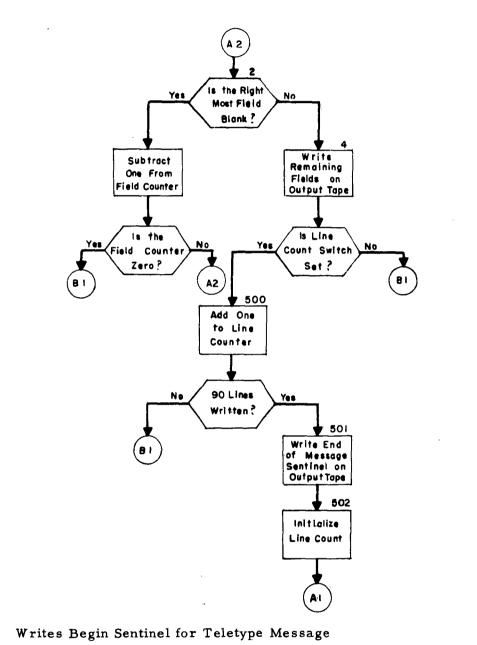
It is generally expected that the TELTYP program will be used immediately following this program or, at least, before logical tape eleven is wrapped up by the system.

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Reads Next Input Card

### 9.1 XYZLAR, Look Angle Report From x, y, z, Coordinates

## 9.2 Function

The program is used for predicting the position of a satellite in terms of the tracking coordinates of a particular station. The program is especially useful for deep space probes.

The position of the satellite is defined in terms of the right ascension, declination, azimuth, elevation, and slant range at the prediction time, for a particular station. The elevation and illumination angles of the sun are also computed to determine if the satellite is visible.

Input options exist to consider the restrictions imposed by the type of observing equipment used. No data will appear for a requested prediction time unless these limitations have been met.

The ephemeris data, used in the calculations of the look angles, may be read from an ephemeris tape previously written by either the Unified Encke Differential Correction Program (accuracy, approximately ten lunar distances), or by the Interplanetary Program. If the ephemeris tapes are not available, the data may be read in from punched cards. Predictions for more than one station may be based on the same ephemeris data.

### 9.3 Input

An input set consists of a standard station card for the observing station, a request card, and the ephemeris information from an ephemeris tape or punched cards containing the ephemeris data followed by a blank card. Each input set results in a look angle schedule for the station specified. If look angle schedules are required for the same satellite but different observing stations, additional pairs of cards, composed of a station card and a request card, may be added to the input deck. If new ephemeris data is to be entered from punched cards, the station card of the next input set must contain a negative station number. Any number of input sets may be entered. The last set must be followed by a blank card which is in addition to the blank card following the ephemeris data cards.

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The format of the request card follows: columns 1 to 4 contain the year of the desired predictions; columns 5 to 7 contain the day of year; columns 8 and 9 contain the hour; columns 10 and 11 contain the minutes; columns 12 to 17 contain the base, or starting time, for the desired predictions.

The time increment used is determined by the ephemeris data. Columns 25 to 34 contain the maximum time increment or the time range for predictions. This time is expressed in minutes from the base time of the request. Any ephemeris data beyond this time range will be ignored. If columns 25 to 34 are blank, the program will calculate predictions for all ephemeris data supplied. Program option switches are specified in columns 44 to 47. A 1 punch in the respective column will set an internal switch. A 1 punch in column 44 indicates that the ephemeris data will come from the ephemeris tape for this set of input. If column 44 is blank the program expects ephemeris data from ephemeris cards. A 1 punch in column 45 indicates that only visible passes are desired. If a 1 is punched in column 46, negative elevations will be acceptable. A 1 is punched in column 47 if punched ephemeris cards are desired as output.

### 9.3.1 Ephemeris Cards

The ephemeris cards contain the time increment from the base time specified on the request card and the inertial geocentric coordinates of the satellite. All data is in floating point. Columns 1 to 14 contain the time increment, columns 15 to 28 contain the x-coordinate, columns 29 to 42 contain the y-coordinate, and columns 43 to 56 contain the z-coordinate in earth radii. Ephemeris cards are ordered by increasing time increment. A satellite identification card with the alphanumeric satellite name punched in columns 1 to 16 must precede the ephemeris cards.

### 9.3.2 Ephemeris Tape

The ephemeris tape can be obtained from two sources: 1) the Unified Encke Differential Correction Program and; 2) the Interplanetary Program. Both programs were written by Aeronutronic a division of the Ford Motor Company. In the former case, logical tape 10 is the desired tape and in the latter, the required tape is logical 9. The format for both tapes follows:

- Block 1. Contains the alphanumeric satellite name in the first two words.
- Block 2 thru N. Contains a time increment from the base time and x, y, z, x, y, z in the inertial geocentric coordinate system at that particular time. Each value is contained in one full computer word and is in the floating point format. Eighteen groups of these seven values plus two zero words make up each one hundred and twenty-eight word block. The tape is considered to be terminated when the first word of a seven word set is filled with Z's.

### 9.4 Output

If a 1 is punched in column 47 of the XYZLAR request card, the ephemeris data will be punched on cards in the same format as the input ephemeris data when the binary ephemeris tape is used for input. The satellite name appears on the first card.

Printed output contains all TELTYP control functions, therefore, transmission of the output is possible. The first line is the alphanumeric satellite name. The second line is the comment LOOK ANGLES FOR followed by the alphanumeric station name. The third and fourth lines are heading lines describing the output data for each prediction time. Each data line contains the day of year, hour, minute, and fraction of minute of the search point, the predicted right ascension, declination, azimuth, and elevation in degrees and slant range in kilometers. These quantities define the position of the satellite. The elevation and illumination angles of the sun complete the data line and determine if the satellite will be visible at the time of the prediction.

### 9.5 Processing

XYZLAR predicts the right ascension, declination, azimuth, and elevation angles for a given satellite as well as the elevation and illumination angles of the sun at the requested prediction time. Restrictions on the observing capability of the station, specified on the request card, are considered. Any restrictions that are not met will cause computations to cease for that particular prediction time,

and no printed output will appear. The program will begin processing the next ephemeris position. In addition, when and if the maximum time increment is exceeded, the program will ignore the remaining ephemeris data of that input set. The next request is then processed.

If an input set utilizes ephemeris cards, all of the ephemeris data is written in binary on logical tape seven before processing begins. However, an input set may originally include an ephemeris tape mounted on logical seven and in that case, processing will begin immediately.

The program adds the specified base time from the Request card, and the time increment from the ephemeris data to obtain the time of the search point. The sidereal time at Greenwich at prediction time and the coordinates of the station in a fixed system are found in order to compute the slant range. The right ascension and declination of the satellite, the sidereal time at the station, the hour angle, zenith distance, and elevation angle are computed for the prediction time.

If negative elevation angles are acceptable, the program continues on to compute the azimuth. If, however, negative elevation angles are not acceptable, a test is made and if the angle is found to be negative the program returns to process the next group of data.

Visibility at prediction time is determined by computing the elevation and illumination angles of the sun. If the elevation angle is less than -4 degrees, and the illumination angle is greater than -4 degrees, then the satellite is visible. If only visible passes are required, these conditions must be met.

Output quantities at each prediction, or search point, time include the time, right ascension, declination, azimuth, elevation, and slant range of the satellite, and the elevation and illumination angles of the sun.

The program then processes the next group of data. If the station number is positive, the program computes another search ephemeris for the same satellite in the manner stated above, but for this new station. If the station number is found to be negative new ephemeris information is assumed. If the station card is blank, control is returned to the executive program.

Punched card output, if specified, is under control of a PATH switch which eliminates multiple punching of the same ephemeris data. This switch also controls the production of an ephemeris tape from a set of ephemeris cards. This switch equals one (1) for the first time through the program for a particular set of ephemeris data, and two (2) for each subsequent pass through the program for this same data. It can be reset to one by reading a negative station number. Thus, ephemeris cards are punched or an ephemeris tape made only on the first pass through the program for that particular set of ephemeris data.

# 9.5.1 Error Message

If the program is unable to read the ephemeris tape, the comment EPHEMERIS TAPE TROUBLE will be written on the output tape. In addition, the comment PROGRAM TERMINATED DUE TO EPHEMERIS TAPE TROUBLE will appear on the Flexo, and control will return to the executive routine.

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9.6	For	mulation	Û
	1.	$\phi' = \tan^{-1} (.99329985 \tan \phi)$	
	2.	$R = [.9966443/(100670015 \cos^2 \phi)^{1/2} + H/6378174] 6378.174$	
	3.	Compute station coordinates in rotating geocentric system.	
		$x_{stat} = R \cos \lambda \cos \phi'$	
		$y_{stat} = R \sin \lambda \cos \phi'$	
		$z_{stat} = R \sin \phi'$	
	4.	Convert starting or base time to days and fractional days.	
		$t_a = days + hours/24 + minutes/1440 + secs/86400$	
	5.	Find time of computed point.	
		$t_c = t_a + t_i$	
		If $t_i > t_{max}$ return to process next input set.	
	6.	$\mathbf{r} = \mathbf{x}^2 + \mathbf{y}^2 + \mathbf{z}^2$	
	7.	$\theta_{\rm G} = \theta_{\rm o} + .98564735 t_{\rm d} + 360.985647 t_{\rm f}$ $0 \le \theta_{\rm G} < 360^{\rm o}$	
·····	8.	Compute station coordinates in inertial geocentric system	
		$\overline{\mathbf{x}} = \mathbf{x}_{stat} \cos \theta_{G} - \mathbf{y}_{stat} \sin \theta_{G}$	
		$\overline{y} = x_{stat} \sin \theta_G + y_{stat} \cos \theta_G$	
		$\overline{z} = z_{stat}$	
	9.	$\rho = \left[ \left( \mathbf{x} - \overline{\mathbf{x}} \right)^2 + \left( \mathbf{y} - \overline{\mathbf{y}} \right)^2 + \left( \mathbf{z} - \overline{\mathbf{z}} \right)^2 \right]^{1/2}$	
		where x, y, and z are geocentric coordinates from the ephemeric data.	S
	10.	$\cos a = (x - \overline{x}) / [(x - \overline{x})^2 + (y - \overline{y})^2]^{1/2}$	
	•	$a = \tan^{-1} \left[ \left( y - \overline{y} \right) / \left( x - \overline{x} \right) \right]$	
		If cos a negative, $a = a + \pi$ . Otherwise, $0 \le a \le 360^\circ$	
	11.	$\sin \delta = (z - \overline{z})/\rho$	
		$\delta = \sin^{-1} (\sin \delta)$	

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12.	$\theta_{ST} = \theta_{G} + \lambda$
13.	$HA_s = \theta_{ST} - a  0 \le HA_s < 360^{\circ}$
14.	$\beta = \cos^{-1} (\sin \phi \sin \delta + \cos \phi \cos \delta \cos HA_s)$
15.	$h = \pi/2 - \theta$ If negative elevations acceptable, go to 16. If negative elevations not acceptable and $h < 0$ : Go to 29.
16.	sin Az = $(\cos \delta \sin HA_s/\sin \beta)$ cos Az = $(-\sin \delta \cos \phi + \cos \delta \sin \phi \cos HA_s)/\sin \beta$ Az = $\tan^{-1} (\sin Az/\cos Az)$ If cos Az < 0, Az = Az + $\pi$ . $(0 \le Az < 360^{\circ})$
17.	Begin visibility computations. $I_{\odot}(t) = L_{\odot} + .98564735 t_{c} + 1.91665 sin (.98564735 t_{c} - c_{14})$
18.	$a_{\Theta} = \ell_{\Theta}(t) - 2.46682 \sin(2\ell_{\Theta}(t))$
19.	$\delta_{\varphi} = \tan^{-1} (.4336608 \sin a_{0})$
20.	$\theta(t) = \theta_G + \lambda$
21.	$h_{\Theta} = \sin^{-1} \left[ \sin \phi \sin \delta_{\Theta} + \cos \phi \cos \delta_{\Theta} \cos (\theta(t) - a_{\Theta}) \right]$
22.	$a_{gc} = tan^{-1} (y/x)$
23.	$SD_{gc} = sin^{-1} (z/r)$
	$\xi = \cos^{-1} (-\cos (SD_{gc}) - \cos \delta_{\Theta} \cos (a_{\Theta} - a_{gc}) - \sin (SD_{gc}) \cdot \sin \delta_{\Theta})$
24.	$\eta = \sin^{-1}(1/r)$ If only visible passes desired, go to 25. Otherwise, go to 27.
25.	If $h_{o} > -4$ degrees, pass not visible: Go to 29.
26.	I = $\xi - \eta$ If I < - 4 degrees, pass not visible: Go to 29. Otherwise Go to 28.
27.	Ι = ξ - η

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- 28. Output of time, a,  $\delta$ , Az, h,  $\rho$ , h<sub> $\varphi$ </sub>, I.
- 29. Pick up next group of ephemeris data, Go to 5.

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9.7 Glossary

Location	Symbol	Meaning
AA	×stat	x coordinate of station in rotating geocentric system
BB	y <sub>stat</sub>	y coordinate of station in rotating geocentric system
C14	°14	Difference between longitude of sun and argument of perigee of sun
C2PI	2π	Constant 6.28318530
CAPR	R	Radius of earth at station
CAZ	cos Az	Cosine of azimuth angle of satellite
СС	<sup>z</sup> stat	z coordinate of station in rotating geocentric system
CRA	Cos a	Cosine of right ascension of predicted sighting
CRAGC	cosagc	Cosine of a <sub>gc</sub>
CXI	cos ξ	Cosine of §
CZD	сов β	Cosine of zenith angle
ELEVA	h	Elevation angle of sat. (predicted)
ETA	η	Intermediate calculation for illumination
FMIN		Fraction of minute of computed point
HITE	Н ~	Height of station
HSUNJ	h.	Elevation of sun
IDAY		Day of month of computed point
IDCI		Symbolic name for input tape
IDCO		Symbolic name for output tape
IHM		Hour, min, fractional minute of computed point
IPUNCHC		Symbolic tape name for intermediate punch card tape
ISSW4		l, Ephemeris tape 0, Ephemeris cards
ISSW5		1, Only visual passes desired 0, All <b>pass</b> es desired
ISSW6		l, Only positive elevations desired 0, All elevations desired

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ISSW7	·	<ol> <li>punch cards from ephemeris tape</li> <li>don't punch cards</li> </ol>
ISW		l, for first pass through program for a particular set of ephemeris data
		2, for each subsequent pass through program for same ephemeris data
NSTA		Station number
ox	l <sub>o</sub> (t)	Longitude of the sun at prediction time
PHI	φ	Geodetic latitude of station
PHIP	<b>ب</b>	Geocentric latitude of station
PI	π	Constant: 3.14159265
RAD		Constant: degrees/radian
RAV	r	Distance of satellite from center of earth
S	HAs	Hour angle of satellite
SAZ	sin Az	Sine of azimuth angle of satellite
SD		Day of year of base time from Request card
SDEC	sin δ	Sine of declination .
SDGC	$SD_{gc}$	Intermediate calculation
SH	_	Hour of base time from Request card
SHSUNJ	$\sin h_{\Theta}$	Sine of elevation of sun
SLRANG	ρ	Slant range
SM		Minutes of base time from Request card
SS		Seconds and fraction of seconds from Request card
STIME	ta	Base time of predictions
STNM1, STNM STNM3	M2	Alphanumeric station name
ТА	ti	Time increment from ephemeris data
	t <sub>c</sub>	$t_i + t_a$ Time of computed point
TD	td	Prediction time in days
TDECS	tan $\delta_{\boldsymbol{\varphi}}$	Tangent of declination of sun
TF	t <sub>f</sub>	Prediction time in fractional days
тн	t <sub>h</sub>	Hour of prediction time
THETG	θG	Sidereal time at Greenwich at prediction time in degrees
THETST	θst	Sidereal time at the station
THTGT	θ(t)	Sidereal time at station

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TMAX	t <sub>max</sub>	Maximum time increment from base time
TPHIP	tan $\phi'$	Tangent of geocentric latitude of station
TRA	tan a	Tangent of right ascension of predicted sighting
x	x	Geocentric x coordinate of satellite in earth radii from ephemeris data
XAZ	Az	Azimuth angle of satellite
XBAR	x	x coordinate of station in inertial geocentric system
<b>X</b> DECS	δ	Declination of sun
XI	£	Intermediate quantity
XIA	I	Illumination angle (sun).
XL	λ	Longitude of station
XLSUN	L	Celestial longitude of sun at start of year
XRASUN	a o	Right ascension of sun
XTHETG	θ <sub>G</sub>	Sidereal time at Greenwich at prediction time, in radians
XZD	β	Zenith angle
Y	У	Inertial geocentric y coordinate of satellite in earth radii, from ephemeris data
YR		Year of base time of predictions
YBAR	У	y coordinate of station in inertial geocentric system
Z	2	Inertial geocentric z coordinate of satellite in earth radii, from ephemeris data
ZBAR	Z	z coordinate of station in inertial geocentric system
ZDEC	δ	Predicted declination of satellite
ZRA	a	Right ascension of predicted sighting
ZRAGC	agc	Right ascension of satellite

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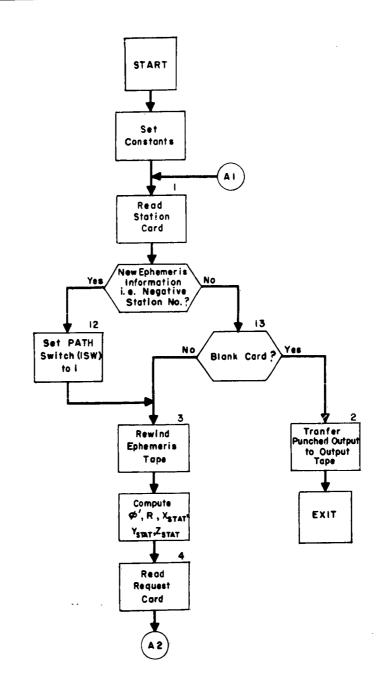
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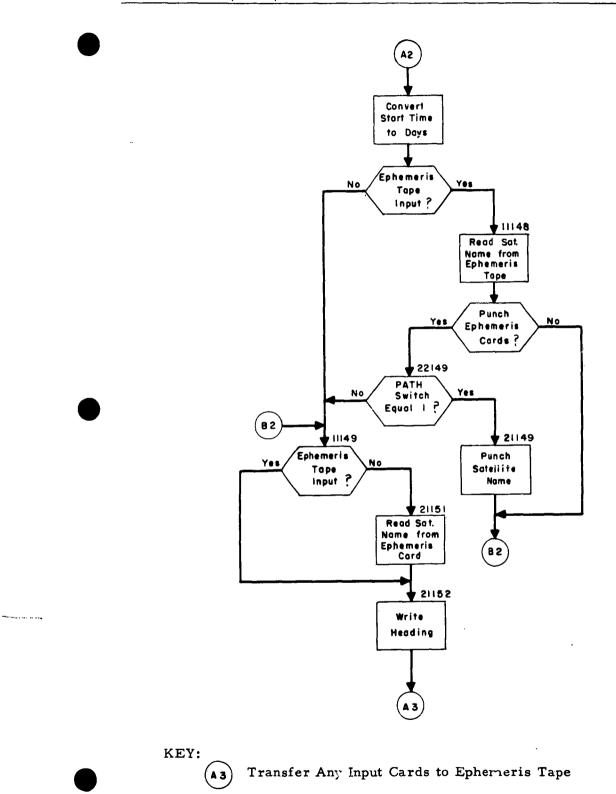




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Obtain Ephemeris Data

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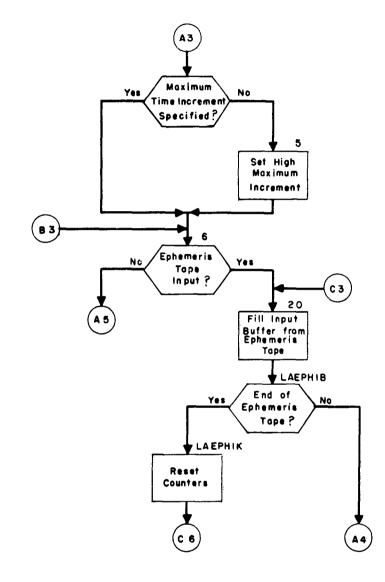


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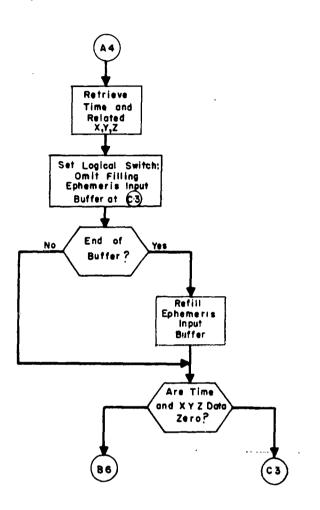


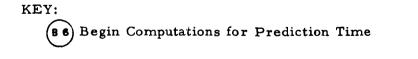
KEY: A 5 Process Ephemeris Cards A 4 Retrieve Time and Related x, y, z (C6) Obtain Next Input Set

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XYZLAR

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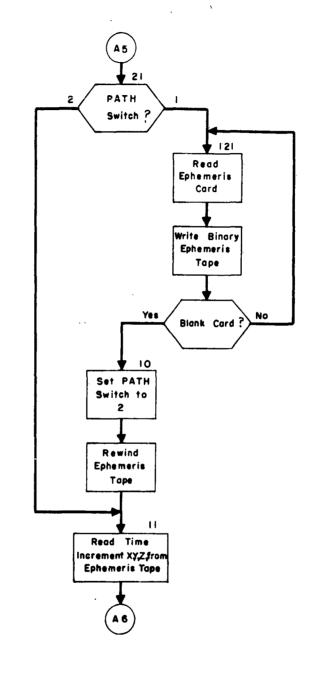




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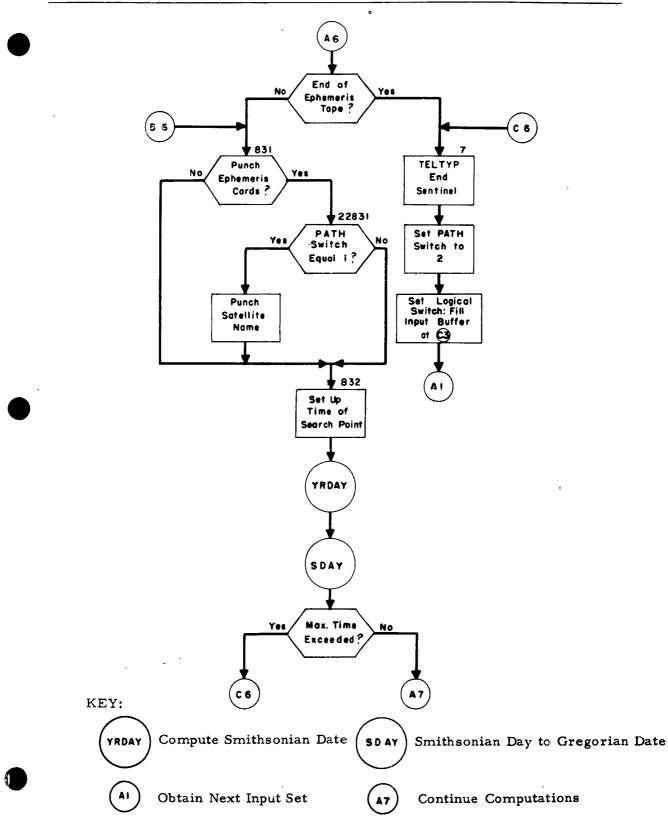
-

) Begin Computations for Requested Prediction Time



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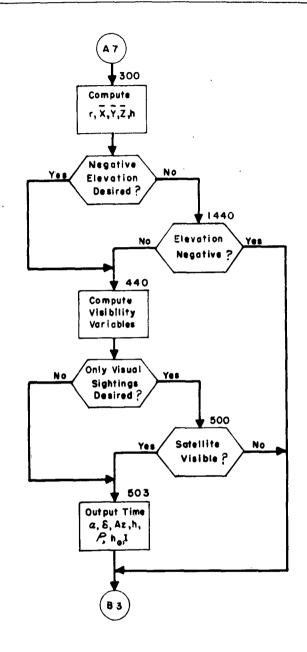
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#### 10.1 POSE, Point Search Ephemeris

# 10.2 Function

When a radar picks up an unknown object, one may have for some reason, the suspicion that it could have been an unidentified satellite. If this is the case, one likes to search for it. All orbital elements, however, are completely unknown. Disregarding the recession of the line of the nodes and the precession of the line of apsides, which is permissible over one or two revolutions, once can expect that the object (if it should be a satellite) will come close to the same point in inertial space again, one revolution later.

The program computes a search ephemeris for a given station for this point in space.

# 10.3 Input

The input consists of: (1) a request card, (2) a standard observation card, (3) one or more standard station cards and (4) an option card.

The information contained on the request card is as follows:

Cols.	1-5	T-start in minutes (DDDD.)		
		the smallest orbital period expected		
		(usually 80 mins.)		
Cols.	6-8	$\Delta t$ in minutes (DD.) - the time increment added		
		to the orbital period (usually 5 mins.)		

- Cols. 9-13 T-stop in minutes (DDDD.) the largest orbital period expected.
- Col. 14 Iout the output option desired
  - 1 = long form
  - 0 = short form

(explained in output section)

A decimal point should be punched within fields one, two and three.

#### 1-10-1

The option card is used to initiate the read-in of a new request deck or to terminate the program. If a -1 is punched in cols. 1-4, the program will read in a request card, observation card, station card(s) and an option card. If a zero or blank field is read from the option card the program will be terminated.

# 10.4 Output

The output consists of a look-angle schedule for each station requested. The long form of output includes the following data:

- a) Station information: latitude, longitude and height
- b) Observed time in days, hours, minutes and seconds (GMT)
- c) Observed elevation, azimuth and slant range
- d) Value of  $\theta_G$  sidereal time at Greenwich
- e) Value of  $\theta_{s}$  sidereal time at the station
- f) Distance of station from center of the earth (CAPR)
- g) Geocentric station latitude
- h) Station 's Cartesian coordinates where x and y are in the equatorial plane and x points to the Greenwich meridian.
- i) Right ascension and declination of station
- j) Cartesian coordinates of station 's position where x and y are in the equatorial plane and x points to the vernal equinox.
- k) Cartesian coordinates of the point or satellite where x and y are in the equatorial plane and x points to the vernal equinox.

In addition to the above the short and long forms of output list the look-angle information:

- a) Time of crossing day, hour and minutes of Zebra time.
- b) Elevation and azimuth angles
- c) Slant range in kilometers

The sentinels required by the TELTYP program will be supplied to enable transmission of the short form, or look-angle section, of the output.

#### 1-10-2

## 10.5 Processing

The first request card, observation card and one station card are read at the start of the program. The x, y, and z coordinates of the station and the object are then calculated. These values are written on the output tape if a long form of output is requested.

The assumption made is that the object will appear in the same place one revolution later. Therefore, the look-angle, or search ephemeris quantities are calculated for the time t, which is equal to the observation time plus the smallest orbital period expected. If the point is visible at the station the quantities are written out on tape.

The orbital time is then updated by the time increment and compared to the largest orbital time expected. If this period is within the limits, the search ephemeris is computed for this new time. The process continues until the time limits have been exceeded.

After the search ephemeris has been computed for the largest orbital period expected the next station card is read into core. If the station number is greater than zero, the look-angles for this same object are computed for this new station. A negative station number indicates that a new request deck is to be read into core and the processing initiated for another point in space. A blank or zero <sup>(2)</sup> station number terminates all processing and the program returns to the executive program.

# 10.5.1 Error Messages

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1. THE SLANT RANGE IS MISSING.

The program exits to the executive routine.

SUBROUTINE ERROR EXIT FROM OCTAL \_\_\_\_\_. Subroutine or irrecoverable input-output error. Exits to executive program if the GO option is taken. The message is retyped if the STOP option is specified. A dump should be taken if possible.

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10.6	Formulation	
	a) Initialization	
1.	$\theta_{\rm G} = \theta_{\rm o} + .98564724 \cdot t_{\rm D} + 360.98565^{\rm o} \cdot t_{\rm FR}$	
2.	$\theta_s = \theta_G + \lambda_E$	
3.	$t_{g} \phi' = .99329985 \cdot t_{g} \phi$	
4.	$\xi = \mathbf{R} \cdot \cos \boldsymbol{\ell} \cdot \cos \boldsymbol{\phi}'$	
5.	$\eta = \mathbf{R} \cdot \sin \ell \cdot \cos \phi'$	
6.	$\zeta = \mathbf{R} \cdot \sin \phi'$	
7.	$\beta = 90 - h_{o}$	
8.sir	$h \delta = \cos \beta \cdot \sin \phi + \sin \beta \cdot \cos \phi \cdot \cos Az$	ſ
9. sir	$n S = -\sin \beta \cdot \sin Az/\cos \delta$	
10.	$a = \theta_s - S$	
11.	$\Delta \mathbf{x} = \boldsymbol{\rho} \cdot \cos \alpha \cdot \cos \delta$	
12.	$\Delta y = \rho \cdot \sin \alpha \cdot \cos \delta$	
13.	$\Delta z = \rho \cdot \sin \delta$	
14.	$X = \zeta \cdot \cos \theta_G - \eta \cdot \sin \theta_G$	
15.	$Y = \zeta \cdot \sin \theta_G + \eta \cos \theta_G$	
16.	Ζ = ξ	
17.	$\mathbf{x} = \mathbf{X} + \Delta \mathbf{x}$	
18.	$\mathbf{y} = \mathbf{Y} + \Delta \mathbf{y}$	
19.	$z = Z + \angle z$	$\cap$

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	b) Computation of look-angle
20.	$t = t_{obs} + t_{start}$
21.	$\theta_{G} = \theta_{o} + C_{3} \cdot t_{d} + C_{2} \cdot t_{fr}$
22.	$\theta_{s} = \theta_{G} + \lambda_{E}$
23.	$X = \zeta \cdot \cos \theta_G - \eta \sin \theta_G$
24.	$Y = \zeta \cdot \sin \theta_G + \eta \cos \theta_G$ (Note that Z is a constant)
25.	$\Delta \mathbf{x} = \mathbf{x} - \mathbf{X}$
26.	$\triangle y = y - Y$ (Note that $\triangle z$ is a constant)
27.	$\rho = \sqrt{\Delta x^2 + \Delta y^2 + \Delta z^2}$
28.	$\sin \delta = \Delta z / \rho$
29.	$\cos \delta = \sqrt{1 - \sin^2 \delta}$
30.	$\cos \alpha = \Delta x / \rho \cos \delta$
31.	$S = \theta_s - \alpha, \qquad 0 \leq S \leq 360$
32.	$\cos\beta = \sin\phi \cdot \sin\delta + \cos\phi \cdot \cos\delta \cdot \cos S$
33.	$\sin Az = (-\cos \delta \cdot \sin S) / \sin \beta$
34.	$\cos Az = (\sin \delta \cdot \cos \phi - \cos \delta \cdot \sin \phi \cos S)/\sin \beta$
35.	$h = 90^\circ - \beta$
36.	$\mathbf{t} = \mathbf{t} + \Delta \mathbf{t}$
37.	If $t \leq t_{stop}$ : Go to 21



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10.7	Glossary		<b>O</b>
	Location	Symbol	Meaning
	AAI		First digit (may be overpunched) of satellite no. from observation card
	AZ	Az	Azimuth angle
	AZD		Azimuth of computed point in degrees
	AZ1		Azimuth of original observation in degrees
	Ci	c <sub>i</sub>	.99329985 conversion coefficient used to compute φ
	C2	°2	360.98565 rotation rate of the earth in degrees/solar day
	C3	c3	.98564724 rotation rate of the earth in (deg/solar day) -360°
	C4	C <sub>4</sub>	1440 minutes/day
	CAPR	R	Radius of the earth at the station
	CAPX	x	x coordinate of station x, y in equatorial plane;
	CAPY	Y	y coordinate of stationy points toward the
	CAPZ	Z	z coordinate of station vernal equinox
	COSDEC	cos δ	Cosine of declination
	COSLAT	cos φ	Cosine of station latitude
	COSZD	cos β	Cosine of zenith angle
	C2PI	2 π	6.2831853072
	DA		Day of observation
	DECL	δ	Declination
	DELT	$\Delta t$	Time increment
	DELX	$\Delta \mathbf{x}$	x coordinate - increment from station to satellite
	DELY	∠ y	y coordinate - increment from station to satellite
	DELZ	riangle z	z coordinate - increment from station to satellite
	EARTHR		6378.174 km/e.r.
	ELEV	ho	Elevation angle of observation
	ELEVD	h	Elevation angle of computed point
	ETA	η	y coordinate of station (x, y in equatorial plane; x points to Greenwich meridian)

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IOUR		Hour at which point is computed
IRANGE		Slant range for output point
IYR		Last digit of year of observation
IDAY		Day at which point is computed
KSTA		Station number
МО		Month of observation
PI	π	3.1415926536
RA	a	Right ascension of satellite
RANGE	ρ	Slant range of computed point
RHO		degrees/radian = 57.29578
S	S	Hour angle
SINAZ	Sin Az	Sine of azimuth angle
SINDEC	Sin δ	Sine of declination
SINLAT	Sin φ	Sine of station latitude
SINZD	sinβ	Sine of zenith angle
SNI, SN2		
SN3, SN4		Station name
STALAT	¢,	Station latitude
STALATC	¢	Geocentric latitude of station
STALON		Station longitude in degrees
STH		Station height
STLORA	λ <sub>E</sub>	Station longitude in radians
Т	t	Time at which ephemeris is computed
TD	t <sub>D</sub>	Day of year of observation
T DAYS	td	Integer part of t
T FR	t <sub>fr</sub>	Fractional part of t
TH	_	Hour of observation
THETAO	θ <sub>o</sub>	Greenwich Sidereal time at the beginning of year
THETAS	θs	Siderial time at station
THETAG	θG	Greenwich sidereal time
THE TAG1		Siderial time at Greenwich in degrees
ТМ		Minute of observation
TMIN		Minute at which point is computed
TOBFR	<sup>t</sup> fr	Time of observation in fractional part of day

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TOBS	tobs	Day of year of observation
TS	0.00	Seconds of observation
TSTART	t <sub>start</sub>	Minimum period
TSTOP	t stop	Maximum period
x	x	$\mathbf{x}$ coordinate of satellite
XAT		Last 2 digits of satellite number from observation card
XI	Ę	x coordinate of station (x, y in equatorial plane; x points to Greenwich meridian)
Y	У	y coordinate of satellite
YEAR		Year of observation
YR		Last digit of year of observation

z coordinate of satellite

 $\beta$  Zenith angle

 $\mathbf{z}$ 

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z coordinate of station(x, y in equatorial plan x points to Greenwich meridian)

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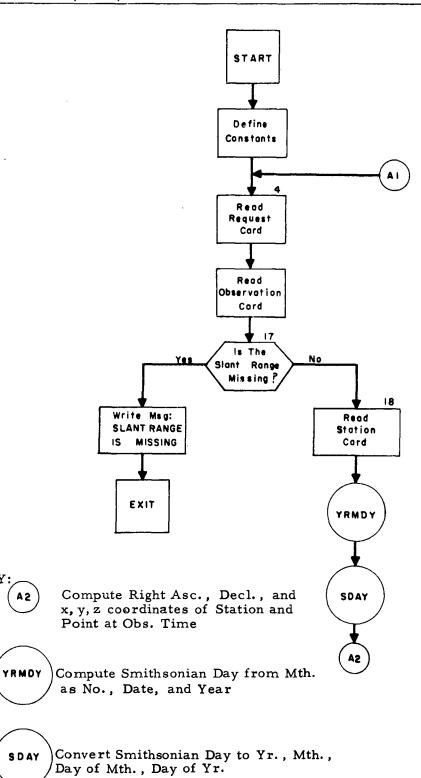
1.00

ZETA

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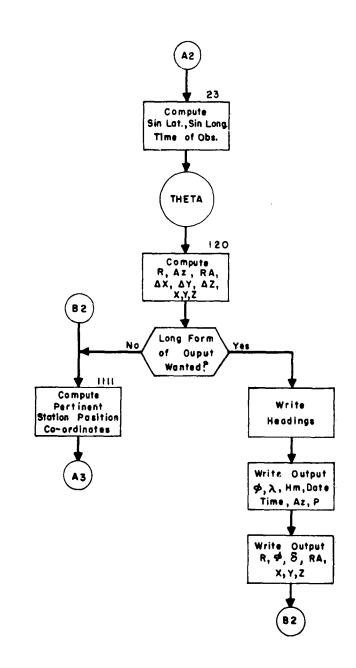
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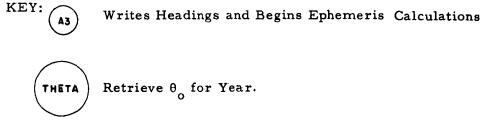
KEY:



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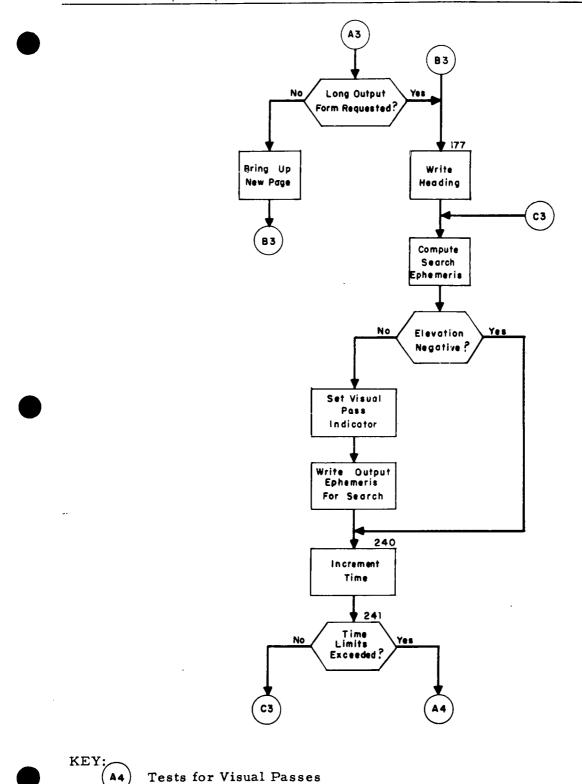


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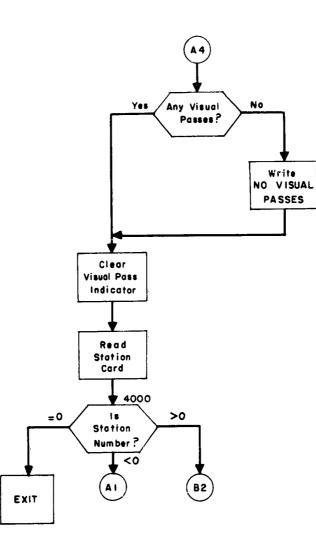
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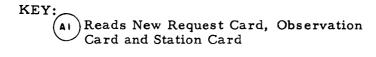


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(B2) Calculates Search Ephemeris for Same Satellite but New Station

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#### 11.1 ORPS, Orbital Plane Search

### 11.2 Function

ORPS is used for predicting search times for observing satellites at a given observing station. Limitations of the observing station, imposed by the type of equipment used, are considered. For example, the results must meet the criteria for maximum slant range, minimum elevation angle, and visibility.

The program computes the XYZ search point coordinates of a satellite from which its position may be defined in terms of right ascension, declination, azimuth, elevation, and slant range. The elevation and illumination angles of the sun are also computed at the search point or prediction time.

#### 11.3 Input

A set of input consists of a standard six or seven card element set, a standard station card for the observing station, and a request card. The set may be repeated as often as desired, but the last set must be followed by a blank card to terminate the program.

The format of the request card is as follows:

 columns 1-10 contain the requested start time, in day of year, for predictional data for a given observing station and satellite; 2) columns 11-20 contain the time increment in minutes between search points; 3) columns 21-30 contain the stop time in day of year; 4) the search azimuth angle in degrees of the station is contained in columns 31-40; 5) columns 41-47 contain the maximum slant range in kilometers acceptable to the observing station; 6) columns 48-57 contain the minimum elevation angle in degrees acceptable to the observing station;
 7) if only visible passes are desired, a one (1) is punched in column 58.

Quantities specified in the above paragraph depend upon the limitations of the equipment type used by the observing station.

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### 11.4 Output

ORPS may be used with the TELTYP program if teletype transmission of the output is desired. Therefore, all TELTYP control functions are included with the output.

The first line is the heading "ORBITAL PLANE SEARCH AND XYZ LOOK ANGLE PROGRAM". The next line is the satellite and element numbers of the satellite for which the predictions were requested. The next line is the comment "ORBITAL ELEMENTS". This indicates that the nodal elements used in the computations appear in the next three rows, five columns to a row. The elements that are printed out by row are the day of year of epoch, fractional day of epoch, nodal period at epoch, rate of change of period in days/ rev<sup>2</sup> at epoch, rate of change of c in days/rev<sup>3</sup>, first and second derivatives of the nodal period, the right ascension of the ascending node at epoch, first and half second derivatives of the right ascension, the argument of perigee at epoch, first and half second derivatives of the argument of perigee, eccentricity, and the inclination.

The next three lines consist of the parameters specified on the parameter input card. The start time in days of year, the time increment in minutes, and the stop time in days of year are printed. The search azimuth of the station in degrees and the criteria for maximum slant range in kilometers and minimum elevation angle in degrees are printed.

The year constants are printed including the sideral time at the start of the year, the longitude of the sun, and the difference between the longitude of the sun, and the argument of perigee of the sun.

The next line is a heading line, and the station name, latitude, longitude, and height appear directly below. The satellite and element numbers are printed next.

The main output of the program consisting of the time of search point, predicted position, and visibility quantities is preceded by two heading lines. The output appears in order of increasing time of search point in increments specified by the parameter input card.

1-11-2

Each data line includes the year, month, day, hour, and minute of the search point, right ascension, declination, azimuth, elevation, and slant range of the satellite, and the elevation and illumination angles of the sun. Unless the specified criteria for maximum slant range, minimum elevation angle, and visibility are met, no output for the time of search point will apppear.

#### 11.4.1 Comments with Output

1. PROGRAM ASSUMES PN = PA FOR SATELLITE NO. XXX. This comment indicates that the nodal period at epoch in days is missing from the sixth card of a seven card element set. The anomalistic period is used in its place.

2. ELEMENT CARDS ARE OUT OF ORDER FOR SAT NO XXXX ELEM NO XXXX. This comment indicates that a card was out of order or missing in an element set. The program returns to read another element set. SATELLITE XXX CARD XXX OUT OF ORDER appears simultaneously on the Flexowriter.

# 11.5 Processing

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ORPS predicts the right ascension, declination, azimuth and elevation angles for a given satellite and the elevation and illumination angles of the sun, in a specified time range, for a given station. The limitations of the observing station are considered. The program consists of two sections, the Orbital Plane Search section and the XYZ Look Angle section.

The Orbital Plane Search section computes the xyz point coordinates from the satellite elements and station coordinates. The geocentric latitude of the station, difference in longitude from station to search point, longitude at the search point, radius of earth at station, and the x, y, z coordinates of the station in a rotating system are computed. The time of search point is converted from day of year to month, day, hour, and minute. The sidereal time at Greenwich at the search point time, sidereal time at the search point, right ascension of ascending node at search point, argument of latitude of satellite in the orbit plane, the argument of perigee at the prediction time, the true anaomaly, semi-major axis at epoch, anomalistic period at the prediction time, semi-major axis at the predicted time, eccentricity at the prediction time, and the distance of the satellite from the center of the earth are computed. Then the x, y, z search point coordinates are found. 1-11-3

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The XYZ Look Angle section defines the position of the satellite in terms of right ascension, declination, azimuth, elevation angle, and slant range from the x, y, z search point coordinates. The x, y, z coordinates of the station in a fixed system are found to compute the slant range. If the predicted slant range is greater than the maximum specified slant range, the program increments the time of search point and returns to the Orbital Plane Search section to compute the x, y, z search point coordinates for the incremented time. Otherwise, the right ascension of the satellite, declination of the satellite, sidereal time at the station, hour angle of the satellite, zenith angle, and elevation angle are computed for the search point or prediction time. If the specified minimum elevation angle is greater than the predicted elevation, the program increments the time of search point and returns to the Orbital Plane Search section. Otherwise, the azimuth is computed.

The program determines if the satellite will be visible at search point time by computing the longitude of the sun at the prediction time, right ascension, declination, and elevation of the sun, and the angle  $\eta$  (see illustration). If only visible passes are required, the program tests the elevation and illumination angles of the sun. If the former is less than -4 degrees and the latter is greater than -4 degrees than the pass will be visible. If not, the program increments the time and returns to the Orbital Plane Search section.

Output quantities at each search point time include the time, right ascension, declination, azimuth, elevation, and slant range of the satellite, and the elevation and illumination angle of the sun. The program then increments the time of search point, and returns to the Orbital Plane Search section unless the stop time has been exceeded.

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11.6.1	Formulation - Orbital Plane Search Section			
1.	$\phi = \tan^{-1} (.99329985 \tan \phi)$			
2.	$i_{sp} = \cos^{-1} (\sin Az \cos \phi)$ $0 \le i_{sp} < 180^{\circ}$			
3.	$\Delta \lambda = \cos^{-1} (\cos Az/\sin i_{sp}) \qquad 0 \leq \Delta \lambda < 360^{\circ}$			
4.	$\lambda_{sp} = \lambda - \Delta \lambda$ $0 \leq \lambda_{sp} \leq 360^{\circ}$			
5.	$R = 6378.145 (.998320047 + .001683494 \cos 2\phi000003549 \cos 4\phi + .000000008 \cos 6\phi) + H/1000$			
6.	$\mathbf{x}_{stat} = R \cos \lambda \cos \phi'$			
	$y_{stat} = R \sin \lambda \cos \phi'$			
	$z_{stat} = R \sin \phi'$			
7.	$S t_c = t_a$			
8.	$\theta_{\rm G} = \theta_{\rm o} + .98564735 t_{\rm d} + 360.985647 t_{\rm f}  0 \le \theta_{\rm G} \le 360^{\rm o}$			
9.	$\Omega_{\rm sp} = \Theta_{\rm G} + \lambda_{\rm sp}  0 \le \Omega_{\rm sp} < 360^{\rm O}$			
10.	$\Omega_{t} = \Omega + \dot{\Omega} (t_{c} - T_{o}) + \frac{1}{2} \dot{\Omega} (t_{c} - T_{o})^{2} \qquad 0 \leq \Omega_{t} < 360^{\circ}$			
11.	$\Delta \Omega = \Omega_{\rm sp} - \Omega_{\rm t} \qquad 0 \le \Delta \Omega < 360^{\rm o}$			
	If $ \Delta \Omega  \ge 10^{-31}$ go to 14.			
	If $ \Delta\Omega - 180^{\circ}  \leq g_{\circ}$ to 14.			
12.	$\cos \gamma = \cos i \cos i_{sp} + \sin i \sin i_{sp} \cos \Delta \Omega$			
	Find proper quadrant of $\gamma$			
13.	$\sin\mu = (\sin \Delta \Omega \sin i_{sp}) / \sin \gamma$			
	Go to 15.			
14.	$\sin\mu = \sin\phi' \sin i$			
	Go to 16.			

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15. 
$$\mu = \tan^{-1} (\sin \Delta \Omega / \cos \Delta \Omega \cos i - \sin i / \tan i_{sp})$$
  
 $\cos \mu = \sin \mu / \tan \mu$  Go to 17.  
16.  $a_{st} = \theta_G + \lambda$   
 $\cos \mu = \cos (a_{st} - \Omega_t) \cos \phi$   
 $\mu = \tan^{-1} (\sin \mu / \cos \mu)$   
17. If  $\cos \mu$  is negative  $\mu_0 = \mu + \Pi$ , otherwise  $\mu_0 = [\mu]_{(0, 360)}$   
18.  $\omega_t = \omega + \dot{\omega} (t_c - T_0) + \frac{1}{2} \ddot{\omega} (t_c - T_0)^2 \quad 0 \le \omega_t < 360^\circ$   
19.  $V = \mu_0 - \omega_t$   $0 \le V < 360^\circ$   
20.  $a_0 = (P_a / .058672947)^{2/3}$   
21.  $Pa_t = P_a + \dot{p} (t_c - T_0) + \frac{1}{2} \ddot{p} (t_c - T_0)^2$   
22.  $a_t = (Pa_t / .058672947)^{2/3}$   
23.  $e_t = 1 - (a_0/a_t) (1 - e)$   
24.  $r = a_t (1) - e_t^2)/(1 + e_t \omega sV)$   
25.  $x = 6378.145 r (\cos \mu_0 \cos \Omega_t - \sin \mu_0 \sin \Omega_t \cos i)$   
 $y = 6378.145 r (\sin \mu_0 \sin i)$   
11. 6. 2 Formulation - XYZ Look Angle Section  
26.  $\overline{x} = x_{stat} \sin \theta_G + y_{stat} \cos \theta_G$   
 $\overline{z} = z_{stat}$   
27.  $\rho = [(x - \overline{x})^2 + y - \overline{y}^2 + (z - \overline{z})^2]^{1/2}$   
If  $\rho > \rho_{max}$ , go to 47.

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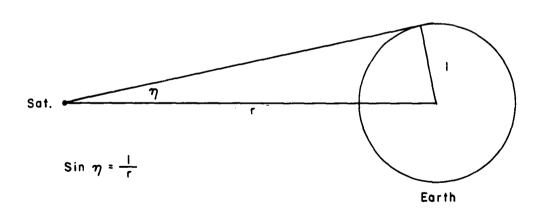
28. 
$$\cos a = (\mathbf{x} - \overline{\mathbf{x}})/[(\mathbf{x} - \overline{\mathbf{x}})^2 + (\mathbf{y} - \overline{\mathbf{y}})^2]^{1/2}$$
  
 $a = \tan^{-1}[(\mathbf{y} - \overline{\mathbf{y}})/(\mathbf{x} - \overline{\mathbf{x}})]$   
29. If cosanegative  $a = a + \Pi$ , otherwise  $a = [a]_{(0, 360)}$   
30.  $\sin \delta = (\mathbf{z} - \overline{\mathbf{z}})/[(\mathbf{x} - \overline{\mathbf{x}})^2 + (\mathbf{y} - \overline{\mathbf{y}})^2 + (\mathbf{z} - \overline{\mathbf{z}})^2]^{1/2}$   
31.  $\theta_{st} = \theta_G + \lambda$   
32.  $HA_s = \theta_{st} - a$   $0 \le HA_s < 360^\circ$   
33.  $\mathbf{z}_{\beta} = \cos^{-1}(\sin \phi \sin \delta + \cos \phi \cos \delta \cos HA_s)$   
34.  $h = \Pi/2 - \mathbf{z}_{\beta}$   
If  $h_{\min} > h$ , go to 47.  
35.  $\sin Az = (\cos \delta \sin HA_s)/\sin \beta$   
 $\cos Az = (-\sin \delta \cos \phi + \cos \delta \sin \phi \cos HA_s)/\sin \mathbf{z}_d$   
 $Az = \tan^{-1}(\sin Az/\cos Az)$   
If .cos Az negative, Az  $= Az + \Pi$   $0 \le Az < 360^\circ$   
36.  $t_{0}$  (t)  $= L_{0} + .98564735 t_{a} + 1.91665 \sin (.98564735 t_{a} - C_{14})$   
37.  $a_{0} = t_{0}(t) - 2.46682 \sin (2t_{0}(t))$   
38.  $\delta_{0} = \tan^{-1}(.4336608 \sin a_{0})$   
39.  $\theta(t) = \theta_G + \lambda$   
40.  $h_{0} = \sin^{-1}[\sinh \delta_{0} + \cosh \phi \cos \delta_{0} \cos (\theta(t) - a_{0})]$   
41.  $a_{gc} = \tan^{-1}(y/x)$   $0 \le a_{gc} < 360^\circ$   
42.  $SD_{gc} = \sin^{-1}(z/6378.145 r)$   
 $\xi = \cos^{-1}(\cos SD_{gc} \cos \delta_{0} \cos(a_{0} - a_{gc}) - \sin SD_{gc} \sin \delta_{0})$   
43.  $\eta = \sin^{-1}(1/r)$   
If all passes desired go to 45, otherwise test elevation of sun (h\_{0}). If h\_{0} > -4 degrees, pass not visible, go to 47.

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- 44.  $I = \xi \eta$ Test illumination angle of sun (I). If I < -4 degrees, pass not visible, go to 47. If  $I \ge 4$  degrees go to 46.
- 45.  $I = \xi \eta$
- 46. Output time,  $a, \delta$ ,  $Az, h, \rho, h_0$ , I.
- 47. Increment time:

 $t_c = t_c + \Delta t$ If  $t_c \le t_e$  return to 8, Orbital Plane Search Section. Otherwise return to start of program to process next set.



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# Glossary

Location	Symbol	Meaning
А	а	Semi-major axis
	t <sub>c</sub> -t <sub>o</sub>	Start time minus epoch time
AA	x <sub>stat</sub>	x coordinate of station in rotating system
ALST	a <sub>st</sub>	Sidereal time at the station
AO	ao	Semi-major axis at epoch
AT	at	Semi-major axis at prediction time
AZ	Az	Search azimuth angle
BB	Ystat	y coordinate of station in rotating system
С	С	Rate of change of period at epoch in days/rev $^2$
C2PI	2π	Constant, 6.28318530
C14	C <sub>14</sub>	Difference between longitude of sun and argu- ment of perigee of sun
CAL		Month of computed point
CAPR	R	Radius of earth at station
CAZ	cos Az	Cosine of astronomical azimuth of object
CC	<sup>z</sup> stat	z coordinate of station in rotating system
CDL	$\cos \Delta \lambda$	Cosine of delta longitude
CMUO	cos μ	Cosine of $\mu$
CRA	cos a	Cosine of right ascension
CRAGC	cosagc	Cosine of agc
CXI	cos ξ	Cosine of ξ
CZD	β	Cosine of zenith angle
D	d	Rate of change of c in days/rev <sup>3</sup>
E	е	Eccentricity
ELEVA	h	Elevation angle
ET	e <sub>t</sub>	Eccentricity at prediction time
ETA	η	See illustration
FNIOM		Constant 10 <sup>-31</sup>
GI	i	Inclination
HITE	н	Height of station
HMIN	h <sub>MIN</sub>	Minimum elevation angle in degrees
HSUNJ	<sup>h</sup> ⊙	Elevation of sun

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ORPS

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ID		Element card number
IDAY		Day of month of computed point
IDCI		Input tape
IDCO		Output tape
IELNO		Element number
IHR		Hour of computed point
IMIN		Minute of computed point
ISATNO		Satellite number
IYR		Year of computed point
MP		Satellite number for Flexo error message
NIX		Switch set to 1 if element card out of order
NOE		E to check against column 80 for 7 card set
OMEGT	ω <sub>t</sub>	Argument of perigee at prediction time
OMGAT	$\Omega_t$	Right ascension of ascending node at search time
OMGSP	Ω <sub>sp</sub>	Sidereal time at the search point
ox	$I_{\odot}(t)$	Longitude of the sun at prediction time
PA	$P_a$	Anomalistic period at epoch
PDDOT	p	Second derivative of nodal period(days/rev. <sup>3</sup> )
PDOT	P	First derivative of nodal period(days/rev. <sup>2</sup> )
PHI	ф	Latitude of station
PHIP	φ <b>'</b>	Geocentric latitude of the station
PI	π	Constant, 3.14159265
PMAX	<sup>ρ</sup> MAX	Maximum slant range in km.
PN	Pn	Nodal period at epoch
PT	$P_{a_t}$	Anomalistic period at prediction time
R		Constant, km/earth radius
RA	Ω	Right ascension of ascending node at epoch
RAD	••	Constant, degrees/radian
RADDOT	1/2Ω	$1/2$ second derivative of $\Omega$ in deg/day <sup>2</sup>
RADOT	Ω	First derivative of $\Omega$ in deg/day
RAV	r	Distance of satellite from center of the
		earth
S	HAs	Hour angle of satellite
SAZ	sin Az	Sine of astronomical azimuth of object
SD	tds	Smithsonian days at time of computed point

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SDEC	$\sin \delta$	Sine of declination
SETA	sin η	Sine of angle η
SHSUNJ	$\sin h_{\odot}$	Sine of elevation of sun
SLRANG	ρ	Slant range
SMUO	sin µ	Sine of argument of latitude of the satellite
STEP	∆t	Time increment in minutes
STNM1, STI STNM3,STN	•	Alphanumeric station name
ТА	ta	Start time (day of year)
	tc	Time of computed point
TANMU	tan µ	Tangent of angle $\mu$
TD	td	Prediction time in days
TDECS	$\tan \delta_{\Theta}$	Tangent of declination of sun
TE	t <sub>e</sub>	Stop time (day of year)
TF	tf	Prediction time (fraction of day).
THETG	θG	Sidereal time at Greenwich at prediction time in degrees
THETO	θ <sub>O</sub>	Sidereal time of Greenwich at start of year
THETST	θ <sub>st</sub>	Sidereal time at station
THTGT	$\theta(t)$	Sidereal time of Greenwich at time t
TO	То	Day of year of epoch
TOD	to	Integral day of year of epoch
TOF		Fractional day of epoch
TPHIP	$tan \phi'$	Tangent of geocentric latitude of station
TRA	tana	Tangent of right ascension of predicted sighting
v	v	True anomaly
w	ω	Argument of perigee at epoch
WDDOT	1/2 హి	$1/2$ second derivative of $\omega$ in deg/day <sup>2</sup>
WDOT	ů	First derivative of $\omega$ in deg/day
x	x	x coordinate of station in fixed system
XAZ	Az	Azimuth of object.
XBAR	x	x coordinate of station in fixed system
XCDL	Δλ	Difference in longitude from station to search point.

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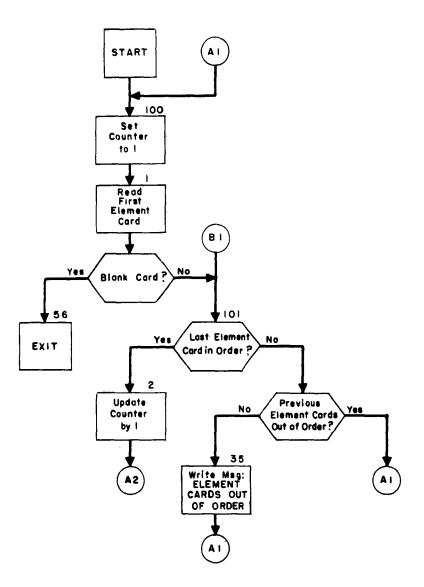
XCGAM	Υ	Intermediate quantity used in computation of SMUO
XDECS	δo	Declination of sun
XI	٤	Angle § intermediate
XIA	I	Illumination angle
XISP	isp	Intermediate quantity used in computing CDL
$\mathbf{X}\mathbf{L}$	λ	Longitude of station
XLSUN	$L_{\odot}$	Celestial longitude of sun at beginning of year
XRASUN	°⊙	Right ascension of sun
XTHETG	θG	Sidereal time at Greenwich at prediction time in radians
XZD	β	Zenith angle
Y	у	y coordinate of satellite
YBAR	<b>y</b>	y coordinate of station in fixed system
YLSP	$\lambda_{sp}$	Longitude at the search point
YR	-F	Year of epoch
Z	q	z coordinate of satellite
ZBAR	z	z coordinate of station in fixed system
ZDEC	δ	Declination of satellite
ZMUO	μο	Argument of latitude of satellite in orbit plane
ZOMG	$\Delta \Omega$	$\Omega_{sp} - \Omega_t$
ZRA	a	Right ascension of satellite
ZRAGC	agc	Right ascension of satellite

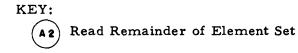
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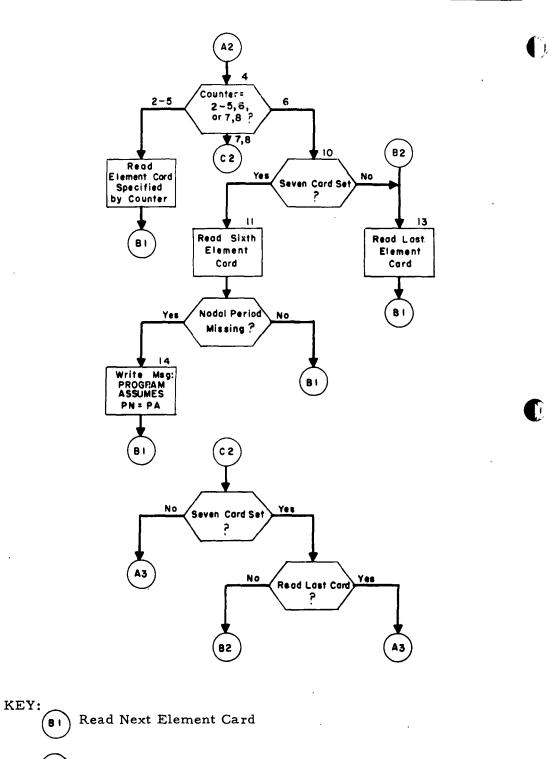
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Read Station Card

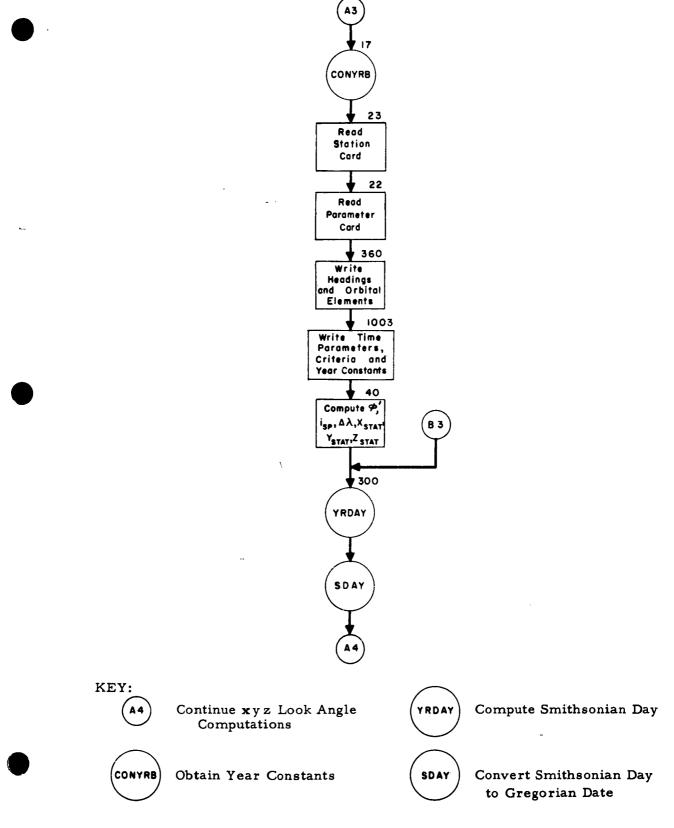
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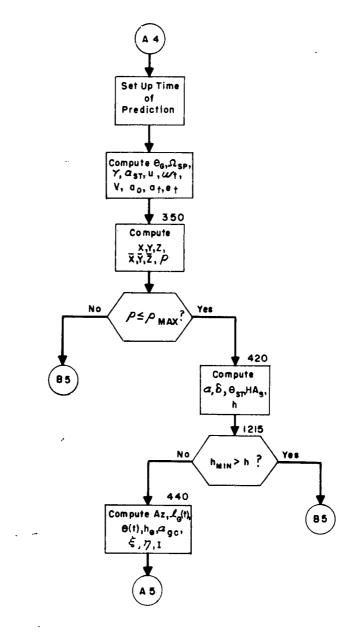
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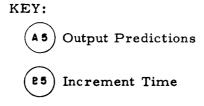
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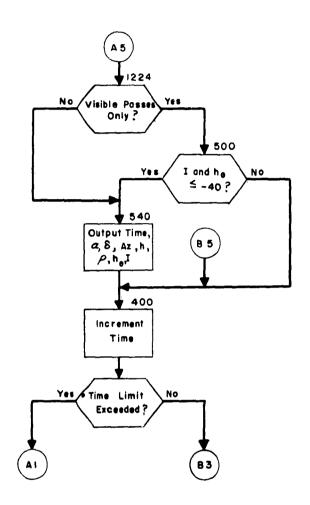


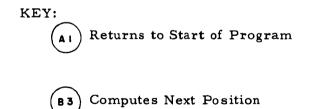
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12.1 ASUM, Observation Acquistion File Summary

12.2 Function

This program will provide a complete listing of the current Observation Acquistion File.

12.3 Input

The input is comprised entirely of the information written on the A-File of the SEAI tape.

12.4 Output

The output consists of the sensor information required to obtain look angles for a given satellite. The output is sequential according to satellite number and includes the following:

Satellite number Sensor number Sensor name Pass code (all or visual) Format request (short or complete) Type (all, 3-point, Baker-Nunn.) Minimum azimuth Maximum azimuth Maximum azimuth Maximum elevation Maximum range Grid, or step, size in minutes

# 12.5 Processing

The program reads the Sensor File to obtain the required sensor information. The Acquisition File is then located and decoded. Sensor information is retrieved from the sensor storage area when the need arises. If the information is not available the following message is printed:

#### NOT IN FILE

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The data is written onto the output tape under control of counters which allow fifty lines of output per page. Output is single spaced. It is in satellite number order and in the sensor number order called for by the Acquisition File. A double space separates the satellites.

12.5.1 Error Messages

When tape reading difficulties occur the following message will appear on the flexo:

TAPE TROUBLE

TYPE GO TO RETRY JOB. TYPE STOP TO TERMINATE.

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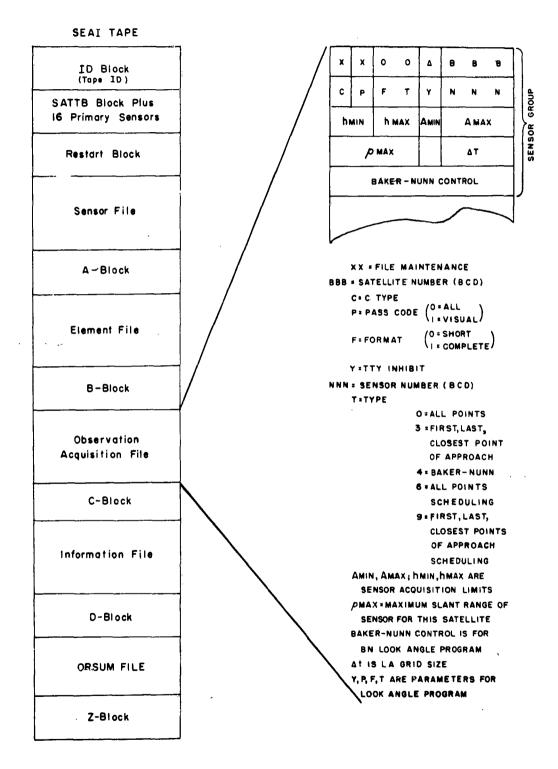
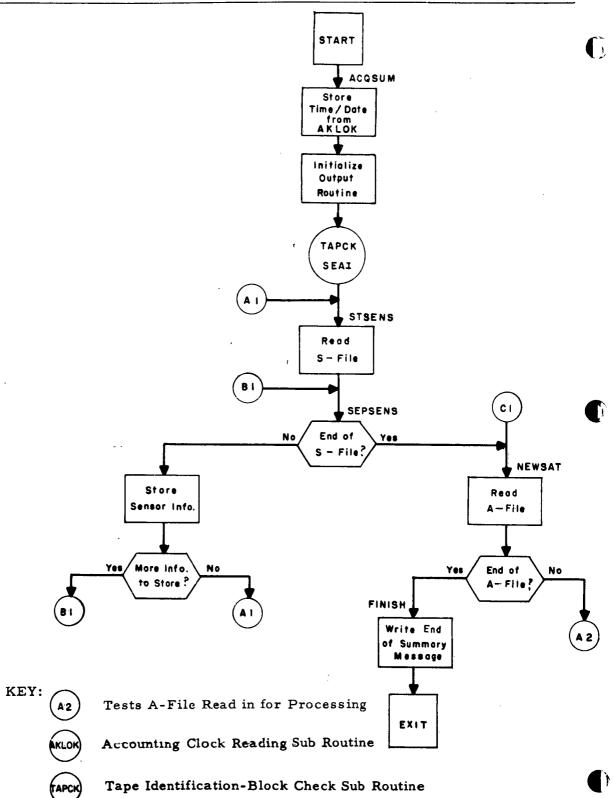


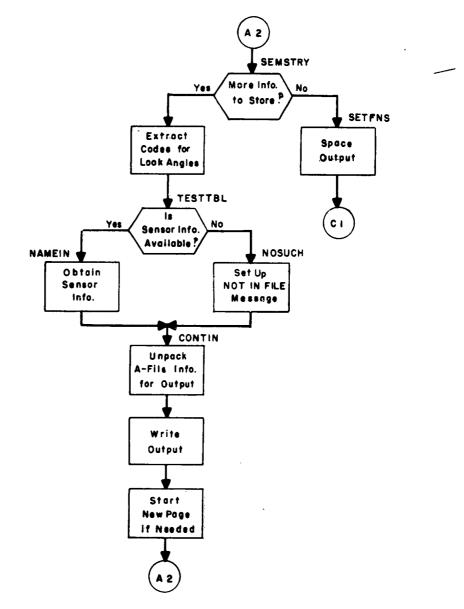
Fig. 12. 1 1-12-3

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## 1-12-4





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Returns to Read More from A-File

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- 13.1 SSUM, Sensor File Summary
- 13.2 Function

This program prepares a complete listing of the Sensor File.

13.3 Input

The input is comprised entirely of the Sensor File willien on the SEAI tape.

13.4 Output

A listing of the sensor data contained in the S-File is provided. The following information is included as part of the output:

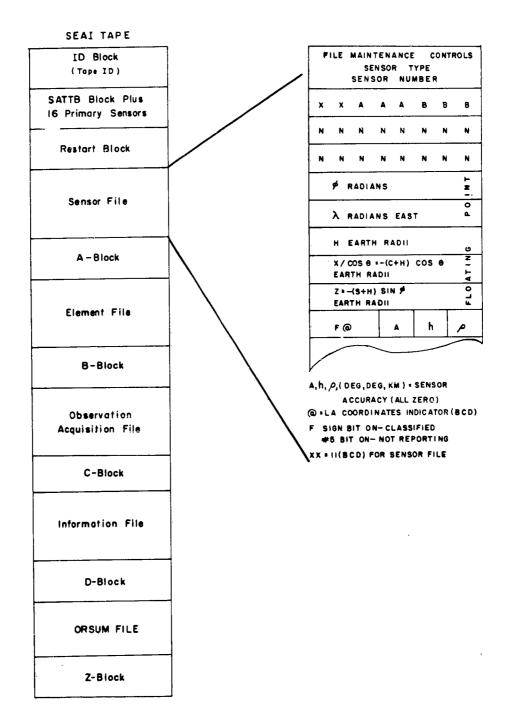
Sensor number Sensor name Latitude in degrees Longitude in degrees Height in earth radii  $X/\cos \theta$  or  $-(C+H) \cos \theta$  in earth radii Z or  $-(S+H) \sin \phi$  in earth radii Accuracy digit for azimuth, elevation and range Classification code Sensor type

#### 13.5 Processing

The SEAI tape is read to locate the Sensor File. After locating this file the program reads 1386 words into core storage. If the ending sentinel is found before the buffer area is filled, the program will process as much information as it has been able to retrieve. If more information is still to be read from the Sensor File the program returns for a second pass after processing the first buffer full of data.

Counters are used to provide for nine groups of five sensors for each page of printed output. Headings and a page count are written on the top of each page. The order of the S-File determines the order of the printout.

#### 1-13-1





## 1-13-2

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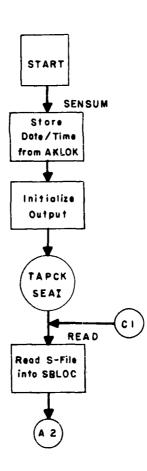
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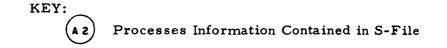
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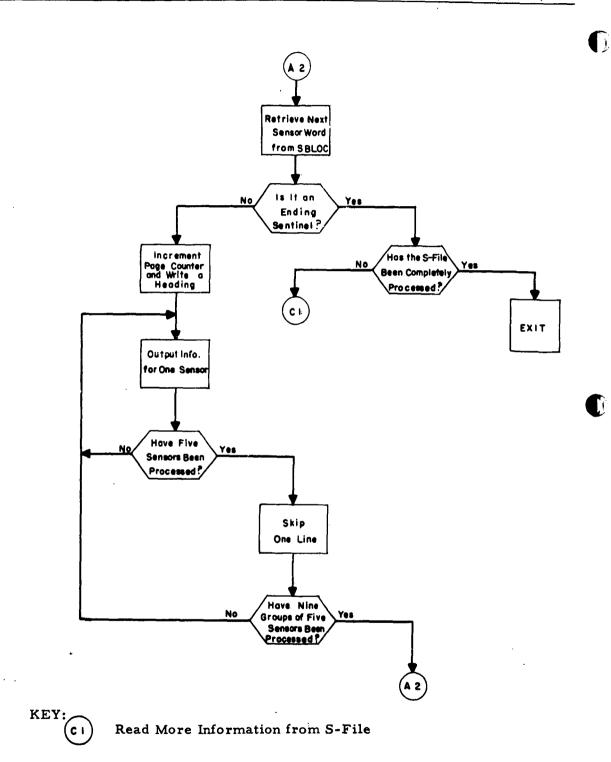
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- 14.1 ISUM, Information File Summary
- 14.2 Function

The program generates a listing of all or specified sections of the Information File of the SEAI tape. A "box score" or tally of the number of satellites still orbiting can be obtained.

### 14.3 Input

The input consists primarily of the I-File of the SEAI tape. Control cards included in the Job deck contain the information required to set up the options available in the program. The following data is punched in cols. 1 - 3:

- a) ALL used to obtain a complete listing of the I-File.
- b) BOX used to obtain a tally of the number of satellites orbiting.
- c) NNN a three digit satellite number for each satellite, required if the ALL option is not used.
- d) END this signals the end of the control data and is always required.

### 14.4 Output

The I-File of the SEAI tape is written for off-line printing. Information printed out includes:

Satellite number Satellite name Launch date Launch site Booster country Payload country

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### 14.5 Processing

The control information described in the input section is read from the system data tape (logical zero). If all satellites are requested, the entire I-File is read and prepared for output purposes. If individual satellites are requested, then the I-File is scanned and only the information desired is written onto the output tape. If a box-score of the orbiting satellites is called for, then this is written out following the printout of the I-File.

### 14.5.1 Error Messages

If irrecoverable tape errors are encountered while reading the SEAI tape the following message is written on the Flexo:

PROGRAM TERMINATED DUE TO POOR BLOCK MARKS, SPROCKET ERROR, S1 OR S2 ERROR, OR DISABLED UNIT ON LOGICAL 4.

A parity error will cause a single reread and a second failure will result in the program issuing a minus one read order with the following comment:

MINUS ONE READ ORDER ISSUED TO OVERRIDE PARITY ERROR





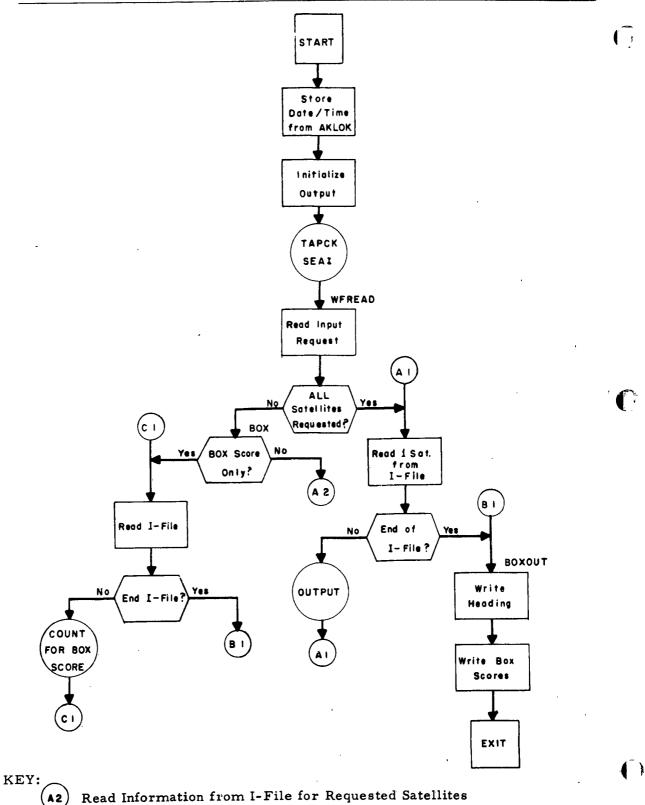
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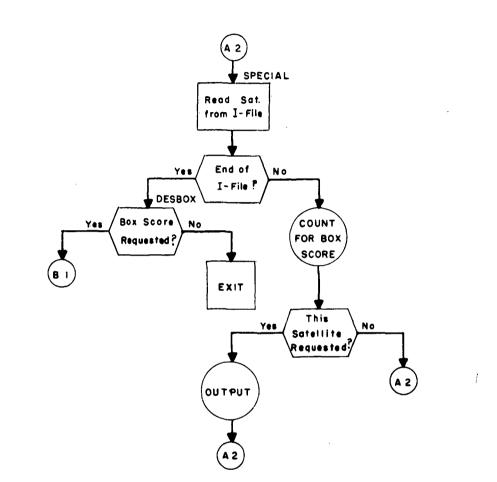
	/
ID Block	2 SATELLITE NUMBER (XCNNN)
(Tape ID)	3
SATTB Block Plus	4 COMMON NAME
16 Primary Sensors	6 LAUNCH DATE
	7 LAUNCH SITE
Restart Block	8
	9
	I O BOOSTER COUNTRY
	I PAYLOAD COUNTRY
Sensor File	1 2 MISSION OR DESCRIPTION
	I S SHAPE
A-Block	I 6 LENGTH
· · · · · · · · · · · · · · · · · · ·	IT HEIGHT
Element File	
	20 MEAN DRAG ( ,DDD)
	21 VARIANCE ( .DDD)
	22 RADAR CROSS SECTION
B-Block	23 VARIANCE ( D.D)
Brock	24 MEAN REFLECTION ( D.D)
	24 MEAN REFLECTION ( D.D) 25 VARIANCE ( D.D) 26 TUMBLING DATE & RATE
	26 TUMBLING DATE & RATE
Observation	L 27 TUMBLING MODE
Acquisition File	28 STABILIZATION
	29 3 30 MANEUVER CHARACTERISTICS
	- 30 MANEUVER CHARACTERISTICS
C-Block	31 TRANSMITTING FREQUENCIES ( .D.D.)
	32
	33
	34
Information File	35
	36 RECEIVING FREQUENCIES
	37
	38
D-Bluck	39 DECAY DATE
D-DIOCK	40 DETERMINED
	41 LIFETIME IN YEARS (
	42 HELIOCENTRIC ELEMENTS (. DDDDD) E
ORSUM FILE	43 ( .DDDDD) A
	44 ( .DDDDD) Q
	45 ( .DDDDD) 02
Z-Block	46 ( . DDDDD) J.
E DIVCK	47 ( .DDDD) P

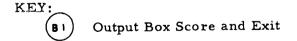
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(Subroutine Output)



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### 15.1 <u>TELTYP</u>, Magnetic Output Tape to Teletype Tape Conversion

# 15.2 Function

The TELTYP Program is used to convert an output tape, written by other programs, to Baudot code. The magnetic output tape used as input to this program must contain the sentinels described in the input section. These sentinels are used to identify the messages to be converted by the TELTYP program. The output tape is searched for the beginning sentinels. Conversion then proceeds from BCD to Baudot code until an end of message sentinel is located. With the completion of one message, succeeding units are sought, identified and converted.

After completely processing the input tape the TELTYP program writes the teletype code onto the input tape. This new information follows the last output previously written onto that tape. A special data select character is used to mark this output as information to be punched from magnetic tape to 5-level paper tape via the Universal Buffer Controller (UBC).

# 15.3 Input

Input to the TELTYP program is comprised entirely of the output tape (logical 11) previously written by any one program, or several programs, processed by the computer.

The TELTYP program must be called upon to perform its function immediately following the programs desiring this optional output. The output tape must not be wrapped up or rewound by the operator or system under which the programs are functioning. If the tape has been rewound, the teletype conversion will not be performed and a message will be written on the Flexo indicating the status of the input tape.

The following sentinels will be searched for on the input tape which must be addressable as logical tape 11 (eleven):

1. Data select 1

2. 14 B's

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- 3. \* \$
- 4. A block of U's.

### 15.3.1 Description of the functions of the sentinels.

#### 1. Data select one

This sentinel is searched for at the beginning of each line on the input tape. If a data select one character is found, printed output is assumed and the processing continues. If any other data select character is present, other types of output are indicated and the line is ignored. The next line is then brought in for processing.

2. BBBBBBBBBBBBBBB (14 B's)

This sentinel marks the beginning of a message to be processed into Baudot code. There will be as many sentinels of this type on the input tape as there are messages to be converted.

3. \* \$ (an asterisk followed by a dollar sign)

This signals the end of the message being converted. The message is wrapped up when this sentinel is found and the sentinel described under 3.1.1 is then searched for by the program.

4. 128 words (1 block) of U's

This sentinel is written on the input tape by the TELTYP program itself before the input tape is rewound and processing is started. It signals the end of the input tape and all processing ceases once it has been found.

When this sentinel has been located the program will rewind the intermediate tape used to store teletype output and copy its contents onto the input tape. The program then exits to the executive program.

### 15.4 Output

The output consists of the required conversion to Baudot code and is written on logical tape 11 (eleven) following the information previously written on that tape. Processing of other programs requiring tape 11 (eleven) for output, but not requiring teletype conversion, can follow the termination of this program.

The 5-level paper tape output can be obtained through use of of the UBC if data select 4 is used.

### 15.5 Processing

The TELTYP program first tests the input tape to determine if data is available for conversion to Baudot code. If the tape is in a rewound status a comment is made on the Flexo and control is returned to the executive program.

After it has been determined that data is available, a block of U's are written on the tape following the last piece of information recorded. The tape is then rewound and processing starts.

A search is made for an output line of information written under data select one. Once this requirement has been met, an attempt is made to locate the second sentinel (14 B's). If the B's are not located the program returns to the search for data select one. If the B's are located before the end sentinel (one block of U's) is reached, conversion of the message which follows is initiated by the setting of a logical switch.

In converting to Baudot code each character is examined. Numeric and alphanumeric characters are converted directly upon entry to a dispatch table. The appropriate shifts are determined and placed in the output buffer as required. Illegal characters are treated as blanks and a space code is supplied to the output buffer. Line feed can be called for by the insertion of an 8-5 punch (octal 15). Two carriage returns and a line feed are supplied by the program at the end of each line of information. An octal 32 (asterisk) is treated the same as any end-of-line indicator (octal 77). Each new line is examined for the data select one character.

The character by character evaluation and conversion continues until an end of message indicator is met. Two indicators are used: (a) an asterisk followed by a dollar sign, or (b) an absolute stop code which normally signifies the end of an output

#### 1-15-3

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tape. The message will be terminated by either signal. Letter shifts will be used to fill out the output area and an end of message indication will be given to the operator via the Flexo. The program then returns to search for the next message on the input tape.

This process continues until the block of U's is found. At that time the intermediate tape, which is used to store the converted message, is rewound and copied onto the input tape. The block of U's is written over by the information being added to the tape. An exit is then made to the executive program.

### 15.5.1 Error Indications

- 1. The following messages will be typed on the Flexo and a return will be made to the Exec.:
  - (a) TAPE FOR STANDARD TTY CONVERSION ALREADY REWOUND. JOB TERMINATED.
  - (b) NON-RECOVERABLE ERROR IN READING INPUT TAPE. UNABLE TO RESTORE TO ORIGINAL POSITION. REMOVE WITHOUT WRAPUP. JOB TERMINATED.

The first message was discussed under the input and processing sections. The second message is self-explanatory.

2. Other error messages are as follows:

This comment will be made only if the computer jumps to location zero or three. A stop-go option is given. Go will send control to the executive program. Stop will print out the message again. The operator is requested to give a dump only if the computer is in a non-interruptable mode.

(b) ERROR IN READING INPUT TAPE. TYPE GO TO ACCEPT OR STOP TO TERMINATE.

This is typed out whenever parity or sprocket errors are encountered.

#### 1-15-4

(c) BAD INTERMEDIATE TAPE. CHANGE TAPE AND TYPE GO.

Converted messages can be lost when the tape change is made.

(d) SCRATCH OR INPUT TAPE DID NOT REWIND, TYPE GO TO TRY AGAIN OR STOP TO TERMINATE JOB.

This message is self-explanatory.

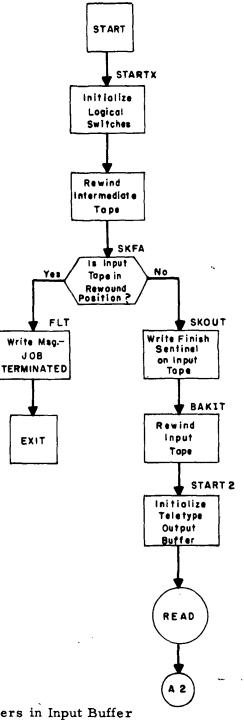
(e) EM

This indicates that the end of message indicator has been encountered.

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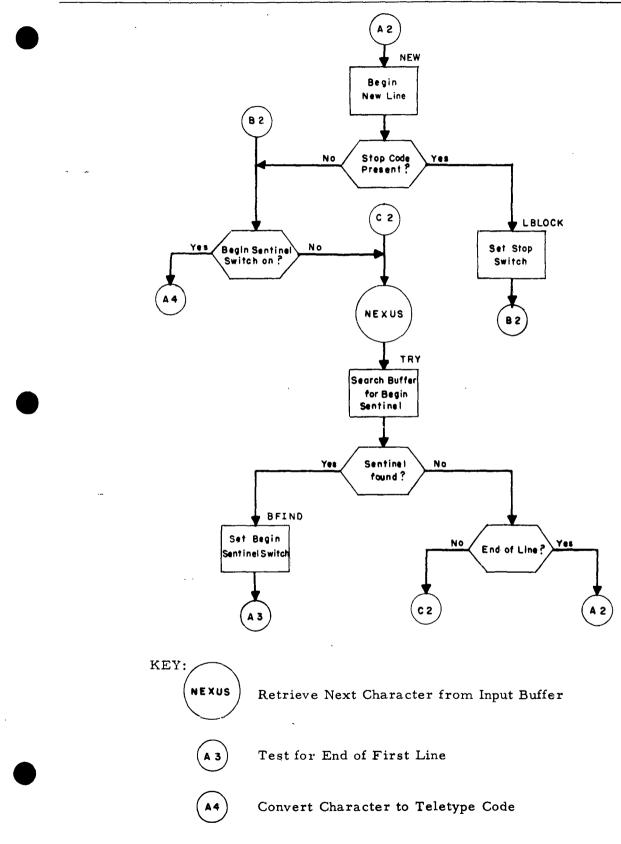


READ

Test Characters in Input Buffer

Fill Input Buffer from Input Tape

### 1-15-6



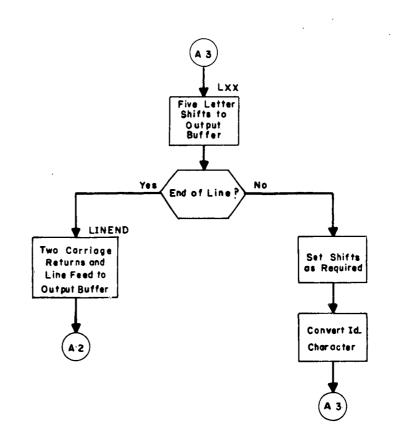
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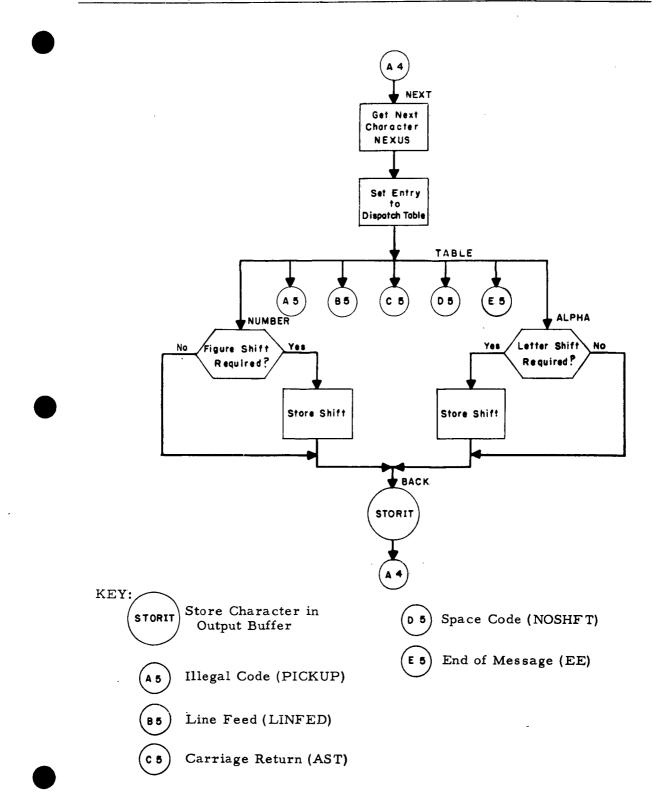




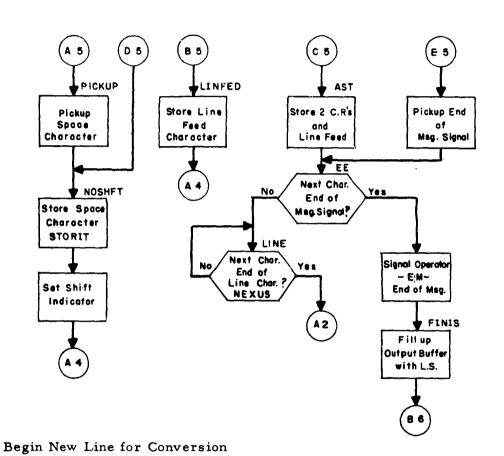
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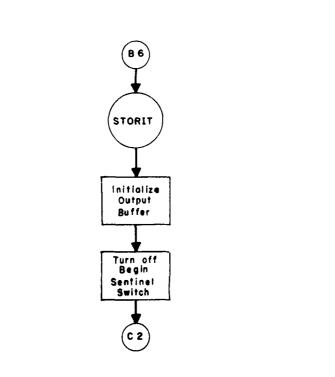
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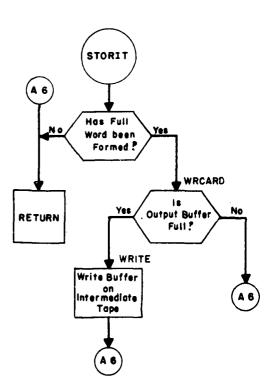
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Process Next Character

Write Output Buffer

(Subroutine Storit)





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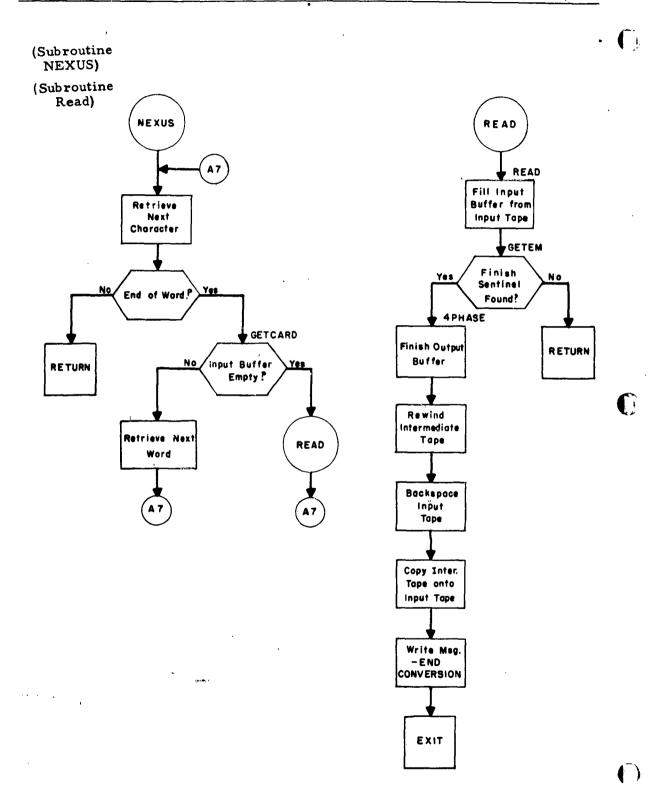


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Search for Start of Next Msg.

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# 1-15-12

### 16.1 BMEWSPT, BMEWS Paper Tape Conversion

### 16.2 Function

BMEWSPT scans output messages from the DIP computer in search of possible SPADATS messages. Other messages of variable length are ignored. Any possible SPADATS message is examined for consistency. If the message appears to be valid, certain control bits are stripped, and the message is converted to standard observation format and put out as a binary tape. This binary output tape is in the format of a system  $TTY \triangle IN$  tape and is used as direct input to ORCON.

### 16.2.1 Definitions

The following terms will be used throughout the BMEWSPT description:

- BMEWS Character The format of this character changes during preparation of magnetic input tape. See BMEWS word. It originally consists of 6 information bits preceded by a parity (P) bit during transmission to SPADATS.
- BMEWS Group Consists of three 24 bit BMEWS words (input tape format). There are three BMEWS groups in a SPADATS message.
- BMEWS Message Any message output by the DIP which may or may not be a SPADATS message.
- BMEWS Word Consists of three BMEWS characters. The format of a BMEWS word changes during preparation of the magnetic input tape. Details appear in that section. The format of a BMEWS word is shown in Figure 16.1.

1-16-1

### Fig. 16.1

BMEWS Word Bit Number	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24
Chad-less Paper Tape	<b>*</b>	I	I	I	I	Ι	Ι	¥ P	I	I	Ι	I	Ι	Ι	<b>*</b> P	Ι	Ι	I	Ι	I	Ι		$\langle \rangle$	
Five Channel Paper Tape	Ρ	E	I	I	I	Ι	I	I	I	Ι	Ι	I	Ι	I	I	I	Ι	I	Ι	I				
Input Tape	S	P	E	I	I	I	S	I	I	Ι	I	I	S	I	I	I	I	I	S	I	I	I	I	I

Chad-less Paper Tape - Original paper tape received by SPADATS from the DIP. BMEWS words appear as three 7 bit characters. Control Bit - An S, P, or E bit.

- DIP Computer used to process all BMEWS data. Q-point data (observations) are outputted on chad-less paper tape and received by SPADATS.
- E-bit Error bit. The second bit of the 20 bit BMEWS word appearing on five channel paper tape. Set to one if transmission errors are detected between the DIP and SPADATS.

Filler Words - Philco 8 character words of all filler characters (32)<sub>8</sub>.

- Five Channel Paper Tape Intermediary between chad-less paper tape and magnetic input tape. BMEWS words appear as four 5 level characters.
- I-bit-Information data bit. Any bit contained in a SPADATS message. except for control bits.
- P-bit Parity bit. The first bit of the 20 bit BMEWS word appearing on five channel paper tape.
- P -bit Parity bit. Original parity bit associated with a BMEWS character.
- S-bit Sixth level bit. The sixth level (most significant) bit added to each of the four 5 level characters on the five channel paper tape when preparing the magnetic input tape.
- SPADATS Message Three groups of three BMEWS words each of which contain Q-point data.
- TTY△IN Tape System input tape for teletype data which is processed by ORCON Tape ID is 70TTY△IN

### 1-16-2

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TYPE 25 Standard Observation - Observation in standard format from a station using a moving beam antenna with T tracking capability. Data readout is automatic.
Zebra - Time - Greenwich Mean Time.

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#### 16.3 Preparing Magnetic Input Tape

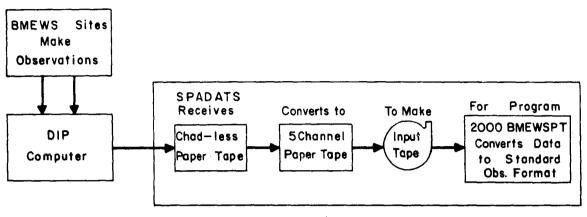


Fig. 16.2

SPADATS receives chad-less paper tape which is direct output from the DIP computer. The chad-less paper tape is converted by equipment to five channel paper tape by decoding BMEWS words.

BMEWS words are decoded into four 5-level characters. The P\* bit of each of the three BMEWS characters is checked and deleted. The remaining 18 bits become the least significant of the 20 on the five channel paper tape. If the check of the P\* bits detected transmission errors, the E-bit is set to one. A new parity bit is generated and becomes the P-bit. An odd parity for the 20-bit group results.

In preparing the input tape from the paper tape, the reader must be in 5-level binary mode. The equipment adds an S-bit to each of the four 5-level characters. Therefore, the required six bit character will appear in computer storage. At least one block of filler words must be added via the simulator.

### 16.3.1 Zebra Time from Flexowriter

Since SPADATS messages do not contain the zebra year, month, and day of the Q-point data, these times must be obtained before conversion to standard observation format. Normally, there is a restriction that input data cannot be more than one day old. Then the year, month, and day may be obtained from the accounting clock. However, if the accounting clock is not working properly, or is set to local time, or if the input data is more than one day old, the correct zebra time must be obtained from the flexowriter.

#### 1-16-4

The program first interrogates the accounting clock and converts the time to BCD. The time is typed out in the format of YR-MM-DD-HHMM ZEBRA. If the time is correct, the operator types a carriage return and the program starts processing the messages.

If the time is incorrect the operator types the requested YR-MM-DD-HHMM followed by a carriage return. The corrected time is then typed out for approval. See 16.4.1-2 for further details.

### 16.3.2 Input Tape

A BMEWS word will appear on the prepared input tape and in computer storage as 24 bits. Since there are 9 BMEWS words in a SPADATS message, a total of at least 4 1/2 Philco words are required.

The input tape may or may not contain messages of varying lengths other than SPADATS messages. The program will ignore extraneous data. A SPADATS message consists of three BMEWS groups. Each BMEWS group contains 72 bits. The format of each group including control bits is illustrated in tables following this section. Note, however, that a BMEWS message in core does not necessarily start at the beginning of a Philco half word. The beginning character of a BMEWS message may be any one of the 8 characters in a Philco word, therefore, the program examines each character separately.

If the control bits are ignored a BMEWS message will appear on the input tape and in core storage as shown in Figures 16.3 and 16.4.



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Group Number	Number Bits	Information Data	Representation or Data Range
ł	3	SPADATS Identifier	IOO(Binary)
	3	Group Count	000(Binary)
	2	Site Number	01 or 10
			(Binary)
	14	Sequence Number	0-16383
	5	Time - Hours	0-23
	6	Time – Minutes	0-59
	6	Time - Seconds	0-59
	4	Time-Fractional Seconds	0-15/16
	5	Credence	Ignored
	6	No. Consolidated Reports	0-63
2	3	SPADATS Identifier	IOO(Binary)
	3	Group Count	001 (Binary)
	2	Site Number	OI or IO
	14	Sequence Number	0-16383
	1	Azimuth Sign O=+,l=-	O or I
	2	Radians	0-3
	14	Fractional Radians	0-1
	1	Range Rate Sign 0=+,  =-	O or I
	3	Nautical Miles Per Second	0-7
	11	Fractional NM/Sec	
3	3	SPADATS identifier	IOO(Binary)
	3	Group Count	010(Binary)
	2	Site Number	OI or iO
	14	Sequence Number	0-16383
	12	Range Nautical Miles	0-4095
	I.	Fractional NM	0 or 5
	2	Elevation Angle Radians	0~3
	14	Fractional Radians	0-1
	3	Blank	000(Binary)

Fig. 16.3

1-16-6

### BMEWSPT

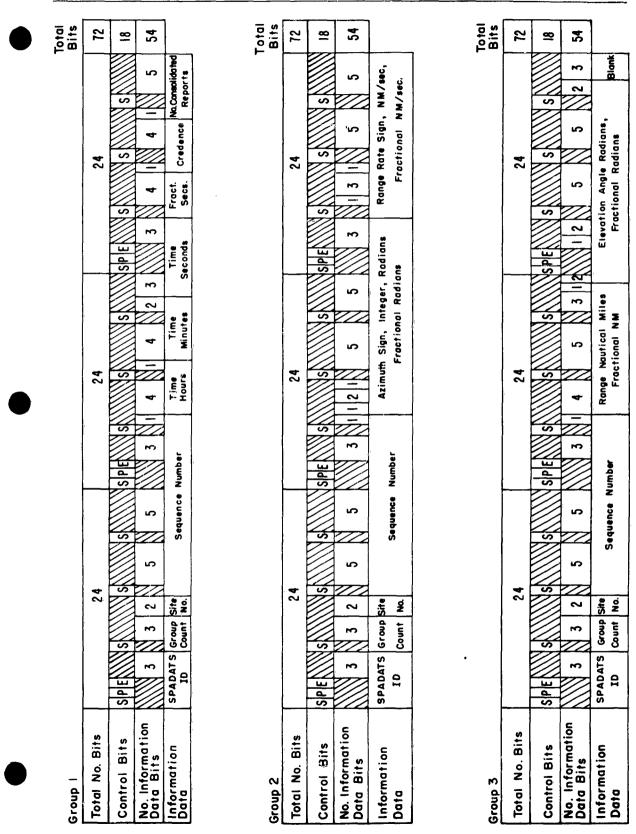
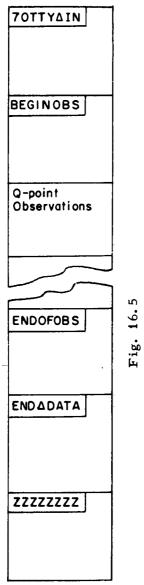


Fig. 16.4

### 16.4 Output

BMEWSPT output is a magnetic tape containing Q-point data converted to standard observation format. The output tape is in the system TTY $\Delta$ IN tape format described below and is suitable for input to ORCON. Unused portions of all blocks are filler words.



The first word of the first block is the tape ID,  $70TTY \triangle IN$ . This indicates to the executive routine that the tape is an ORCON input tape.

The first word of the second block is the ORCON control word BEGINOBS. This indicates to ORCON that the following input will be in standard observation format.

The Q-point data converted to type 25 standard observations are written 12 observations per block. There are no restrictions on the number of blocks used for the observations which are all classified unknowns.

The first word of the second to last block is the ORCON control word ENDOFOBS. This is a signal for ORCON indicating the end of the standard observation input.

The first word of the next to last block is the ORCON control word END $\triangle$  DATA. This indicates to ORCON that there is no more input from the TTY $\triangle$ IN tape.

The last block contains 120 words of Z's

Each generated observation will occupy the equivalence of ten Philco computer words on the output tape. Each observation will be in the standard observation format shown in figure 16.6.

#### 1-16-8

BMEWSPT

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Fulico Word		-	•		2	3	4		5 6
Card Columns 1-3 4,5 6-9 10 11-15	1-3	4,5	6-9	01	11-15	16-24	25 - 30	25-30 31-37	38 - 44
No. BCD Characters	3	2	4	-	2	6	9	1	7
Information	Sat. No. Type	Type	ta tior	∢ບບ	A CC Date	Time (Z)	Elevation	Azimuth	Slant Range Km.
Representation -00 25 xxxx to YMMDD or Data Range 9	00 -	25	XXXX	ර° ග	AMMDD	HHMMSS. sss	D D.dddd	DDD. dddd	КККК. 44

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Philco Word	9	~	 8	6	01	
Card Columns	45 - 53	┝━╉╌ 			73-78	
No. BCD Characters	6				9	
Information	Range Rate KM/Sec.	e			Sequence Number	
Representation or Data Range	K K. KKKKKK	×			XXXXX	

Fig. 16.6

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### 16.5 Processing

BMEWSPT obtains the zebra year, month, and day of observation from the accounting clock and/or from corrections entered via the Flexowriter. The two header blocks containing the tape identification  $70 \text{TTY} \Delta \text{IN}$  and the ORCON control word BEGINOBS as their first words respectively are written on the output tape. The input data is then transferred to the data storage area. If the data storage area does not contain any zero characters, the message, "MAY BE CODE MODE" is typed on the Flexo. This indicates that the input tape may have been prepared from the five channel paper tape under code mode, which ignores blanks, rather than binary made which converts blank paper tape to zero characters. This is only a possible error message.

The program then searches for a SPADATS message. Each character in core is examined until one is found which has 100 as the last three bits. This is tentatively assumed to be the SPADATS identifier. The next three information bits are checked and should indicate the first group count, 000. If they are, the bits corresponding to the SPADATS identifiers and group counts for the second and third groups are examined. If they agree with the expected values, the site and sequence numbers from the first group are compared with the corresponding values in the second and third groups. If there is agreement and if the site number is 1 or 2, it is assumed that a SPADATS message has been located. If any test fails, the search for the SPADATS identifier continues with the character following the original character with 100 as its last three bits.

After a SPADATS message is found, the Q-point data is converted to a type 25 standard observation. The sequence number is converted to BCD. The site number is converted to station number by a table look up. The zebra time is converted to BCD. Year, month, day are obtained from the accounting clock or Flexowriter while hours, minutes, and seconds are obtained from the Q-point data. The credence is ignored.

1-16-10

The number of consolidated reports is used in the determination of the accuracy. The first three of the six bits form a binary number, 0-7. This number is subtracted from 7 to give the accuracy. A resultant accuracy of 0 corresponds to a large number of reports. An accuracy of 8 or 9 may result from the detection of one or more parity errors, respectively. An error count is made up of the sum of the 9 E-bits and the number of bad parity checks. A non-zero error count results in an 8 or 9 accuracy.

If the azimuth is positve, there is a direct conversion from radians to circles. If it is negative, the magnitude is substracted from  $2\pi$  and converted to circles.

Elevation is converted to circles. If the result is greater than 0.25, one-half circle is added to the azimuth and the elevation is subtracted from 0.5. The resultant elevation and azimuth values are converted to degrees.

Range and range rate are converted to kilometers and kilometers/second respectively. Because the output range rate has nine characters, the integer and fractional parts are processed separately.

The resultant observations are written out on the output tape, 12 observations per block. After all data has been processed, and all observations written on the output tape, two blocks containing the ORCON control words ENDOFOBS and END $\Delta$  DATA as their first words respectively are written on the output tape. The last block written on the output tape contains 120 words of z's. The output tape is now ready to be processed by ORCON.

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#### 16.4.1 Flexowriter Messages

#### 1. MOUNT SCRATCH ON UNIT 7 AND INPUT ON UNIT 0.

This message is typed at the start of the program. After typing this message, the program goes to the STOP-GO routine. The program will begin when the operator types GO.

#### 2. YR-MM-DD-HHMM ZEBRA

The zebra time at the start of the program is typed for the operator's approval. One or more of the zebra groups may be altered by the operator. The groups are separated by hyphens. See 16.3.1.

#### a. TRY AGAIN

This message is typed if the operator makes an error in typing in the requested time. Program waits for another type in of the time and retests.

### b. ILLEGAL DATE-TIME. RETRY

Typed if an illegal group in the zebra time is found such as a month > 12. The operator will then type in a new zebra time which will be tested.

- 3. WRITE ERRORS
  - READ ERRORS
  - PROC ERRORS

Appropriate comment is typed if tape errors occurred and could not be corrected. Control returns to the executive routine after typeout.

#### 4. INPUT MAY BE CODE MODE

This message indicates that the paper tape may have been converted to the magnetic input tape under code mode instead of binary mode. The program tests one block for zero characters. If no zero characters are found, the above message is typed on the Flexo. The program then continues normally.

1-16-12

Punched card output, if specified, is under control of a PATH switch which eliminates multiple punching of the same ephemeris data. This switch also controls the production of an ephemeris tape from a set of ephemeris cards. This switch equals one (1) for the first time through the program, for a particular set of ephemeris data, and two (2) for each subsequent pass through the program for this same data. It can be reset to one by reading a negative station number. Thus ephemeris cards are punched, or an ephemeris tape made, only on the first pass through the program for that particular set of ephemeris data.

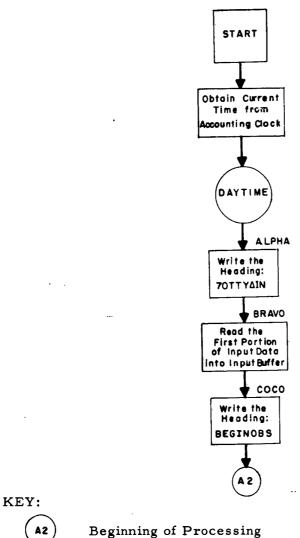
### 9.5.1 Error Message

If the program is unable to read the ephemeris tape, the comment EPHEMERIS TAPE TROUBLE will be written on the output tape. In addition, the comment PROGRAM TERMINATED DUE TO EPHEMERIS TAPE TROUBLE will appear on the Flexo, and control will return to the executive routine.

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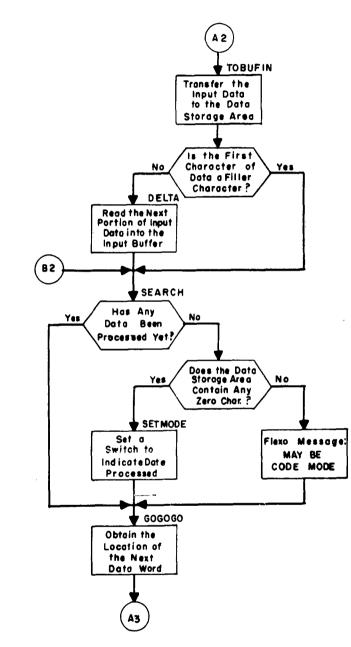


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Beginning of Processing



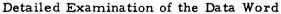
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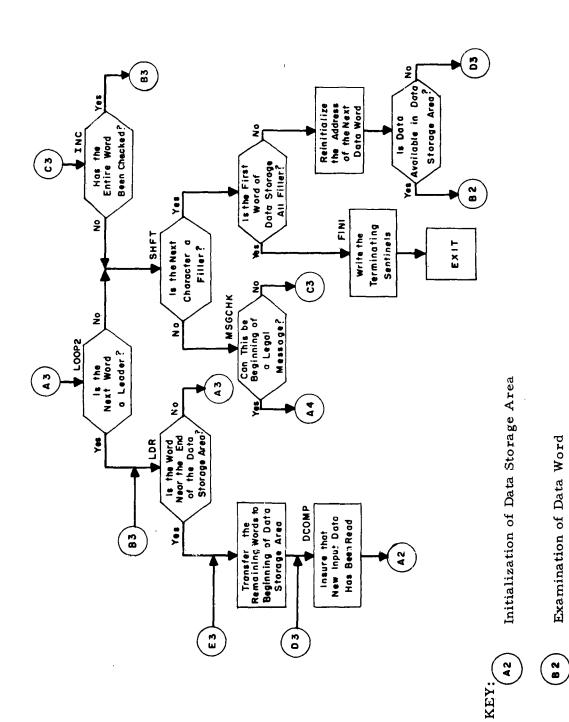
Examination of Possible Message

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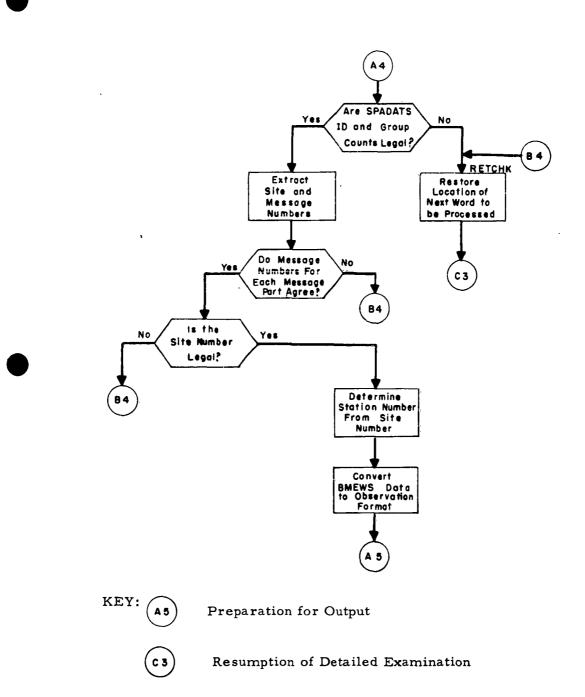
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**Examination of Data Word** 

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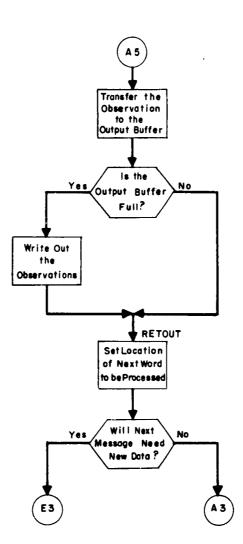
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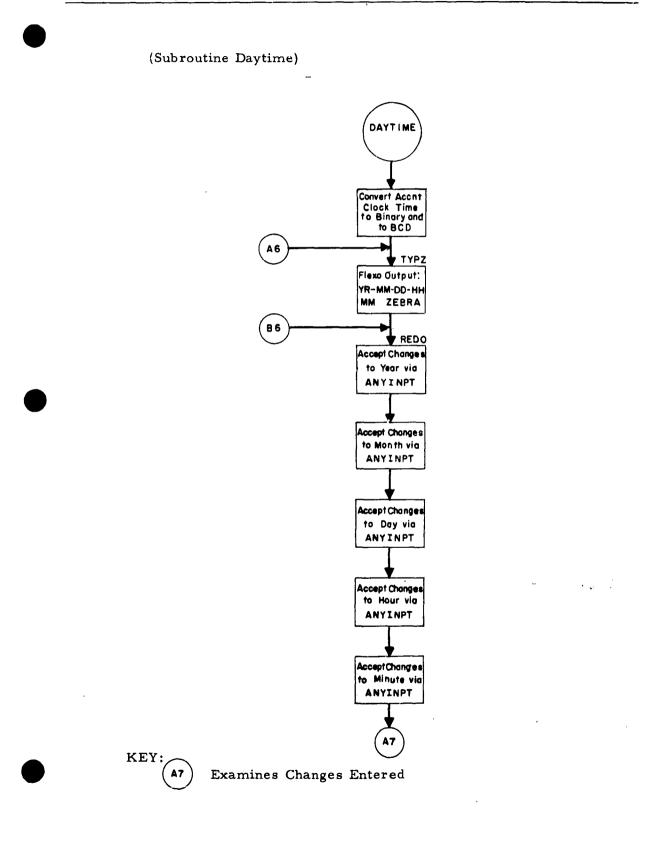
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Detailed Examination of the Data Word

Transfer of New Data to Data Storage Area

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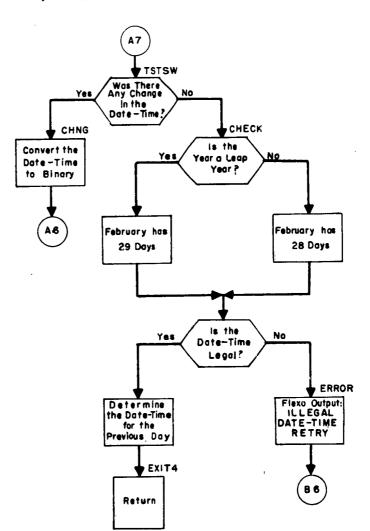
BMEWSPT



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## (Subroutine Daytime)



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Output of Date-Time to Flexowriter

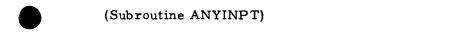
Acceptance of New Date-Time from Flexowriter

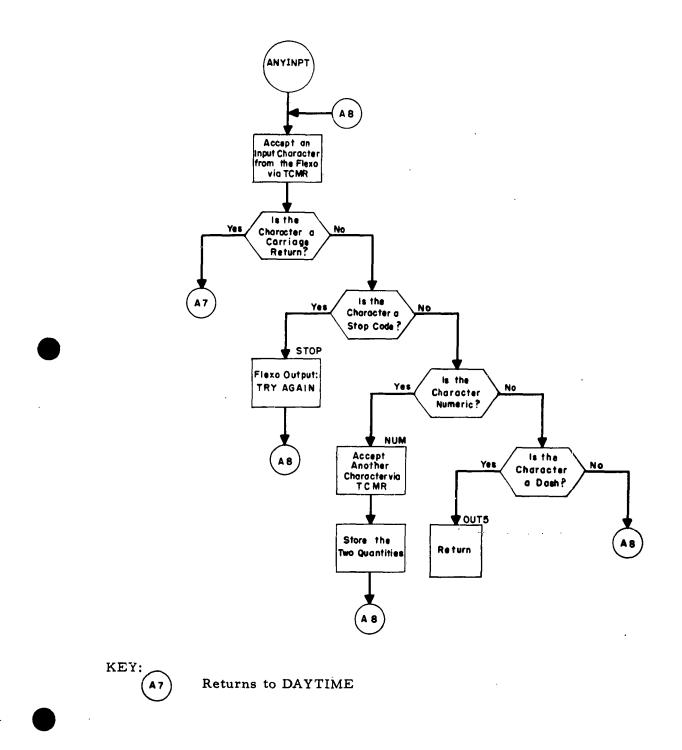
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# SECTION 2 SUBROUTINES AND FUNCTIONS

This section contains descriptions of subroutines and functions which are used by many of the programs described in Section 1. All, with the exception of XYZSB, have been written with the linkage required by FORTRAN or ALTAC. The description of XYZSB, which was written by Aeronutronic, a division of the Ford Motor Company, has been included to serve as a convenient reference within this volume.

The functions described are used to perform intermediate calculations and produce a single valued result. The subroutines perform a series of calculations and may produce more than one result.

Number	Name	Title	<u>Type</u>
1	CONYRB	Obtain year constants	S
2	NHOLY	Remove overpunch (fix. pt.)	F
3	PROPR	Properize argument (0 to $2\pi$ )	F
4	SDAY	Gregorian date from Smithsonian day	S
5	SHOLY	Restore Hollerith sign	F
6	SMITH	Obtain theta Greenwich	F
7	UNHOLY	Remove overpunch (flt. pt.)	F
8	YRDAY	Convert to Smithsonian day from yr., day of yr.	S
9	YRMDY	Convert to Smithsonian day from yr., mth., day	S
10	XYZSB	Analytical integration routine	S

1.1 CONYRB, Year Constants (Subroutine)

1.2 Purpose

CONYRB is used to obtain Theta Greenwich, the celestial longitude of the sun, and the difference between the longitude of the sun and the argument of perigee of the sun at the start of a year.

1.3 Use

Call CONYRB (IX, X1, X2, X3, X4, X5, X6, X7, X8)

1.4 Input

IX must be a fixed point integer equal to the last digit of the year. No further arguments are necessary for entry to the sub-routine.

# 1.5 Output

Upon exit from the subroutine, the arguments represent the following constants:

- IX (Fixed point) = last digit of the year.
- X1 (Floating point) = year.
- X2 (Floating point) = Theta Greenwich at start of year.
- X3 (Floating point) = celestial longitude of the sun.
- X4 (Floating point) = .017202789
- X5 (Floating point) = difference between the longitude of the sun and argument of perigee of the sun.
- X6 (Floating point) = .0334502.
- X7 (Floating point) = .043053055.
- X8 (Floating point) = . 43365539.

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# 2.1 NHOLY, Remove Hollerith Overpunch (Function)

# 2.2 Purpose

NHOLY is used to remove an overpunched sign and to convert the argument to the absolute value of the fixed point integer.

2.3 <u>Use</u>

NHOLY (IX)

2.4 Input

The argument (IX) contains any left justified BCD character which may include an overpunched sign.

# 2.5 Output

NHOLY yields the absolute value of the fixed point integer contained in the left most BCD character of the original argument.

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3.1 PROPR, Properize Argument (Function)

3.2 Purpose

PROPR is used to properize an angle, i. e., find the value of the angle between 0 and  $2\pi$ . If x is any angle in radians, PROPR = [X] (0,  $2\pi$ )

3.3 Use

PROPR (X)

3.4 Input

The angle (X) must be in floating point radians. There is no restriction on its value.

## 3.5 Output

PROPR yields the value of the angle (X) in the range  $0 \le x < 2\pi$ .

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4. 1	SDAY, Gregorian Date from Smithsonian Day (Subroutine)		
4.2	Purpose		
	SDAY is used to convert the Smithsonian day number to		
	Gregorian day, month, and year.		
4.3			
4. 3	Use		
	Call SDAY (SDATE, YEAR, YRDATE, MO, HMO, DATE)		
4.4	Input		
	SDATE is the floating point Smithsonian day number. No		
	further arguments are necessary for entry to the subroutine.		
4.5	Output		
	Upon exit from the subroutine the arguments represent the		
	following quantities:		
	SDATE (Floating point) = Smithsonian day number.		
	YEAR (Floating point) = year, i.e., 1963		
	YRDATE (Floating point) = day number in year.		
	MO (Fixed point) = month number.		
	HMO (Alphanumeric) = month (3 BCD characters) abbreviation, i.e., JAN.		
	DATE (Floating point) = day number in month.		
4.6	Formulation		
	F = Smithsonian day -15018		
	$y = \frac{F}{365, 25} + 1900$		
	[x] = integer function of x		
	A = [365.25y + .9]		
	year day = $F - A$		
	If year day $< 0$ , $y = y - 1$ recompute A and year day.		
	B = 365.25y + .9 - A		
	=		

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C \begin{cases} = 1 \text{ if } B - 8 \ge 0 \text{ (leap year add 1 day)} \\ = 0 \text{ if } B - .8 < 0 \end{cases}
J = \text{year day} - 59 - C
If J \le 0, h = 30. 1; otherwise h = 32. 3 - C
month = \frac{\text{year day}}{30.6} + 2
E = [30.6 \text{ month} - h]
If year day \le E subtract 1 from the month and recompute E.
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Otherwise day number in month = year day - E.

5.1 SHOLY, Restore Hollerith Sign (Function)

5.2 Purpose

SHOLY is used to interpret the original overpunched BCD argument in terms of an algebraic + or -.

5.3 Use

SHOLY (X, Y)

5.4 Input

SHOLY requires two arguments. The first argument (X) is the overpunched BCD character. The second argument (Y) is the unsigned value of the first argument (X) in floating point.

## 5.5 Output

SHOLY yields the signed floating point value of the second argument.

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# 6.1 SMITH, Obtain Theta Greenwich (Function)

6.2 Purpose

SMITH is used to obtain Theta Greenwich in degrees for the start of the year.

6.3 Use

SMITH (X)

6.4 Input

The argument (X) is the last digit of the year expressed in floating point.

6.5 Output

SMITH yields the proper Theta Greenwich (X) in degrees.

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# 7.1 UNHOLY, Remove Hollerith Overpunch (Function)

# 7.2 Purpose

UNHOLY is used to remove an overpunched sign and to convert the argument to an unsigned floating point number.

## 7.3 <u>Use</u>

UNHOLY (X)

# 7.4 Input

The argument (X) contains any left justified BCD character which may include an overpunched sign.

# 7.5 Output

UNHOLY yields the absolute value of the floating point number contained in the left-most BCD character of the original argument. UNHOLY uses the function NHOLY.

8.1	YRDAY, Smithsonian Day from Year and Day of Year (Subroutine)
8.2	Purpose
	YRDAY is used to convert the Gregorian year, and day
	number to Smithsonian day.
8.3	Use
	Call YRDAY (YRDATE, YEAR, SDATE)
8.4	Input
	YRDATE is the floating point day number of the year and
	YEAR is the floating point year, i.e., 1963.
8.5	Output
	Upon exit from the subrouting the arguments represent the
	following quantities:
	YRDATE (Floating point) = day number of year.
	YEAR (Floating point) = year.
	SDATE (Floating point) = Smithsonian day.
8.6	Formulation
	[x] = integer function of x
	y = year - 1900
	A = [365.25y + .9]
	d = day number of year
	c = 15018, Smithsonian day number of January 1, 1900
	Smithsonian day = $A + d + c$ .
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9.1	YRMDY, Smithsonian Day from Year, Month, Day (Subroutine)
9.2	Purpose
	YRMDY is used to convert the Gregorian year, month, and
	day to Smithsonian day.
9.3	Use
	Call YRMDY (MO, DATE, YEAR, SDATE)
9.4	Input
	MO is the fixed point integer month number, DATE is the
	floating point day number of the month and YEAR is the floating
	point year, i.e., 1963.
9.5	Output
	Upon exit from the subroutine the arguments represent the
	following quantities:
	MO (Fixed point) = month.
	DATE (Floating point) = day number of month.
	YEAR (Floating point) = year.
	SDATE (Floating point) = Smithsonian day.
9.6	Formulation
	[x] = integer function of x
	y = year - 1900
-	m = month of year
	d = day of month
	A = [365.25y + .9]
	B = 365.25y + .9 - [365.25y + .9]
	h $\begin{cases} = 32.3 \text{ if } B8 \geq 0 \text{ (leap year, add 1 day)} \\ = 31.3 \text{ if } B8 < 0 \end{cases}$
	c = 15018, Smithsonian day number of January 1, 1900
	Smithsonian day = $[30.6m - h] + d + A + c$ .
<del></del> .	2-9-1

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# 10.1 XYZSB, Analytical Integration Routine

This routine was written by Aeronutronic. A complete description of the subroutine, including equations, appears in the Aeronutronic Publication U-1691, pages 4-78.

# 10.2 Perpose

XYZSB computes the predicted position, r, and velocity, r, at some given time, t. The <u>N</u>, <u>M</u> orbital elements at some epoch time, t<sub>o</sub>, must be supplied upon entry to the routine (c. f. Fig. 10.1).

# 10.3 Calling Sequence

	(10, Z
TIXZ	20, 2
	30, 2
JMP	XYZSB

## 10.4 Input

Index register two (XR2) must be set equal to 10, 20 or 30 upon entry. The significance of this setting will be explained under the output section (10.5) and will be referred to as the exit flag.

The following quantities must be available in the indicated locations upon entry into this routine: (See Assign Table).

Location	Symbol	Contents
Т	t - t <sub>o</sub>	Time interval in minutes since epoch time, $t_0$
AO	a ,	Semi-major axis at t
QO	۹ <sub>0</sub>	Perigee distance at t
EO	e	Eccentricity at t
XNO	n o	Mean angular motion at t

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AXNO Components in orbit plane at t of  $\underline{a} = \underline{eP}$ AYNO Ω **XNODEO** Longitude of ascending node at to in radians L<sub>o</sub> XLO Mean longitude at t radians c'' С Drag coefficient (-4/3)c" XM4CO3 COSI  $\cos i = W_{-}$ Cosine of the inclination k\_L\_so XKELSO  $N_{\pi} = 0$ XNX+2  $M_z = \sin i$ XMX+2  $k_{p} = 0.07436574$ XKE a constant  $k_e^{-}\sqrt{\mu} Ja_e^{-2} = 0.120717 \times 10^{-3}$ , a constant XJGRCF

Index register two (2) must be set equal to 10, 20 or 30 upon entry. The significance of this setting will be explained below under Output and will be referred to as the exit flag.

#### 10.5 Output

Q

1. (XR2) = 10, 20 or 30

If the exit flag is equal to 10, 20 or 30 the following quantities are available upon exit from this routine in the indicated locations:

Location	Contents	Location	Contents
А	a	ESQ	e <sup>2.</sup>
E	e	XNODE	$\Omega$ (rad)
р	р	XNODOT	$\Omega$ (rad/min)
RTP	√́Р	RTA	$\sqrt{a}$
		XN	n

2. (XR2) = 30

If the exit flag is equal to 20 or 30 the following is also output:

#### 2-10-2

Location OMGDT	<u>Contents</u> ŵ (rad/min)	<u>Location</u> RTESQ	$\frac{\text{Contents}}{\sqrt{1 - e^2}}$
ÓMGAS	$\omega_{\rm s}$ (rad)	XL	L (rad)
AXN	a xN	U	U (mod 2 <b>π) (rad</b> )
AYN	<sup>a</sup> yN		

3. (XR2) = 30

# If the exit flag is equal to 30 the following is also output:

Location	Contents	Location	Contents
SINO	$\sin \Omega = N_{y}$	UX	υ
COSO	$\cos \Omega = N_{x}$	UY	$\mathbf{U}_{\mathbf{y}} \left\{ \underline{\mathbf{U}} \right\}$
WX	Wx	UZ	U <sub>z</sub>
WY	Wy	VX	v
XMX	M <sub>x</sub>	VY	$v_{y}^{-}$ $\underline{v}$
XMY	M <sub>v</sub>	VZ	vz
DENM	$(1 + \sqrt{1 - e^2}) \sqrt[2]{1 - e^2}$	х	×
COSEO	$\cos (E + \omega)$	Y	y <u>r</u>
SINEO	$\sin (E + \omega)$	Z	z
ESINE	e sin E	RDOT	ŕ
ECOSE	e cos E	RVDOT	rv
R.	r	XDOT	×Ì
AR	a/r	YDOT	ý <u>ř</u>
SINU	sin u	ZDOT	żJ
COSU	cosu		

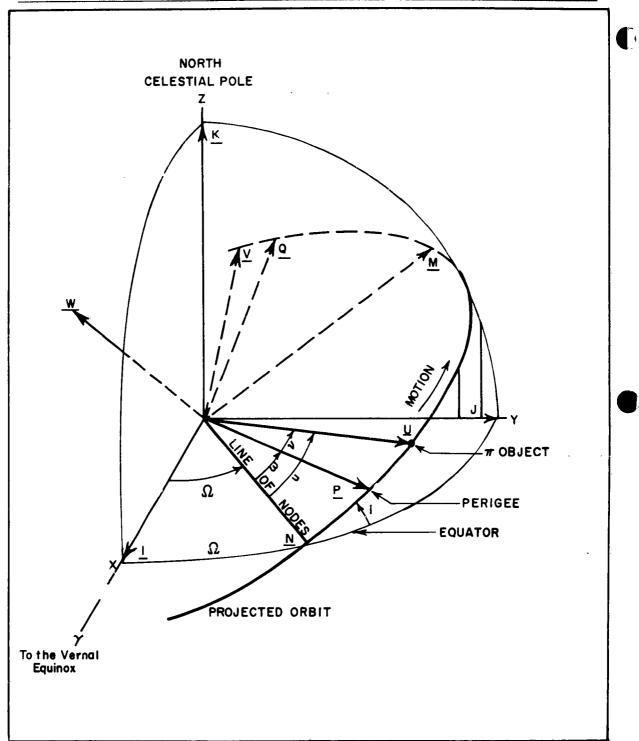


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PROJECTION OF ORBIT ON CELESTIAL SPHERE, WITH ORIENTATION UNIT VECTORS AND ANGLES DISPLAYED

Fig. 10.1

Element Card Formats:

All programs accept a standard 7-card element set as input. The formats for these cards appear on the following pages. Those columns which are shaded are not used. In all floating point numbers, the decimal point may be placed anywhere in the field. Numbers requiring an exponent indicator such as  $c_a$ ,  $d_a$ , and  $c_n$ , must be in one of the following formats:

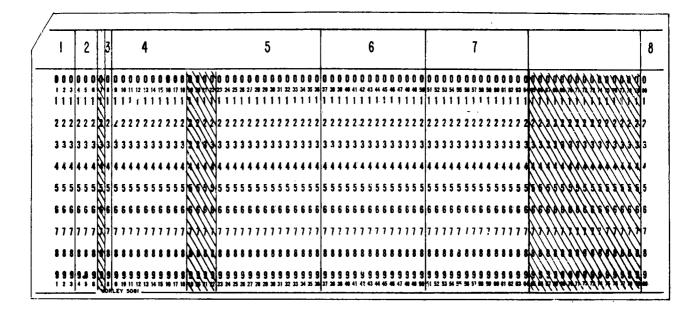
 $\pm$  . DDDDDDDDE  $\pm$  XX

- + . DDDDDDDD $\Delta\Delta$  + X
- $\pm$  , DDDDDDDD  $\pm$  XXX

All programs except ASUM, SSUM, ISUM, PSR, and RESPLT will also accept a 6-card element set. Fields which differ from those of the 7-card element set are noted in the rightmost column of the card format description.

0

## Element Card (1 of 7)

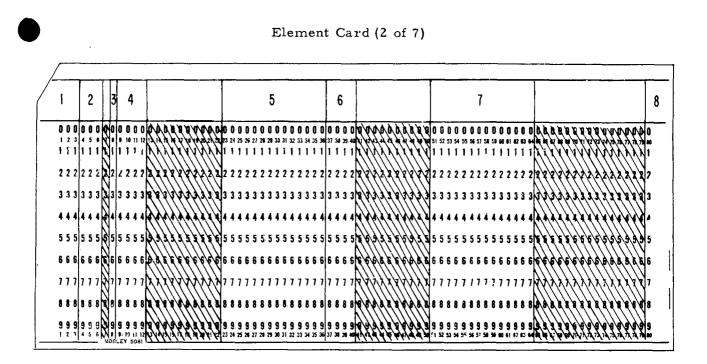


Field	<u>Columns</u>	7-Card Element Set	6-Card Element Set
1	1 - 3	Satellite Number, Right Justified	Same
2	4-6	Element Set Number, Right Justified	Same
3	8	Card Number = 1	Same
4	9-18	Satellite Name, Alpha- numeric	Same, except in Col. 9-20
5	23-36	Epoch Revolution Number, (N <sub>o</sub> ), Right Justified	Same
6	37-50	Eccentricity, (e)	Same
7	51-64	Inclination, (i)	Same
8	80	Card Type = E	Blank

# Element Card (1 of 7)

# 3-1-2

Elements



Field	Columns	7-Card Element Set	6-Card Element Set
1	<b>1 -</b> 3	Satellite Number, Right Justified	Same
2	4-6	Element Set Number, Right Justified	Same
3	8	Card Number = 2	Same
4	9-12	Epoch Year	Same
5	23-36	Epoch Time, (T <sub>o</sub> ), in Days of Year	Same
6	37-40	Not Used	Epoch Year of Decay
7	51-64	Mean Orbital Longitude, (L <sub>o</sub> ), in.Degrees	Epoch Time of Decay in Days of Year
8	80	Card Type = E	Blank

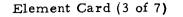
Element Card (2 of 7)

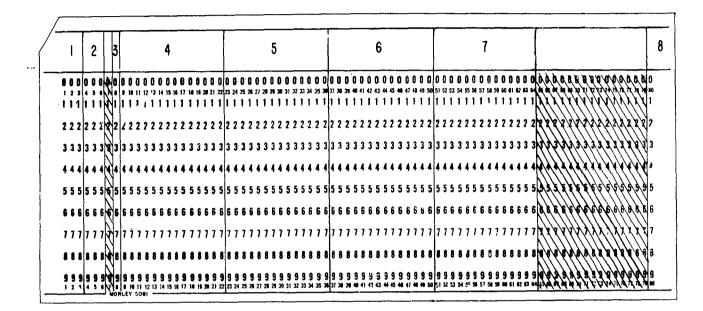
# 3-1-3

Elements

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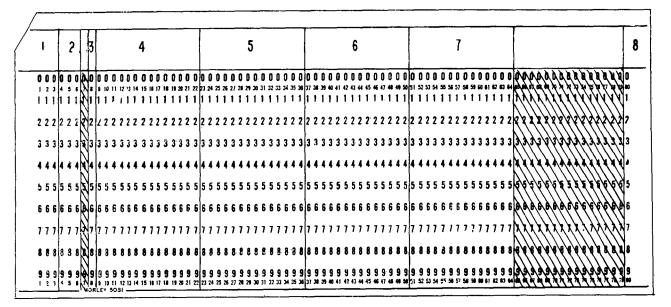


Field	Columns	7-Card Element Set	6-Card Element Set
1	1-3	Satellite Number, Right Justified	Same
2	4-6	Element Set Number, Right Justified	Same
3	8	Card Number = $3$	Same
4	9-22	Anomalistic Period at Epoch, (P <sub>a</sub> ), in Days/ Rev.	Nodal Period at Epoch, (P <sub>n</sub> )
5	23-36	Right Ascension of Ascending Node at $T_0^{,}$ ( $\Omega_0^{,}$ ), in Degrees	Same
6	37-50	Argument of Perigee at $T_o$ , ( $\omega_o$ ), in Degrees	Same
7	51-64	Perigee Distance, (Q <sub>0</sub> ) in Earth Radii	Same
8	80	Card Type = E	Blank

Element Card (3 of 7)



# Element Card (4 of 7)

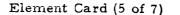


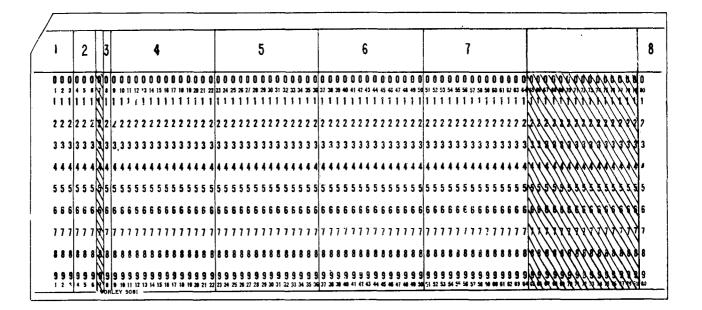
Field	Columns	7-Card Element Set	6-Card Element Set
1	1-3	Satellite Number, Right Justified	Same
2	4-6	Element Set Number, Right Justified	Same
3	8	Card Number = 4	Same
4	9-22	Rate of Change of P <sub>a</sub> , (c <sub>a</sub> ), in Days/(Rev.) <sup>2</sup> , <u>+</u> .DDDDDDDDE-XX	Rate of Change of $P_{n}$ , $(c_{n})$ , + D. DDDDDDDDE-XX <sup>n</sup>
5	23-36	First Derivative with Respect to Time of Right Ascension of Ascending Node, $(\Omega_0)$ , in Deg/ Day	Same
6	37-50	First Derivative with Respect to Time of the Argument of Perigee, $(\dot{\omega}_{0})$ , in Deg/ Day	Same
7	51-64	Not Used	First Derivative with Respect to Time of Perigee Distance $(Q_0), \pm D. DDDDDDDE-XX$
8	80	Card Type = E	Blank

Element Card (4 of 7)

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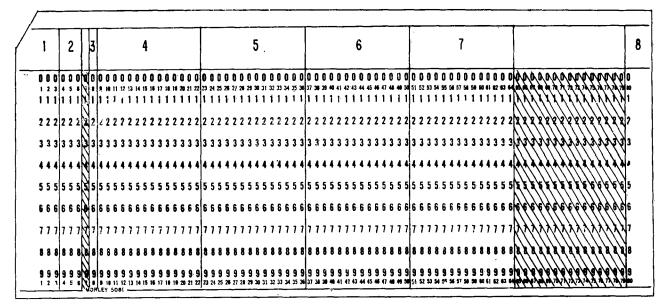
Field	Columns	7-Card Element Set	6-Card Element Set
1	1-3	Satellite Number, Right Justified	Same
2	4-6	Element Set Number, Right Justified	Same
3	8	Card Number = 5	Same
4	9-22	Rate of Change of c <sub>a</sub> , (d <sub>a</sub> ), in Days/(Rev) <sup>3</sup> , <u>+</u> D. DDDDDDDDE-XX	Second Derivative with Respect to Time of the Rate of Change of $P_n$ , $\dot{c}_n$ ), <u>+</u> D. DDDDDDDE-XX
5	<b>23-</b> 36	Not Used	One-Half of Second Derivative with Respect to Time of RA, $(1/2 RA_0)$ , <u>+</u> D. DDDDDDD <b>E</b> -XX
6	37-50	Not Used	One-Half of Second Derivative with Respect to Time of $\omega$ , $(1/2 \ \omega_{0})$ , <u>+</u> D. DDDDDDDE - XX
7	51-64	Not Used	One-Half of Second Derivative with Respect to Time of $Q_0$ , $(1/2 \ddot{Q}_0)$ , <u>+</u> D. DDDDDDDE-XX
8	80	Card Type = E	Blank

Element Card (5 of 7)

## 3-1-6

Elements

Element Card (6 of 7)

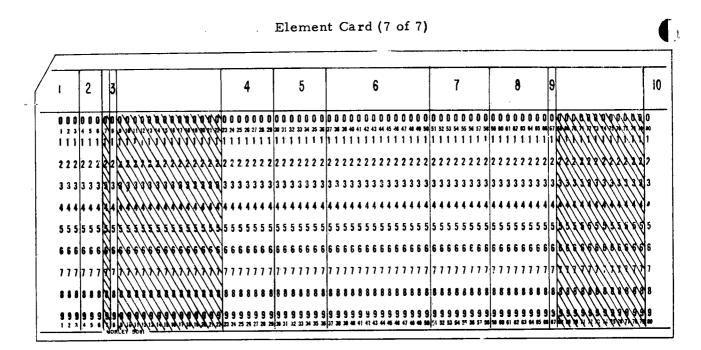




Field	Columns	7-Card Element Set	6-Card Element Set
1	1-3	Satellite Number, Right Justified	Same
2	4-6	Element Set Number, Right Justified	Same
3	8	Card Number = $6$	Same
4	9-22	Semi-Major Axis, (a), in Earth Radii	Limit of $\Delta T_n$
5	23-36	Nodal Period at Epoch, (P <sub>n</sub> ), in Days/Rev.	Limit of $\triangle RA_n$ in Columns 23-32 Limit of $\triangle H$ in Columns 33-36
6	37-50	Rate of Change of P <sub>n</sub> , (c <sub>n</sub> ) in Days /(Rev.) <sup>2</sup> <u>+</u> .DDDDDDDDE-XX	Standard Brightness
7	51-64	Not Used	Phase Angle Coefficient
8	80	Card Type = E	Blank

Element Card (6 of 7)

Elements



Field	Columns	7-Card Element Set
1	1-3	Satellite Number, Right Justified
2	4-6	Element Set Number, Right Justified
3	8	Card Number = 7
4	23-29	Initial Revolution Number, (N <sub>i</sub> ), Right Justified
5	30-36	Final Revolution Number, $(N_f)$ , Right Justified
6	37-50	Bulletin Expiration Time in YMMDDHHMMSS.SS, Decimal Point in Column 48
7	51-58	Root Mean Square of Error, in Kilometers, (RMS)
8	59-66	Number of Observations Used to Determine RMS
9	67	i-Stop { Blank = Correct i i = Do not correct i
10	80	Card Type = E

# Element Card (7 of 7)

#### 3-1-8

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Sensor

Station or Sensor Card

I	±	00	2 D I	D		ŧp	D	3 D 0	) D	0 0	D	D	4 0 0	D	0		6	;		7		8	3						9	)						10		1	2					13	j										•	
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5555	5 :	55	5 !	5	5	55	5	55	i 5	5 !	5 5	5 3	55	5	5 5	5	5	5 5	i 5	5	55	5	5	55	5	55	5.	55	5	55	5	55	5	55	i 5	5 !	5 5	5	5	55	5	5 (	55	5	5	5 5	5 5	5 5	5	Þ		Ņ	Ì	Ņ	5	5
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	Field No.	Columns	Description
	1	1 - 4	Sensor Number
	2	5-11	$\phi^{o} \left( \begin{array}{c} + N \\ - S \end{array} \right)$ - Latitude(Assumed Decimal Between Cols. 7-8)
	3	12-19	$\lambda^{\circ} \left( \begin{array}{c} + W \\ - E \end{array} \right)$ - Longitude (Assumed Decimal Between Cols. 15-16)
	4	20-25	H (meters) - Altitude (Assumed Decimal After Col. 25)
	5	26	Classification
	6	27-30	Sensor Type
	7	31 - 34	Previous Sensor Number
	8	35-36	Number Within Sensor Complex
	9	37-54	Name
i.	10	55-56	Equipment type
	11	57	Continent
	12	58-59	Country or State
	13	60-72	Comments
	14	79	Look Angle Coordinate Indicator
	15	80	Card Type (code type = S Punch 0, 2)

Station or Sensor Card

# 3-2-1

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FIELD 2	-	8	b b		9i 1: 91
SAT TYPE NG 12 3 4 8	STATION A NUMBER C 7 8 3 10 11	DATE DATE 1 12 0 14 15 16	STATIONA C DATE TIME(2) NUMBER C V M M O, D M M M G 515 3 3 6 7 8 3 10 11 12 0 14 18 16 17 18 19 20 21 22 23 24	(1) ELEUNTION AZINUTH AZINUTH (2) RANGE RATE AF K K K K K K K K K K K K K K K K K K	0055 0057 0057 00 00 00 00 00 00 00 00 00 00 00 00 00
FIELD	COLUMNS				
-	1-3	Sotellite n	Sotellite number, column 1 conto	contains a minus averpunch if this observation is classified. Ö ar Ö artoverpunches are not allowed.	
2	4-5	Equipment type	type		
£	6-9	Station numb <del>ar</del>	unber		
4	10	Accuracy	Accuracy of signal strength (See	(See Part 2 of Observation Card)	
ŝ	11-15	Date		•	4) <del>6</del> – 107
9	16-24	Time (Z)			
7	25-30	Elevation	or declination, depend	Elevation or declination, depending on punch in column 31	
8	31-37	Azimuth or	r right ascension, with	Azimuth or right ascension, with possible + or – minus overpunch in column 31. Minus overpunch means fields 7 and 8 are declination and right ascension.	
6	38-44	Slant range	Slont range, in kilometers		
0	45-53	Range rate,	Range rate, in kilometers per sec	second; or maximum frequency shift in cycles per second; depending on punch in column 54.	
11 2	3	Zero puncț	h or space means range	Zero punch or space means range rate in field 10; 1 punchmeans field 10 is moximum frequency shift.	
120	55-57	Brightness	Brightness at observation time	When colume 58 contrine a minus curret. Hen columes	
. 1 <b>2</b> 4	58-59	Maximum brigh <del>iness</del>	brightness	55-57 contain the radio cross section in meters	
. 12c	(9-09	Minimum brightness	prightness	59-63 contain the frequency, with decimal point between 60 and 61.	
<b>12</b> d	62-63	Time interval	Iov		
13	° 64-65	Date or line number	ne number		
*	69-99	Message number	umber		
15	R	Equinax (S	Equinox (See Part 3 of Observation Card)	vation (card)	
16	73-78	Observatio	Observation number, as assigned by ORCON	ed by ORCON	
12	62	Switch ind	Switch indicator used by manual system	udi system	
18	8	Card type.	. 0 means unknown; 1-	Card type. O means unknown; 1-9 coded according to the Association Status as determined in Report Association.	

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Observation

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3-3-1

#### COLUMN 10 (ACCURACY)

Either accuracy or signal strength may be indicated in column 10, coded according to the following:

If Type, in columns 4 and 5, is 31 or greater, column 10 contains signal strength. If Type is 30 or less, column 10 contains accuracy.

Code Figure	Accuracy	Signal
0	Normal observations made under fair conditions.	Signal strength good, reliable measurement.
1	Observations slightly under par due to outside interference (e.g. some clouds, reduced visibility).	Signal fair.
2	Observations only poor due to outside interference.	Signal weak, results poor.
3	Only estimates possible (malfunction of instrument, observation time span too short).	Signal questionable.
4	Doubtful observation, unable to verify either object or instrument behavior. Observations should be considered only as tentative.	

#### COLUMNS 55 - 63 (CROSS SECTION-FREQUENCY/ MAGNITUDE)

The block containing columns 55 through 63 is a dual purpose block where cross section and frequency, or magnitude and time interval are indicated. In order to specify cross section and frequency, a minus is used in column 58. No sign is used in column 58 when this block contains magnitude and time interval.

- A. Cross section, given in square meters, is listed in columns 55 through 57. To indicate less than one square meter cross section, use appropriate numbers and a minus in column 55 thus in effect putting a decimal point before column 55. For larger values where three digits would not be sufficient, use a plus in column 55 to represent ten times the indicated value (adding a zero to the value listed).
- B. Frequency in megacycles, is listed in columns 58 through 63 with the decimal point understood to be located between columns 60 and 61. In rare cases it might be desirable to increase the range of frequency given either side of the decimal point. To do this, use a minus in column 63 to move the point one place to the left, or a plus in column 63 to move the point one place to the right.

Observation Card (Part 2 of 3)

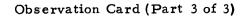
#### 3-3-2

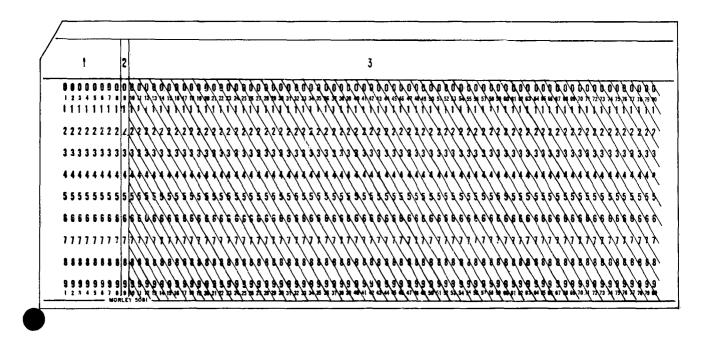
. . .

# COLUMN 70 (EQUINOX)

Column 70 contains year of Equinox as specified by the following:

0 = year of date 1 = 1900 2 = 1920 3 = 1950 4 = 1975 5 = 2000 6 = 1850 7 = 1855 8 = 1875 9 = 1960



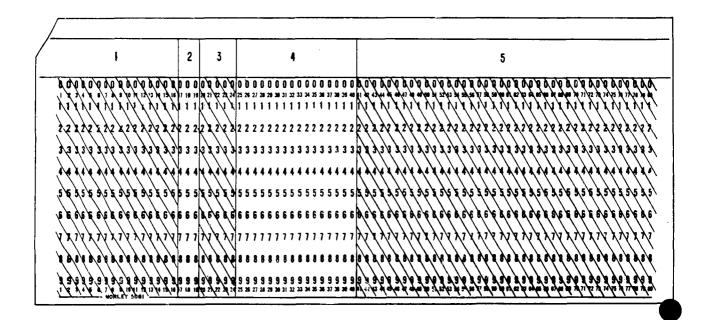


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First Card of Schedule Tape Input Deck

Schedule Tape ID Card

#### 3-4-1

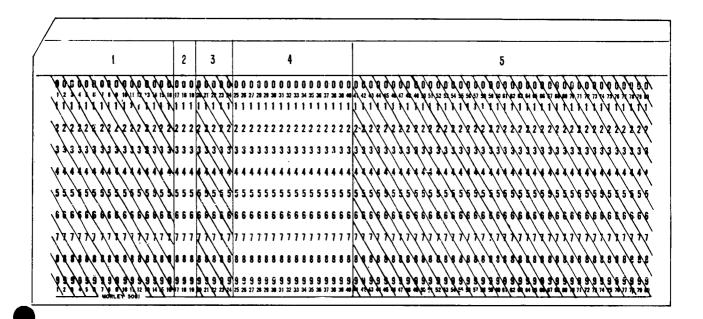


Field	Cols.	
1	1 - 16	Blank
2	17 - 19	Job
3	20 - 24	Blank
4	25 - 40	Comments
5	41 - 80	Blank

First Card of Each Job Deck

# JOB Card

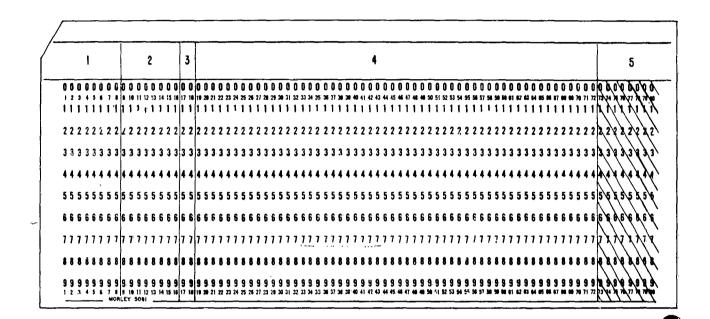
## 3-4-2



Field	Cols.	
1	<b>1 -</b> 16	Blank
2	<b>17 - 1</b> 9	REM
3	20-24	Blank
4	25 <b>-</b> 40	Comments
5	41 - 80	Blank

Optional Comment Card

REM Card

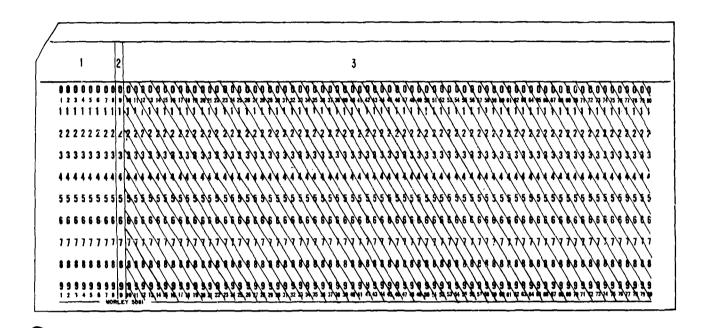


Field	Cols.	
1	<b>1</b> – 8	SPSJOB $\triangle \Delta$ , Left Justified
2	9 - 16	Program ID, Left Justified
3	17	Input Option } See Option Table
	18	Input Option Output Option } See Option Table
4	17 - 72	Comments
5	73 - 80	Blank

Used with Schedule Tape Mode Programs See Operating Summary Table for List of SPSJOB Cards

SPSJOB Card

END CASE

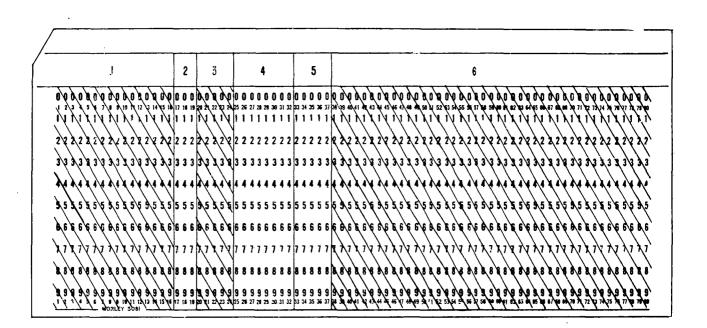


Field	Cols.	
1	1 - 8	$END\Delta CASE$
2	9	11, 8, 2 Punched
3	10 - 80	Blank

Used with Schedule Tape Mode Programs

END CASE Card





Field	Cols.	
i	1 - 16	Blank
2	17 - 19	RUN
3	20 - 24	Blank
4	25 - 32	Program ID, 8 Characters or Less - next field
		may follow directly
5	33 - 37	, DATA if Data Follows
6	38 - 80	Blank

Used to Run Manual Schedule Tape Mode Programs

See Operating Summary Table for List of RUN Cards

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2 2 2 2 7 2 7 2 7 2 7 2 7 2 7 2 7 2 7 2
1 × 1 × 1 × 1 × 1 × 1 × 1 × 1 × 1 × 1 ×
, , , , , , , , , , , , , , , , , , ,
e k ve k e k e k e k e k e k e k e k e k
4 fq \$14 \$14 \$14 \$14 \$14 \$14 \$14 \$14 \$14 \$14
5 6 6 5 x 5 x 5 x 5 x 5 x 5 x 5 x 5 x 5
e fe de la cé de de la cela cé de la cela ce la cela cela de la cela cela c
נא נא ער ער אר אר ארא גר גר גר גר ער ארא גר גר גר גר גר ארא אר אר אר אר גר גר גר אר ארא גר ארא גר אר
0,00,00,00,00,00,00,00,00,00,00,00,00,0
6

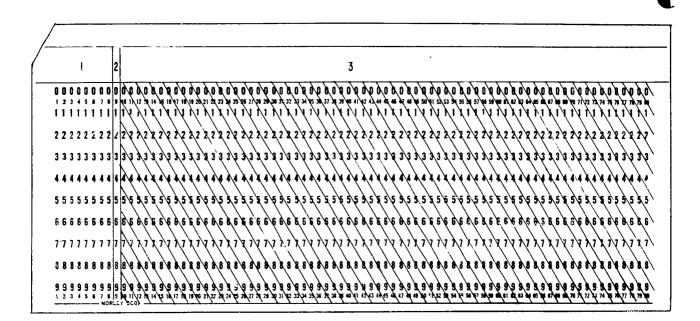


0018.	
1 - 16	Blank
17 - 23	ENDDATA
24 - 80	Blank
	17 - 23

~ 1

## Required if RUN Card Used

END DATA Card

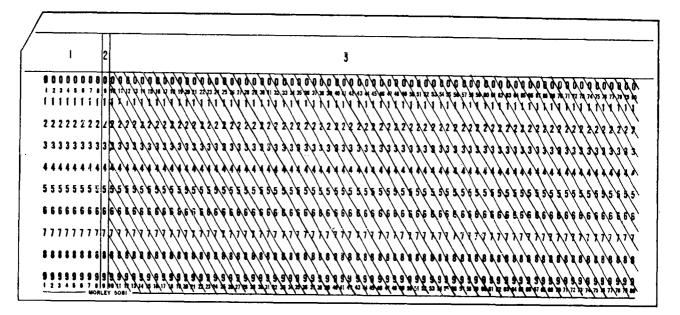


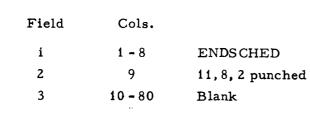
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Required to Terminate Each Job Deck

END OF JOB Card







Last Card of Schedule Tape Input Deck

Schedule Tape END Card

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RODTX, RODTX, RODTZ, ROMTZ, RPLLOD, RPLLOD, RPLLOD, RSU3A, RSU3A, RSU3A, RSUBA, RSUBU, RSUBU, RTEAN, RTETAN,	RTETYPE, RTESQ, RTP, RUDUT, RVDUT, RYN, SATNO, SATNO, SATNA, SATTG, SBLOC, SLOC, SLOC, SLOC,	SINI, SINEO, SINO, SINO, SINU, SINU, SINU, SSINU, SSTROB,	SSUBY, SSUBZ, STABZ, STABD, STARZ, STNM, STNM2, STOPGO, STPER, SUM,	SUM2, SYSE SYSEOT, SYSEOT, SYSEOK, SYSEOK,
ZZZZZZZZZZZZZZZZZZZZZZZZZZZZZZZZZZZZZZ	A A A A A A A A A A A A A A A A A A A	7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7 7	Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z Z	A S S C N A S S S N A S S S N A S S S S S S S S S S S S S S S S S S S

20 20 COEFF USED IN COMP C SUB DELTA N/N FOR RANGE RATE OES IN DC COEFF USED IN COMP C SUB DELTA N/N FOR RANGE RATE OBS IN DC COEFF USED IN COMP C SUB DELTA A SUB XN FOR RANGE RATE OBS IN DC COEFF USED IN COMP C SUB DELTA A SUB XN FOR RANGE RATE OBS IN DC COEFF USED IN COMP C SUB DELTA A SUB YN FOR RANGE RATE OBS IN DC COEFF USED IN COMP C SUB DELTA A SUB YN FOR RANGE RATE OBS IN DC U-UO AT EPOCH, U = MEAN ARG OF LAT MEAN ARG OF LAT AT EPOCH MEAN ARG OF LAT AT EPOCH MEAN ARG OF LAT AT EPOCH OCEFF USED IN COMP C SUB DELTA A/A FOR RANGE AND ANGLE OBS IN DC COEFF USED IN COMP C SUB DELTA A/A FOR RANGE AND ANGLE OBS IN DC COEFF USED IN COMP C SUB DELTA UN FOR RANGE AND ANGLE OBS IN DC COEFF USED IN COMP C SUB DELTA UN FOR RANGE AND ANGLE OBS IN DC COEFF USED IN COMP C SUB DELTA D/A FOR RANGE AND ANGLE OBS IN DC COEFF USED IN COMP C SUB DELTA D/A FOR RANGE AND ANGLE OBS IN DC COEFF USED IN COMP C SUB DELTA A SUB YN FOR RANGE AND ANGLE OBS IN DC COEFF USED IN COMP C SUB DELTA A SUB YN FOR RANGE AND ANGLE OBS IN DC COEFF USED IN COMP C SUB DELTA A SUB YN FOR RANGE AND ANGLE OBS IN DC COEFF USED IN COMP C SUB DELTA A SUB YN FOR RANGE AND ANGLE OBS IN DC COEFF USED IN COMP C SUB DELTA A SUB YN FOR RANGE AND ANGLE OBS IN DC COEFF USED IN COMP C SUB DELTA A SUB YN FOR RANGE AND ANGLE OBS IN DC COEFF USED IN COMP C SUB DELTA A SUB YN FOR RANGE AND ANGLE OBS IN DC Y COMP OF V BAR (UNIT VECTOR DIRECTED TO ODJ) RATE OF CHARGE OF TRUE ANOMALY Y COMP OF V BAR (UNIT VECTOR DIRECTED TO ODJ RATE OF CHARGE OF TRUE ANOMALY Y COMP OF V BAR (UNIT VECTOR NOR PLANE PERPENDICULAR TO UBAR) Y COMP OF W BAR (UNIT VECTOR IN ORB PLANE PERPENDICULAR TO UBAR) Y COMP OF W BAR (UNIT VECTOR PERPENDICULAR TO UBAR) Y COMP OF W BAR (UNIT VECTOR PERPENDICULAR TO UBBIT PLANE) Y COMP OF W BAR (UNIT VECTOR PERPENDICULAR TO UBBIT PLANE) Y COMP OF W BAR (UNIT VECTOR PERPENDICULAR TO UBBIT PLANE) Y COMP OF W BAR (UNIT VECTOR PERPENDICULAR TO UBBIT PLANE) Y COMP OF W BAR (UNIT VECTOR PERPENDICULAR TO UBBIT PLANE) Y COMP OF W BAR (UNIT VECTOR PLANE PERPENDICULAR TO N 2 SIGMA SUB 1 (USED IN DC AUTO REJ) CONSTANT •6667 CONSTANT 3 PI OVER 2 X COMP OF R BAR (VECTOR DIRECTED TO 05J) X COMP OF R BAR DUT AT EPOCH TAPE TABLE I/O ROUTINE TAPE TABLE TAPE TABLE TAPE TABLE TAPE ID CHECK ROUTINE TAPE ID CHECK ROUTINE TAPE TABLE SÉT-UP SWITCH THETA DOT GREENWICH SIDEREAL TIME AT THETA GREENWICH SUB ZERO THETA = SIUEREAL TIME T SUB N COMPUTING L FLOATING .07436574 TTY OUTPUT SWITCH LA TYPE EXECUTIVE SWITCH MEAN ARG. OF LAT. CM/EARTH RADII **INCLINATION** YEAR OF TO CONSTANT USED IN 0 SwI TCH SUB 2 PI TEMP EMP M/1155 M/17+145 M/555 M/555 M/55665 M/4335705 M/43335 M/35705 M/35705 M/34465 M/34775 M/37235 M/32235 M/55545 M/35503 M/33035 M/33045 M/35045 M/35055 M/36015 M/33075 M/33135 M/33135 M/3702\$ M/3642\$ M/3643\$ M/3331\$ M/3316\$ M/3667\$ M/3670\$ M/3670\$ M/3445\$ M/3555\$ M/37225 M/43345 M/36655 M/36655 M/3664\$ M/33345 M/3226\$ 2666\$ È TN, TOTOBER, ТАРСК, ТАРЕМАҮ, SYSNAME, SYSNO, SYSTAE, SYSTAL, XIMFS0, XISTSG, X20V3, X3PI02, XKE. XKERIM. XKMPER. TTYT, TTYTAP, ТЖОРІ, ТҮРЕ, ТҮРЕS«, KELSO, USUBA. USUBA. USUBN. USUBU. UTIME. TENM6, THUOT, THGR. UD0TC, UD0TN, UD0TU, UDTXN, XINCL. THTA. uZ, VUUT, VMAG, (DOT ; • 00Mr тоү, TEM, •NXU -х., UYN, •07 •۲۰

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<pre>AMAUTICAL MILES/HLURD (EARTH RADII/KEMIN) AMAUTICAL MILES/HLURD (EARTH RADII/KEMIN) 1+(1=+*2)**1/2, t = ECCENTRICITY FEAN FEAN LONG AT EPOCH X COMP OF L BAR SUB O (UGS UNIT VECT FROM USS TO OBU) 2 COMP OF L BAR SUB O (UGS UNIT VECT FROM USS TO OBU) 2 COMP OF L BAR SUB O (UGS UNIT VECT FROM USS TO OBU) 2 COMP OF L BAR SUB O (UGS UNIT VECT FROM USS TO OBU) 2 COMP OF L BAR SUB O (UGS UNIT VECT FROM USS TO OBU) 2 COMP OF L BAR SUB O (UGS UNIT VECT FROM USS TO OBU) 2 COMP OF L BAR SUB O (UGS UNIT VECT FROM USS TO OBU) 2 COMP OF L DAR, UNIT VECT FROM OBS TO OBU MEAN LONGINUS 2 COMP OF COMP L DAR, UNIT VECT FROM OBS TO OBU DES UNIT VLCT FROM OBS TO OBU MEAN LONGINUS 2 COMP OF COMP L DAR, UNIT VECT FROM OBS TO OBU MEAN LONGINUS -1:3333 FLOATING -1:5 FLOATING -1:5 F</pre>	Y COMP UF K BAR (VECTOR UIRECTED TO OBJ) Y CGMP OF R BAR DUT Y COMP OF R BAR DUT EXECUTIVE SEQUENCE SWITCH Z COMP OF R BAR (VECTOR JIRECTED TO OBJ) Z COMP OF R BAR (VECTOR JIRECTED TO OBJ) Z COMP OF R BAR DOT FLOATING POINT ZERO CURRENT PROBAR BAR (VECTOR DIRECTED TO OBJ) CURRENT PROBAR BAR (VECTOR DIRECTED TO OBJ) CURRENT PROBAR BAR (VECTOR FROM OBS TO 2ENITH) Y COMP OF Z BAR (UNIT VECTOR FROM OBS TO ZENITH) Z COMP OF Z BAR (UNIT VECTOR FROM OBS TO ZENITH) Z COMP OF Z BAR (UNIT VECTOR FROM OBS TO ZENITH) Z COMP OF Z BAR (UNIT VECTOR FROM OBS TO ZENITH)
MA MA MA MA MA MA MA MA MA MA MA MA MA M	M/440/L8 M/33325 M/33325 M/43445 M/43445 M/33335 M/434645 M/33335 M/334665 M/334665 M/34675 M/34675 M/34675 M/34675 M/34675
	YWRAPUP, YDOT, YDOT, YSELECT, YWRAPIU, Z, ZERO, ZERO, ZERO, ZSUBX, ZSUBX, ZSUBX, ZSUBZ, ZSUBZ, ZYEAR,
X X X X X X X X X X X X X X X X X X X	A A A A A A A A A A A A A A A A A A A

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#### **IOPS AND PROC**

#### ERROR PROCEDURES AND COMMENTS

Irrecoverable Errors:

- 1) IRRECOV ERROR
- 2) PROGRAM ERR PROC (See PROC ERRORS Table)
- 3) PROGRAM ERR IOPS (See IOPS ERRORS Table)

Irrecoverable errors comments are accompanied by a typeout of the contents of the A register and Q register:

A REG =

Q REG =

The program will then jump to location 3.\*

**Recoverable Errors:** 

	) PARSPROC ERR	Sprocket or parity error
2	2) S1, S2 ERR	Start or end block sentinal error
	3) RING TAPEXX	No write ring on unit XX
4	) LOCAL TAPEXX	Tape unit XX not available
5	) NO TAPE XX	Rewind issued – tape not available

Recoverable errors are accompanied by the STOP-GO option, wherein the operator may retry the job by typing GO or may type STOP to effect a jump to location 3.\*

\* Location 3 will generally cause the typeout of the Flexo comment: SUBROUTINE ERROR EXIT FROM OCTAL XXXXX, LOAD

COMPACT INTO UPPER CORE AND TAKE DUMP.

Following the message, a stop or go loop controls the console such that if STOP is typed the message will be repeated. Typing GO returns control to the executive program terminating the job. If possible, a dump should always be taken.

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# **IOPS ERRORS**

A REG	Q REG	MEANING
I/I	Unit TI5 or T23	Attempt to use a tape which has not been defined in the IOUNITS statement
2/1	Unit TI5 or T23	Attempt to read a tape described in the IOUNITS statement as "Write Only".
3/1	Unit TI5 or T23	Attempt to write on a tape described in the IOUNITS as "Read Only".
4/1		Attempt to read from a binary tape more words than there are in the record.
5/1		The checksum of the blocks after reading does not equal the checksum written with the block. This is applicable to binary tape only.
6/1		Attempt to store, for printing,more than 120 printable characters for a given line.
7/1		Attempt to read from a binary tape which is not in the correct format.
8/l		The program was run even though an error occurred at compilation time.
9/1		Probably machine error; or the program has destroyed some IOPS coding (2GNC).
10/1		More characters are requested from a card than were described in the record size parameter in the IOUNITS statement.

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11/1		Same as 9/1. (BIWTR)
12/1	Last character processed T5	Illegal character or illegal exponent field in floating to floating (E) input conversation.
3/	Last character processed T5	Illegal character or illegal exponent field in fixed decimal to floating (F) input conversion.
14/1	Last character processed T5	Illegal character in fixed decimal to integer (D) input conversion.
15/1	Last character processed T5	Illegal character or exponent field in integer to integer (I) input conversion.
16/1		Rewind order to a unit defined as paper tape.

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# PROC ERRORS

A REG =	Q REG =	Meaning
Zeros	Location of macro-instruc- tion in left address	List full
Location of macro-instruc- tian being checked in left address; 1/1T47 (D/1)	Normal exit of check macro-instruction in left address	Not in list
Location of macro-instruction being checked in left address; 1/1T46 (D/2)	Normal exit of check macro-instruction in left address	Checking non-magnetic tape order with magnetic tape check macro- instruction
Location of macro-instruction being checked in left address; 1/1T45 (D/4)	Normal exit of check macro-instruction in left address	Checking too much (B= NBP of original order.)
Location of macro-instruction being checked in left address; 1/1T43 (D/16)	Normal exit of check macro-instruction in left address	Checking out of sequence
Unit <sup>#</sup> T23; 1/1T41 (D/64)	Location of macro-instruc- tion being checked in left address	Write or read only

# SPSJOB OPTION TABLE

PSR:	
SPSJOB	PSR (I) (O)
(I) :	Col.17 0 = Parameter Cards
<b>(O)</b> :	Col.18 0 = Print (D.S. 1) and punch paper tape (D.S.4) 1 = Print only (D.S. 1)
RESPLT:	
SPSJOB	RESPLT (I) (O)
(1) :	Col.17 0 = Observation Cards 1 = Observation and Element Cards 2 = Observation and Element Number Cards 3 = Observation and Sensor Cards 4 = Observation, Element, and Sensor Cards 5 = Observation, Element Number, and Sensor Cards
<b>(O)</b>	Col.18 0 = Print (D.S.I) and punch cards (D.S.2)

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PROGRAM		INPUT		INTERIM		OUTPUT	INDEX
Name	Tape Setup	Input Deck (SPS)	Special Tapes	Tape Usoge	Form	Contents	
ASUM	- s 0 0 0	1) JOB 2) RUN ASUMB2 3) ENDOFJOB@	@s= seal	None	⊕ O= Print (DS 1)	Contents of Acquistion File	Description: Section 1 Program 12 Input/Output: Section 5 Program 12
BMEWSPT	©©©© ©©©© ©©©	1) JOB 2) RUN BMEWSPTB 3) ENDOFJOB@	Ans = DIP Tape	None	() = Input for ORCON () = Punched () (Write AII)	TTY∆IN Tape Containing Standard Observation Data	Description: Section 1 Program 16 Input/Output: Section 5 Program 16
CCOE*	© 0 0 − 0 0 0 − 0 0 0 − 0	<ol> <li>JOB</li> <li>RUN CCOEB2, DATA</li> <li>RUN CCOEB2, DATA</li> <li>Standard Element Set</li> <li>Request Card(s)</li> <li>Blank Card</li>     &lt;</ol>	oue Z	() X = Scratch Used for Data Input	(D = 0) Print (DS 1)	Position and Velocity Components of a Sat. at Specified Intervals	Description: Section 1 Program 5 Input/Output Section 5 Program 5

31 Ď, 1 = Input, O = Output, S = Special, X = Scratch, (N)= Write King, <u>LN</u> = \* = IELTYP Program May Follow, [= Repeat if Necessary, () = Optional

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Name Tape Setup ISUM () X ISUM () X ISUM () X X ISUM () X X X X X X X X X X X X X X X X X X X	Input Deck (SPS)	Second	Tame I have			
		Tapes	after adei	Form	Contents	
	<ol> <li>JOB</li> <li>RUN ISUMB2, DATA</li> <li>RUN ISUMB2, DATA</li> <li>(ALL) Control Cord</li> <li>(BOX) Control Cord</li> <li>(NNN) (Sat. No., Used in Absence of ALL Card)</li> <li>END Control Card</li> <li>ENDDATA</li> <li>ENDDATA</li> </ol>	ی SEAI SEAI	©X = Scratch Used for Data Input	0 = 0 = Print Print (DS 1)	Contents of Information File with Optional Box Score	Description: Section 1 Program 14 Input/Output: Section 5 Program 14
	<ol> <li>JOB</li> <li>RUN LOCVECB2, DATA</li> <li>Request Card(s) (500 mox.)</li> <li>7 Punch in Col.8</li> <li>Element Set(s) (500 mox.)</li> <li>Blank Card</li> <li>ENDDATA</li> <li>ENDDATA</li> </ol>	Nane	© X = Scratch Used For Data Input	() O= Print (DS 1) (DS 1) Punch 5-jevel Paper Tape (DS 4)	Vector Coordinates for Lockheed	Description: Section 1 Program 4 Input/Output: Section 5 Program 4
MAKETAPP* () 0	<ol> <li>JOB</li> <li>RUN MAKETAPE, DATA</li> <li>RUN MAKETAPE, DATA</li> <li>(TELEFORM) Control Card</li> <li>Cards to be Converted</li> <li>Cards to be Converted</li> <li>FINDATA Control Card</li> <li>ENDDATA</li> <li>ENDDATA</li> </ol>	a Z	@X = Scratch Used for Data Input	(DS 1)	Alphanumeric Card Data Transferred to Tape, with Sentinals Required by TELTYP	Description Section 1 Program 8 Input/Output: Section 5 Program 8
SYMBOL KEY: 1= Input, 0= Output, 5= Soecial, X= Scratch, (N=Write Ring, A * = TELTYP Program May Follow, [= Repeat if Necessary, ()= Optional	it, O=Output, S≠Special, X=Scratch, (M=Write Ring, M =No Write Ring,	ite Ring.	=No Write Ring	@ =11.8.2 Punch	Punch	

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PROGRAM		INPUT		INTERIM	б	OUTPUT	INDEX
Name	Tape Setup	Input Deck (SPS)	Special Tapes	Tape Usage	Form	Contents	
ORPS		<ol> <li>JOB</li> <li>JUN ORPSB2, DATA</li> <li>RUN ORPSB2, DATA</li> <li>Element Set</li> <li>Startion Card</li> <li>Request Card</li> <li>Blank Card</li> <li>Blank Card</li> <li>ENDDATA</li> <li>ENDDATA</li> </ol>	None	© X = Scratch Used for Data Input	(D) 0 = Print (DS 1)	Search Point Data for Satellite Sighting	Description: Section 1 Program 11 Input/Output: Section 1 Program 11
POSE*	© os − s ×	<ol> <li>JOB</li> <li>IN POSEB2, DATA</li> <li>RUN POSEB2, DATA</li> <li>Request Card</li> <li>Observation Card</li> <li>Station Card</li> <li>Option Card</li> <li>Option Card</li> <li>ENDDATA</li> <li>ENDDATA</li> </ol>	None	<ul> <li>x =</li> <li>Scratch</li> <li>Used for</li> <li>Data Input</li> </ul>	(j) O = Print (DS 1)	Search Ephemeris of a Point in Space for a given Station	Description: Section 1 Program 10 Input/Output: Section 5 Program 10
PRE PINT*	©-©©_ © 2 = 2 × ×	<ol> <li>JOB</li> <li>JUN PREPINTB, DATA</li> <li>RUN PREPINTB, DATA</li> <li>Request Cards (500 max.)</li> <li>2 Punch in Col.7</li> <li>Time Cards</li> <li>7 Punch in Col.8</li> <li>7 Punch in Col.8</li> <li>8 Blank Card</li> <li>8 Blank Card</li> <li>9 ENDDATA</li> <li>10 ENDOFJOB (2)</li> </ol>	ene Z	() X = Scratch Used for Data Input	∰int Print (DS 1)	Satellite Situation Report from Nodal Elements	Description: Section 1 Program 7 Input/Output: Section 5 Program 7

SYMBOL KEY: I = Input, O = Oųtput, S = Special, X = Scratch, (M) = Write Ring, (M) = No Write Ring, @ =11,8,2 Punch \* = TELTYP Program May Follow, [= Repeat if Necessary, () = Optional

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SPS Operating Summary

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×			и - с и и - с и - с и и	in centre 1 o centre 1
INDEX		Description: Section 1 Program 1 Input/Output: Section 5 Program 1	Description: Section 1 Program 2 Input/Output: Section 5 Program 2	Description: Section 1 Program 6 Input/Output: Section 5 Program 6
OUTPUT	Contents	Situation Reports (Position or Satellite), for all Satellites at a given time	Residuals. Observations Reduced to Last Nodal Crossing	Residuals of At and Vector Mognitude Versus the Revolution, Punched Cards to be Used on EAI Data Plotter
O	Form	<ul> <li>(1) O =</li> <li>Print</li> <li>(DS 1)</li> <li>(11 O =</li> <li>Punch 5-level</li> <li>Paper Tape</li> <li>(DS 4))</li> </ul>	()) 0 = Print (DS 1)	<ul> <li>(1) O =</li> <li>Print</li> <li>(DS 1)</li> <li>(1) O =</li> <li>Punch</li> <li>Cards</li> <li>(DS 2)</li> </ul>
INTERIM	Tape Usage	eve Z	<ul> <li>X =</li> <li>Scratch</li> <li>Used for</li> <li>Dato Input</li> <li>Bato Input</li> <li>Scratch</li> <li>Untagged</li> <li>Obs.</li> </ul>	е Z
1	Special Tapes	دی) s = seal	As = . Station Tape (Optional)	€ SEAI SEAI
INPUT	Input Deck (SPS)	<ol> <li>JOB</li> <li>REM</li> <li>SPSJOB PSR (I) (O)</li> <li>Parameter Card (Time Request)</li> <li>(Parameter Card (s))</li> <li>ENDΔCASE @</li> <li>ENDOFJOB @</li> </ol>	<ol> <li>JOB</li> <li>REDUCTB2, DATA</li> <li>RUN REDUCTB2, DATA</li> <li>Option Switch Card</li> <li>Punch in Col.8</li> <li>Flement Sets (250 max.)</li> <li>Punch in Col.8</li> <li>Runch in Col.8, Tolerance Limit in Col.7</li> <li>Banch in Col.8, Tolerance Limit in Col.7</li> <li>Doservation Cards (no max.)</li> <li>Punct in Col.79</li> <li>Runch in Col.79</li> <li>Rundrata</li> <li>RUDATA</li> <li>RUDDATA</li> </ol>	<ol> <li>JOB</li> <li>REM</li> <li>S REM</li> <li>SPSJOB RESPLT (I)(O)</li> <li>4) (Element Cards, if Input Options 1,2,4, or 5)</li> <li>(Sensor Cards, if Input Options 3,4, or 5)</li> <li>(Sensor Cards</li> <li>Observation Cards</li> <li>DACASE @</li> <li>ENDOFJOB @</li> </ol>
	Tape Šetup	@  0 0 0	ୁଇଁହିଛିଚିତି କାହା	-@@= 0 2 - 2
PROGRAM	Name	Š	REDUCT*	RESPLT

5 . 3) ğ \* TELTYP Program May Follow, [= Repeat if Necessary, () = Optional

# SPS Operating Summary

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INTERIM OUTPUT INDEX Tape Usage Form Contents	(D) X =     (D) O=     Predicted     Description:       Scratch     Print     Look-Angles     Section 1       Used for     (DS 1)     tor Sotellite     Program 9       Data Input     (D) O=     at Specified     Input/Output:       (D) C=     times, for a     Section 5       (D) C=     times, for a     Section 5       (D) C=     times, for a     Section 5       (D) C=     times, for a     Program 9       Used for     (DS 2))     tertion     Program 9       Ordes     branch     given     Section 5       (D) C=     tertion     branch     Bara for       (D) X =     Ephemeris     branch     Program 9       Data Storage     Data Storage     Section 5	
Special Tapes	As a first term of the second	
Input Deck (SPS)	<ol> <li>JOB</li> <li>RUN XYZLARB2, DATA</li> <li>Station Card</li> <li>Station Card</li> <li>Request Card</li> <li>(Identification Card)</li> <li>(Ephemeris Cards)</li> <li>(Request Card)</li> <li>(Request Card)</li></ol>	
Tape Setup	\$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$ \$	
Name	XYZLAR	

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SPS Operating Summary .

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	Special Tape Usage Form Contents Tapes	OPTION II     Description:       JOB     JOB       JOB     RUN ROCB2, DATA       None	(4) S =     None     (1) O =     Contents of     Program 13       SEAI     Print     Sensor File     Input/Output:       SEAI     (DS 1)     Program 13	(1)     I =     (3)     X =     (1)     O =     Teletype     Description:       Output     Scratch     Punch     Conversion     Scratch     Punch     Section     Section       Tape     Storage for     5-level     of Contents     Program 15     Program 15       from     Messages     Paper Tape     of (1)     1     Section 5       Programs     Transfer to     (D54)     of (1)     1     Section 5       Programs     Transfer to     (D54)     of (1)     1     Section 5
TUPUI	Tape Input Deck (SPS) Setup		1 5 1) JOB (1) 2) RUN SSUMB2 (1) 0 (1) 0	1 5 1 2 RUN TELTYP () 108 () 2 RUN TELTYP () 10 () 10 () 10 () 10 () 10 () 10 () 10 () 108 () 108
PROGRAM		, to Q	SSUM	TELTYP

SYMBOL KEY: 1 = Input, 0 = Output, 5 = Special, X = Scratch, () = Write Ring, () = No Write Ring, @ =11,8,2 Punch \* = TELTYP Program May Follow, [= Repeat if Necessary, () = Optional

SPS Operating Summary

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#### SYSTEM OPERATING MODES

The definition of terms as used in the descriptions included in the Input-Output section under Input Deck are as follows:

#### A. Schedule Tape Mode

This term implies that the program is completely SPS system oriented. The assign deck listed in the Assign Table has been used to bind the program to the system. An SPSJOB card must be used to call the program from the system library tape. All elements, observations, sensors, as well as the parameter cards are brought in by the executive routine after consulting the input-output options contained on the SPSJOB card and an internal Job Table. Only system supplied routines are used in the execution of input-output requirements. Upon completion the program will exit to the system location called EXEND. Toggle 24 must be on and 1-23 off to run under this mode.

#### B. Manual Schedule Tape Mode

The programs of this type require a RUN card to call the program from the system library tape. If data follows, it will be moved to logical tape zero (0) by the executive program. The program will read data from this tape.

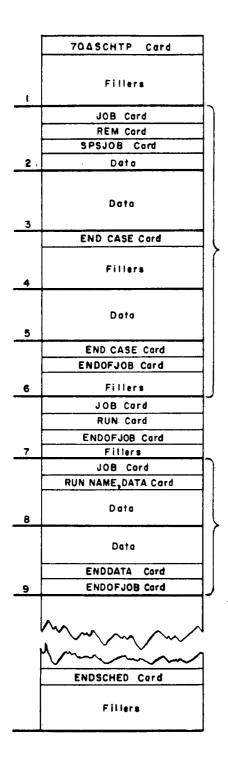
PROC and IOPS, which are supplied at compilation time by the ALTAC compiler, are, with few exceptions, the input-output routines used. These input-output routines have been modified and are interruptable by the real-time devices available. Variable formatting of input-output requests, as used in FORTRAN and ALTAC, as well as Macro capabilities, have been maintained as a result of the modifications made.

The system assign deck is not used and the program contains the minimum amount of information required to bind it to the system. Upon completion of the program a return is made to the location called MANEXIT.

Toggle 24 must be on and 1-23 off to run under this mode.

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#### SCHEDULE TAPE FORMAT



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Each new job deck starts at the beginning of a block, to simplify interrupt.

Blocks 2-6 show the layout of an SPSJOB run in Schedule Tape Mode.

Block 7 shows a job run under the Manual Schedule Tape Mode, without data.

Blocks 8-9 show a job run under the Manual Schedule Tape Mode, with data.



#### 5-0-2

(Postion S	PSR ituation Report (SITRPT) )
INPUT DECK: S	chedule Tape Mode SPSJOB card required with the following options available:
<b>.</b>	<pre>a) Input options (col. 17) 0 = parameter cards</pre>
	<ul> <li>b) Output option (col. 18)</li> <li>0 = teletype output required</li> <li>1 = inhibit teletype output</li> </ul>
INPUT DATA:	<ol> <li>SEAI tape</li> <li>First parameter card is a request card (P in col. 80). See layout.</li> </ol>
	3) Up to eleven parameter cards (P in col. 80) which will be treated as comment cards for Part I of Satellite Situation Report.
OUTPUT:	Logical tape 11 contains:
	1) Position Situation Report if option zero requested.
	2) Satellite Situation Report if option one requested.
	3) Both if option two requested.
	4) Teletype output if indicated on SPSJOB card.

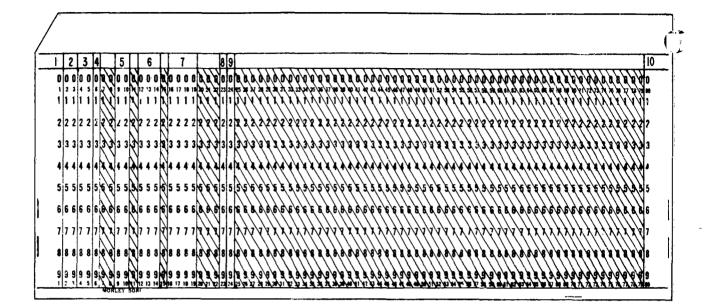
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Field	Cols.	
1	1	Output Option
		<ul> <li>0 Position situation Report</li> <li>1 Satellite Situation Report</li> <li>2 Both Reports</li> </ul>
2	2-3	Hour at Which Report Supplies Position
3	4-5	Minutes at Which Report Supplies Position
4	6	Z Punch
5	9-10	Day of Month at Which Report Supplies Position
6	12-14	First Three Letters of Month
7	16-19	First Three Letters of Month
8	23	Satellite Situation Report Output Code:
		Blank or 0 Suppress Print of Debris 1 Print All Satellites
9	24	Position Situation Report Output Code:
		Blank or 0 Perigee and Apogee in statute miles.
		1 Perigee and Apogee in Kilometers
10	80	P Punch

## Request Card

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# POSITION SITUATION REPORT

PAGE 1

HE FOLLDWING INFORMATION COMPUTED FOR 14302 02 OCT 1962

						THE FOL	OLLOWING	INFORMA	INFORMATION COMPUTED	UTED FOR	1430Z	02 OCT 19	1962			
80	띡		ם ש	I	LAT	NO NO	INCL	PERIOD	APOGEE	PERIGEE	REV	i sub N	RA SUB N	L SUB N	ECC	CLASS
Ż	ANN		Z	0	53	DĒC	DEG	NIH	X	X	02	DAYS	فيتا	DEG		
	LPHA	_	~	46	55	136	33.20		. 9</th <th>•</th> <th>19640</th> <th>475,549</th> <th>98.</th> <th>0./</th> <th>e D</th> <th></th>	•	19640	475,549	98.	0./	e D	
	ETA		Ŧ	03	23	267	34.24		3045.0	652.9	17845	275.558	1.79.1	3 <b>2.</b> 5	.190	
	LPHA		41	75	24	329	32,87		291.	-	15191	275,592	15.	8.1	<b>°</b>	
	LPHA		÷1	36	40	166	32.91		• 2 4 •	•	14112	275,539	<b>6</b> 9	56.6	cn ا	
	ETA		-	<b>1</b>	11	170	34.27		000	രി	12439	279,531	• •	87.1	20	
	<b>4 1</b>		<b>41</b>	5	40	119	33.02		711.	n i	12306	275.542	÷	89 ° • • • •	ഹ	
	014		*1	25	<b>9</b>	58	50.30		٠	2	15445	¢75.556	87.	283./	~	
	014		<b>₩</b> .	33		170	5n . 4		063.	n.	15468	. 75.578	71.	307.1	2	
	LPHA				H	-	CENTRIC		1	,	1	1				
	ETA		*1		5	•	48.57		\$7.		32	5, 57	÷		•00•	
	ETA		*1		23	147	48.37		52.	4	35	5.59	50.	67.	¥00.	
	AMMA		#ł		8	28	51.29		24.	÷	36	5.55	20.	51.	.020	
	PSUN		-	26	•62	404	64.94		300.0	240.4	13600	ذ؟ذ، د7۲	152,9	57.5	•00•	
	PSLN		÷1		•	331	64.79		02.	4	4	5.59	35.		.016	
	PSLN					POSIT	ION UNC									
	ETA		-		<b>1</b> 8	1.03	33.04		. 66.	483.3	13147	٠ <b>،</b>	53.	53.3	.001	
	T A				11	253	69°99		0 > 3.	613.5	11792	5	40.	72,3	.031	
	T A				46	294	60°9		052°	612.9	11794	<u>ہ</u>	40.	79.3	.030	
	T.A				<b>₽</b>	76	66.05		036.	610.3	11820		27.	97. Q	.030	
	014		-	80	4.4	240	4/,20		528.	1482.8	9636	ຄ		96.	.003	
	0TA			76	27	206	47.45		<b>.</b> 89.	1500.1	9535	<u>ب</u>	54.	30.	.012	
	OTA			70	19	208	47.22		1692.5	1512.5	9521	275,598	2,2	224.0	.011	
	07.4				47	129	47.26		689.	1530.7	9507			.80	.010	
	ņ				ň	27	28.33		220.	962.8	9809	·0	27.	5.	.017	
	P				13	85	28,30		<b>1</b> 8/.	963.2	9844			28.	.015	
	**				<b>1</b> 3	4	49.91		240.	433.5	8940	r.	76.	49.7	.117	
	H				12	252	40.04		218.	424.9	8955	0	69.	45.7	.116	
	ы				47	260	48.50		735.7	615.9	9946	, n	16.	<b>*</b>	• 0 <b>0</b> 8	
	н				30	229	48.51		728.	613.1	9955	<u>б</u>	10.	e	• 0 0 8	
	H				50	212	49.39		2063.2	409.6	5 <b>4</b> 0 6	·••	٠	1 <b>14</b> ° P	011.	
	LPHA				18	38	97,37		539.6	470.4	9237	n.		60.	.005	
	<b>H</b> 1				-	4	48.52		128.7	616.4	1566	••••			.008	
	H.					000				0.120	1946	<b>.</b>				
						- ē			20100	**/0*	1-74	•	1.0	, ,	• 0.05	
				-	4				417.	55.	7235	5.58	5	4	104	
						225	39.05		-		7211		62		100	
	PSLN		183	10	14	101	80.81		5.743	197.3	9216	275.602	125.0	ć.102	000	
61 0	DELTA	4	ن <b>∀</b> €			POSIT	ION UNCE	ERTAIN								
	APPA		860			<u> </u>	ON CNC									
	AMMA		600	•	<b>-5</b> 2	168	51,26		752.9	178.5	13395	275.555	152.0	58.7	.018	
	ETA		101	•	00	12	48.50		2		64471	520.672	20	~	0	
	•		105	-	•	187	50.48		23	417.6	6165	275.558	79.	•	-	
	5		107	•		338	28.02		4	4	6984	275,588	59.	2	<b>ع</b> ۵	
	<u>⊃</u> .		112		Fi	~ ~	<b>CENTRIC</b>									
	2			1	H u	MELIU	CENTRIC				4 7 1	t L				
		Ŧ	6112	•	•	7.7	67°25		/•278	7 · • > 0	6 <b>9</b> 701	0+6*6/7	C*411		900.	

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# REDUCT (Nodal Crossing Reduction)

INPUT DECK: Manual Schedule Tape Mode

INPUT DATA:

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The first card of the input data is an option switch card controlling the processing and the output. This card is followed by the standard element sets, station cards (if any), and observation cards, with each group preceded by a lead card.

The deck should be in the following order:

- 1. Option Switch card:
  - Col. 1: a numeric punch (1-9) will cause the Station Tape(logical 7) to be inhibited.
  - Col. 3: a numeric punch, (1) inhibits the interim tape (logical 8), (2) eliminates tolerance limit tests.

Col. 8: a numeric punch causes all observations

- to be treated as unknowns, reducing all against the available element sets.
- Col. 9: a numeric punch modifies headings, spacings, and format length to allow subsequent processing by TELTYP to obtain output for teletype transmission.
- Col.10: a numeric punch eliminates a re-reduction of unknown obs. which have remained untagged after a reduction against the selected tolerance limits. If the rereduction is not eliminated, the untagged obs., which have been written out on the interim tape, are tested against open tolerance limits ( $\Delta t= 99.0$  days,  $\Delta RA=$ 15.0 degrees,  $\Delta h = 99999.0$  kilometers).

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- Col.11: a numeric punch directs the program to use the perigee distance (q term) of the element cards. If this term is not used, perigee distance will be computed by the formula: q = a (1-e).
- 2. Element lead card: a 7 punch in col. 8.
- 3. Standard 6 or 7 card element sets (up to 250).
- 4. Station lead card: a 9 punch in col. 8.
- 5. Standard station cards (up to 750).\*
- 6. Observation lead card: and 8 punch in col. 8 with the tolerance limit key punched in col. 7:<sup>\*\*</sup>

	Tolerance Table									
Col. 7	∆t, in days	$\triangle$ RA, in degs.	$\Delta H$ , in kms.							
0 or blank	.002	20.0	200							
1	1.0	5.0	500							
2	.003	360.0	10000							
3	.002	360.0	10000							
4	.001	360.0	10000							
5	.003	5.0	300							
6	.05	2.5	200							

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- 7. Observation cards (any number).
- 8. A blank card to terminate the preceding group. group of observation cards.
- 9. End of input card: a 9 punch in col. 79.
- \* The Station Tape (logical 7) may be used in place of, or in addition to, the card groups 4 and 5.
- \*\* Observation lead cards (any number) may be inserted to vary the tolerance limits applied to the reduction of the following observation cards.

# OUTPUT:

Output begins with the satellite inventory. Seven groups, composed of satellite number and element number, are printed per line.

The next output is the satellite number (first line), selected elements (two lines), and the tolerance limits used in the observation output following (one line).

Observation output will occur in one of three formats, depending on which part of the program has done the reduction, as in the following table:

General Reduction Output:

ID = 16 digit observation identification number comprised of the satellite number, the last digit of year of observation, month day, hour, minute, seconds, and hundreths of seconds.
N = Epoch revolution number
U = Argument of latitude of satellite in orbit plane
T SUB N = Computed time of nodal crossing
DELTA T = Time residual in days (difference between predicted and observed)
PHI S = Computed latitude of subsatellite point
L S = Longitude of subsatellite point
RA N = Right ascension of ascending node in degrees
DEL RA = Right and observed)
H(KM) = Height in kilometers

DEL H = Height residual (difference between predicted and observed)

TYPE = VIS for visual observations

RDR for radar observations

B-N for Baker-Nunn observations

- ELEM = Element number
  - STA = Station number

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Doppler Reduction:

ID = Observation identification

N = Epoch revolution number

T SUB N = Computed time of nodal crossing

DELTA T = Time residual in days (difference between predicted and observed)

- D = Arc distance in nautical miles from station to subsatellite point.
- H = Elevation, in degrees
- S = Slant range, in kilometers
- ELEM = Element number
  - STA = Station number

### Direction Finder Reduction:

ID = Observation identification N = Epoch revolution number T SUB N = Computed time of nodal crossing DELTA T = Time residual in days (difference between predicted and observed) H = Elevation, in degrees S = Slant range, in kilometers H(KM) = Object height, in kilometers ELEM = Element number STA = Station number

If the output is to be processed by TELTYP for teletype transmission, a modified version of the above output will be printed. In general, each line, as described above, will be broken into two lines of output. Each line will be followed by an asterisk (\* ), and the abreviation, DOP, for doppler observations, or DF, for direct finder observations will follow the station number.

		EDUCTION TEST				
	RUN R	EDUCTB2,DATA				
1 11 7		TOGGL	E CARD			
51 19 160 IOTA 51 19 21961	-	2300• 10885429	0.01095429 1	47•199999 48•10885429		E E
51 19 3 51 19 4 51 19 5-0.		-3.08994973	211.88786888 2.97470850- 1259364E-06-		32	Ë E E
51 19 6	0.0	B2U5393 5.	0345715E-009			E
60 19 1 60 XI 1 60 19 21961		1390• •94619608	0•11134295 1	49.975999 50.94619608		E
			355.09534836	1.076729		E
60 19 4-1.86070		-3.35409367	2.78489476-			E
60 19 5-0. 60 19 6	-2•	1083826E-05 1	•7505843E-05-	-0•		E
73 6 161 BETA		200.	0.00638922	64.886558	53	-
73 6 21961 73 6 3 0.00		•55804205 19 43•76211357	61 60•99729776	47.55804205 1.035650		
73 6 4-7.4575		-3.65672696	-0.42809739-		09	
73 6 5 0.			.8967934E-04-	-0•		
73 6 6 -0. 77 4 11961 GAN		0• -0 90•	-0• 0•00817470	-0. 65.010977	75	0
77 4 21961		•61976957 19		48.61976957		
			20.37677312	1.029883	44	
77 4 4-2.1226 77 4 5-6.1808		-3.68851939 69530705-03-5	-0.47019580- .9853653E-04-			
		0.00001 -0	-0.	-0.	0	0
9		0				
0039+276191+099: 0135+512139+000		-	REDO TEXAS RNHAM ENGLANE	143 0 401		
0002+425561-288	5997 67	0 9 000201BI	LLERICA MASS	121		
0300+107330+0610	500U 213	5 510C30000TR	INIDAD	159		
-582103000102210	0158270002	9000016800000	18980	MSG 1	1877 -27	-150-24
410002010208		0	0000000		06878 35	-80-66
		0700015400002 0700015400000		MSG MSG 2	1877 -21	-120-24
0 4220039010223		• • • • • • • • • • • • • • • •		1100 2		15-8
060 10135010220					1401 -16	
060 10135010220 073220039010223			11380	010	1401 -16 313 24	-25-19 -33- 6
073220039010223						-16- 6
077220039010223	1655190001	0400014020000	08540		188 37	-4- 4
58						
051 100690102230				1105	2373 -16	
2103000102210 051 100690102230		0700015400000 158003600000	18520	MSG_2 030	1877 -21 2373 -16	
-51 100690102230				090	2373 - 16 2373 - 16	-

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Wolf Research and Development Corporation

BBBBBBBBBBBBBBBB PN MISSING SO PA USED FOR SAT 60 LLEM 19

 SATELLITE INVENTORY FOLLOWS.\*

 51
 19/
 60
 19/
 73
 6/
 77
 4/
 \*

 05810221
 01582700
 NO
 ELEMENTS IN
 SYSTEM(1).
 3CC\*

UO S FULLOW\* UNK\*\*\*\*\*\*\*\* 6878 0 \* 2189 39.00631465 0.00538\* 05110208 00221500 1256. 21.81 3099. 19. 25 DOP\* 1267 39.00684022 0.02643\* 06010208 00221500 52 DOP\* 824, 27,84 2022. 19. 63 38.99523414 0.00552\* 07310208 00221500 7425. 6. 25 DOP\* 4179. -33.18 -65 38.99127295 0.02301\* 07710208 00221500 5885 4. 3216, -24,20 25 D0p\* UNK\*\*\*\*\*\*\*\* Mitt 05110221 01592900 2348 225.08 52.03144556 -0.01601\* 286.79 \_94.19 19322.3 17807.4 RDR -31,301 320,43 19. 300 \* 1404 222./2 52.03242214 -0.00864\* 06010221 01592900 290.35 -77.61 19322.3 17408.9 RDR -31.301 320.43 19. 300 🛓 272 324.99 52.02669415 -0.02319\* 07310221 01592900 -31,301 320.43 157.62 30.31 19322.3 19057.3 RDR 6. 300 \* 07710221 01592900 145 325,03 52.02714149 0.00002\* 157,52 4.05 19322.3 19117,7 RDR -31,301 320,43 ٩. 300 \* 1877 - M\* UNK\*\*\*\*\*\*\*\* 2348 180.54 52.04140303 -0.00005+ 05110221 01592900 -0.394 303.74 304.01 -76.97 1116.7 -421.4 RDR 19. 300 \* 06010221 01592900 1404 180.51 52.04354025 0.00248\* 304.04 : -63.92 1116.7 -1093.3 RDR \*0.394 303.74 19. 300 \* 272 359,56 52,02072186 -0.02916\* 07310221 01592900 124,56 -2.75 1116.7 871.8 RDR -0.394 303.74 300 . 6. 07710221 01592900 145 359,56 52,02126079 -0.00586\* -0.394 303.74 124.56 -28.92 1116.7 923.7 RDR 4. 300 \* 00410223 15084950 NO ELEMENTS IN SYSTEM(1) 39×

5-2-6

REDUCT

60\* 1390 50,94620 0.078205 -0.186-007 0.000+000 1.21164 0.11134\* 11.636 -3.354 -0.0000 355.095 2.785 0.0000 112.62 49.98\* \* 0.00200 20.00000 200.00000\* 06010220 19424540 1401 81.38 51.80628582 -0.00016\*

00010200 17404040			~ ~U•VUCIO*		
		-0.25 1176.3		19.	135 🖈
06010220 19424540	1401	81.38 51.80628582	2 -0.00016*		
49,214 359.07	作,59	-0.25 1176.3	0.0 VIS	19.	135 *

77\* 90 48.61977 0.062087 -0.212-005 -0.618-008 1.03837 0.00817\* 166.095 -3.689 -0.0047 20.377 -0.470 -0.0006 89.41 65.01\* \* 0.00206 20.00006 200.00000\* 07710223 16551900 188 155.88 54.67049040 0.00037\* 21.739 265.66 143.54 -0.04 208.0 -7.6 RDR 4. 39 \* END OF RUN \*\*\*\*\*\*\*

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ROC

## (Radar Orbit Computation)

INPUT DECK: Manual Schedule Tape Mode

INPUT DATA: Two options for input are accepted by the program: 1) radar observations, 2) geocentric rectangular coordinates and velocities.

Option 1:

1. Request Card: Col. 1-3

Number (odd) of observation cards to be read, from 3 to 999. 6-17 Alphanumeric satellite

name or identification, for output.

18-35 Date of computation

41 1 punch request element cards as output. 42 l punch prints x, y, z, x, y, z for the mean of the observations.

72-80 Are blank

2. Standard Station Card

- 3. Standard Observation Cards, in order of increasing time. (any odd number from 3 to 999)
- 4. 9 punch in col. 79: card used to end data deck.

Option 2:

1. Request card:

Alphanumeric satellite name or identification, for output.

- 18-35 Date of computation
  - l punch requests element cards as output

Must be blank

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Col. 6-17

72-80

ROC

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2.	Rectangular	Coordinate	Card:
<b>.</b>	recengular	oooramate	Oaru.

Col.	1-14*	x, in km.
	15-28*	y, in km.
	29-42*	z, in km.
	43-46	Year of epoch

3. Velocity Component Card:

-		•		
Cal	1-14*		n km/	000
CO1+	T = 1.4 .	X	п кш/а	500.

15-28\* y, in km/sec.

29-42\* z, in km/sec.

43-56\* Day of year of obser-

vation of coordinates

4. 9 punch in col. 79: card used to end data deck.

\* The field must have a decimal punch.

OUTPUT:

The output consists of the normal Radar Orbit Computation plus two options as selected by the input request card.

The Radar Orbit Computation is comprised of headings (program title, station name, alphanumeric satellite identification, and the date) and output. Each datum of output follows its label: semi-major axis in km., earth radii, right ascension of ascending node in degrees, radius vector in km., eccentricity squared, nodal period in days and minutes, argument of perigee in degrees, velocity in km./sec., perigee distance in km. and in earth radii, time of last perigee pass in days, inclination in degrees, apogee in km. and earth radii, and the time of the last nodal crossing in days.

Output Option A, which is possible only with input option (1), produces a printing of the geocentric rectangular coordinates x, y, z (headed XX = , YY = , ZZ =) and the velocity components (headed XDER = , YDER = , ZDER =). The time of the observation is given under the heading TIME =. Output Option B produces punched cards in the standard 6 card element set format. The data punched are the satellite number, element card number, satellite name, eccentricity, inclination, year of epoch, time of epoch, nodal period at epoch, right ascension, argument of perigee, and perigee distance. All other fields of the standard set are left blank.

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If teletype transmission is desired, the program TELTYP may be used.

JOB ROCE2 RUN ROCE2,DATA 27 ROC TEST MARCH 17,1961 001+426175+0714922 1560 5 000101MFLOSTONE M	- C -			
10317014434770 87700262300001217006602		517	7	5-5
30210001 1031/014440290 9000026310000118200660		17	~	8-5
30210001 1031/014446710 3800026450000114300657		17	4	6-5
30210001 1031/01445256010640026490000169800651		-5179	9	38-59
30316001 1031701445835011170026560000106300554		517	~	2-5
JUZIVUU 1031/014504/2012050056 /40000102200637		17	4	8-5
30210001 1031701451092012980026820000098200631		517	Δ	1-5
30210001 1031/01451688013/3002691000094800622		517	~	6-5
30210001 1031/01452298014610027100000090900612		517	4	0-5
30210001 1031/01452901015/6002720000087200600		17	9	4-5
JUZIVUU 1U31/01453524016/3002736000083300587		517	0	6-5
30210001 1031/01454135017870027590000079800576		517	4	2-5
30210001 1031/01454741019060027730000076500559		517	9	4-5
30210001 1031/01455347020200027970000073200539		517	ഹ	6 – 5
30210001 1031/01455960021570028230000070000516		517	4	2-5
30210001 1031/01460564022800028450000067000492		517	9	9-5
30210001 1031/0146118202411002877000063900463		517	S	6-5
30210001 1031/0146178502552002910000061100431		517	ŝ	7-5
30210001 1031701462378026880029440000058700394		17	9	8-5
3UZ100UL 1031/01462991028290029660000056700362		17	ഹ	1-5
30210001 1031701463596029510030300000054400310		517	ഹ	5-5
30210001 103170146418903074003078000052800256		17	S	2-5
30210001 1031701464810032020031340000051500199		17	ŝ	0-2
30210001 1031701465432033110031960000050600137		17	4	6-5
30210001 1031701470033033600032550000049600074		17	Ŝ	2-5
30210001 1031701470652033730033200000049600005		17	4	2-5
30210001 103170147125603373003388000005000058		17	m	ŝ
ENDDATA				11111111
ENDOFJOBR				

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SATELLITE NO: MILLSTONE MASS RADAR OREIT COMPUTATION ROC TEST SATEIIJTE

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MARCH 17,1961 THE BELOW INFORMATION COMPUTER FOR

0.063122711 MINUTES 90.896704110 PERIOD SEMI-MAJOR AXIS X

VELOCITY KM75EC 7.761152966 ARGUMENT OF PERIGEE DEGREES 07.409955800 RIGHT ASCENSION DEGREES 73,093154430 6696.710489300 Eárty radii 1.049941643

TIME OF NODE 1)4YS 76,062922770

TIME OF PERIGEE Days 76,014993400

274.090458100 Earth Radit 1.042973186

НX PERIGEE

INCLINATION Degrees 51.470701640

APOGEE Km 362,982519000 Earth Radii 1,056910100

FADIUS VECTOR Км 6656.862561000

ECCENTRICITY SOUARE ECCENTRICITY 0.006636994

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LOCVEC

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# LOCVEC

# (Vector Coordinates for Lockheed)

INPUT DECK:	Man	ual Schedule Tape	Mode
INPUT DATA:	(1)	Request cards -	as many as required up to 500.
		cols. 1-3	SPADATS sat. no.
		ccl. 9	a) 1st digit of STC sat. no.
	,		b) 9 if SPADATS sat. no. used
		cols. 10-12	Sat. no.
		cols. 13-20	Name of sat. (alphanumeric)
		cols. 21-24	Country of origin
	(2)	Col. 8 - a nume	ric punch (1-9) which signals the end
		of the request d	eck.
	(3)	Standard six or	seven card element sets. Up to 500
		sets can be read	d into core.
	(4)	One blank card	terminates the element card read in
		process.	,
OUTPUT:	Logi	ical tape 11 contair	ns x, y, z in feet and x, y, z, in
	ft/s	ec. to be printed.	Five-channel paper tape is available
	for	teletype transmiss	sion to the CDC 1604 computer at
	Sunn	yvale Tracking C	enter.

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		0			
	66.814169118 287.584307021 1.136606549				
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	•0090423448 5•386727540 -•693309309	+856-9 137•51			
	• 00904 35 • 386 - • 6933	•0721074839-•5262754886 54392072522037•			
AŢA	0 10 10 10	39- • 3920		•-	·
82 82•0	509 66878: 15355 96759	0748 54			
LOCVECB2 LOCVECB2,UATA	182•1869666 286•999153 -2•426967	•07,21			
<u> </u>	6 7 19 6 7 19	r-4	TA		
JOB RUN OMIC	1 0975323 36414-9	9779291	ENUDATA		

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LOCVEC

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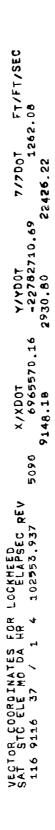
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ENDOFJOBR

5-4-2

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## CCOE

Manual Schedule Tape Mode.

### (Cartesian Coordinates from Orbital Elements)

INPUT DECK:

OUTPUT:

INPUT DATA:

- TA: (1) Set of Standard Elements (6 or 7 cards).
  - (2) Request cards as many as required:
    - Cols. 1 10: Start time (days of year).
      - 11 20: Time increment (days of year).
      - 21 30: End time (days of year).
        - 31: Blank, output in kilometers per sec.
          - 1, output in earth radii per min.
  - (3) One blank card
  - (4) The above may be repeated any number of times. To terminate the program a second blank card must be used.

Logical tape 11 contains the position and velocity components with the output in kilometers and kilometers per second or in earth radii and earth radii per minute. The information is in order of increasing time (in days) for each time step during the range requested.

#### 5-5-1

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D D D O O O O O O O O O O O O O O O O O	(DAYS) - I - RADII, BL - KM		
11 12 13 14 15 18 17 18 19 28 21 22 23 24 25 2	1 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0		
19 15 20 14 10 10 11 10 10 10 10 10 14 10 10 10		x 8 8 8 9 9 9 9 9 8 8 8 8 8 9 8 9 8	VYYYY A A A A A A A A A A A A A A A A A
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		* * * * * * * * * * * * * * * * * * * *	199999999999999
		888888888888888888888888888888888888888	\$

Columns:

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1 - 10	Begin Time, In Days,	DDDDD.DDDDD
11 - 20	Step Time, In Days,	DD. DDDDDDDD
21 - 30	End Time, In Days,	DDDDD. DDDDD
31	Blank Means Output In Ki	lometers Per Sec.
	1 Means Output In Earth	Radii Per Min.

Note: Decimal points may be punched anywhere in each of the first three fields.

Request Card CCOE

# 5-5-2

JOB CCOEr 2 RUN . CCOEB2,DATA 0.09786993 4172 1 58 ALPHA 11620. 35.20999976 48.54090166 1961 4172 21961 48.54090166 4172 3 0.07414181 209.17569106 104.01010047 1.05549147 7.:2633275-.. 4172 4-1.2992010E+07 -4.00362795 4172 5-0. -2.4287259E-04 J.6286715E-04- . 4172 6 1.00000000 360.000009999 -:--C. 0 ( 60.00000.00100000 60.01000 e

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ENDDATA

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REEFENELEER+BBBRDs.	AND VELOCITIES	FOR 58 ALPHA	ELEMENTS 1724
TIME	X	Y	X KM +
DAYS	XDOT	TOOT	ZDOT KM/SED#
60,00000	5280,734	5011,540	
	#5.151638	3.709421	
60,00100	4816,630	5512,215	-1803,102*
	+5,598174	3,252776	- 017801804
60.00200	4315,338	5571,853	-1724,595+
	•6.012291	2,759113	-5.656223.
¢0.00300	3779.884	5787.371	-1433,9774
~~~~~	-6,388631	2,230950	
50 0040a			
60.00400	3213.765	5955,922	• g1 28 , 43 24
60 00000	•6.721775	1.471480	•3.3107824
60,00500	2620,941	6074,970	-2405.148+
	•7.006355	1,084616	-4,089341+
00 <b>0</b> 000	2005,829	6142,346	-2401.353*
	-7,237188	0.475013	-R:036218+
<u>60 * 00700</u>	1373,280	6150,317	-2894,3624
	-7,409424	-0,151938	#¥,5526554
60 <b>.</b> 00 <b>8</b> 00	728,547	6119.644	-#101,618+
	.7,518712	-0.790139	-8+240463+
60.00900	77,225	6019.644	-3260 748 s
	-7.541371	-1 432857	-1,902039+
60.01000	-574,800	5668,238	-3429,611+
	•7 534568	-2,072828	-1,540379.
END+5			

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# 5-5-4

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RESPLT

(Residual Plot)

INPUT DECK: Schedule Tape Mode

An SPSJOB card is required. The following options of data input and output are specified in columns 17 and 18:

a) Input option, column 17:

Option	Observation Data	Element Data	Sensor Data
0	Observation Cards	SEAI Tape	SEAL Tape
1	Observation Cards	Element Cards	SEAI Tape
2	Observation Cards	Element Number Cards	SEAI Tape
3	Observation Cards	SEAI Tape	Sensor Cards
4	Observation Cards	Element Cards	Sensor Cards
5	Observation Cards	Element Number Cards	Sensor Cards

b) Output option, column 18:

0 = print (DS. 1) and punch cards (DS. 2)

### INPUT DATA:

Element data and sensor or station data may come either from standard cards or from the SEAI Tape depending on the input options chosen. The number of input cards is limited by the size of the O, E and S blocks of the system.

Observational data always originates from standard observation cards, which have been ordered by satellite number and contain an association indicator in col. 80.

If untagged observations (a zero sat. no.) are to be reduced, the first of the untagged observations must contain the satellite number of the element set against which the observations are to be reduced. If this procedure is inconvenient, a card may be inserted preceding

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the untagged observations. This card must contain the satellite number to be used and an association indicator in col. 80. In this case, the sensor number must be blank.

The association indicator is necessary on all observation cards to indicate whether the observation is an angles only or a radar observation.

Association Status	Observation Kind	Association Category
1	Radar	Associated
2	Radar	Doubtful
3	Radar	Unassociated
4	Angles Only	Associated
5	Angles Only	Doubtful
6	Angles only	Unassociated
9	Radar	Special Unassoc.

#### OUTPUT:

Two types of output are provided: 1) the printed output of the reduced observations, and 2) the punched card output of the reductions, with pen commands, required as input to the EAI Data Plotter.

- The printed output consists of the reduced observations, sorted by revolution number, for a given satellite. Headings are printed for the data columns which contain: the satellite number, the observation number, the revolution number, the time (minutes/100) since epoch, the vector magnitude in kilometers, the revolution number (N/1000) since epoch, the element number, the association indicator, and the station number.
- The punched card output of the reduction is used as input on an EAI Data Plotter which is assumed to have line plotting capabilities. The required

## 5-6-2

RESPLT

commands are punched in column 55. The first card of output will stop the plotter (7 in col. 55). This is done to enable the operator to change graph paper and position the origin. The cards which draw the axes follow the stop command (six cards). The graph labeling cards are next and cause the plotter to draw the characters S, E, R and T. These characters are followed by the satellite number, element number, epoch revolution number and the epoch time in days. A stop command follows the above data (7 in col. 55). (The deck may be broken at this point if the axes and labeling is not desired). The data cards containing the information to be plotted follow the stop command and continue until the next stop command is reached.

Two types of plots can be obtained from the data cards: 1) the time difference versus the revolution, and 2) the vector magnitude versus the revolution. Two IBM 523 board wiring diagrams have been supplied in Section 1 under the program description. The scaling is such that both plots can be obtained on the same piece of graph paper. To achieve this result, the IBM 523 board must be changed when the final stop command is executed. The deck should be broken as described above and the data cards reprocessed. It is recommended that a symbol pen or color change be used in order to distinguish the two plots.

The graph scaling is as follows:

- 1) x represents the epoch revolution number
- 2) each centimeter represents 20 revolutions

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- 3)  $y_0$  is the origin of either a) the vector magnitude, or b)  $\Delta t$ .
  - a) each centimeter represents 500 kilometers
     and under b) each centimeter represents 5
     minutes.
- 4)  $x_0$  should be 5 centimeters from the left edge of the lined graph paper and  $y_0$  should be 9 centimeters from the bottom.
- 5) full scale is then:
  - a)  $x_{R} = -100$  to +400 revolutions from epoch

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- b)  $y_{VM} = 0$  to 4500 kilometers
- c)  $y_{+} = -45$  to +45 minutes.

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RESPLT

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034119	5											
v34119	61.0499	039926										
034119	7											
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			235.398		292.7250							
			- 4.640						_			
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	61.1052	13.445										
064072		100440										
	161 ALP			4505	00593788	836	097.365	1.26	· / 7			
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034250	03902040	52344315	25137000	660000	113340	50440	00000000	0	0 0	03452		4335
034250	03902040	52344255	30147000	672600	1:9770	40310	0000000	0	0 0	03452		4335
			57149000									4335
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			71212000				000000000					4334
			77217000				00000000					4334
			62220000				00000000					4333
			67219000				00000000					4333
			73220000				00000000					4333
			79227060				00000000					4333
634250	03902040	52342435	64225000	1107000	53070	150	00000000	0	0 0	03452		4333
			69227000			3000	00000000	0	0 0	03452		4333
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			44296000 44296000				000000000					2477
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5 9 5 5 5 1 7 9 9 9 9 9 9 9 9 9 9 9 9 9 9 9 9 9 9	342500390204100843435 342500390204100843435 702500390204090543495 702500390204090543495 702500390204090543435 702500390204090543315 702500390204090543315	0702500390204060633J15 0702500390204060632555 0702500390204060632495 0702500390204060632375 0702500390204060632375 0702500390204060632195 END CASER END CASER END CASER

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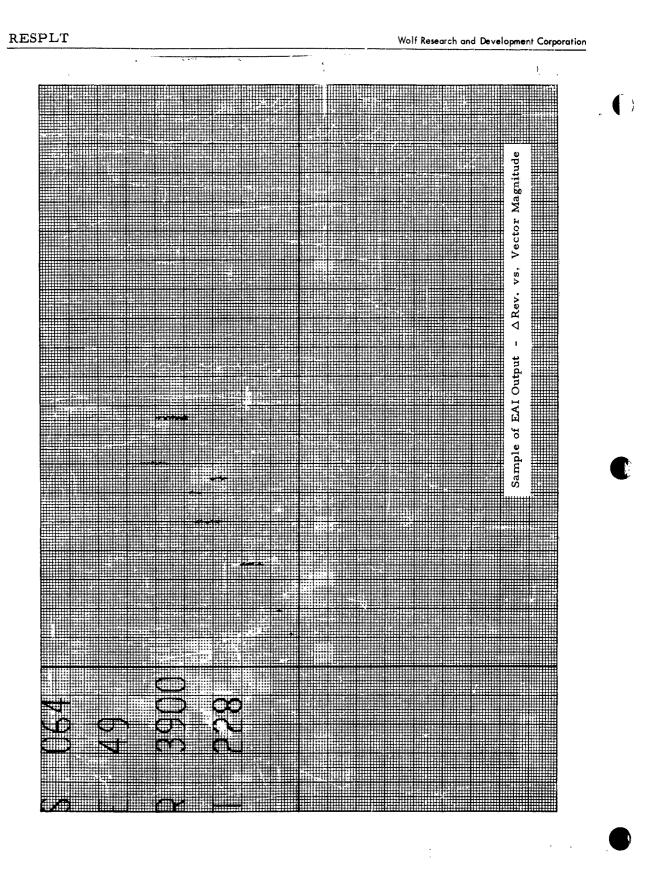


SATNO	OBSNO	REV	DT(MIN)	VMAGN	NREV		ELNO	ASTAT	STA
034	43334	10731	0012	56	,007	8	119	1	039
034	43335	10731	0010	48	.007	8	119	1	039
034	43336	10731	<b>⊳.</b> 001.0	48	.007	8	119	1	039
034	43337	10731	00.1.1	52	.007	8	119	1	039
034	43338	10731	0010	48	.007	8	119	1	039
034	43339	10731	0009	41	.007	8	119	1,	039
034	45340	10731	-,0009	43	.007	8	119	ĺ	039
034	43341	10731	0012	54	.007	8	119	1	039
034	43343	10731	<b></b> 0010	49	.007	8	119	1	039
034	43344	10731	0012	54	.007	8	119	1	039
034	43345	10731	0011	51	.007	8	119	1	039
0.34	43346	10731	0013	62	.007	8	119	1 `	039
034	45347	10731	0012	53	.007	8	119	1	039
034	43348	10731	-,0015	72	.007	8	119	2	039
034	43349	10731	0013	60	. 007	8	119	1	039
034	43350	10731	0014	66	.007	8	119	1	039
034	43351	10731	0012	58	.007	8	119	1	039
034	43352	10731	0017	78	.007	8	119	1	039
034	43353	10731	-,0017	76	.007	8	119	1	039
034	43354	10731	<b>*•0013</b>	57	.007	8	119	1	039
034	43355	10731	0015	72	.007	8	119	1	039
034	43356	10731	+,0016	72	.007	8	119	1,	0.39
034	45357	10731	0016	75	.007	8	119	1	039
034	43358	10731	-,0010	47	.007	8	119	1	039
034	24798	10800	.0268	1194	.076	8	119	1	039
034	24797	10800	.0269	1195	.076	8	119	1	039
034	24796	10800	.0268	1189	.076	8	119	1	039
034	24795	10800	.0269	1191	.076	8	119	1	039
034	24794	10800	.0268	1185	.076	8	119	1	039
034	24793	10800	.0268	1180	.076	8	119	1	039
034	24792	10800	.0271	1188	.076	8	119	1	039
034	24791	10800	,0269	1174	.076	8	119	1	039
034	24790	1,0800	.0268	1166	.076	8	119	1	039
034	24789	1,0800	.0269	1161	.076	8	119	1	039
034	24788	10800	.0267	1147	.076	8	119	1	039
034	24787	10800	.0268	1142	,076	8	119	1	039
034	24786	10800	.0267	1132	.076	8	119	1	039
034	24778	10800	.0262	1001	.076	8	119	1	039
034	24777	1,0800	.0263	988	.076	8	119	1	039
034	24776	10800	.0263	973	.075	8	119	1	039
034	24775	10800	,0262	964	.075	₽ ₽	119	1	039
034	24774	1,0800	.0262	959	.076	8	119	1	039
034	24773	10800	.0261	961	.075	8	119	1	039
034	24772	1,0800	.0261	971	.076	8	119	1	039
034	24770	10800	.0260	1014	.076	8	119	1	039
034	24769	10800	.0258	1045	.076	8	119	1	039
034	24768	10800	.0258	1090	.076	8	119	1	039
SATNO	ORSNO	REV	DT(MIN)	VMAGN	NREV		ELNO	ASTAT	STA
070	06973	6515	0003	36	.010	8	63	2	039
070	06971	6515	0004	28	.010	8	63	1	039
070	06970	6515	0005	40	.010	8	63	2	039
070	06969	6515	0005	37	.010	8	63	1	039
070	06968	6515	+.0005	26	.010	8	63	1	039
070	-06967	6515	0006	37	.010	8	63	1	039
070	06966	6515	0007	40	.010	8	63	2	039
070	06962	6515	0008	40	.010	8	63	1	039
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(Satellite Situation Report from Nodal Elements)

INPUT DECK: Manual Schedule Tape Mode The input deck should be in the following order: INPUT DATA: Request Cards (up to 500) in the following format: 1. Cols. 1 - 3: Satellite number 7 Comment indicator defined as follows 0. no comments on this card 1, read comments contained in cols. 21 - 629 - 20: Satellite name 21 - 62: Comments to be printed if col. 7 contains a 1 punch. (A max. of 50 sats. may have comments) 2. Request deck terminator: a 2 punch in col. 7. 3. Time cards (times for which reports should be issued) in the following format: Cols. 1-4 Year 6 - 7 Month number 9 - 10 Day of month 11 - 17 Hours and minutes (HHMM. MM) 20 Output unit indicator, defined as follows: 0, output in statute miles 1, output in kilometers Time Deck terminator: a 7 punch in col. 8. 4. 5. Standard 6 or 7 card element sets. Only the sets required are stored, therefore more than 500 sets may be included in the input deck. Input deck terminator: a blank card. 6. OUTPUT: Logical tape 11 contains the satellite name and number, element number, latitude, longitude, inclination, period, apogee distance, perigee distance, revolution number, time of node, right ascension of node, longitude of node,

PREPINT

5-7-1

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eccentricity, height of satellite, and course. The distances may be in kilometers or statute miles. The dimensions are determined by the option selected on the Time card (col. 20).

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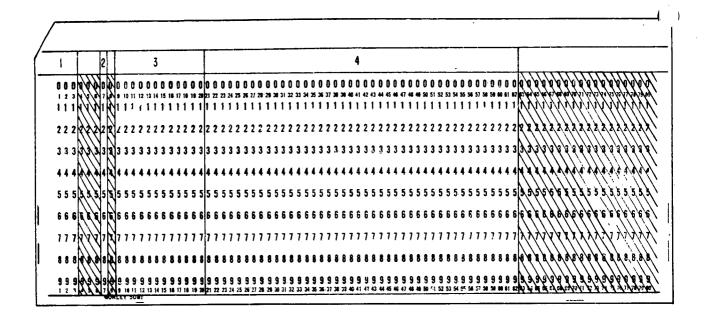
Field	Cols.	
1	1-4	Year
2	6-7	Month
3	9-10	Day
4	11-17	Hours and minutes
5	20	l, output in kilometers
		0, output in statute miles

Time Card PREPINT

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5-7-3

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Field	Cols.	
1	1-3	Satellite number
2	7	l, cards contain comments
		0, no comments
3	9-20	Satellite name
4	21-62	Comments

Request Card PREPINT

## 5-7-4

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JOB TEST PREPINT RUN PREPINTS, DATA 58 ALPHA 4 58 BETA 16 1 58 BETA 2 5 59 ALPHA 11 1 59 ALPHA 12 2 INSUFFICIENT OBSERVATION 1 59 DELTA 15 59 ETA 20 59 IOTA 22 1 23 59 IOTA 2 1 59 MU IN HELIOCENTR ORBIT 450.0D 1.317AU .9766AU 112 60 EPSILN 3 36 1 60 EPSILN 4 INSUFFICIENT OBSERVATION 37 43 60 ZETA 1 45 60 ETA 1 60 ETA 2 46 60 ETA 3 47 1 49 60 IOTA 102 61 LAMBDA 2 107 61 NU 115 61 OMICRON1 61 RHO 4 167 163 61 SIGMA 1 61 UPSILON 170 61 OMEGA 1 182 186 61A BETA 185 61 SIGMA 3 2 1961 10 11 1200 1961 10 11 1200 1961 10 11 1430 . . 1 1961 10 11 1430 1 7 0.09500069 33.18999958 14740. 4214 1 58ALPHA 4214 21961 280.80691970 1961 280.80691970 4214 3 201.61903020 107.42337540 1.05476451 4214 4 -4.96605333 7.42265104 0. 4214 5 0. -3.780903E-004 5.651232E-004 U. 0.07371200 -1.968034E-007 421 6 4214 7 0.18984776 34.24499,89 5 78 158 BETA 2 13950. 

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 278.25430390 1961 278.25430390 94.50760521 1.10183595 4.41471416 0. -9.897075E-005 1.446693E-004 0. 5 78 6 1.00000000 30.00000 0. 0. 11149 1 59 ALPHA 1 1108). 0.16288486 32.84826871 282.87323970 11149 21961 282.87323970/ 1961 i1149 3 0.08698856 32.53621422 190.73210390 1.08936014 11149 4 4.736775E-008 -3.51236494 5.28651193 0. 11149 5 0. ... 4.308964E-005-6.485486E-005 0. 0.01525878 10.00000000 11149 6 1.00000000 30.00000 12110 1 59 ALPHA 2 32.66509133 10100. 0.18352360 279.61322930 1961 279.61322930 ···· 12110 21961 12110 3 0.08993200 255.21911570 214.33844310 → 1.08631547 12110 4 1.999180E-008 -3.30525151 4.99311847 0. 1.568518E-005-2.369501E-005 0. 12110 5 0.

PREPINT



ENDDATA

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5-7-7

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# MAKETAPE

# (Make Input Tape for TELTYP)

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INPUT DECK: Manual Schedule Tape Mode.

- INPUT DATA: (1) TELEFORM punched in cols. 17-24. Used if the message is to broken into 90-line segments.
  - (2) Data cards to be converted to teletype.
  - (3) FINDATA punched in cols. 17-24. Used to terminate the data deck.

OUTPUT: Logical tape 11 is input to TELTYP. See writeup of program. The tape can be printed using data select one. The output will consist of a listing of the input data cards with the sentinels required by the TELTYP program.

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S BACK

THE QUICK BROWN FOX JUMPED OVER THE LAZY DOG

FIN DATA ENDDATA

HE QUICK BROWN FOX JUMPED OVER THE LAZY DOG

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MAKETAPE, DATA OVER THE LAZY DOG S BACK 0 1 2 RUN

MAKETAPEB2

JOB

ENDOFJOBR

### XYZLAR

(Look Angle Report from x, y, z, Coordinates)

INPUT DECK: Manual Schedule Tape Mode

INPUT DATA:

The input data is composed of sets, or units, each of which consists of a standard station card, a request card, and optional ephemeric data. Ephemeris data is required only in those data units which initiate a new vehicle or satellite output. The source of ephemeris data is either punched cards or an ephemeris tape (unit 7).

The order of cards for an input data unit using ephemeris cards and initiating a new satellite is as follows:

- (1) Standard Station Card
- Request card specifying time parameters for the look-angle computations and input/output options.

Cols. 1-4	year	)
Cols. 5-7	day	base time of
Cols. 8-9	hour	desired look-
Cols. 10-11	minutes	angle
Cols. 12-13	seconds	predictions
Cols. 14-17	thousandths of see	
Cols. 25-34	maximum increm	ent of time from base
	time, for desired	look-angle predic-
	tions. When blan	k, all ephemeris
	data will be used	to generate predictions.
		<b>5</b>
Col. 44	0 = ephemeris fol	;
Col. 44 Col. 45	0 = ephemeris fol l = only output of	lows on cards
	-	lows on cards
	l = only output of	lows on cards visual sightings
	<pre>l = only output of desired 0 = output of all s</pre>	lows on cards visual sightings
Col. 45	<pre>l = only output of desired 0 = output of all s</pre>	lows on cards visual sightings ightings
Col. 45	<pre>1 = only output of desired 0 = output of all s 1 = include negativity in output</pre>	lows on cards visual sightings ightings
Col. 45	<pre>1 = only output of desired 0 = output of all s 1 = include negativity in output</pre>	lows on cards visual sightings ightings ve elevation sightings tion sightings only in cutput

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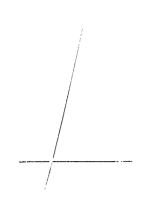
- (3) Vehicle identification card containing the alphanumeric vehicle name in columns 1-16.
  (4) Ephemeris cards, in order of increasing time increment over base time.
  Cols. 1-14 time increment, in minutes, from base time
  Cols. 15-28 x coordinate of vehicle position
  Cols. 29-42 y coordinate of vehicle position
  Cols. 43-56 z coordinate of vehicle position
- (5) Blank card to terminate the ephemeris cards.
- (6) Blank card to terminate program.

This set or unit of input data (1-5) will generate one schedule of look-angles for the vehicle data contained on card types 3 and 4. Additional schedules for different stations are obtained by adding station and request cards, in pairs, after the blank card (5).

To obtain schedules for vehicles or satellites in addition to that of the first input unit, another ephemeris card set (1-5) is required. A minus sign (-) must be punched preceding the station number (cols. 1-4)of the station card. Schedules for different stations result from using station and request cards, in pairs, as above.

The order of cards for an input data unit using the ephemeris tape is as follows:

- (1) Standard Station Card.
- (2) Request card, as above, with two exceptions:
  - Col. 44 1 = ephemeris tape to be used Col. 47 1 = punch ephemeris data 0 = no ephemeris card output
- (3) Blank card to terminate program.



5-9-2

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Prediction angles for additional stations for a given vehicle are obtained by adding additional station card-request card pairs before the blank card (3).

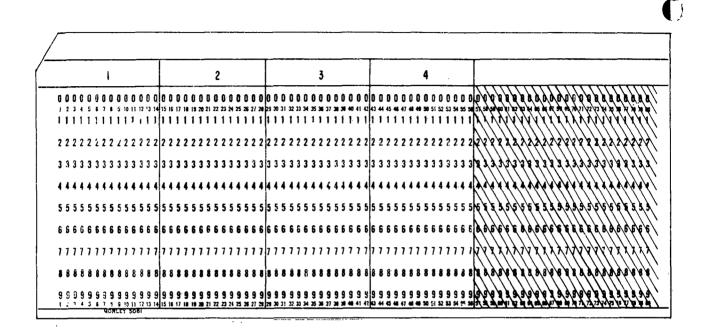
OUTPUT:

The basic output is a schedule of predicted lookangles for a particular station at specific times. The report headings are the satellite name; the label LOOK-ANGLES FOR, followed by the station name; and headings for the sighting coordinates and time data.

Each data line consists of a day of year, hour, minute, and fraction of minute of the search point; the predicted right ascension, declination, azimuth, and elevation in degrees; the predicted slant range in kilometers; and the elevation and illumination angles of the sun. The sun data indicates if the satellite will be visible at the time and location of prediction.

In addition to the look-angle data output, the program will also produce punched cards containing the satellite's ephemeris data in the format of the ephemeris card input. The punched cards may be obtained only when the ephemeris information originates from ephemeris tape input. A 1 punch in column 47 of the first request card for a given satellite will produce one set of ephemeris card output.

5-9-3

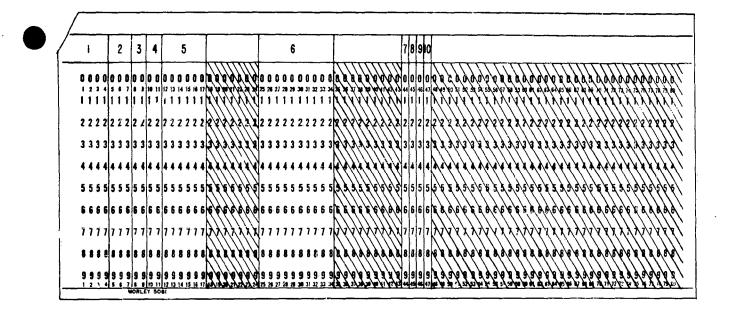


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Field	Cols.	
1	1-14	Time increment, in minutes
2	15-28	x coordinate of vehicle position
3	29-42	y coordinate of vehicle position
4	43-56	z coordinate of vehicle position

Ephemeris Card

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Field	Cols.	
1	1-4	Year
2	5-7	Day
3	8-9	Hour
4	10-11	Minutes
5	12-17	Seconds
6	25-34	Time maximum
7	44	l, ephemeris tape data input
		0, ephemeris cards data input
8	45	l, visual sightings only
		0, all sightings desired
9	46	l, negative elevations permitted
-		0, only positive elevations desired
10	47	l, ephemeris cards punched
		0, no punched cards

Request Card

## XYZLAR

# 5-9-5

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JOE					
RUN 0026+428481-2859292		CATA 2601 SCHENEC TAD'	YNY GE132	( 01 (	
1962239071919.000	4040 9 002	1		6016	45
MARINER JPL 3		-	-		
0.000000+000 0.12					
0.3000000+002-0.21					
0.6000000+002-0.35					
0.900000+002-0.46					
0.1200000+003-0.56					
0.1800000+003-0.74					
0.2100000+003-0.82					
0.2400000+003-0.89	993919+001-0.78	83390+001-0.2	107805+001		
0.2700000+003-0.97	741469+001-0.88	51568+001-0.2	178128+001		
0.300000+003-0.10					
0.3300000+003-0.11					
0.3600000+003-0.11			· · - • · -		
0.3900000+003-0.12 0.4200000+003-0.13					·
0.4500000+003-0.13					
0.4800000+003-0.14					
0.510000+003-0.15	519158+002-0.16	09307+002-0.20	529574+001		
0.5400000+003-0.15					
0.5700000+003-0.16					
0.6000000+003-0.17					
0.6300000+003-0.17 0.6600000+003-0.18					
0.6900000+003-0.18				•	
0.7200000+003-0.19					
0.7500000+003-0.20					
0.7800000+003-0.20					
0.8100000+003-0.21					
0.8400000+003-0.21	· · - · · · · · · · · · · · · · · · · ·				
0.8700000+003-0.22 0.9000000+003-0.23					
0.9300000+003-0.23					
0.9600000+003-0.24		-			
0.9900000+003-0.24					
0.1020000+004-0.25	047638+002-0•30	11204+002-0.3	363452+001	1	
0.1050000+004-0.26					
0.1080000+004-0.26					
0.1110000+004-0.27 0.1140000+004-0.27					
0.1170000+004-0.28					
0.1200000+004-0.28					
0.1230000+004-0.29					
0.1260000+004-0.30					
0.1290000+004-0.30					
0.1320000+004-0.31	118506+002-0•37	96457+002-0•3	747441+001		
0337+189167+1556833	2 1272 411	COUTH DOINT		031061	4.0
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<pre>MangLER JPL 3 * Durk ANGLES JPL SCHEMECTADY AY * Day TIPE K.A. DEC A21h ELE' RANGE SUNS TLLUMI 2 UFG. DFG AAG. ANG. ANG. 39 719.2 124.943 42.163 119.920**2.7.9 86*4127.9 28.0* 24.73 21 10.3 53 -32.0* 24.73 24.1 110.3* 23.6 A49.12 124.943 42.163 119.920**2.7.9 86*4127.9 28.0* 24.73 24.1 110.3* 24.73 24.1 110.3* 24.7 1</pre>	REAB	нвавааевы	888				
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$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	LOOK	ANGLES F	OH SCHENE	CTADY NY	÷		
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	DAY	TIME	. K . A .	OEC	AZIM ELE!	RANGE	SUNS TLLUMI*
$ \begin{array}{c} 3 & 7 & 7 & 0 & 3 & 2 & 1 & 2 & 4 & 9 & 4 & 2 & 1 & 1 & 2 & 2 & 0 & 2 & 7 & 4 & 0 & 4 & 4 & 4 & 1 & 1 & 0 & 3 & 2 \\ 7 & 7 & 7 & 3 & 2 & 1 & 6 & 7 & 5 & 3 & 2 & 1 & 5 & 7 & 5 & 7 & 1 & 5 & 1 \\ 7 & 7 & 7 & 1 & 1 & 1 & 3 & 1 & 0 & 0 & 1 & 1 & 1 & 3 & 1 \\ 7 & 7 & 1 & 1 & 3 & 2 & 2 & 1 & 1 & 5 & 1 & 2 & 1 & 5 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 5 & 1 & 1$				DEG	ANG, ANG,		ELEV NATION+
236740322167353321473-24.1110.3236A40322006.49320.5519.575-71.61310.0-10.9111.5236A40.32204.122-17.8915.155-74.4741434-15.4110.8236140.32214.122-17.8915.155-74.4741434-15.4110.8236140.32214.122-17.8915.155-74.47414.47-5.4110.8236140.32214.448-13.45630.385-44.9174.725.3107.52361110.32224.448-13.45647.165-51.178.2721.1010.810.752361140.32224.406-11.1934.122-44.99.860671.21.7116.810.62361210.32225.048-11.7904.122-44.99.86071.21.7116.310.62361210.32225.048-11.7904.122-44.99.86071.21.7116.310.62361210.32225.048-11.7006.537-47.5113.445.4242.510.62361210.32226.44-11.7113.465.4510.5510.610.52361210.32227.548-0.35992.164-22.07117.46.35.610.6236141.32227.548-2.599.44.35+11.7113.464.55.610.6236140.32227.548-2.599.44.35+11.6114.46.914.615.72361	239		: 24.943	-42.163		8941.	-27.9 28.0*
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	230		185.735			21473	-24 1 110 3*
$ \begin{array}{c} g_{36} & g_{6}, g_{2} & 206, 493 & -20, 551 & 9, 515 + 4, 474 & 41 + 41 + 41 + 41 + 41 + 41 + 41 + 4$							-19.9 111.5*
$ \begin{array}{c} 936 & 010, 32 & 214, 122 & +7, 829 & 15, 115 + 4, 330 & 50, 80 & -10, 4, 119, 42 \\ 936 & 1019, 32 & 219, 438 & +14, 474 & 31, 191 + 4, 43 & 50, 847, -5, 4, 118, 64 \\ 936 & 1149, 32 & 229, 438 & +14, 474 & 31, 191 + 54, 378 & 66061, -0, 1, 108, 01 \\ 936 & 1149, 32 & 224, 169 & +12, 458 & 47, 163 + 51, 17 & 82321, 10, 8, 107, 55 \\ 936 & 1149, 32 & 224, 169 & +11, 719 & 54, 122 + 46, 94 & 89, 441, 164, 210, 64 \\ 936 & 179, 32 & 225, 138 & +11, 099 & 60, 534 + 42, 342 & 96761, 21, 716, 33 \\ 936 & 1749, 32 & 225, 638 & +10, 670 & 66, 537 + 17, 543 & 103701, 227, 2146, 0 \\ 936 & 1149, 32 & 227, 093 & -9, 713 & 77, 226 + 73, 299 & 117426, 32, 27, 2146, 0 \\ 936 & 1449, 32 & 227, 093 & -9, 713 & 77, 226 + 73, 207 & 117426, 34, 74, 6105, 105, 64 \\ 936 & 1449, 32 & 227, 058 & -9, 713 & 77, 226 + 73, 349 & 117126, 37, 6, 105, 64 \\ 936 & 1449, 32 & 227, 952 & -4, 879 & 91, 636 + 11, 171 & 136484 & 50, 6 & 105, 14 \\ 936 & 1440, 32 & 227, 952 & -4, 879 & 91, 636 + 11, 171 & 136484 & 50, 6 & 105, 14 \\ 936 & 1440, 32 & 227, 764 & -9, 264 & 101, 184 & -1, 216 & 1486427 & 57, 2 & 1144, 74 \\ 936 & 1740, 32 & 227, 764 & -7, 848 & 114, 116 & 10, 448 & 141797 & 57, 51 & 144, 774 \\ 936 & 1740, 32 & 227, 764 & -7, 486 & 114, 1864 & -1, 216 & 1486727 & 57, 2 & 1144, 74 \\ 936 & 1740, 32 & 229, 145 & -7, 466 & 127, 348 & 1, 73791 & 52, 7 & 1144, 74 \\ 936 & 1140, 32 & 229, 246 & -7, 486 & 131, 116 & 10, 446 & 141797 & 554, 1144, 74 \\ 936 & 1140, 32 & 229, 248 & -7, 165 & 135, 144 & 794 & 555 & 1144, 794 & 557 & 1144, 74 \\ 936 & 2149, 32 & 229, 248 & -7, 165 & 135, 144 & 794 & 794 & 557 & 1144, 74 \\ 936 & 2149, 32 & 229, 248 & -7, 165 & 135, 144 & 794 & 79761 & 354 & 1144, 24 \\ 936 & 2149, 32 & 229, 248 & -7, 165 & 135, 144 & 794 & 79761 & 354 & 1144, 24 \\ 936 & 2149, 32 & 229, 248 & -5, 151 & 154, 144 & 794, 249 & 7794 & -254 & 1144, 24 \\ 936 & 2149, 32 & 229, 248 & -5, 154 & 154, 144 & 249, 177, 24, 81 & 1144, 24 \\ 936 & 2149, 32 & 229, 136 & -6, 254 & 199, 116 & 90, 172 & 290, 932 & 8, 3133, 99 \\ 246 & 146, 32 & 229, 136 & -6, 25$		- •					
$ \begin{array}{c} 236 & 046 & 322 & 27 & 297 & -15 & 969 & 22 & 615 + 51 & 443 & 584 & 74 & -55 & 4108 & 083 \\ 236 & 1149 & 32 & 221 & 448 & +13 & 556 & 49 & 383 + 4 & 941 & 747/2 & 5.3 & 107 & 55 \\ 236 & 1149 & 32 & 227 & 899 & +12 & 458 & 47 & 163 + 51 & 17 & 82221 & 10.8 & 107 & 09 \\ 236 & 1249 & 32 & 225 & 038 & +11 & 099 & 6n & 594 + 42 & 342 & 96761 & 217 & 106 & 39 \\ 236 & 1249 & 32 & 225 & 038 & +11 & 099 & 6n & 594 + 42 & 342 & 96761 & 217 & 106 & 39 \\ 236 & 1249 & 32 & 225 & 038 & +11 & 099 & 6n & 594 + 42 & 342 & 96761 & 217 & 106 & 39 \\ 236 & 1249 & 32 & 225 & 038 & +11 & 099 & 6n & 594 + 42 & 342 & 96761 & 217 & 106 & 39 \\ 236 & 1249 & 32 & 227 & 584 & -10 & 570 & 573 & -573 & 1 & 10748 & 326 & 5105 & 69 \\ 236 & 1249 & 32 & 227 & 584 & -10 & 570 & 973 & 77 & 226 & 27 & 39 & 11724 & 37 & 65 & 105 & 59 \\ 236 & 1249 & 32 & 227 & 584 & -8159 & 91 & 636 + 1171 & 136 & 184 & 500 & 515 & 27 \\ 236 & 1249 & 32 & 227 & 584 & -8159 & 91 & 636 + 31171 & 136 & 184 & 500 & 515 & 117 \\ 236 & 1249 & 32 & 227 & 584 & -836 & 101 & 964 & 515 & 134 & 604 & 104 & 99 \\ 236 & 1249 & 32 & 227 & 548 & -836 & 101 & 964 & 51 & 15474 & 57 & 2144 & 78 \\ 236 & 1249 & 32 & 229 & -836 & 104 & 105 & 598 & 51 & 154074 & 57 & 2144 & 78 \\ 236 & 1249 & 32 & 229 & -842 & -7 & 641 & 116 & 511 & 436 & 431 & 7169 & 57 & 2144 & 78 \\ 236 & 1240 & 32 & 229 & -242 & -7 & 241 & 127 & 704 & -35 & 179977 & 49 & 2144 & 42 \\ 236 & 1240 & 32 & 229 & -26 & -7 & 466 & 137 & 112 & 729 & 125721 & 436 & 2144 & 12 \\ 236 & 1240 & 32 & 229 & -26 & -7 & 146 & 1377 & 12 & -918 & 5121 & 778 & 1378 & 1144 & 22 \\ 236 & 1240 & 32 & 229 & -26 & -7 & 146 & 1377 & 12 & 918 & 107761 & 35 & 4 & 144 & 22 \\ 236 & 1240 & 32 & 229 & -26 & -7 & 146 & 1377 & 12 & 918 & 107761 & 35 & 4 & 1144 & 22 \\ 236 & 1240 & 32 & 229 & -26 & -5 & 613 & 1577 & 4178 & 2239 & 916 & -223 & 1144 & 2239 \\ 236 & 1240 & 32 & 229 & 226 & -56 & -574 & 255 & 146 & -128 & 210776 & -228 & 103 & -728 & 103 & -728 & 103 & -728 & 103 & -728 & 103 & -728 & 103 & -728 & 103 & -728 & 103 & -728 & 103 & -728 & 103 &$							
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$ \begin{array}{llllllllllllllllllllllllllllllllllll$	230	1119.32		-12,458			
$ \begin{array}{c} 3 & 3 & 2 & 2 & 5 & 6 & 4 & -1 & 5 & 7 & 0 & 6 & 6 & 5 & 37 & -7 & 5 & 3 & 1 & 1 & 1 & 27 & 2 & 1 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0 & 0$	200	1149.32	224,067				
$ \begin{array}{c} g_{30} & 1349, 32 & 226, 525 & -10, 113 & J_2, 050 + 22, 5^{-2} & 110, 433, 32, 5 105, 6 \\ g_{30} & 1349, 32 & 227, 958 & -9, 713 & J_7, 225 + J_7, 3+9 & 117426, 37, 6 105, 6 \\ g_{30} & 1449, 32 & 227, 958 & -9, 744 & 84, 941 + 6, 655 & 130, 60 & 46.9 & 105, 2 \\ g_{30} & 1449, 32 & 227, 958 & -9, 764 & 84, 941 + 16, 655 & 130, 60 & 46.9 & 105, 2 \\ g_{30} & 1449, 32 & 228, 287 & -4, 552 & 94, 337 + 5, 6-2 & 142, 355 & 54.0 & 104, 9 \\ g_{30} & 1449, 32 & 228, 764 & -8, 256 & 101, 084 & -0, 216 & 1486, 27 & 55.4 & 104, 8 \\ g_{30} & 1740, 32 & 229, 154 & -7, 848 & 111, 116 & 10, 448 & 141, 92 & 55.9 & 104, 5 \\ g_{30} & 1740, 32 & 229, 154 & -7, 646 & 112, 364 & -0, 443 & 173, 99, 52, 7 & 104, 4 \\ g_{30} & 1740, 32 & 229, 244 & -7, 321 & 128, 704 & 50.9 & 179, 977, 492 & 104, 3 \\ g_{30} & 1740, 32 & 229, 244 & -7, 321 & 128, 704 & 50.9 & 179, 977, 492 & 104, 3 \\ g_{30} & 1740, 32 & 229, 244 & -7, 1016 & 143, 177 & 12, 978 & 197, 179, 197, 492 & 104, 3 \\ g_{30} & 1940, 32 & 229, 248 & -6, 475 & 131, 334 & 40, 02 & 107, 135, 44 & 104, 2 \\ g_{30} & 2149, 32 & 229, 248 & -6, 740 & 160, 356 & 6, 4-1 & 204, 92 & 30, 2 & 104, 0 \\ g_{30} & 2149, 32 & 229, 136 & -6, 254 & 199, 116 & 0, 644 & 120, 408 & 30, 2 & 104, 0 \\ g_{30} & 2149, 32 & 229, 136 & -6, 6254 & 199, 116 & 0, 642 & 210, 779 & 24.8 & 104, 0 \\ g_{30} & 2149, 32 & 229, 136 & -6, 548 & 199, 499 & 0, 3.3 & 2 & 204, 92 & 30, 2 & 104, 0 \\ g_{30} & 2149, 32 & 229, 136 & -6, 548 & 199, 499 & 0, 3.3 & 2 & 209, 138 & 103, 9 \\ g_{30} & 2249, 32 & 229, 136 & -5, 652 & 259, 136 & -6, 244 & 126, 972 & 236, 126 & -77, 7 & 103, 74 \\ g_{40} & 132 & 229, 136 & -5, 643 & 249, 649 & 7, 241 & 249, 748 & -77 & 103, 74 \\ g_{40} & 149, 32 & 229, 136 & -5, 643 & 249, 647 & 229, 148 & -26, 9 & 103, 64 \\ g_{40} & 149, 32 & 229, 138 & -5, 643 & 241, 736 & 26, 574 & 256, 144 & -127, 7 & 103, 74 \\ g_{40} & 149, 32 & 229, 138 & -5, 643 & 249, 736 & -544 & 226, 964 & -77 & 73, 976 & -36, 1 & 103, 74 \\ g_{40} & 149, 32 & 229, 138 & -5, 643 & 249, 647 & -3312 & 296, 567 & -326, 7 & 103,$	230	1219.32	225.038	-11,099		96761.	
$ \begin{array}{c} g_{30} & 1349, 32 & 226, 525 & -10, 113 & J_2, 050 + 22, 5^{-2} & 110, 433, 32, 5  105, 6 \\ g_{30} & 1349, 32 & 227, 958 & -9, 713 & J7, 226 + J7, 349 & 117426, 37, 6  105, 6 \\ g_{30} & 1449, 32 & 227, 958 & -9, 744 & 84, 941 + 16, 615 & 130, 60 & 46.9  105, 2 \\ g_{30} & 1449, 32 & 227, 958 & -9, 759 & 91, 836 + 11, 171 & 136484 & 50, 8  105, 1 \\ g_{30} & 1549, 32 & 228, 287 & -8, 759 & 94, 337 + 5, 6-2 & 142, 35, 54.0  104, 9 \\ g_{30} & 1549, 32 & 228, 764 & -8, 256 & 101, 094 & -0, 216 & 148, 27, 56.9  104, 59 \\ g_{30} & 1740, 32 & 229, 104 & -7, 848  111, 116  10, 448 & 161, 92, 55.9  104, 59 \\ g_{30} & 1740, 32 & 229, 104 & -7, 646 & 116, 551 & 55, 50.9  173, 99, 52, 7  104, 48 \\ g_{30} & 1740, 32 & 229, 244 & -7, 321  128, 704 & 50.9  179, 197, 492, 2104, 32 \\ g_{30} & 1740, 32 & 229, 244 & -7, 321  128, 704 & 50.9  195, 721, 414, 48 \\ g_{30} & 1940, 32 & 229, 244 & -7, 116  143, 177  22, 978  197, 492, 2104, 32 \\ g_{30} & 1940, 32 & 229, 248 & -6, 475  151, 434  40, 072  197, 55.4  104, 104, 22 \\ g_{30} & 1940, 32 & 229, 248 & -6, 475  151, 434  40, 072  197, 55.4  104, 104, 22 \\ g_{30} & 2149, 32 & 229, 248 & -6, 475  151, 434  40, 072  197, 55.4  104, 104, 22 \\ g_{30} & 2149, 32 & 229, 248 & -6, 475  151, 434  40, 072  216, 332  549, 120  556  64, 471  204, 092  30, 2144, 40 \\ g_{30} & 2149, 32 & 229, 136  -6, 254  199, 116  40, 644  216, 704  103, 313, 04 \\ g_{30} & 2219, 32 & 229, 136  -6, 254  199, 116  40, 642  2210, 732  228, 130  410, 072  236, 229  210  32  229, 136  -5, 573  224, 529  164  -5, 038  244, 526  174  414  224, 274, 22  9103, 8 \\ g_{30} & 2349, 32  229, 136  -5, 652  255, 549  721  226, 2498  -77, 7103, 74 \\ g_{40} & 10, 32  229, 198  -5, 653  244, 626  7, 4114  249, 748  -25, 0103, 74 \\ g_{40} & 10, 32  229, 198  -5, 643  249, 734  265  544  226, 674  -226, 0103, 74 \\ g_{40} & 149, 32  229, 198  -5, 653  244, 626  7, 411  249, 299, 99, -17, 5103, 74 \\ g_{40} & 149, 32  229, 198  -5, 653 $	239	1249.32	225.848	-10.570	66,537-17,513		27.2 106.0+
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	239		226.525			110483.	32,5 105.8*
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$\begin{array}{cccccccccccccccccccccccccccccccccccc$	239	1649.32				-	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	590	1719.32	229,054				
230 $1^{p}40$ , $32$ $229$ , $242$ $-7$ , $321$ $12^{p}$ , $70^{k}$ , $25$ , $h \cdot 9$ $179397$ , $40$ , $210^{4}$ , $22$ 240 $10^{1}0$ , $32$ $229$ , $259$ $-7$ , $165$ $136$ , $614$ $29$ , $298$ $10^{1}6733$ , $40.4$ $10^{4}$ , $22$ 239 $2019$ , $32$ $229$ , $248$ $-6$ , $875$ $151$ , $434$ $-6$ , $073$ $197861$ , $35.4$ $11.4$ , $12$ 239 $2019$ , $32$ $229$ , $229$ $-6$ , $740$ $160$ , $356$ $46$ , $4^{-1}$ $204092$ , $30$ , $2$ $104$ , $09$ 230 $2140$ , $32$ $229$ , $229$ $-6$ , $740$ $160$ , $356$ $46$ , $4^{-1}$ $204092$ , $30$ , $2$ $104$ , $09$ 230 $2140$ , $32$ $229$ , $120$ $-6$ , $4875$ $170$ , $641$ $-0$ , $644$ $216704$ , $19$ , $3$ $103$ , $99$ 230 $2210$ , $32$ $229$ , $136$ $-6$ , $254$ $199$ , $110^{1}$ $90^{107}$ $2296432$ , $8, 3$ $103, 99$ 230 $2210$ , $32$ $229$ , $136$ $-6$ , $254$ $199$ , $110^{1}$ $90^{107}$ $2296432$ , $8, 3$ $103, 99$ 230 $2310^{1}, 32$ $229, 116$ $-6$ , $144$ $208, 239$ , $36, 972$ $236^{0}26$ , $2, 9$ $103, 99$ 230 $2349, 32$ $229, 116$ $-6, 1038$ $216, 740^{1}$ , $44, 14$ $2425776$ , $-7, 7$ $103, 79$ 240 $10^{1}, 32$ $220, 119$ $-5, 936$ $224, 263$ , $30, 64$ $229176$ , $-7, 7$ $103, 79$ 241 $10^{1}, 32$ $220, 119$ $-5, 743$ $238, 3264$ , $221^{16}$ $2429176$ , $-7, 7$ $103, 79$ 241 $10^{1}, 32$ $229, 160^{1}, 5743$ $248, 3264$ $27164^{1}, 24900^{19}, -22.0$ $103, 79$ 241 $10^{1}, 32$ $229, 198$ $-5, 433$ $241, 736$ , $26, 5^{14}$ $226748^{1}, -12, 7$ $103, 79$ 241 $10^{1}, 32$ $229, 313$ $-5, 442$ $255, 549$ $7, 2^{1}, 2424^{2}, 2704^{2}, -296, 103, 6^{1}, 244^{2}, 296, 09^{1}, -22.0, 103, 7^{1}, 244^{1}, 310, 32$ $229, 389$ $-5, 413$ $265, 744$ $10^{2}, 744^{1}, 27064^{1}, -26, 0$ $103, 6^{1}, 240^{1}, 340, 32$ $229, 648$ $-5, 114, 265, 113, 512$ $2474^{2}$ $270641^{1}, -36, 510^{1}, 32$ $210, 667$ $-5, 2573$ $270, 959$ $-8, 770^{1}, 3102^{2}, 11, -36, 510^{1}, 32$ $240, 510, 32, 229, 648$ $-5, 117$ $281, 461^{1}, 10, 459$ $316342^{1}, -37, 1103, 58$ $240, 510, 32, 220, 648$ $-5, 117, 281, 461^{1}, 970, 73$ $329676^{2}, -34, 1103, 58$ $240, 740, 32, 230, 208$ $-4, 938$ $290, 689^{-4}, 4511$ $3364114^{-3}, -31, 51$	239	1749.32	229.145	-7,661	116,551 45,558		
230 $1^{p}40$ , $32$ $229$ , $242$ $-7$ , $321$ $12^{p}$ , $70^{k}$ , $25$ , $h \cdot 9$ $179397$ , $40$ , $210^{4}$ , $22$ 240 $10^{1}0$ , $32$ $229$ , $259$ $-7$ , $165$ $136$ , $614$ $29$ , $298$ $10^{1}6733$ , $40.4$ $10^{4}$ , $22$ 239 $2019$ , $32$ $229$ , $248$ $-6$ , $875$ $151$ , $434$ $-6$ , $073$ $197861$ , $35.4$ $11.4$ , $12$ 239 $2019$ , $32$ $229$ , $229$ $-6$ , $740$ $160$ , $356$ $46$ , $4^{-1}$ $204092$ , $30$ , $2$ $104$ , $09$ 230 $2140$ , $32$ $229$ , $229$ $-6$ , $740$ $160$ , $356$ $46$ , $4^{-1}$ $204092$ , $30$ , $2$ $104$ , $09$ 230 $2140$ , $32$ $229$ , $120$ $-6$ , $4875$ $170$ , $641$ $-0$ , $644$ $216704$ , $19$ , $3$ $103$ , $99$ 230 $2210$ , $32$ $229$ , $136$ $-6$ , $254$ $199$ , $110^{1}$ $90^{107}$ $2296432$ , $8, 3$ $103, 99$ 230 $2210$ , $32$ $229$ , $136$ $-6$ , $254$ $199$ , $110^{1}$ $90^{107}$ $2296432$ , $8, 3$ $103, 99$ 230 $2310^{1}, 32$ $229, 116$ $-6$ , $144$ $208, 239$ , $36, 972$ $236^{0}26$ , $2, 9$ $103, 99$ 230 $2349, 32$ $229, 116$ $-6, 1038$ $216, 740^{1}$ , $44, 14$ $2425776$ , $-7, 7$ $103, 79$ 240 $10^{1}, 32$ $220, 119$ $-5, 936$ $224, 263$ , $30, 64$ $229176$ , $-7, 7$ $103, 79$ 241 $10^{1}, 32$ $220, 119$ $-5, 743$ $238, 3264$ , $221^{16}$ $2429176$ , $-7, 7$ $103, 79$ 241 $10^{1}, 32$ $229, 160^{1}, 5743$ $248, 3264$ $27164^{1}, 24900^{19}, -22.0$ $103, 79$ 241 $10^{1}, 32$ $229, 198$ $-5, 433$ $241, 736$ , $26, 5^{14}$ $226748^{1}, -12, 7$ $103, 79$ 241 $10^{1}, 32$ $229, 313$ $-5, 442$ $255, 549$ $7, 2^{1}, 2424^{2}, 2704^{2}, -296, 103, 6^{1}, 244^{2}, 296, 09^{1}, -22.0, 103, 7^{1}, 244^{1}, 310, 32$ $229, 389$ $-5, 413$ $265, 744$ $10^{2}, 744^{1}, 27064^{1}, -26, 0$ $103, 6^{1}, 240^{1}, 340, 32$ $229, 648$ $-5, 114, 265, 113, 512$ $2474^{2}$ $270641^{1}, -36, 510^{1}, 32$ $210, 667$ $-5, 2573$ $270, 959$ $-8, 770^{1}, 3102^{2}, 11, -36, 510^{1}, 32$ $240, 510, 32, 229, 648$ $-5, 117$ $281, 461^{1}, 10, 459$ $316342^{1}, -37, 1103, 58$ $240, 510, 32, 220, 648$ $-5, 117, 281, 461^{1}, 970, 73$ $329676^{2}, -34, 1103, 58$ $240, 740, 32, 230, 208$ $-4, 938$ $290, 689^{-4}, 4511$ $3364114^{-3}, -31, 51$	230	1419.32	229 200	-7.4P6	122,384 20,4*3	173291	- 52,7 <u>1</u> ₩4,4*
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POSE

POSE

(Point Search Ephemeris)

INPUT DECK: Manual Schedule Tape Mode

INPUT DATA:

TA: (1) Request card:

Cols 1 - 5: T-start in minutes (DDDD.) (minimum period expected)

Cols 6-8:  $\triangle t$  in minutes (DD.) (time increment used to increase period)

Cols 9-13: T-stop in minutes (DDDD.)

(maximum period expected)

Col 14 output options:

1 = long form requested

- 2 = short form requested
- (2) Standard Observation Card
- (3) Standard Station Card(s)

(4) Option Card:

Cols 1-4: -1 = read another unit of input data (cards 1-4)

0 or blank = terminate program

OUTPUT:

Logical tape 11 contains a table of look-angles, for each station requested, in either long or short format. The short form of output consists of lookangle information, time of crossing in days, hours, and minutes of Zebra time; elevation and azimuth angles; and slant range in kilometers. The long form includes, in addition, the Cartesian coordinates of the station and object.

The TELTYP program may be used if teletype transmission is required.

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Field	Cols.	
1	1 - 5	T-Start in Minutes
2	6 - 8	$\Delta t$ in Minutes
3	9 - 13	T-Stop in Minutes
4	14	1, Long Form of Output
		0, Short Form Output
	Note:	Decimal points may be punched anywhere

Request Card POSE

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in each of the first three fields.

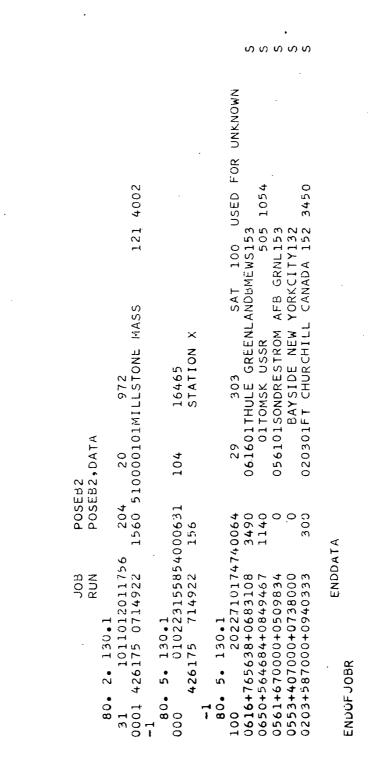
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POSE

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# 5-10-3

#### Wolf Research and Development Corporation

#### POINT SEAPCH PROGRAM

ſ, STATION LATITUDE = 76:5438 LONGITUDE = 291:6892 EAST HEIGHT = 349. METERS URSERVED TIME = 58. DAYS 10. HOURS 17. MINUTES 47.40 SECONDS GMT ELEVATION = 64.0000 ATIMUTH = 2.9000 RANGE = 30310 KM/11 THETAG = 311.22012 DEGREES 1441AS = 242.90932 DEGREES 04PR - 6358.216 KM GEDCENTFIC STATION LATITUDE = 1.33477 RADIANS xT = 549.491 TTA = -1381.569 ZFTA = 6181.929 KM DECL # 1.3509 RA # 0.9951 RADIANS CAPX # -677.104 CAPy # -1323.708 CAPZ # 5181.929 KM THE COORDINATES OF THE POINT IN THE INEPTIAL REFERENCE-SYSTEM APE x = -641.181 y = -1258.236 Z = 6477.534 KH+ \* THILF GRFENLANDBMEWS ALC ZEBRA TIME ELFY AZTH R DAY HR MIN ANG. ANG. 50 11 37.79 25.7 294.0 UNIDENTIFIED DBJECT+ THULF GREENLANDSHEWS RANGES K9+ 584+ 56 11 42.79 24.2 293.A 56 11 47.79 22.2 293.A 511. 538+ 56 11 52,79 21 293 4 58 11 57.79 20 2 293 4 6554 593. 56 12 2.79 19.3 293.A 721+ 750. 56 12 7,79 18,7 293,7 293.9 58 12 12.79 17.4 779+ 88 12 17.79 16.3 294.1 804+ 50 12 22 79 15 4 294 3 58 12 27 79 14 1 294 5 835\* 863. \*5 650 . TÓPSK USSR 630 UNIDENTIFIED DBJECT# ZEBRA TIME ELEV AZIM RANGE. Day HR MIN ANG: ANG. KYA NU VISUAL PASSES\* \*法 BEREBERBERE BERER 561 UNIDENTIFIED DEUECT+ SUNDRESTROM AFR GANL 555 ZEBRA TIME PLEV AZTM DAY HR MIN ANG: ANG. RANGE Кме NU VISUAL PASSES+ \* %

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#### ORPS

### (Orbital Plane Search)

INPUT DECK: Manual Schedule Tape Mode

INPUT DATA: The input data is in units consisting of an element set, a station card, and a request card. Any number of units may occur. The last unit must be followed by a blank card to terminate the program. Each unit will consist of:

- (1) Standard element set (6 or 7 cards)
- (2) Standard station card
- (3) Request card:
  - Cols. 1-10: Start time of search (days of year)
    - 11-20: Time increment between search points (minutes)

21-30: End time of search (days of year)

- 31-40: Observing station's search azimuth angle (degrees)
- 41-47: Station's maximum slant range (kilometers)
- 48-57: Station's minimum elevation angle (degrees)

58: 1 punch-visible passes only.

(4) Blank card, will terminate the program. The card should follow the last repetition of (1) to (3).

OUTPUT:

The search point data for a satellite is preceded by the satellite elements, request information, year constants used in the calculations, and the station data. Each line of the search point data consists of the year, month, day, hour and minute; the right ascension, declination, azimuth and elevation in degrees; the slant range in

### 5-11-1

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kilometers; and the elevation and illumination angles of the sun in degrees. The information is in order of increasing time during the range requested. If only visible passes were requested, a check is made and only visible points will appear in the output.

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#### Wolf Research and Development Corporation

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Field	Cols.	
1	1-10	Start Time
2	11-20	Time Increment
3	21-30	End Time
4	31 - 40	Search Azimuth
5	41 - 47	Maximum Slant Range in Kilometers
6	48-57	Minimum Elevation Angle in Degrees
7	58	l, Visual Passes Only

# Request Card ORPS

# 5-11-3

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Wolf Research and Development Corporation

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ORPS52 JOB RUN ORPSB2,DATA 33 13 160 GAMMA 3 33 13 21960 1100.024445879 51.287999 176.37720871 1960 176.37720871 33 13 3 .06497800 291.27649 166.96225 1.04471998 33 13 4-5.3074610E-09- 4.9070522 3.7491586 -4.1834998E-05 33 13 5-0. E- -0. Ε--0. Ε--0. E-33 13 60. 0. 0. 0. 0300+107330+0616000 140 5 030001TRINDAD 159 10. 191.33333 120. 190.916670 4000• 0.

ENDDATA

ENDOFJOBR

# 5-11-4

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ORBITAL PLANE SEARCH AND XY LOUK ANGLE PROGRAM

SAT NO: 33 Unbital elements	EFEM NO. 13*			
U,176UUNUN+003	U.377208/1+800	0.649/8000-001	-0,53074610-008	0;0000000+000
zv.1633017∛−006	9.V0CA0000+000	0.29127649+003	-0.49070522+001	0,00000000000000000
0.16096225+803	0.37+91506+001	8.0000000 <b>+00</b> 0	U.24445879-UU1	0;51287999+002

TIME BEGIN STEP FIME FNF SEARCH AZJMUTH 190.91667 10.00000 191.33333 120.00000

Prak = 4000. HMIN ±

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HELA(U) = 98,674000 LONG SUN = 278,55980 C14 = 3.69265

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	LNDA				.735n	A1.0			:	[4.*				
		T NU	•	33	ELÉM	NO. 1.								
	T1		(2)			56	ARCH	90	FNT		ELEV	ILL*		
	" " U N				ΗA	ល⊾C	Α		ELEV	SH	SUN	ANG +		
	յսլ		55		218.3	-21.0	120	Ú	7.5	2471.	∎Ü.3	6.4+		
	UUL		22	40.	?77.7	-20.7	120	Ú	10.4	2253.	#2.6	6.2+		
	UUL		22	51	276.7	-25.7	120	0	13.7	2037.	=4.d	5.9+		•
	ບບປ	6	23	ŋ	275.3	-24.4	120	0	17.6	1824.	×7.U	5.74		
	υUL	Ð		10	273.1	-22.7	120	0	22.2	1617.	=9.5	5,5+		
	UUL	8	-	20	270.1	-20.4	120	0	27.7	1418.	=11.5	5.2+		
	սնլ	.8	-	3.0	206.0	-17.3	120	0	34.A	1230.	+13.7	5.0+		
	ູມປະ	8		<b>4</b> 1)	201.2	-13.1	120	1	43.8	1061.	=15.9	4.9+		
	νUL	ь	23	<b>5</b> n	252.4	-7.2	120	1	55.5	921.	-18.1	4.7+		
	μUL	9	ŋ	n	242.2	0.5	120	2	70.4	822.	=20.2	4.5+		
	មហដ្ឋ	9	ή	10	229.9	9.5	121	5	87.7	782.	+22.4	4.4*		
	սնլ	9	Û	29	214.5	1/,9	299	8	74.8	809.	=24.5	4.2*		
	JUL	9	O	30	204.0	24.1	299	9	59.5	895.	-26.6	4,1+		
-	JUL	9	Ú	40	193.6	28.0	299	9	47.3	1025.	=28.7	4.0+		
	JUL	9	Ű	50	185.7	30.1	300	Ú	37.9	1185.	-30.7	3,9+	•	· .
	JUL	. 9	1	n	179.8	31.1.	300	0	50.5	1362.	=32.7	3.6+		
	JUL	9	1	10	175.6	31,ó	300		24.9	1551.	=34,7	3.7+	-	• ••
	JUL	9	1	· 20	172.6	3.17			20-2-	1747		3.6*		
	JUL	9	1	30	170.6	31,6	300	0	16.3	1947.	-38.5	3,5*		· · · · · · · · · · · · · · · · · · ·
	JUL	9	- 1 -	-40	-1-692-		_30.0_	0	13-0-	- 21.50				
€ V	JUL	9	1	511	164.3	31,1	300	0	10.1	2354.	=42.2	3.4+		
* · 🖌 U		9.	· • 2 ·	- 1)						-2559	43.9			
· 6v		9	2	10	167.7	30.4	300	Ó	5.2	2764,		3,3+		• •••••
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6 U	JUL	9	2	3 n	168.0	29.7	300		1.1	3174,	-#48 .7	3,3+		

ORPS

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### ASUM

(Acquisition File Summary)

INPUT DECK: Manual Schedule Tape Mode
INPUT DATA: SEAI Tape (A-File)
OUTPUT: The output consists of the information contained in the A-File of the SEAI tape. The output is sequential according to satellite number and includes satellite number, sensor number, sensor name, pass code (all or visual), format request (short or complete), type (all, 3-point, Baker-Nunn), minimum azimuth, maximum azimuth, minimum elevation, maximum elevation, maximum range, and step size in minutes.



SAT     SEMS     SEMSUR     RENSUR     RENSUR     Rensure     COUL     CURMAL     LIMITS       NT       NT     NT     NT     NT     NT     NT     NT     NT     NT     NT     NT       NT     NT     NT     NT     NT     SENSUR     REL     CURMAL     MAX     HTV     HTV     HTV       NT     039     1.48FD0 <texas< td="">     ALL     CUMAL     SHORT     S     JUU     US0     US</texas<>			
039     1.4AFDO TEXAS     AL     UJUU     5     Juuu     5     Juuu       306     030     TATAIDAT     VIS     S404T     Juuu     550     Juu       039     1.44EDD TEXAS     ALL     S404T     Juuu     5     Juu       039     1.44EDD TEXAS     ALL     S404T     Juuu     550     Juu       77     4TTE SAUUS     2.44E     ALL     UJMU     5     Juu       72     4TTE SAUUS     7.5     4LL     UJMU     5     Juu       72     4TTE SAUUS     ALL     UJMU     3     Juu     5       72     4TTE SAUUS     ALL     UJMU     Juu     3     Juu	MITS GRID SIZE	B-N CONTROL	TTY C. Inhib Type
030     1.44EDA TEXAS     ALL     SHAT     J     J     J     J       3DA     TATIDAN     VIS     C394L     J     VUU     350     JUU       723     44TTE SAVUS 3A1E     ALL     C394L     J     UUU     350     JUU       94r     *PATTE SAVUS 3A1E     ALL     C394L     J     UUU     350     JUU       724     *ATTE SAVUS 3A1E     ALL     C394L     J     UUU     JUU     JUU       724     *ATTE SAVUS 3A1E     ALL     C394L     JUU     JUU     JUU       724     *ATTE SAVUS 3A1E     ALL     C394L     JUU     JUU     JUU	ugy v5000 2.0000 ugy v5000 2.0000	74	YES A
72%     41TE     SAUS 341E     ALE     UC     341     341     341     341     341     341       94r     34r     VLS     VLS     VLS     VLS     VLS     341       94r     34r     41     VLS     VLS     VLS     341     341       94r     34r     41     VLS     34r     34r     34r     34r       72%     41     37     34     41     VLS     34r     34r       72%     41     41     VLS     34     41     34r	491 45000 - 3.4000 990 45000 - 5.0000	4	ES 7
72: "4TTE SAVUS JAYE ALL CO <b>42L 3 uuu J</b> ön 72: "4TTE SAVUS JAYE VIS 5407T J uuu Jon	490 1000 2.0000 490 5008 2.0000	2 2	D ON
724 ATTE SAVUS JA4E VIS SHORT J UJU 350	U90_140000 2.4.0000		0
	U90 14000 2.0000	<b>X</b>	40 3
084 72× 447TE SAVDS JA4E ALL S404F J, 940 300 309 084 946 Jafey VIS CJMPL 5 Jun 369 309	494 10000 2.0000 490 43008 2.0000	2 3	20 20 20 20
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ASUM

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Wolf Research and Development Corporation

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5-12-2

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### SSUM

(Sensor File Summary)

INPUT DECK: Manual Schedule Tape Mode .

INPUT DATA: SEAT Tape (S-File)

OUTPUT:

A listing of the sensor data contained in the S-File is provided. The output includes sensor number, sensor name, latitude in degrees, longitude in degrees, height in earth radii,  $x/\cos\theta$  or -(C + H) $\cos\theta$  in earth radii, z or  $-(S + H) \sin\phi$  in earth radii, accuracy digit for azimuth, elevation and range, classification, teletype request code, and sensor type.

5-13-1

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PAGE 1 Sensor TYPE	<b>ሶ ን ን ፕ</b> ኛ	1344N	0 0 र ग ग	オオナ	<b>1</b> 00	ተታ ቦስ	0 0 0 0 0 4 4 0 0	o, ე, Q	o.o.o.o.o
RPTG	****	****	***	****	****	<b>**</b>	****	* *	***
:S10	22272	7.2.2.2.2	72272	22722		2 7 7 7 7 7	22222	z	Z Z Z Z Z
رور #6^	20000 20000 20000	55665 56666	00000	,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,,	000000000000000000000000000000000000000	2 7 2 7 9 2 2 2 2 2 2		00000 000000	00000 00000
ACCURACY J ELU XGA				00000	00 00 00 00 00	39999 9999 9999 9999 9999 9999 9999 99	00000		00000 00000
AZD	000000		10000	000000	000000000000000000000000000000000000000	00000	00 00 00 00 00 00 00	0 0 0 0 0 0 0 0	20000
L EAHLA RAD	-6./36-001 -6./28-001 -5.303-001 -5.303-001 -5.32-631	++++++++++++++++++++++++++++++++++++++	-5.5.4-001 -4.0/5-001 -4.417-001 -4.459-001 -4.215-001	-2.48-801 +5.258-011 -3.2512-001 -6.178-001 +1.3/8-001	-7.488-001 -7./28-001 -2.927-001 -2.927-001 -4./35-601	-9.5-3-001 -5.5-99-001 -7.7.10-001 -1.5.0-001 -1.5.0-001	-8.7.3-001 -2.5.3-001 -5.451-001 -5.456-001 -5.456-001	+5.×46.001 -3.5186011 -5.1524011 +6.1524011 +4.408-011 +1.050-001	+2.0<7-0,1 +5.552-001 +5.457-0,1 -2.758-0,1 -4.442-0,1
2 مال میں 14 مار کا 14 14 مار کا 14		10, -7/0, + 10, -3/0, + 10, -3/0, + 10, -3/0, + 10, -3/0, + 10, -2/0, + 10, -2			19-18-0.5 17-12-5 101-80-5 101-80-5 101-80-8 101-26-6-	10, -15, 10, -61, 10, -62, 10, -62, 10, -62, 10, -62,	10,-17,-17,-17,-17,-17,-17,-17,-17,-17,-17		10,-16, 10,-16, 10,-16, 10,-16, 10,-16,
טא-10-52 11-52 ב ב באאוי אנח	+2,442+33 +1,122+015 1 +2,233+194 +2,423+194			+1.02/16 +9.5/216 +9.7316 +4.733-16 +2.431-16,12 +2.431-16	+2,554 201-727.1 1. 1. 1. 2. 1. 2. 2. 2. 2. 2. 2. 2. 2. 2. 2. 2. 2. 2.	40	L +2,4-1-4U +2,4-1-4 +2,-2,5-4 +7,-1,4 Åy-1,40,1+	+1.25+-146 u -2 +2.5-5-64 +5.9-42	+7.6.<2-1136 +8.10/-1245 +1.627-094 +1.627-094 +7.637-307
05-10 1000000000000000000000000000000000	- /. 1 /. 000 - / 1 + 000 	, , , , , , , , , , , , , , , , , , ,	, , , , , , , , , , , , , , , , , , ,	+2, +11,+00K +2, +4,+00K +2, 1,57+00K +2, 157+00K +2, 152+00K +2, 152+00C		+ 1, 1, 1, 5 + 0 0 + 1, 2 + 0 0 2 1, 2 + 0 0 2 	·	++, \$< \$+00. +>, \$1.5+00. +>, \$1.5+00. +/, Yu4+00. +2, J/400. +2, J/400. +2, J/400.	10,4/1/ 10,4/20 10,4/20 10,4/20 10,4/20 10,4/20 10,4/20 10,4/20 10,4/20 10,4/20 10,4/20 10,4/20
L'ÉTTUDE L'ÉGRÉES	++,251+001 -1,255+001 +3,570+031 +5,242+001 +5,242+001	- 5. L10+UJ1 + 5. 746+0UL + 5. 357+0UL - 1. 345+012 - 1. 345+012 + 2. 751+001	+,,257+041 +2,355+041 +2,355+041 +2,355+041 +2,355+041 +2,702+041	+ L. 299+001 - 5. L94+001 - 70+001 - 70+001 - 7. 370+001 - 7. 374+090	++. \$80+091 		+ ), +28+0 )	- 2, 539+041 + 2, 278+041 + 5, 345+041 + 5, 345+041 - 4, 522+041 - 5, 226-041	- 4. 677+041 - 2. 652+001 - 3. 614+001 - 5. 714+001 - 1. 714+041 + 2. 294+041
אראייייא איייאיד אראי ארגייז, אראיייאר זיייד אראיי	TLESTIN 4555 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-1115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115 1115-115	1 144 (2) :	.5156 314750 1407404 41767 0048 4557 001 14260 3297100 513	сиватти ихт «([Lan)» 1005 Анз «AUT 4NT «ACDA 4 <sup>1</sup> т. ()AVT5 «ACDA 4 <sup>1</sup> т. ()AVT5 »ATRTG7 т.М. 14 5		LARVEST	· · · · · · · · · · · · · · · · · · ·	1:14, 113,450 17,444,471 14,6 50147 - 14,6 19,81,24,51,54,54 19,81,24,51,54,54	ATZ 1.2.4 MP 1 ATZ 1.1.4 MP 2 ATZ 1.1.4 LEAN ATZ 1.1.4 LEAN Contrait 1.7 MP 2 Contrait 1.7 MP 2 Contrait 1.7 MP 2 MP 2 MP 2 MP 2 MP 2 MP 2 MP 2 MP 2
8 805 11 805 11	107 107 108 108 107 107 107 107 107 107 107 107 107 107	55 75 77 77 77 77 77 77 77 70 70	5 C F 4	11 4 4 4 4 4 4 4 4 4 4 7 0 8 7 0 8 7 0 8	5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5 5	<u>у с о с с с с с с с с с с с с с с с с с </u>	2 + 2 + 4 0 4 4 4 4 0 + 0 K 4	4 - 0 4 4 4 - 0 4 4 4 - 0 4 5	4 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0

5-13-2

#### ISUM

### ISUM

### (Information File Summary)

INPUT DECK:

INPUT DATA:

(1) SEAI Tape (I-File)

Manual Schedule Tape Mode

- (2) Request cards with any of the following entries punched in columns 1-3:
  - (a) Three decimal digit satellite number.
     Any number of cards may be used to request specific satellites.
  - (b) ALL. The ALL punch requests all satellite information, replacing request card type (a).
  - (c) BOX. The BOX punch will provide a box score of still orbiting satellites as additional output.
  - (d) END. The END punch signals the end of request deck, and must be used as part of the input deck.

#### OUTPUT:

The I-File of the SEAI Tape is written for off-line printing. Information printed out includes satellite number, satellite name, launch date, launch site, booster country and payload country. . .

#### INFORMATION SUMMARY

09-10-62

0927.8

SAT. NUMBER CLASSIFICATION CODED SAT. NAME INTERNAT'L CODE COMMON NAME LAUNCH DATE LAUNCH SITE LAUNCH AGENCY BOOSTER COUNTRY PAYLOAD COUNTRY MISSION WEIGHT (KG) SHAPE LENGTH (MTR) HEIGHT (MTR) WEDTH (MTR) DIAMETER (MTR) MEAN DRAG VARIANCE RAD X-SEC (6 MT) VARIANCE MEAN REFLECT. VARIANCE TUNB DATE & RATE TUMBL'TNG MODE STABTLIZATION NANEUV. CHAR. TRANS. FREQS

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RECEIVING FREQS

DECAY DATE DETERMINED LIFFTIME (YRS) HELIO. ELMS. E A 01 Q2 1 P

084

6106 61ZETA DISCOVERER21 18FEB 61 VANDENBERG AFB

2 2 2 A

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### INFORMATION SUMMARY

SAT. NUMBER CLASSIFICATION CODED SAT. NAME INTERNAT'L CODE COMMON NAME LAUNCH DATE LAUNCH SITE LAUNCH AGENCY BOOSTER COUNTRY PAYLOAD COUNTRY MISSION WEIGHT (KG) SHAPE LENGTH (MTR) HEIGHT (MTR) WEDTH (MTR) DIAMETER (MTR) MEAN DRAG VARIANCE RAD X-SEC (6 MT) VARIANCE MEAN REFLECT. VARIANCE TUMB DATE & RATE TUMBL'TNG MODE STABTL IZATION NANEUV. CHAR. TRANS. FREQS

#### RECEIVING FREQS

DECAY DATE DETERMINED LIFFTIME (YRS) HELIO. ELMS. E A

09-10-62

0927.8

084 61.06

61ZETA DISCOVERER21 18FEB 61 VANDENBERG AFB

TELTYP

### TELTYP

### (Teletype Output Conversion)

INPUT DECK: Manual Schedule Tape Mode.

INPUT DATA: Tape 11 is scanned for sentinels marking the beginning and ending of messages to be converted to teletype code. All appropriately marked messages will be converted.

Sentinels used:

- \* \$: the asterisk followed by a dollar sign marks the end of a message to be converted.

OUTPUT:

. . . . . . . ....

Tape 11 contains the data converted to Baudot code following output previously written on the tape.

# BMEWSPT

# (BMEWS Paper Tape)

INPUT DECK: Manual Schedule Tape Mode.

INPUT DATA: Tape 0 should contain the Q point data from the DIP as described in the writeup.

OUTPUT: Tape 7 will contain observation information in the standard format required for input to ORCON.

## 5-16-1

## BMEWSPT

(BMEWS Paper Tape)

- INPUT DECK: Manual Schedule Tape Mode.
- INPUT DATA: Tape 0 should contain the Q point data from the DIP as described in the writeup.

OUTPUT: Tape 7 will contain observation information in the standard format required for input to ORCON.